

# **Magnetic Mirror Effect in Magnetron Plasma:**

## Modeling of Plasma Parameters

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## 1.1 Density Function

$f(x, y, z, v_x, v_y, v_z, t)$  :

number of particles at position  $(x, y, z)$  at time  $t$  with velocities between  $v_x$  and  $v_x + dv_x$ ,  $v_y$  and  $v_y + dv_y$ ,  $v_z$  and  $v_z + dv_z$

Notation:

$f(\mathbf{r}, \mathbf{v}, t) := f(x, y, z, v_x, v_y, v_z, t)$

$$f(\mathbf{r}, t) = \int_{all\ v_x} dv_x \int_{all\ v_y} dv_y \int_{all\ v_z} dv_z f(\mathbf{r}, \mathbf{v}, t)$$

Also written as

$$\int_{all\ \mathbf{v}} d^3\mathbf{v} f(\mathbf{r}, \mathbf{v}, t) \quad \text{or} \quad \int_{all\ \mathbf{v}} d\mathbf{v} f(\mathbf{r}, \mathbf{v}, t)$$

## Density function

Normalized density function denoted  $\hat{f}(\mathbf{r}, \mathbf{v}, t)$

$$\int_{all \mathbf{v}} d\mathbf{v} f(\mathbf{r}, \mathbf{v}, t) = 1$$

Average velocity

$$\bar{\mathbf{v}} = \int_{all \mathbf{v}} d\mathbf{v} \mathbf{v} f(\mathbf{r}, \mathbf{v}, t)$$

Other parameters

Root Mean Square velocity  $v_{rms}$

average absolute velocity  $|\bar{\mathbf{v}}|$

average velocity in  $z$  direction  $\bar{v}_z$

## 1.2 Maxwell Boltzmann Distribution

Density Function

$$\widehat{f_M}(\mathbf{r}, \mathbf{v}, t) = \left(\frac{m}{2\pi KT}\right)^{\frac{3}{2}} \exp\left(-\frac{v^2}{v_{th}^2}\right) \quad (1)$$

$$v_{th}^2 = \frac{2KT}{m}$$

Features of the Maxwell Boltzmann Distribution (Maxwellian)

$$v_{rms} = \sqrt{\frac{3KT}{m}}, |\bar{v}| = 2\sqrt{\frac{2KT}{\pi m}}, |\bar{v}_z| = \sqrt{\frac{2KT}{\pi m}}, \bar{v}_z = 0$$

## 2. Equation of Motion

Derivative in 1 dimension

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} \frac{dx}{dt} + \frac{\partial f}{\partial v_x} \frac{dv_x}{dt}$$

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} v_x + \frac{\partial f}{\partial v_x} a_x$$

If there are no external disturbances and interactions

$$\frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} v_x + \frac{\partial f}{\partial v_x} a_x = 0$$

which simply means

$$\frac{df}{dt} = 0 \tag{2}$$

## Equation of Motion

Derivative in 3 dimensions

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} \frac{dx}{dt} + \frac{\partial f}{\partial v_x} \frac{dv_x}{dt} + \frac{\partial f}{\partial y} \frac{dy}{dt} + \frac{\partial f}{\partial v_y} \frac{dv_y}{dt} + \frac{\partial f}{\partial z} \frac{dz}{dt} + \frac{\partial f}{\partial v_z} \frac{dv_z}{dt}$$

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} v_x + \frac{\partial f}{\partial v_x} a_x + \frac{\partial f}{\partial y} v_y + \frac{\partial f}{\partial v_y} a_y + \frac{\partial f}{\partial z} v_z + \frac{\partial f}{\partial v_z} a_z$$

which is often written as

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \mathbf{a} \cdot \partial_{\mathbf{v}} f$$

If there are no external disturbances and interactions

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \mathbf{a} \cdot \partial_{\mathbf{v}} f = 0$$

which is again just equation (2)

## 2.2 Vlasov and Boltzmann equations

**Vlasov equation:**

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \frac{q}{m} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \partial_{\mathbf{v}} f = 0 \quad (3)$$

Lorentz force is substituted for the acceleration term in the 3 dimensional equation of motion without external disturbances and interactions

**Boltzmann equation:**

$$\frac{df}{dt} = \left( \frac{\partial f}{\partial t} \right)_c \quad (4)$$

Collision terms are included to describe the change in the density function

Examples: Coulomb force, Collision with neutral particles

## 3.1 Magnetic Mirror - Conserved quantities

Magnetic Moment : Adiabatic Invariant

$$\mu = \frac{mv_{\perp}^2}{2B_z}$$

Kinetic energy

$$KE = \frac{1}{2}mv^2$$

Writing  $v_{\perp} = v \sin\theta$

$$\frac{\mu}{KE} = \frac{\sin^2\theta}{B_z}$$

is conserved



## Magnetic Mirror

I am still a bit confused on how to use this for a dynamic velocity distribution to get an expression for reflected and lost flux.

**TO DO 1**

**So things after here in Magnetic mirror are not entirely clear.**

$$\frac{\sin^2 \theta}{B_z} = \frac{\sin^2 \theta_0}{B_{z0}}$$

a

$$\sin^2 \theta = \frac{B_z}{B_{z0}} \sin^2 \theta_0$$

Since

$$\sin^2 \theta \leq 1 \qquad \frac{B_z}{B_{z0}} \sin^2 \theta_0 \leq 1$$

The reflected fraction is given by

$$\begin{aligned}
 f_{loss} &= \frac{\int_{loss\ cone} f(\mathbf{r}, \mathbf{v}, t) \, d\mathbf{v}}{\int_{all\ \mathbf{v}} f(\mathbf{r}, \mathbf{v}, t) \, d\mathbf{v}} \\
 &= \frac{\int_0^{2\pi} d\phi \left[ \int_0^{\theta_0} \sin\theta d\theta + \int_{\pi-\theta_0}^{\pi} \sin\theta d\theta \right] \int_{all\mathbf{v}} \frac{f(\mathbf{r}, \mathbf{v}, t)}{4\pi v^2} \, dv}{\int_0^{2\pi} d\phi \int_0^{\pi} \sin\theta d\theta \int_{all\mathbf{v}} \frac{f(\mathbf{r}, \mathbf{v}, t)}{4\pi v^2} \, dv} \\
 &= 1 - \cos\theta_0
 \end{aligned}$$

## 4.1 Approximations to Maxwellian

The uniform distribution describes a collection of particles where a particle chosen at random is equally likely to have any velocity in the range.

$$f_0 = \begin{cases} \frac{1}{2v_a} & -v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$

which is already normalized. The Dirac delta-like distribution would describe all the particles having the same velocity.

$$f_\delta = \begin{cases} 1 & v = v_a \\ 0 & \text{else} \end{cases}$$

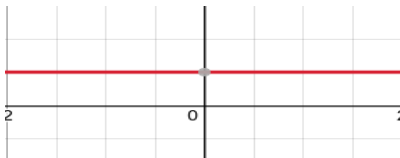


Figure 1:  $f_0$

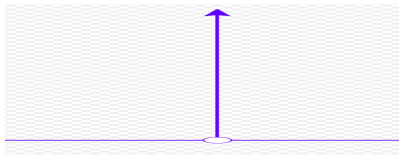


Figure 2:  $f_\delta$

$f_0$  generated using desmos graphing calculator.  $f_\delta$  snipped from the image in wikipedia for the Dirac delta function.

- Uniform and delta distributions are not very realistic in a system of large number of particles.
- Parabolic functions in a given range are better and yet simple to work with.

## 4.2 Parabolic density functions as approximation to Maxwellian

$$f_1 = \begin{cases} 1 - \frac{v^2}{2v_a^2} & -v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$
$$f_2 = \begin{cases} 1 + \frac{v^2}{2v_a^2} & -v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$

These functions can be normalized as

$$\int_{v_x=-v_a}^{v_x=v_a} dv_x \int_{v_y=-v_a}^{v_y=v_a} dv_y \int_{v_z=-v_a}^{v_z=v_a} dv_z \, c_0 \left( 1 - \frac{v_x^2 + v_y^2 + v_z^2}{2v_a^2} \right) = 1$$

$$\int_{v_x=-v_a}^{v_x=v_a} dv_x \int_{v_y=-v_a}^{v_y=v_a} dv_y \int_{v_z=-v_a}^{v_z=v_a} dv_z \, c_0 \left( 1 + \frac{v_x^2 + v_y^2 + v_z^2}{2v_a^2} \right) = 1$$

## Parabolic density functions - Normalized

$$\hat{f}_1 = \begin{cases} \frac{1}{4v_a^3} \left(1 - \frac{v^2}{2v_a^2}\right) & -v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$
$$\hat{f}_2 = \begin{cases} \frac{1}{12v_a^3} \left(1 + \frac{v^2}{2v_a^2}\right) & v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$

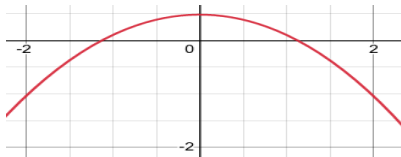


Figure 3:  $\hat{f}_1$

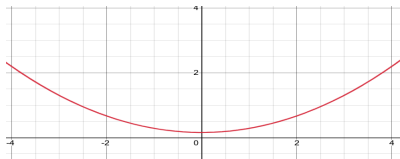


Figure 4:  $\hat{f}_2$

### 4.3 $\hat{f}_1$ , $\hat{f}_2$ and $\widehat{f_M}$ in the collisionless Vlasov equation

Plugging in  $\hat{f}_1$  ,  $\hat{f}_2$  and  $\widehat{f_M}$  in the collisionless Vlasov equation

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \frac{q}{m} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \partial_{\mathbf{v}} f = 0$$

For all three of the density functions

$$\frac{\partial f}{\partial t} = 0, \quad \frac{\partial f}{\partial x} = 0, \quad \frac{\partial f}{\partial y} = 0, \quad \frac{\partial f}{\partial z} = 0$$

as all three have no explicit dependence on position or time. The Vlasov equation then becomes

$$(\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \partial_{\mathbf{v}} f = 0$$

which is

$$(E_x + v_y B_z - v_z B_y) \frac{\partial f}{\partial v_x} + (E_y + v_z B_x - v_x B_z) \frac{\partial f}{\partial v_y} + \\ (E_z + v_x B_y - v_y B_x) \frac{\partial f}{\partial v_z} = 0$$

## $\widehat{f_M}$ in the collisionless Vlasov equation

For the Maxwellian,

$$\widehat{f_M} := \hat{f}(\mathbf{r}, \mathbf{v}, t) = \left( \frac{m}{2\pi KT} \right)^{\frac{3}{2}} \exp \left( -\frac{v^2}{v_{th}^2} \right)$$

$$\frac{\partial f_M}{\partial v_x} = \left( \frac{m}{2\pi KT} \right)^{\frac{3}{2}} \exp \left( \frac{-v^2}{v_{th}^2} \right) \left( \frac{-2v_x}{v_{th}^2} \right)$$

$$\frac{\partial f_M}{\partial v_y} = \left( \frac{m}{2\pi KT} \right)^{\frac{3}{2}} \exp \left( \frac{-v^2}{v_{th}^2} \right) \left( \frac{-2v_y}{v_{th}^2} \right)$$

$$\frac{\partial f_M}{\partial v_z} = \left( \frac{m}{2\pi KT} \right)^{\frac{3}{2}} \exp \left( \frac{-v^2}{v_{th}^2} \right) \left( \frac{-2v_z}{v_{th}^2} \right)$$

plugging in, the equation becomes

$$\begin{aligned} & \left( \frac{m}{2\pi KT} \right)^{\frac{3}{2}} \exp \left( \frac{-v^2}{v_{th}^2} \right) \left( \frac{-2}{v_{th}^2} \right) [v_x (E_x + v_y B_z - v_z B_y) + \\ & v_y (E_y + v_z B_x - v_x B_z) + v_z (E_z + v_x B_y - v_y B_x)] = 0 \end{aligned}$$



## $\widehat{f}_M$ in the collisionless Vlasov equation

which gives

$$v_x E_x + v_y E_y + v_z E_z = 0$$

which can also be written as

$$\mathbf{v} \cdot \mathbf{E} = 0 \quad (5)$$

which says that the particles move perpendicular to the electric field which is characteristic of the Lorentz force.

For the parabolic functions,

$$\hat{f}_1 = \begin{cases} \frac{1}{4v_a^3} \left(1 - \frac{v^2}{2v_a^2}\right) & -v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$
$$\hat{f}_2 = \begin{cases} \frac{1}{12v_a^3} \left(1 + \frac{v^2}{2v_a^2}\right) & v_a \leq v \leq v_a \\ 0 & \text{else} \end{cases}$$

## $\hat{f}_1$ and $\hat{f}_2$ in the collisionless Vlasov equation

$$\frac{\partial \hat{f}_1}{\partial v_x} = \begin{cases} \frac{1}{4v_a^3} \left( \frac{-2v_x}{2v_a^2} \right) & -v_a < v < v_a \\ \text{undefined} & v = -v_a, v = v_a \\ 0 & \text{else} \end{cases}$$

$$\frac{\partial \hat{f}_1}{\partial v_y} = \begin{cases} \frac{1}{4v_a^3} \left( \frac{-2v_y}{2v_a^2} \right) & -v_a < v < v_a \\ \text{undefined} & v = -v_a, v = v_a \\ 0 & \text{else} \end{cases}$$

$$\frac{\partial \hat{f}_1}{\partial v_z} = \begin{cases} \frac{1}{4v_a^3} \left( \frac{-2v_z}{2v_a^2} \right) & -v_a < v < v_a \\ \text{undefined} & v = -v_a, v = v_a \\ 0 & \text{else} \end{cases}$$

$$\frac{\partial \hat{f}_2}{\partial v_x} = \begin{cases} \frac{1}{12v_a^3} \left( \frac{2v_x}{2v_a^2} \right) & -v_a < v < v_a \\ \text{undefined} & v = -v_a, v = v_a \\ 0 & \text{else} \end{cases}$$

$$\frac{\partial \hat{f}_2}{\partial v_y} = \begin{cases} \frac{1}{12v_a^3} \left( \frac{2v_y}{2v_a^2} \right) & -v_a < v < v_a \\ \text{undefined} & v = -v_a, v = v_a \\ 0 & \text{else} \end{cases}$$

$$\frac{\partial \hat{f}_2}{\partial v_z} = \begin{cases} \frac{1}{12v_a^3} \left( \frac{2v_z}{2v_a^2} \right) & -v_a < v < v_a \\ \text{undefined} & v = -v_a, v = v_a \\ 0 & \text{else} \end{cases}$$

In the range  $v \in \mathbb{R}^n [-v_a, v_a]$  the equation becomes  $0 = 0$  which is trivial. The equation becomes undefined for  $v = v_a$  and  $v = -v_a$  but we can ignore that for now. In the interesting range of  $v \in (-v_a, v_a)$ , the equation becomes

$$\frac{1}{4v_a^3} \left( \frac{-2}{2v_a^2} \right) [v_x (E_x + v_y B_z - v_z B_y) + v_y (E_y + v_z B_x - v_x B_z) + v_z (E_z + v_x B_y - v_y B_x)] = 0$$

for  $\hat{f}_1$  and

$$\frac{1}{12v_a^3} \left( \frac{2}{2v_a^2} \right) [v_x (E_x + v_y B_z - v_z B_y) + v_y (E_y + v_z B_x - v_x B_z) + v_z (E_z + v_x B_y - v_y B_x)] = 0$$

for  $\hat{f}_2$  which both give

$$v_x E_x + v_y E_y + v_z E_z = 0 \quad \text{or} \quad \mathbf{v} \cdot \mathbf{E} = 0$$

$\hat{f}_1$  and  $\hat{f}_2$  behave similar to  $\hat{f}_M$  when plugged into the Vlasov equation.

## 4.4 Plugging into the expression for Magnetic Mirror

Since the quantity  $\frac{\sin^2 \theta}{B_z}$  is conserved in a magnetic mirror, it is a good exercise to calculate the average value or expectation of

$$\sin^2 \theta, \langle \sin^2 \theta \rangle = \left\langle \frac{v_{\perp}^2}{v^2} \right\rangle = \left\langle \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right\rangle$$

For the Maxwellian  $\widehat{f_M}$ ,

$$\left\langle \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right\rangle = \left( \frac{m}{2\pi K T} \right)^{\frac{3}{2}}$$

$$\int_{v_x=-\infty}^{v_x=\infty} dv_x \int_{v_y=-\infty}^{v_y=\infty} dv_y \int_{v_z=-\infty}^{v_z=\infty} dv_z \left( \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right) \exp \left( -\frac{v_x^2}{v_{th}^2} \right) \\ \exp \left( -\frac{v_y^2}{v_{th}^2} \right) \exp \left( -\frac{v_z^2}{v_{th}^2} \right)$$

for  $\hat{f}_1$

$$\left\langle \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right\rangle = \frac{1}{4v_a^3}$$
$$\int_{v_x=-v_a}^{v_x=v_a} dv_x \int_{v_y=-v_a}^{v_y=v_a} dv_y \int_{v_z=-v_a}^{v_z=v_a} dv_z \left( \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right)$$
$$\left( 1 - \frac{v_x^2}{2v_a^2} - \frac{v_y^2}{2v_a^2} - \frac{v_z^2}{2v_a^2} \right)$$

for  $\hat{f}_2$

$$\left\langle \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right\rangle = \frac{1}{12v_a^3}$$
$$\int_{v_x=-v_a}^{v_x=v_a} dv_x \int_{v_y=-v_a}^{v_y=v_a} dv_y \int_{v_z=-v_a}^{v_z=v_a} dv_z \left( \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right)$$
$$\left( 1 + \frac{v_x^2}{2v_a^2} + \frac{v_y^2}{2v_a^2} + \frac{v_z^2}{2v_a^2} \right)$$

Integrals are **Very Difficult!** to evaluate because of the denominators.

So an approximation  $\left\langle \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right\rangle = \frac{\langle v_x^2 + v_y^2 \rangle}{\langle v_x^2 + v_y^2 + v_z^2 \rangle} + c_0$  can be done where  $c_0$  is an error term.

For  $\hat{f}_M$ ,  $\langle v_x^2 + v_y^2 \rangle = \frac{2KT}{m}$  and  $\langle v_x^2 + v_y^2 + v_z^2 \rangle = \frac{3KT}{m}$  so

$$\langle \sin^2 \theta \rangle = \left\langle \frac{v_{\perp}^2}{v^2} \right\rangle = \left\langle \frac{v_x^2 + v_y^2}{v_x^2 + v_y^2 + v_z^2} \right\rangle = \frac{\langle v_x^2 + v_y^2 \rangle}{\langle v_x^2 + v_y^2 + v_z^2 \rangle} + c_0 = \frac{2}{3} + c_0$$

For  $\hat{f}_1$ ,  $\langle v_x^2 + v_y^2 \rangle = \frac{22}{45} v_a^2$  and  $\langle v_x^2 + v_y^2 + v_z^2 \rangle = \frac{11}{15} v_a^2$  so

$$\langle \sin^2 \theta \rangle = \frac{2}{3} + c_0 \text{ and}$$

For  $\hat{f}_2$ ,  $\langle v_x^2 + v_y^2 \rangle = \frac{98}{135} v_a^2$  and  $\langle v_x^2 + v_y^2 + v_z^2 \rangle = \frac{49}{45} v_a^2$  so

$$\langle \sin^2 \theta \rangle = \frac{2}{3} + c_0$$

- $\hat{f}_1$  and  $\hat{f}_2$  behave similar to  $\widehat{f_M}$  when plugged into the Vlasov equation
- $\hat{f}_1$  and  $\hat{f}_2$  behave similar to  $\widehat{f_M}$  when plugged into the expression for  $\langle \sin^2 \theta \rangle$
- **Hence!**  $\hat{f}_1$  and  $\hat{f}_2$  are nice distribution functions to work with as an approximation to  $\widehat{f_M}$ .






## 5. Vlasov - Poisson

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## 6. Particle in a cell

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## References

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