

1. The following is the outcome of using a tabula find on a particular page that is giving me trouble while trying to get the data from the bottom table. Notice that it includes junk paragraphs as part of the table.

Table 1 (continued)																
Meteorite	Cr	Co	Ni	Cu	Ga	Ge	As	Sb	W	Re	Ir	Pt	Au	P	S	G
Norrmundsdorf	23	4.80	59.9	130	56.7	172	7.74	87	1.08	-49	0.077	8.0	0.907	3.0		
Judith	28	4.78	58.9	123	57.1	163	8.36		0.92	-30	0.045	7.1	0.970	2.6		
Norrey	26	4.51	57.2	117	56.0	160	8.74		0.89	-20	0.029	4.9	0.996			
Trachyte Flat	38	4.89	57.8	112	55.5	132	9.84		0.82	-40	0.021	5.2	1.030	5.0		
Sikhote-Alin	31	4.79	60.3	117	53.7	158	9.40		0.74	-30	0.024	5.5	1.040			
Elphinstone	22	4.36	60.4	119	55.8	157	9.71	86	0.74	-100	0.026		1.065	4.6		
Ainsworth	32	4.84	61.4	115	54.0	142	9.25		0.64	-1.5	0.022	4.1	1.068	7.6		
Four Zgaid	22	4.97	58.1	113	55.5	153	9.91		0.72	-40	0.021	5.3	1.078			
Derrick Peak	28	4.76	63.6	113	52.9	135	9.31	94	0.56	-20	0.018	4.1	1.084			
Lake Murray	26	5.00	60.3	106	52.7	141	10.2		0.56	-25	0.015	2.7	1.087	5.0		
NWA320	15	5.11	52.1	118	51.6	130	11.1		0.54	-33	0.017	3.7	1.089			
Silver Bell	15	5.26	58.9	98	45.7	111	12.3		0.36	-30	0.012	1.9	1.252	8.0		
Summit	17	4.84	65.5	116	50.4	115	11.4		0.42	-30	0.014	1.9	1.267			
Santa Luzia	15	5.05	60.4	107	46.2	110	11.7	100	0.33	0.714	0.013	2.9	1.290	9.0		
Sao Julian	16	5.19	57.8	94	45.8	106	12.5	79	0.40	0.765	0.012	2.3	1.350	9.6		

Concentrations are $\mu\text{g/g}$ except P, Co, and Ni (mg/g), and Sb and Re (ng/g). P data are from Buchwald (1975). Re data <50 ng/g (and not upper limits) are from Cook et al. (2004). North Chile data are for the Filomena specimen.

the Au:Fe value is too high because of sample treatment, we feel that 4000 represents a more accurate estimate of the IIAB Ir range. Note that the ranges in Os and Re are still larger, ~10,000 and ~6000, respectively (Cook et al., 2004).

We used modal integration to determine the S and P contents of several large sections of IIAB irons. These are tabulated together with data extracted from Buchwald (1975) in Table 2. With the exception of Carver, all of our data are based on photographic images. There is an important difference between our P values (bold in Table 2) and those of Buchwald; our values reflect only massive schreibersite whereas Buchwald reported bulk P values that include the small schreibersite inclusions called rhabdites as well as P dissolved in the metal. Our Carver specimen has suffered shock shear and FeS melting, with the result that FeS boundaries are frequently ragged and less sharp than optimum for integration.

3. Results: element-Au trends in group IIAB

Scatter diagrams for 12 elements plotted against Au are shown in Figs. 1 and 2. In Fig. 1 we show data for six elements that are commonly used for taxonomy. As already well known, the volatile elements Ga (Fig. 1a) and Ge (Fig. 1b) have high concentrations in group IIAB; for Au <0.9 $\mu\text{g/g}$ there is no resolvable slope in either distribution, an observation potentially useful for assessing the initial nonmetal content of the magma. High-S contents are expected to yield D_{Ga} and D_{Ge} values $\gg 1$ (e.g., Chabot, 2004).

In all iron meteorite groups the As and Au concentrations are closely related; log As-log Au diagrams show slopes slightly greater than unity and very limited scatter indicating small analytical and sampling errors (Fig. 1c). Our sparse RNAA Sb data show that the Sb-Au trend (Fig. 1d) is similar to the As-Au trend, implying near-identical

Table 2						
Meteorite	Area (cm^2)	P (mg/g)	S (mg/g)	Ir ($\mu\text{g/g}$)	G	Trap fr.
Ainsworth	120 ^a	7.0	1.4	0.024	48	28
Braunau	~200 ^a	2.4	0.8	11.3	15	8
Carver	468		1.7	11.3	14	3
Coahuila	720		0.85	16.1	11	4
Derrick Peak	420	3.5 ^b	2.6	0.018	46	30
Guadalupe y Calvo	860		1.1	42.6	2	13
Hex River Mtns.	1680		1.8	4.13	28	2
Keen Mountain	225	2.3	3.0	12.1	14	7
Mount Joy	4700		1.3	0.423	39	7
North Chile	1600	3.0	1.0	3.43	28	3
Santa Luzia (Rio)	1800	6.6 ^c	12.3	0.013	48	70
Santa Luzia-RSC	75	6.1 ^c	4.8	0.013	48	70
Sikhote-Alin	>300	4.6	2.8	0.024	48	28

The Ir concentration is given as a rough measure of the degree of crystallization G; the penultimate column gives a better estimate based on our fits to all Ir-Au and Ir-As data. The final column gives the trapped fraction based on the Ir-Au and Ir-As fits. See text for details.

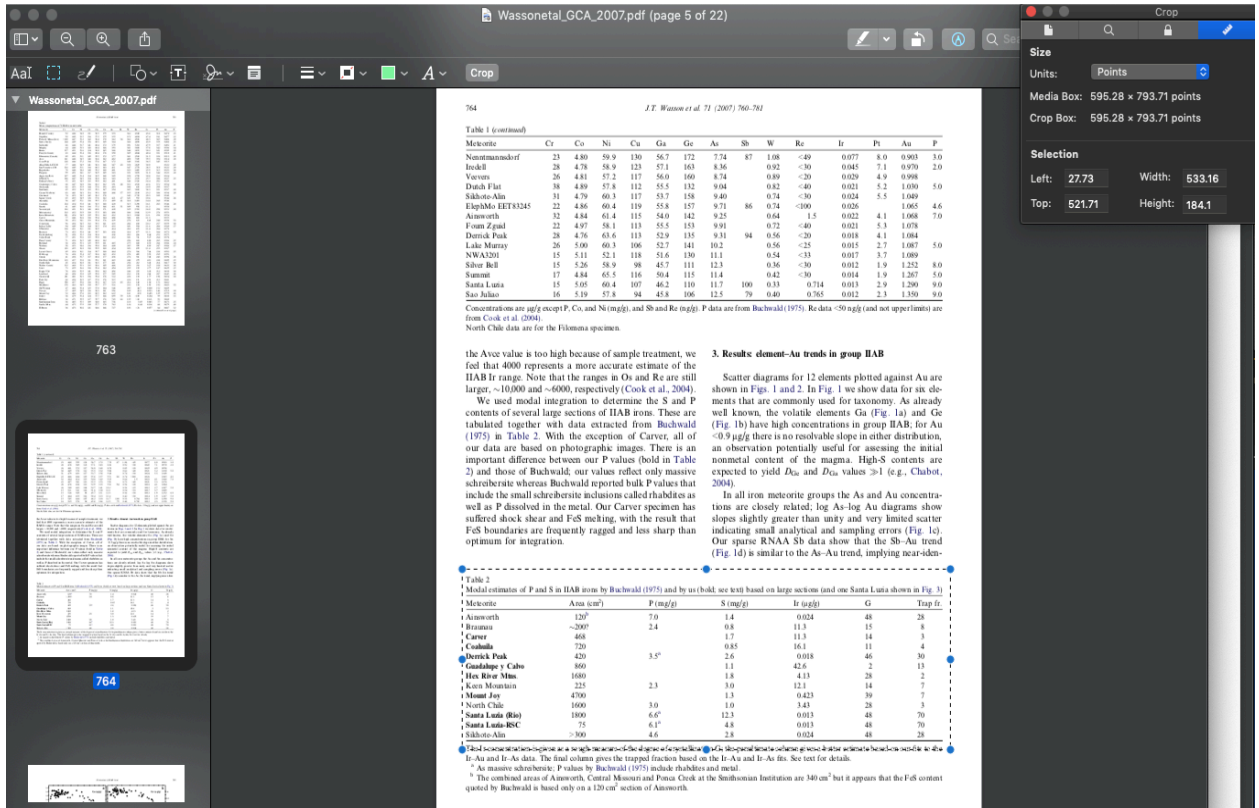
^a As massive schreibersite; P values by Buchwald (1975) include rhabdites and metal.

^b The combined areas of Ainsworth, Central Missouri and Ponca Creek at the Smithsonian Institution are 540 cm^2 but it appears that the FeS content quoted by Buchwald is based only on a 120 cm^2 section of Ainsworth.

2. The following is a snapshot of the data gotten from the table when using tabula in this way, with no coordinate intervention. Notice all the data combined into one cell.

IIAB Ir range. Note that the ranges in Os and Re are still larger, 10,000 and 6000, respectively (Cook et al., 2004), shown in Figs. 1 and 2. In Fig. 1 we show data for six elements that are commonly used for taxonomy. As already well known, the volatile elements Ga (Fig. 1a) and Ge (Fig. 1b) have high concentrations in group IIAB; for Au <0.9 $\mu\text{g/g}$ there is no resolvable slope in either distribution, an observation potentially useful for assessing the initial nonmetal content of the magma. High-S contents are expected to yield D_{Ga} and D_{Ge} values $\gg 1$ (e.g., Chabot, 2004).						
In all iron meteorite groups the As and Au concentrations are closely related; log As-log Au diagrams show slopes slightly greater than unity and very limited scatter indicating small analytical and sampling errors (Fig. 1c). Our sparse RNAA Sb data show that the Sb-Au trend (Fig. 1d) is similar to the As-Au trend, implying near-identical						
Table 2						
Modal estimates of P and S in IIAB irons by Buchwald (1975) and by us (bold; see text) based on large sections (and one Santa Luzia shown in Fig. 3)						
Meteorite Area (cm^2) P (mg/g) S (mg/g) Ir ($\mu\text{g/g}$) G Trap fr.						
Ainsworth	120b	7.0	1.4	0.024	48	28
Braunau	200 ^a	2.4	0.8	11.3	15	8
Carver	468		1.7	11.3	14	3
Coahuila	720		0.85	16.1	11	4
Derrick Peak	420	3.5 ^b	2.6	0.018	46	30
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Sikhote-Alin	>300	4.6	2.8	0.024	48	28

- The following is a snapshot of me manually getting the coordinates from the pdf page in question, using the Apple Preview tool.



- The following is a snippet of my code, where I feed the coordinates of the table into Tabula.

```
print(sys.argv[1])

pdf_name = "pdfs/" + str(sys.argv[1])
pdf_name = "pdfs/Wassonetal_GCA_2007.pdf"
pagenum = int(sys.argv[2]) - 1
pagenum = 5

page_in = open(pdf_name, 'rb')
page_reader = PyPDF2.PdfFileReader(page_in)

page_in.close()

tables_rec_from_page = read_pdf(pdf_name, output_format="dataframe", encoding="utf-8", multiple_tables=True,
                                pages=pagenum, silent=True, area=[521.71, 27.73, 705, 560])

print(tables_rec_from_page)
```

- Lastly, the following is a snapshot of the data gotten with my code.

```
[
  0      0      1      2      3      4      5      6
0  Meteorite Area (cm2)  P (mg/g) NaN S (mg/g) Ir (lg/g) G Trap fr.
1  Ainsworth 120b      7.0 NaN  1.4  0.024 48 28
2  Braunau 2007      2.4 NaN  0.8  11.3 15 8
3  Carver 468      NaN NaN  1.7  11.3 14 3
4  Coahuila 720      NaN NaN  0.85 16.1 11 4
5  aDerrick Peak 420      3.5 NaN  2.6  0.018 46 30
6  Guadalupe y Calvo 860      NaN NaN  1.1 42.6 2 13
7  Hex River Mtns. 1680      NaN NaN  1.8  4.13 28 2
8  Keen Mountain 225      2.3 NaN  3.0 12.1 14 7
9  Mount Joy 4700      NaN NaN  1.3  0.423 39 7
10 North Chile 1600      3.0 NaN  1.0  3.43 28 3
11 Santa Luzia (Rio) 1800      6.6a NaN 12.3 0.013 48 70
12 aSanta Luzia-RSC 75      6.1 NaN  4.8  0.013 48 70
13 Sikhote-Alin >300      4.6 NaN  2.8  0.024 48 28]
```