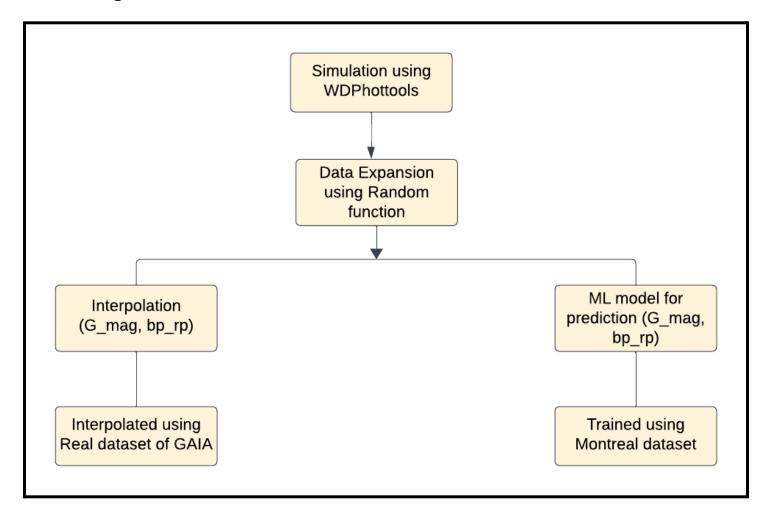
# **Progress on Project**

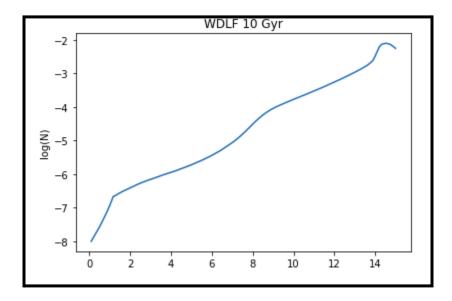
## Flow diagram-



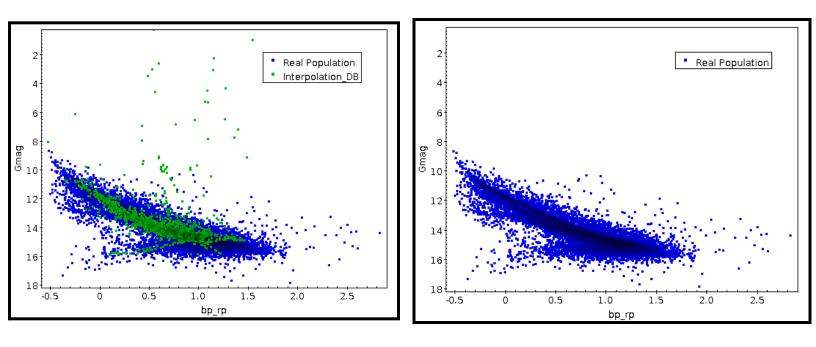
## **DB Population-**

Cooling Model- montreal\_co\_db\_20 IMF Model- C03b MS Model- PARSECz0017 IFMR Model- S09 SFR Model- mode="constant", age=e10

### WDLF-

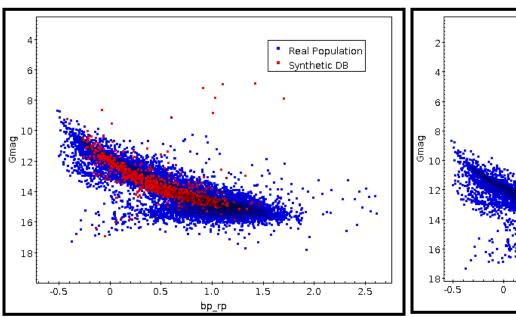


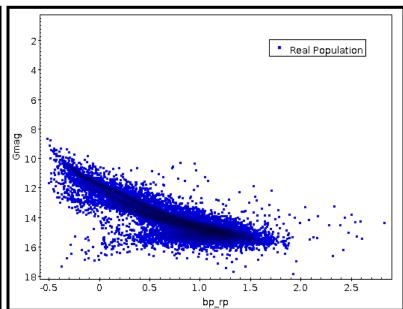
## **Interpolation Result-**



**Inference**- The DB population generated using interpolation does not exactly follow the observed population trend. It faintly follows the DA population trend that is the upper branch.

#### ML model Result-



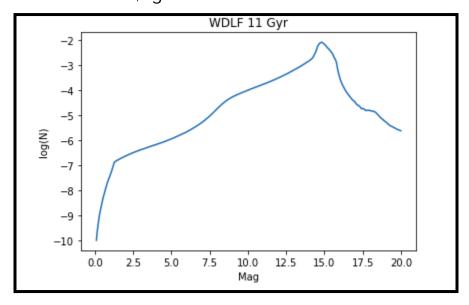


<u>Method</u>- The KNN Model is used in predicting the values of G\_mag, and bp\_rp parameters. The training is performed by the available Montreal dataset.

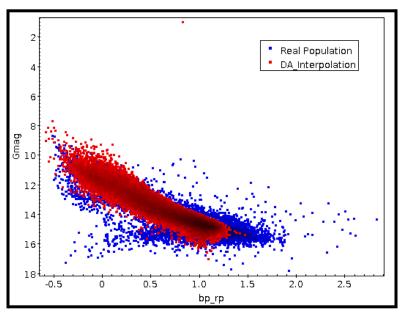
<u>Inference</u>- The ML model result shows that the synthetic DB population follows the observed trend of the DB population.

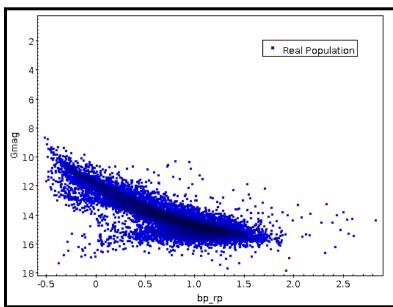
## **DA Population-**

Cooling Model- montreal\_co\_da\_20
IMF Model- Co3b
MS Model- PARSECzoo17
IFMR Model- So9
SFR Model- mode="constant", age=1.1e10



### **Interpolation Result-**

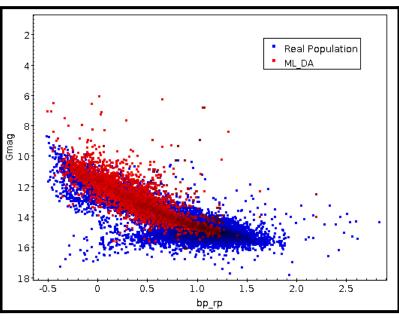


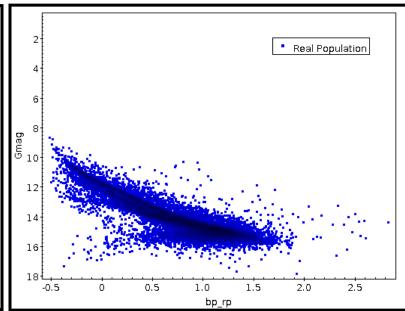


### Inference-

The interpolated synthetic dataset of DA population follows the trend of observed population. Though the branch is not clearly visible but most of the data points lie in the zone.

### **ML Result-**





### Inference-

The ML generated synthetic data of DA stars is more consistent with the observed data as the separate branch is clear in this case.