

12.1 The interaction of matter with radiation

- Equations

- Planck relationship: $E = hf$

- Einstein photoelectric equation: $E_{\text{kinetic}} = hf - \Phi$

- Bohr orbit energies: $E_n = \frac{-13.6 \text{ eV}}{n^2}$

- Conservation of angular momentum: $mvr = \frac{nh}{2\pi}$

- Probability density: $P(E) = \frac{1}{E}e^{-\frac{E}{kT}}$

- Momentum relationship: position-momentum $\frac{mv}{E} = \frac{h}{\lambda}$

- Energy-time: $\Delta E \Delta t \geq \frac{h}{4\pi}$

- The photoelectric effect

- Demonstration of the photoelectric effect

- the photoelectric effect can be demonstrated using the gold-leaf electrometer.

- In this experiment, a clean sheet of zinc should be mounted on an electrostatic plate and the sheet charged negatively by connecting to a high negative potential.

- When radiation from infrared to UV is shown, the leaf will only collapse back down when UV light is shown on the negative plate.

- This is because of the photoelectric effect.

- Explanation of the photoelectric effect

- Einstein explained the photoelectric effect in the following ways:

- light consists of particles known as photons.

- The energy of an individual photon is given by $E = hf$

- The total energy of the photon is given by $E = nhf$

- Each photon interacts with a single electron.

- There is a minimum light frequency, called "threshold frequency (f_0)", below which no electrons are emitted.

- Energy is needed to overcome the attractive force that act on the electron within the metal

- This energy is given by the quantity known as the work function (Φ).

- This further energy supplied to the electron will become kinetic energy of the emitted electron.

- The emitted electron is often called a photoelectron.

- Increasing the intensity of light simply increases the number of photons incident per second.

- They won't increase the energy of the photoelectrons.

- Explaining the gold leaf experiment

- The zinc sheet has a certain work function to pull the electrons off the metal.

- The photons incident on the zinc sheet must have at least this same amount of energy, if not more, than the work function to pull the electrons.

- UV light has the highest frequency, meaning that it will be able to pull the electrons.

- With very intense visible or infra-red light on the sheet the leaf remains damaged.

- Increasing the intensity of the light while emitting a low frequency will emit 0 electrons.

- This is because the photons won't overcome the Φ .

- The intensity will only increase the number of photons incident per second.

- If the intensity increases it won't affect the leaf falling.

- Electrons are instantly repelled from the metal.

- Meaning the zinc plate is no longer negatively charged.

- Causing the electrostatic force of repulsion to collapse, and the leaf to fall down.

- Placing a sheet of glass between the UV light source and the zinc prevents the leaf from falling.

- Glass will only transmit low energy visible photons, while absorbing high energy UV over.

- If the zinc sheet is charged positively the leaf remains damaged for all wavelengths of radiation.

- Raising the potential of the plate will mean that the force of attraction between the plate and electrons would increase.

- This means that the work function increases as well.

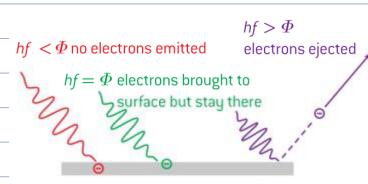
- It's need that the potential will still increase.

- Einstein photoelectric equation

- Einstein expressed the photoelectric effect with the following formula:

$$E_{\text{kinetic}} = hf - \Phi$$

- In the equation " E_{kinetic} " is the maximum kinetic energy of the emitted electrons.



▲ Figure 3 Photoelectric emission and the work function.

- \hbar is planck constant ($6.63 \cdot 10^{-34}$)

- f is the frequency of light.

- ϕ is the work function of the metal.

- everything in the equation is measured in joules.

- although, they can also be expressed in electron volts.

- the energy is written as a maximum because the work function is the minimum energy required to free an electron.

- it turns further into the initial will require more energy to free.

- Worked example

$$- E = 6.6 \cdot 10^{-34} S$$

$$- \phi_0 = \frac{\phi}{\hbar} \\ = \frac{6.6 \cdot 10^{-34}}{6.63 \cdot 10^{-34}} \\ = 1.0 \cdot 10^{-3} J/eV$$

$$- E_{max} = hf \\ = (6.6 \cdot 10^{-34})(1.0 \cdot 10^{-3}) \\ = 6.6 \cdot 10^{-35}$$

$$- E = 2.2 \text{ eV}, \phi = 1.5 \text{ eV}$$

$$(8.6 \cdot 1.6 \cdot 10^{-34}) = hf \\ f = \frac{(8.6 \cdot 1.6 \cdot 10^{-34})}{6.63 \cdot 10^{-34}} \\ = 2.1 \cdot 10^{19} Hz$$

$$- E_{max} = (8.6 \cdot 10^{-34})(2.1 \cdot 10^{19}) \\ = 1.8 \cdot 10^{-14}$$

$$- E_{max} = (8.6 \cdot 10^{-34})(6.63 \cdot 10^{-34}) - (1.8 \cdot 1.6 \cdot 10^{-34})$$

$$= 1.1 \cdot 10^{-34} \\ 1.1 \cdot 10^{-34} = \frac{1}{2} m e^2 \\ U = 6.6 \cdot 10^6 \text{ m s}^{-2}$$

$$- U = 6.6 \cdot 10^6 \text{ m s}^{-2}$$

- Wave theory and the photoelectric effect

- the wave theory for light fails to explain the photoelectric effect because of the instantaneous nature of electron emission.

- it proves that time in one tiny delay from when light is incident on a metal surface to when an electron is free.

- since waves provide a continuous supply of energy, and the intensity of a wave is proportional to the squared amplitude, then with sufficient time, electrons will be expelled, which isn't the case.

- this is because according to classic wave theory, when low intensity electromagnetic radiation of any frequency, given enough time, an electron would be freed from the potential well.

- however this obviously isn't the case as only photons with a frequency above the threshold frequency will free electrons without the delay.

- below this frequency, no electrons are emitted, completely contradicting wave theory.

- since light still has wave-like properties, these being diffraction and interference, then it needs to have wave-particle duality.

- Worked example

$$- f = 1.2 \cdot 10^{19} \text{ Hz}, \phi = 1.8 \text{ eV}$$

$$- E = hf \\ = (6.6 \cdot 10^{-34})(1.2 \cdot 10^{19}) \\ = 8.0 \cdot 10^{-15} \text{ J}$$

$$- E_{max} = hf - \phi \\ = 8.0 \cdot 10^{-15} - (1.8 \cdot 1.6 \cdot 10^{-19}) \\ = 9.12 \cdot 10^{-15} \text{ J} = 9.2 \text{ eV}$$

$$- E_{max} = \phi/m_e \\ = \frac{9.12 \cdot 10^{-15}}{1.6 \cdot 10^{-34}} \\ = 5.7 \text{ V}$$

- De Broglie waves

- de Broglie used the idea of light acting as a wave and a particle to come up with the following idea:

- if something is classically thought to act as a wave, then a species that is thought to be a particle must have wave-like properties.

- this brought him to the "de Broglie wavelength".

$$- \lambda = \frac{h}{p}$$

- where h is the planck constant, and p is the momentum of the species.

- the total energy of any object is its kinetic energy plus its rest energy, shown in the formula below:

$$- E^2 = c^2 + m_e c^2$$

- since photons don't have a rest mass, their total energy is given by:

$$- E = cp$$

- the de Broglie wavelength can be derived from the photoelectric effect and the total energy equation:

- since $E = cp$ and $E = hf = \frac{hc}{\lambda}$ equating them will give:

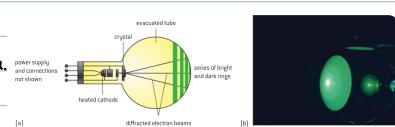
$$- cp = \frac{hc}{\lambda} \\ \lambda = \frac{hc}{cp}$$

- Electron diffraction

- the de Broglie wavelength was experimentally demonstrated by drawing interference minima when a beam of electrons were reflected by a nickel crystal.

- electrons from a heated cathode pass through a thin film of carbon atoms (crystal).

- if the electrons acted as particles, then they would only be slightly deflected by collisions with the carbon atoms.



[a] Figure 6 Electron diffraction tube.

- this would mean that they would form a bright region at the center of the screen.
- the bright regions on the image are where the electrons landed
- the bright regions represent high probability of electrons reaching that spot, while dark areas represent low probabilities.
- the same pattern is made over smaller time interval per second.
- this pattern is very similar to that of the interference pattern of light with a diffraction grating.
- showing that electrons also behave as waves.
- the electrons are accelerated through a potential difference, they gain kinetic energy:
- $E = \frac{1}{2}mv^2 = qV$
- knowing that the electrons don't move at the speed of light, their momentum will equal $p = mv$, meaning $p^2 = (mv)^2$
- $\lambda = \frac{h}{p} = \frac{h}{mv}$
- De Broglie's wavelength, $\lambda = \frac{h}{p} = \frac{h}{mv}$

Worked example

- Electron accelerated through pd of 2000V.

$$\begin{aligned} E &= \frac{1}{2}mv^2 \\ &= qV \\ &= (1.6 \times 10^{-19}) \\ &\quad \frac{1}{2}(1.6 \times 10^{-19} \cdot 1.6 \times 10^{-19} \cdot 2000) \end{aligned}$$

$$\lambda = 3.2 \times 10^{-11} \text{ m}$$

- the diffraction angle (θ) is introduced in the equation $\lambda = d\sin\theta$, where the voltage is measured.
- this is because the wavelength decreases as the voltage increases, mean the electron diffracts less.

Bohr model

- the Bohr model (where electron have fixed amounts of energy), made 3 key assumptions:

Electron in an atom exist in stationary states

- electrons can remain in these states without emitting any electromagnetic radiation
- electrons may move from one stationary state to another by emitting or absorbing a quantum (fixed amount) of electromagnetic radiation.
- if an electron orbiting a nucleus gives up a photon, it can move up to a higher "excited" state.
- the difference between the two energy levels must be equal to the energy of the photon.
- if an excited electron moves to a lower energy state, it will emit a photon.
- it will be equal to the energy difference of the two energy levels.

- the difference in the energy levels is given by the equation: $E = \pm hf$

- angular momentum of an electron in a stationary state is quantized in integral values of $\frac{h}{2\pi}$

- can be represented mathematically by:

$$mvL = \frac{hL}{2\pi}$$

- angular momentum is the (vector) product of the momentum of a particle and the radius of its orbit.

- therefore, for a particle in a circular orbit, the angular momentum will be constant.

- the pattern is corresponds to standing waves.

visualization of electron waves for first three Bohr orbits

electron wave resonance

$$n = 1, \lambda_1 = 2\pi r_1$$

$$n = 2, \lambda_2 = 2\pi r_2$$

$$n = 3, \lambda_3 = 2\pi r_3$$

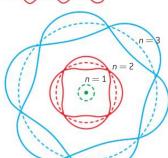


Figure 7 Standing waves in an atom for $n = 2$ and $n = 3$.

Energy in the Bohr orbits

- in the hydrogen atom, the total kinetic and potential energy of the energy levels can be represented by the equation:

$$E = -\frac{13.6}{n^2}$$

- $n = 1$ is the energy state, where $n = 1$ is the ground state, and $n = \infty$ is equal to the ionization/infinitive energy level.

- the energy is negative because the electron is in a potential well.

- due to the attractive force of the positive nucleus on the electron.

- Wheeler example

$$- \frac{mv^2}{2} = \frac{h^2}{2R}$$

- When an electron moves from a higher energy level to a lower one, it will release a photon of the same frequency and wavelength that is found in the hydrogen emission spectra.

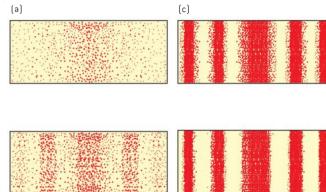
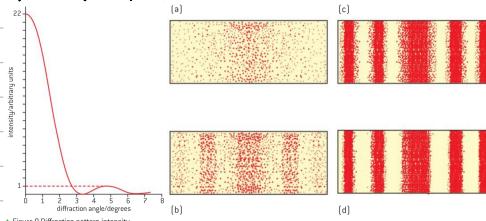
- What happens in practice

- Wave particle duality explains the bright fringes as being areas where there is a high probability of finding a particle.

- These probabilities interfere with one another to produce the expected interference pattern.

- With this pattern, we can determine where we're likely to find an electron, but we're not able to determine its actual location.

- Diagrammatical representation of the interference pattern



- Schrödinger's wave function "ψ" describes the quantum state of particles.

- For a light wave, we observe that $I \propto \psi^2$.

- The wave function shows where we're most likely to find a photon.

- For the wave function, where the square of the amplitude is a maximum, there is the greatest probability of finding a photon.

- Where the wave function is zero, there the probability of finding a photon is zero.

- The square of the amplitude of the wave function ($|\psi|^2$) is proportional to the probability per unit volume of finding the particle.

- This is known as the probability density: $P(\omega) = |\psi|^2 dV$

- $P(\omega)$ is the probability of finding a particle a distance ω from an origin, and dV is the volume considered.

- For the double slit diffraction, in terms of probability density, we assume that the single photon or electron is everywhere until it observes on the screen.

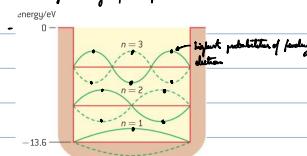
- This is known as the Lorentzian interpretation.

- The Lorentzian interpretation can be summed up to "nothing is real unless it's observed".

- meaning that only the properties of the system measured at that point in time are what are definite.

- In the simplified version of the hydrogen atom, an electron can would be detected somewhere between the nucleus and the outside edge of the atom.

- However by the edge of the potential well.



▲ Figure 12 Electron standing waves in a potential well.

- The potential varies as the inverse of the square of the distance from the nucleus: $V(r) = \frac{k}{r}$.

- So the electron wave would, the probability of finding an electron inside the nucleus or outside the atom are both zero.

- So the wave amplitude is zero.

- The electron has the highest probability of being found half way between the nuclei.

- Wheeler example

- The electron is most likely to be found at "z" or $|V|^2$ is a maximum.

- There is no chance of finding the electron above "z" or the nucleus.

- The Heisenberg uncertainty principle

- When a quantum is deflected, then its path can only be put in terms of a probability wave.

- This outcome falls in line with the Heisenberg uncertainty principle.

- The Heisenberg uncertainty principle is written as $\Delta x \Delta p \geq \frac{h}{4\pi}$.

- The uncertainty principle puts a limit on how precisely we can know the position (Δx) and momentum (Δp) of an electron or photon.

- If Δx is very small, then Δp will be very large, and vice versa.

- Therefore, it's impossible to know the precise position or momentum of an electron.

- If we measure our electron that is free from our electric field, then we can break it as a sine wave since we know its wavelength perfectly.

- This means we will have its commutation property ($[p_i, p_j] = 0$).
- This means that the uncertainty in the position is infinite, and we expect the electron to be spread out over all of space.
- When an electron is diffused, the uncertainty in position will be given by half of the state.

Virtual particle

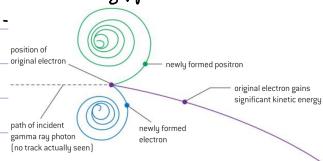
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Pair production and pair annihilation

- Close to an atomic nucleus, where the electric field is very strong, a photon with the right amount of energy can turn into a particle & its antiparticle.
- This process is known as pair production.
- The "right amount" of energy means that the energy of the photon must equal the rest mass of the particle and anti-particle combined.
- This is given by the equation: $E = 2mc^2$
- " m " is the rest mass of the particle. Their value is the same as the antiparticle.
- The reason that they're produced in pairs is to conserve charge, baryon number, lepton number, and strangeness.
- For the pair production of a positron and electron, a gamma ray must have energy equal to 102 MeV.
- They mean photon energy is turned into kinetic energy for the positron and electron.
- Pair production can also occur near an orbiting electron.

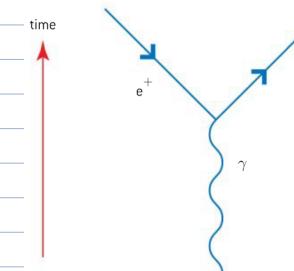
- Although in this case, more energy will be required, as none of the energy will be given to the electron in orbit.

- this is shown in the following figure:



▲ Figure 14 Bubble chamber tracks of electron-positron pair production.

time



▲ Figure 13 Feynman diagram of electron-positron pair production.

- In reality, the gamma ray (photon) doesn't have a "tail", while the produced electron and positron do.
- Even the original electron has a tail.
- The resulting electron gains kinetic energy, meaning that it won't travel back much in a magnetic field.
- It was shown that this type of pair production requires $E_{\text{min}}^2 = 2.048 \text{ GeV}$.

- The equation is: $\gamma + e^- \rightarrow e^- + e^+$

- When a particle meets its antiparticle, they will annihilate.

- This produces a photon with the energy equal to the rest mass of the two particles plus their kinetic energy.

Pair production and the Heisenberg uncertainty principle

- Instead of position and momentum or the Heisenberg uncertainty principle, energy and time are used.

- They're known as conjugate variables.

- The equation in another form $\Delta E \Delta t = \frac{\hbar}{4\pi}$.

- What is interesting is that the threshold energy required for the production of an electron-positron pair can be far from the expected 102 MeV when the photon is near a heavy nucleus.

- If a 102 MeV photon is near a heavy nucleus, then the low energy photon will produce an electron-positron pair.

- If very short time later, the electron and positron collide to produce two photons of 50 MeV.

- This is allowed due to the uncertainty principle.

- During the lifetime of the electron-positron pair, there is an uncertainty regarding the energy.

- If the uncertainty was equal to 1.04 MeV, then we can determine the lifetime of the pair.

$$\Delta t = \frac{\hbar}{4\pi(10^{12} \cdot 1.6 \cdot 10^{-19} \cdot 1.2)} \\ = 3.2 \cdot 10^{-25} \text{ s}$$

- Since the time is so short, we can't detect it occurring.

- If we can't detect it, then quantum mechanics says that there's a chance of it occurring.

Quantum Tunneling

- According to quantum mechanics, a particle's wave function has a finite probability of being everywhere in the universe at the same time.

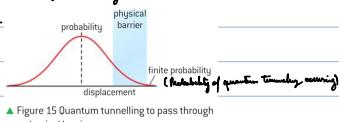
- It may be infinitesimally small, but it's finite nonetheless.

- This means that our electron in the ground state of hydrogen can escape the attraction of the nucleus with 100% chance.

- A particle can essentially "borrow" energy from it surroundings, pass through a barrier, and then return the energy.

- This can be done as long as the time is short.

- This is represented by the following graph:



▲ Figure 15 Quantum tunnelling to pass through a physical barrier.

- Due to quantum tunnelling, two can fuse together two hydrogen at a relatively low temperature.

- To overcome the electrostatic force of repulsion between the two nuclei the hydrogen would have to have (loss) of energy

- This translates to an ionization temp of 10^9 K , which is really in $3 \times 10^9 \text{ K}$.

- Due to the high pressure and small probability of fusion occurring with quantum tunnelling, fusion can occur at lower temperatures.

- Due to the large number of hydrogen nuclei in the sun, a low probability still translates to a lot of fusion being produced.

Wheeler's example

- The wave function refers to Schrödinger wave function [1], where the square of it shows the likelihood of finding an electron in that position.

$$\frac{-2 \cdot 10^{-10}}{r} \cdot 2.3 \cdot 10^{-10} \text{ m}^{-1} = D_p = \frac{e}{4\pi \cdot 8.9 \cdot 10^9} \\ r = \frac{e}{2.3 \cdot 10^{-10} \cdot 2.3 \cdot 10^{-10}} \\ r = 2.0 \cdot 10^{-23} \text{ m}$$

12.2 Nuclear physics

Equations

- Relationship between radius of nucleus and nuclear number: $R = R_0 N^{1/3}$

- Decay equation for number of nuclei at time t : $N = N_0 e^{-\lambda t}$

- Decay equation for activity at time t : $A = A_0 e^{-\lambda t}$

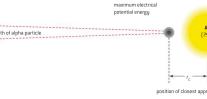
- Range of elastic diffraction first minimum: $\sin \theta = \frac{R}{D}$

Scattering and the nuclear radius

- The method of short approach

- The alpha particles that were back scattered in the gold foil experiment hit the nuclear head on

- Only 1 in 8000 alpha particles were back scattered.



▲ Figure 1 Method of closest approach.

- In the image above, as the alpha particle comes in close to the atom, it will lose kinetic energy and instead gain electric potential energy.

- When the alpha particle has reached the closest point to the nucleus, its kinetic energy is zero and has stopped moving for a second.

- At that point (distance r_0 from the nucleus), equating kinetic energy to electric potential energy gives:

$$-\frac{e^2}{r_0} + \frac{kZz - Ze^2}{r_0^2} = \frac{1}{2} k Z z e^2$$

- k is the Coulomb constant, Z is the proton number of gold, and e is the kinetic energy of the alpha particle.

- The equation is an approximation because the gold nucleus is considered to be a point mass.

- If the alpha particle were to have penetrated the nucleus, then the Coulomb constants would apply.

- This is because the strong nuclear force is dominant within a nucleus.

- Rutherford was able to determine that the radius of an atom could be given by the expression:

$$R = R_0 N^{1/3}$$

- Rutherford was able to come up with this equation by noting that $V \propto N$ (volume is proportional to nuclear number), and the nuclear radius $R \propto N^{1/3}$.

Nuclear density

- If the nucleus is a sphere, the volume can be calculated using the equation:

$$V = \frac{4}{3} \pi R^3 = \frac{4}{3} \pi R_0^3 N^{1/3}$$

- The density of nuclear material will be given by:

$$\rho = \frac{\text{mass}}{V} = \frac{M}{\frac{4}{3} \pi R_0^3 N^{1/3}} = \frac{M}{\frac{4}{3} \pi R_0^3 N^{1/3}}$$

- M is the uniform atomic mass, and N is the total mass of a nucleus of nuclear number N .

- The equation shows that the density of the nuclei is independent of the nucleus.

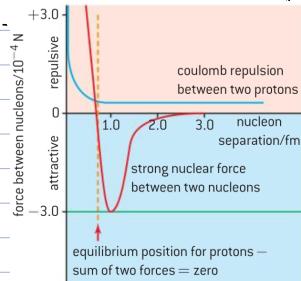
- This is because the density only depends on constants.

$$\rho = \frac{2(1.66 \cdot 10^{-27})}{4 \cdot (1.1 \cdot 10^{-15})^3} = 2.3 \cdot 10^{17} \text{ kg m}^{-3}$$

- In actual, the only agrees with this density is a neutron star.

Predictions from the Rutherford model

- the scattering experiments are limited by energies of the alpha particles emitted by the radioactive sources available to them.
- when the experiment is done with a more energetic alpha particle the Rutherford scattering theory doesn't agree.
- at higher energies the alpha particle can go so close to the nuclei of the atom that the strong force will overcome the electrostatic force of repulsion.
- the direct approach without table or the opposite, one of a nucleus.
- more reliable mass can be obtained with elastic deflection.



▲ Figure 2 Variation of the strong nuclear force and coulomb force with distance.

Electron diffraction

- the electrons are lighter and not heavier, they will not be affected by the strong nuclear force.
- although, they are affected by the strong interaction of the nucleus.
- for light incident on a small nuclear object of diameter D , the angle θ to the first diffraction maximum under with the straight through position ($\theta=0^\circ$) is given by
- $$\sin \theta \approx \frac{D}{\lambda}$$
- the formula is the same for electron diffraction.
- where D is the nuclear diameter, and λ is the de Broglie wavelength of the electron.
- to achieve an appropriate de Broglie wavelength, a 600 MeV electron is used.
- the wavelength of an electron is $\lambda = \frac{h}{p_e}$

practical example

- R of Calcium = 10 fm, 6.5 fm, E = 620 MeV

$$- R = \lambda_0 \cdot q^{\frac{1}{2}} \\ = (1.2 \cdot 10^{-19}) \cdot (40)^{\frac{1}{2}} \\ = 1.1 \cdot 10^{-18} \text{ m}$$

$$- \Delta \theta = \frac{D}{\lambda} \\ = 2.4 \cdot 10^{-18} \text{ rad}$$

$$- \Delta \theta = \frac{D}{\lambda} \\ = \frac{(1.1 \cdot 10^{-18})}{(1.6 \cdot 10^{-18})} \\ = 0.6875 \\ = 16.6^\circ$$

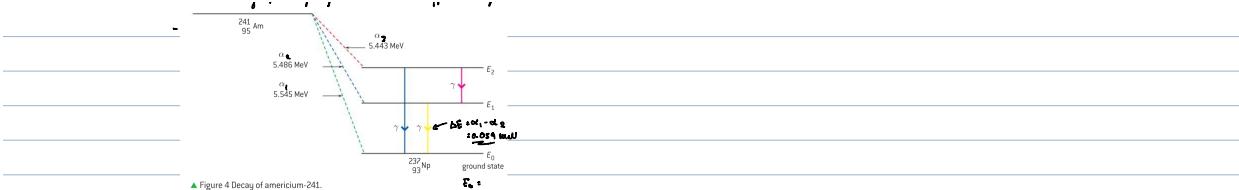
$$- \Delta \theta = \frac{D}{\lambda} \\ = \frac{(1.6 \cdot 10^{-18})(6 \cdot 10^{15})}{(5.0 \cdot 10^{-18})(1.6 \cdot 10^{-18})} \\ = 2.5 \cdot 10^{-18} \text{ rad}$$

Using electrons of higher energy

- when electrons with a greater energy than 600 MeV are used in the scattering experiment, then the collision will no longer be elastic.
- the bombarding electron will have kinetic energy.
- this energy is "converted" into the sum of several mesons when the electron interacts with the nucleus.
- at even higher energies, the electron will penetrate further into the nucleus causing the protons to be scattered.
- this is known as deep inelastic scattering.
- this also provides direct evidence of the quark model of a nucleus.

Energy levels in a nucleus

- much of the evidence of energy levels in a nucleus comes from radioactive decay.
- the emission of β or α or γ radiation usually leaves the daughter nuclei in an excited state.
- to go to the ground state, the nucleus will undergo gamma radiation.
- for example, when a Americium-243 decays to neptunium-237, it can undergo multiple decay paths to reach the ground state.
- the excited decay will be alpha.
- although, the alpha emitted will have different energies.



▲ Figure 4 Decay of americium-241.

- We see in the graph, there are 3 different paths that the particle could decay to the ground state.

- Furthermore, what is also seen is that the energy levels are quantised.

- Providing evidence of nuclear energy levels.

- In addition to which the alpha particle leaves the nucleus in more angular states emitting a gamma ray.

- Alpha particles will be found in the nucleus.

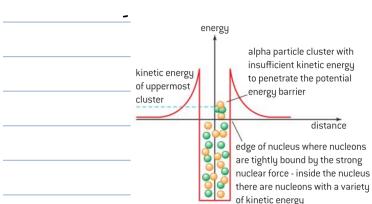
- Leaving two protons and two neutrons will form a cluster.

- The alpha particle will have a kinetic energy which is less than that required to escape the nucleus.

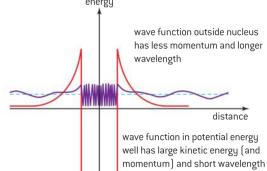
- Although, similarly to how hydrogen atoms, force together in the sun, with significantly less energy than required to overcome the attractive force of attraction, alpha particles can be ejected from the nucleus.

- As long as the time is very short, the uncertainty in the energy can do it.

- What has been determined is that the larger the potential barrier is, and large nucleus, the longer the half-life of the nucleus.



▲ Figure 5 Classical mechanics view of alpha decay.



▲ Figure 6 Quantum tunnelling view of alpha decay.

- When the wave function is at a maximum, then that's the highest chance of tunnelling.

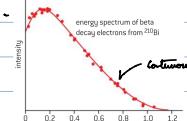
- Meaning that alpha particles with specific energies will be more likely to tunnel.

Negative Beta decay

- An antineutrino will accompany the emitted electron by the β^- .

- Experiments show that β^- decay has a continuous energy spectrum.

- On the other hand, gamma and alpha decay is quantised.



▲ Figure 7 Negative beta-energy spectrum.

- This proves the quantisation of the antineutrinos, because it would be able to measure the energy of they're added together.

The law of radioactive decay

- Radioactive decay is a completely random and unpredictable process.

- What we do know is that the more radioactive nuclei there are, the faster the chance of radioactive decay occurring.

- The decay constant (λ) is defined as: the probability that an individual nucleus will decay in a given time interval.

- The activity (A) is defined as the number of nuclei decaying in a second.

- The units are Bq (Bequerel).

- The activity of N undecayed nuclei is given by the equation:

$$- N = N_0 e^{-\lambda t}$$

- The equation will become: $N = N_0 e^{-\lambda t}$

- This equation can be used to determine the exponential decay of the radioactive particles.

Worked example

$$\begin{aligned} & - 2200 \cdot 2.71 \cdot e^{-\lambda t} \\ & N_t = \frac{2200}{1.35 \cdot 10^{-11}} \\ & N_t = 1.63 \cdot 10^{19} \text{ nuclei} \end{aligned}$$

$$\begin{aligned} & \frac{N_t}{N_0} = \frac{2.71}{1.35 \cdot 10^{-11}} \\ & \frac{N_t}{N_0} = 2.71 \cdot 10^{10} \text{ nuclei} \\ & \frac{\ln \frac{N_t}{N_0}}{t} = 2.71 \cdot 10^{10} \text{ Bq} \\ & = 6.1 \cdot 10^6 \text{ Bq} \end{aligned}$$

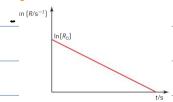
- very short and half-life

- The half-life is defined as the time taken for the number of radioactive nuclei to decrease by half.

$$\text{Therefore, } \tau = \frac{\ln 2}{\lambda}$$

- The equation $n(t) = n_0 e^{-\lambda t}$ can be used as $\frac{n}{n_0} = e^{-\lambda t}$.

- Measuring log half-life



▲ Figure 9 Graph of natural log of the corrected count rate against time.

- When a sample has a very long half-life, such as thorium-232, the following process is used:

- Measure the mass of the sample.

- Calculate the ratio.

- The activity is then obtained by:

$$\text{• } \frac{\text{No. Count ratio} - \text{mass of sample}}{\text{mass of sample}}$$