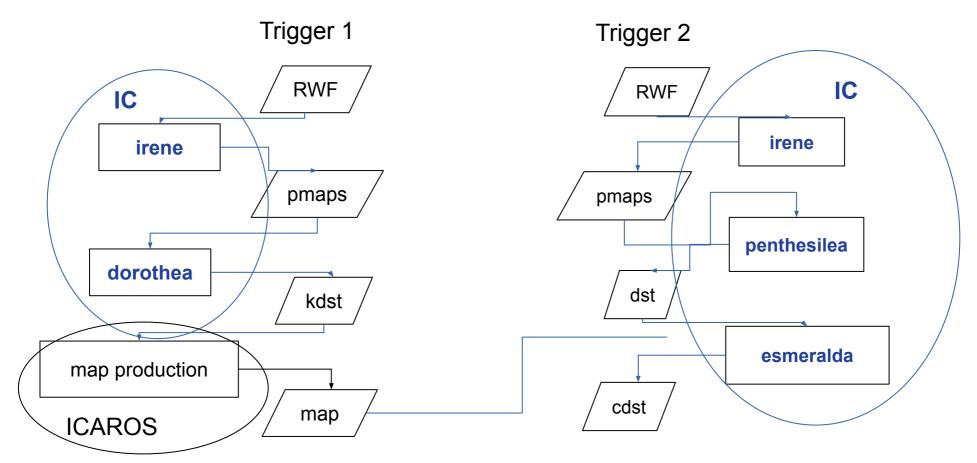
# Data processing algorithms

IC cities

## Official production flow



## Running a city method

After setting up the IC environment

source work\_in\_python\_version 3.7

a city can be run as:

\$ city city\_name config\_file.conf

#### irene 10

**Input** (from the decoder or MC-detsim) – RWF (raw waveforms): a time-ordered signal amplitude for each sensor in ADC.

**Output** – pmaps (Peak maps): a collection of all the peaks and the slices of waveforms belonging to them.

In RWF files the waveforms are in **RD/sipmrwf** and **RD/pmtrwf**.

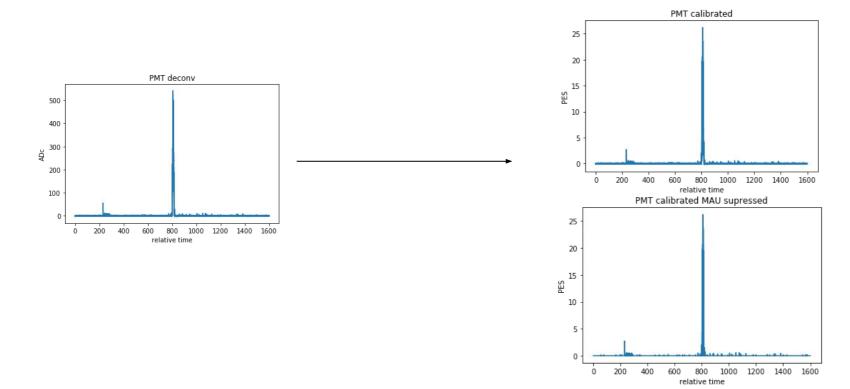
Exercise: Using h5ls on rwf\_example.h5 find the number of events and sensors inside the file. (h5ls -d will print the content of the table)

## **1. PMT pedestal removed and subject to baseline restoration** (per pmt waveform)

Pedestal - the average of the amplitude of the entire waveform; Deconvolution using BLR algorithm (converts bipolar signal to monopolar)



**2. correct wf** - ADC to pes (divide each waveform by the factor read from database)

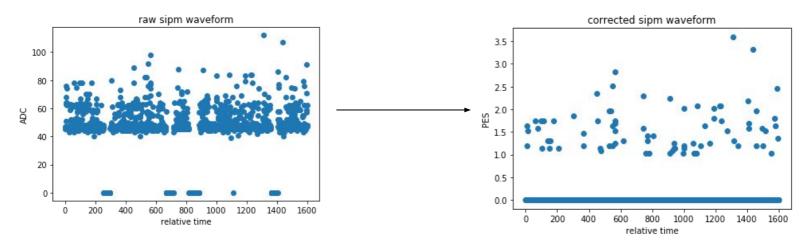


**3. zero suppressed** – find indices of bins where the **sum** of WFS is above given threshold. Two threshold (for S1 candidate use MAU suppressed corrected WF, for S2 candidates use corrected WF)

**4. correct and threshold sipm waveforms** – subtract baseline, convert to pes, set to 0 values below given threshold

Baseline is a mode of a waveform (the most frequent value)

Threshold is either 'common' (1 pes for all sipms), or 'individual' (percentage exclusion from calibration runs)

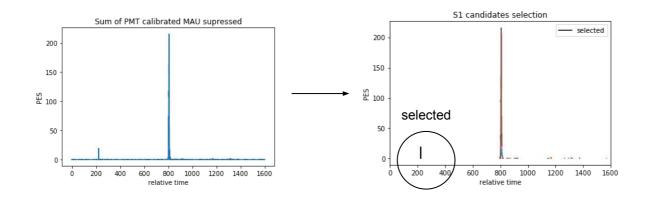


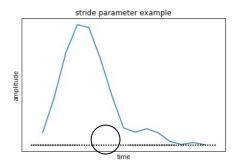
Note: sipm rwf are zero suppressed in FPGA

5. pmap builder: select s1/s2 peaks

#### S1 peak:

- split in disjoint signals (peaks) (using indices from step 3, allowing small gaps)
- search for peaks that satisfy time and length constraint



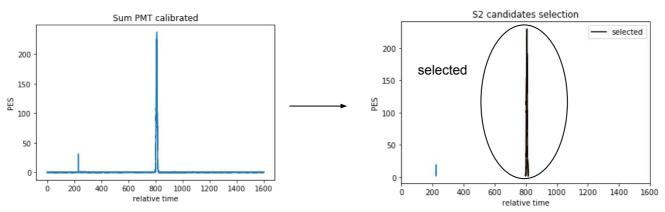


Value of ccwf goes below the threshold

5. pmap builder: select s1/s2 peaks

#### S2 peak:

- split in disjoint signals (peaks)
- search for peaks that satisfy time and length constraint



- pmt signal is rebinned to match sipm binning (from 25 ns to 1 μs)
- include only sipms that have integrated (in time) signal larger than a given threshold (5pes)

#### pmaps

#### In **PMAPS** tables are:

- **S1** contains event, peak id, time, bin width, energy = summed wf over pmt for S1 peaks (per time bin) (summed over PMT waveform)
  - **S1pmt** contains event, peak id, pmt id , pmt signal (per pmt, per time bin)(waveforms)
- **S2** contains event, peak id, time, bin width, energy = summed wf over pmt for S2 peaks(per time bin) (summed over pmt waveform)
  - **S2pmt** contains event, peak id, pmt id, pmt signal (per pmt, per time bin)(waveforms)
  - S2Si contains event, peak id, sipm id, sipm signal (per sipm, per time bin)(waveforms)

Exercise: run irene on rwf.h5 file using provided configuration.

Examine output file and find above-mentioned tables. What is the bin width for S1 and S2 peaks? How many S1 peaks were found in event 1046556? And S2? (Use head, tail or grep piping)

#### dorothea IO

Input (from irene) – pmaps

 Output – kdst (krypton dst): per peak relevant information (time, number of peaks, width, height, position...)

## dorothea algorithm

1. select the peaks that satisfy a given set of conditions on peak width, height and energy. The energy is calculated as a sum of bins that have amplitude greater than some en\_th (usually 0pes for S2 and 0.5 pes for S1)

Events that contain no peaks with the given condition are filtered out

Note: energy (E) refers to signal detected on PMTs and charge (Q) to signal detected on SiPMs (que está mal pero bueno)

### dorothea algorithm

#### 2. build pointlike events

per S1/S2 combination of peaks that passed selection calculate:

- for **S1**: width (in ns), height, energy (above threshold)
- for **S2**: width (in μs), height, energy (above threshold), charge
- Extract xy position of sipms from the database. Find a 2D barycenter of the whole S2 peak (i.e. integrated over peak duration)
- Find Z=1 \* (S2\_time\_at\_max\_energy-S1\_time\_at\_max\_energy) (in mm)
   (Note that drift velocity is just assumed to be 1 mm/µs)

Note: peak\_id is resetted and does not correspond to pmaps peak\_id!

#### kdst

#### **DST/Events** table contains:

event, time, s1\_peak, s2\_peak, nS1, nS2, S1w, S1h, S1e, S1t, S2w, S2h, S2e, S2q, S2t, Nsipm, DT, Z, Zrms, X, Y, R, Phi, Xrms, Yrms

Exercise: run dorothea on the previously produced pmaps using provided config file. Is there any event with more than one S1 and one S2?

### penthesilea IO

Input (from irene) – pmaps

 Output – hdst (hit dst): per energy deposition (hit) relevant information (position, charge, energy...)

### penthesilea algorithm

**1. select peaks that satisfy a given set of conditions** on peak width, height and energy. The energy is calculated as a sum of bins that have amplitude greater than some en\_th (usually 0pes for S2 and 0.5 pes for S1)

Keep events that have one and only 1 S1 (standard configuration but not constrained by the code)

### penthesilea algorithm

#### 2. build hits

- rebin S2 waveform (usually to bins of 2 μs width)
- take S1\_time\_at\_max\_energy of first S1 peak as beginning time
- per S2 peak:
  - Extract xy the position of the SiPMs from the database for all SiPMs in the peak per time bin:
    - runs reco algorithm (finds X, Y, Q position of a hit)
    - distribute the energy (E) of the bin to hits found
    - If no hit was found, the X, Y position set to 0, Q to NN (-99999)

The result is a collection of X, Y, Z, E, Q ... variables per time bin

## penthesilea algorithm

#### 3. create kdst like table

The same procedure as in dorothea.

## reco algorithm (corona)

- Ignores sipms with charge below Qth;
- Select hottest sipm with charge above Qlm;
- 3) Finds first barycenter inside Im\_radius around the hottest sipm;
- 4) Finds second barycenter in new\_lm\_radius around first barycenter
- 5) If there are at least msipm sensors the hit is accepted

To set 1 hit per sipm reco paradigm (used for hdst):

```
Qthr = 5 * pes
Qlm = 5 * pes
Im_radius = 0 * mm
new_Im_radius = 0 * mm
msipm = 1
```

To set global barycenter (used for kdst):

```
Qthr = 1 * pes
Qlm = 0 * pes
lm_radius = -1 * mm
new_lm_radius = -1 * mm
msipm = 1
```

1 hit per sipm means that the XY position of the hit will always correspond to an XY position of the SiPMs

#### hdst

#### **RECO/Events** table:

event, time, npeak, Xpeak, Ypeak, nsipm, X, Y, Xrms, Yrms, Z, Q, E, (Qc, Ec, track\_id, Ep set to -1)

#### Exercise:

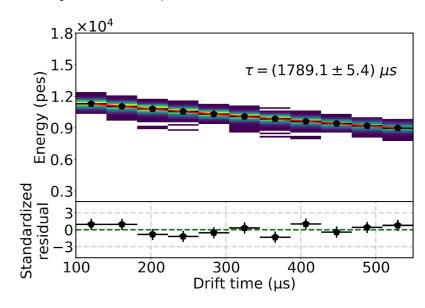
run penthesilea on pmaps.h5 using provided config file.
What is the charge of the first hit of the event 1046556? Why? What's its energy?

### map production

- 1) filter dst to select good kr events
- 2) bin detector in xy (ie you get n\_bins x n\_bins arrays of E, Z)
- 3) per bin fit E, Z to

$$E_{\rm measured} = E_{\rm true} \times E0 \times {\rm e}^{-Z_{\rm measured/LT}}$$

4) Store 2d maps of E0 and LT



#### esmeralda 10

Input (from penthesilea) – hdst (from map\_production) - maps

Output – cdst (corrected hit dst): the same as hdst with additional energy correction

track information: results of running paolina analysis

### esmeralda algorithm

#### 1. correct hits:

- Apply a new charge threshold to the hits. The energy of the hits that don't pass this threshold is redistributed among the hits that have the same time (Z)
  - get rid of failed (NN) hits: energy of those hits is redistributed to the (in Z) closest hits
  - using correction map calculate corrected energy (the units are keV at Kr scale)

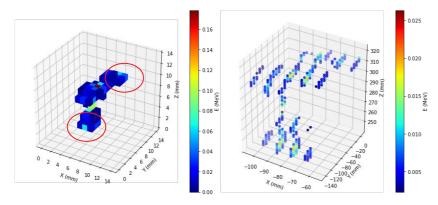
$$E_{\text{corrected}} = E_{\text{measured}}/E0 \times e^{Z_{\text{measured/LT}}}$$

#### This function is run two times:

- 1. for lower threshold used in deconvolution and DNN analysis (5-10pes)
- 2. for higher threshold used for everything else (30-35pes)

### esmeralda algorithm

- 2. run paolina functions (finding voxels and tracks for higher thresholded hits):
  - voxelize event
  - find tracks
  - drop end\_point voxel if its energy is smaller than a given threshold, redistribute its energy
  - find blobs (energy deposition at the end of the tracks)



#### cdst

RECO/highTh and RECO/lowTh

The output of correction procedure, has the same structure as the penthesilea output. (Qc, Ec, track\_id, Ep have meaningful values)

Tracking/Tracks is per track information containing:

```
event, trackID, energy, length, numb_of_voxels, numb_of_hits, numb_of_tracks, x_min, y_min, z_min, r_min, x_max, y_max, z_max, r_max, x_ave, y_ave, z_ave, r_ave, extreme1_x, extreme1_y, extreme1_z, extreme2_x, extreme2_y, extreme2_z, blob1_x, blob1_y, blob1_z, blob2_x, blob2_y, blob2_z, eblob1, eblob2, ovlp_blob_energy (energy of the hits belonging to both end point blobs), vox_size_x, vox_size_y, vox_size_z
```

**Summary/Events** table containing per event information:

```
event, evt_energy, evt_charge, evt_ntrks, evt_nhits, evt_x_avg, evt_y_avg, evt_z_avg, evt_r_avg, evt_x_min, evt_y_min, evt_z_min, evt_r_min, evt_x_max, evt_y_max, evt_z_max, evt_r_max, evt_out_of_map (the XYZ position of event is outside detector limits?; if positive energy and average positions will be nans)
```

DST/Events – copied table from penthesilea output

#### cdst

Exercise: run esmeralda. What is the energy of the first track of event 1046556,? Which energy has the highest energy? Which table has more entries: lowTh or highTh? Why?

Questions?