

Nuclear Physics

Formulas

- Relationship between radius of nucleus and nuclear number: $R = R_0 A^{1/3}$

- Decay equation for number of nuclei at time t : $N = N_0 e^{-\lambda t}$

- Decay equation for activity at time t : $A = \lambda N_0 e^{-\lambda t}$

- Range of electrons diffraction first min: $\sin \theta \approx \frac{\lambda}{D}$

Rutherford scattering and the nuclear radius

- Main reasons of alpha scattering:

- Most of α particles are unaffected

- Some α particles deflected through wide angle

- Some angles turned back in opposite directions

- Interpretations of these results were:

- Most of atom is empty space

- Atom contains small dense regions of electric charge

- These small regions are positively charged.

- Alpha particles which are sent back (backscatter) are alpha particles which have a head-on collision with the gold nucleus.

- Only $\frac{1}{900}$ have a head-on collision.

- As alpha particles move closer to the nucleus its kinetic energy falls and its electrical potential energy increases.

- When the α particle is at its closest to the nucleus, its kinetic energy has fallen to zero, and momentum stopped moving.

- kinetic energy of particle and the electrical potential energy are related by: $E_k = \frac{kZ_0 Z_e}{R_0}$

- $k = \text{proton number} Z$ is proton number of gold (making Z the charge on gold nucleus), and e is alpha particle charge.

- for alpha particles: $R_0 = \frac{kZ^2 e^2}{E_k}$.

- In Rutherford's experiment $E_k = 7.68 \text{ MeV}$ & $Z=79$ meaning R_0 is $2.9 \cdot 10^{-14} \text{ m}$.

- As the volume V of a nucleus must be proportional to the number of its nucleons, $V \propto N^{2/3}$ (N is the nuclear number) and nuclear radius $R \propto N^{1/3}$, giving $R \propto R_0 N^{1/3}$.

- R_0 is Fermi radius with value $1.2 \cdot 10^{-15} \text{ m}$.

Nuclear density

- If the nucleus is spherical its volume would be $\frac{4}{3}\pi R^3 = \frac{4}{3}\pi N A^2 R_0^3$.

- Density of nuclear material will be given by: $\rho = \frac{N_0}{V} = \frac{N_0}{\frac{4}{3}\pi R_0^3} = \frac{3(1.66 \cdot 10^{-27})}{4\pi(1.2 \cdot 10^{-15})^3} = 2.3 \cdot 10^{37} \text{ kg/m}^3$.

- ρ is the uniform atomic mass and N_0 is the total mass of a nucleus of nuclear number N .

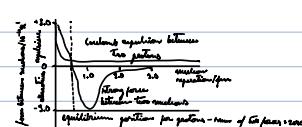
- Density is independent of the number of nucleons in the nucleus.

- The density of a single unified atomic mass is extremely dense, only a neutron star (stellar remnant resulting from gravitational collapse of a massive star), which is made of neutrons.

Reactions from Rutherford scattering

- Using more energetic alpha particles will mean that the strong force will overcome the electrostatic repulsion.

Graphs



Electron Diffraction

- Electrons are lighter (not nuclei) meaning they're affected by the charge distribution & not the strong force.

- High energy electrons have short Debye wavelengths of order 10^{-15} m .

- For light incident on a small circular object of diameter D , the angle θ that the first diffraction minimum makes with the straight-through position (0°) is given by

$$\sin \theta \approx \frac{\lambda}{D}$$

- The intensity of the diffracted beam is seen to be a max in the straight-through position, falling to a min before slightly increasing again.

- The minimum differs from light because it never reaches zero for electrons.

- Given by formula: $\sin \theta \approx \frac{\lambda}{D}$, where D is nuclear diameter & λ is wavelength of electrons.

$$\begin{aligned} \text{Required energy is supposed to be } 663 \text{ eV with wavelength given by: } \lambda &= \frac{3.1 \cdot 10^{-15}}{400 \cdot 10^{-15} \cdot 1.6 \cdot 10^{-19}} \\ &= 3.1 \cdot 10^{-15} \text{ m} \end{aligned}$$

Worked Example

$$R = (1.2 \cdot 10^{-9}) (50)^{\frac{1}{3}} = 2 \cdot \frac{50}{\sqrt[3]{2}}$$

$$R = 4.10 \text{ fm}$$

Radius, diameter $\approx 2 \text{ fm}$

$$\lambda = \frac{6.65 \cdot 10^{-19} (3 \cdot 10^8)}{6.20 \cdot 10^{-16} \cdot 1.6 \cdot 10^{-19}} \quad \theta = \sin^{-1} \left(\frac{3 \cdot 10^8}{6.08 \cdot 10^{-19}} \right)$$

$$\lambda = 2.2 \cdot 10^{-19} \text{ nm} \quad \theta = 19.6^\circ$$

- Assume they're not close enough to the speed of light ($\approx 400 \text{ MeV}$).
Measuring their wavelength is too long.

- Using electrons of higher energies

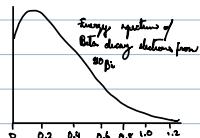
- When electrons of energy greater than 420 MeV are used for the scattering experiment, the electrons are no longer electric electrons (see E_γ).
- The energy & momenta of several muons are emitted from the nucleus.
- At even higher energies the electrons penetrate deeper into the nuclear & scatter off the quarks within protons & neutrons (deep inelastic scattering).

- Energy levels in the nucleus

- Emission of gamma radiation is comparable to the emission of photons by electrons moving between different energy levels (positronium).
- The emission of alpha or beta particles by radioactive nuclei often leaves the daughter nucleus in an excited state.
 - To go back to its ground state the daughter nucleus will emit one or more gamma ray photons.
- For example in americium-241, when it decays it'll emit one alpha particle, being one of a number of possible energies to become neptunium-237.
 - One of the alpha particles will bring the newly formed neptunium-237 to its ground state, while the other will leave it in an excited state.
 - If the nucleus is in an excited state it will emit one or two gamma photons to bring it to its ground state (number of gamma rays depend on which state it is in).
- The energy of the gamma photon can be calculated by using the difference between the energy levels & the ground state.
- ΔE are quantized
- The E_γ of the nucleus in the stars is much lower than the amount required to escape the nucleus
 - Strong force provides E_γ barrier (below electrostatic repulsion) will overcome strong force
 - The higher the E_γ barrier the longer the half-life.
- Directions travelling in culture a radioactive particle passes through a β^- barrier

- Negative Beta decay

- Anti-neutrino accompanies the electron in β^- decay.
- Beta particles have a continuous energy spectrum and aren't discrete single energies like alpha particles & gamma photons
 - β^- energy spectrum



- The antineutrino was discovered as it would violate the law of energy, momentum, and spin angular momentum to be conserved in emission.
- Energy is shared between electrons & antineutrino.

- The law of radioactive decay

- Radioactive decay is a random & unpredictable process
- The decay constant (λ) is the probability that an individual nucleus will decay in a given time interval (of one second/hour/year/etc).
 - Units are: Time⁻¹ (s⁻¹, min⁻¹, h⁻¹, etc.).
- The activity (N) of a sample is the number of nuclei decaying in a second (s^{-1}).
- In the sample of the undecayed nuclei, the activity will be equal to the number of nuclei present times λ .
 - $N = N_0 e^{-\lambda t}$
 - It will always decrease making it negative: $N = -N_0 = \frac{dN}{dt} = -\lambda N_0$.
 - $N = N_0 e^{-\lambda t} \quad \& \quad N = N_0 e^{-\lambda t}$.

- Calculated example

$$\begin{aligned} & 2200 \cdot 2N_0 & - 1.63 \cdot 10^{16} & - 2.7 \cdot 10^{-10} \text{ nm} & - 2.7 \cdot 10^{-10} \cdot 2.26 \cdot 10^{-3} \text{ kg} \\ & N_0 = \frac{2200}{1.95 \cdot 10^{11}} & & \frac{1.63 \cdot 10^{16}}{6.02 \cdot 10^{23}} & \approx 6.1 \cdot 10^{-11} \text{ kg} \\ & N_0 = 1.08 \cdot 10^{14} & & & \end{aligned}$$

- Worked example

$$\begin{aligned} & 10 \cdot 10^{-9} \text{ kg} & t_{\frac{1}{2}} = \frac{0.693}{\lambda} & 6.1 \cdot 10^{-9} \text{ kg} & \lambda: \frac{0.693 \cdot 2.26 \cdot 10^{-3}}{6.1 \cdot 10^{-11}} \approx 0.311 \text{ s} \\ & & = \frac{0.693}{2.1 \cdot 365 \cdot 24 \cdot 3600} & \frac{10 \cdot 10^{-9}}{1.73} = 7.5 \cdot 10^{-9} \text{ nmol} & \lambda: 1.73 \cdot 10^{-9} \text{ s} \\ & & = \frac{0.693 \cdot 10^{-9}}{1.05 \cdot 10^{-11}} & 7.5 \cdot 10^{-9} \cdot 6.02 \cdot 10^{-23} = 4.49 \cdot 10^{-16} & \\ & & = 0.33 \text{ g}^{-1} & \lambda: 10^{-9} \text{ s} & \\ & & & \lambda: (1.05 \cdot 10^{-11}) / (4.49 \cdot 10^{-16}) & \\ & & & = 4.7 \cdot 10^{14} \text{ s} & \end{aligned}$$