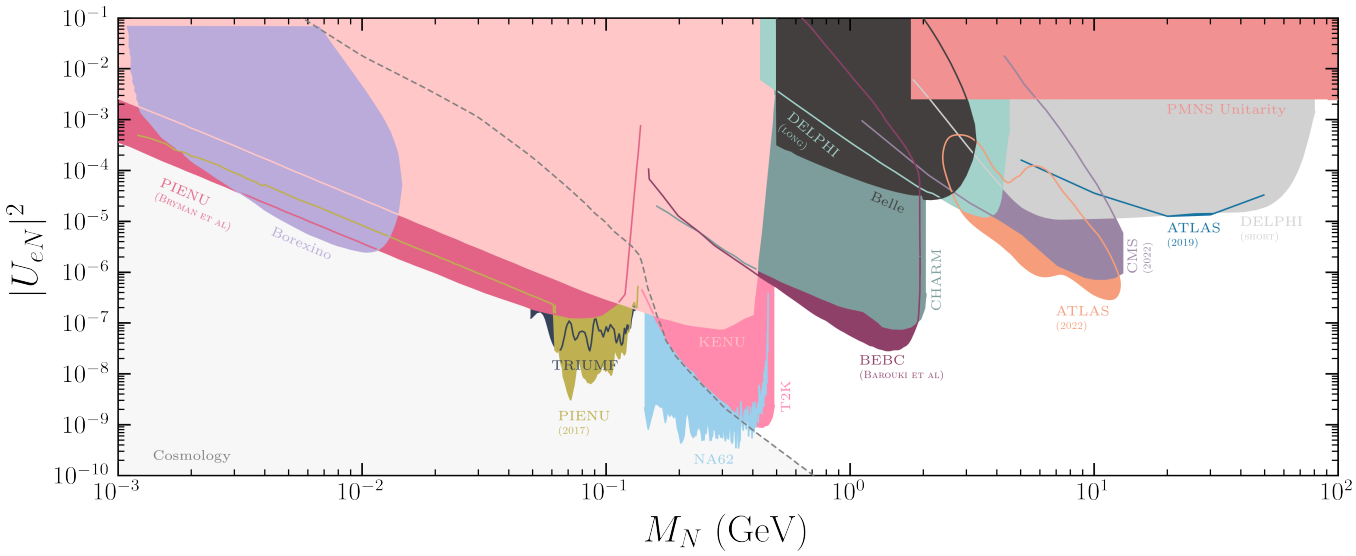


FIG. 1. Constraints on $|U_{eN}|^2$ as a function of the HNL mass $m_N$. Limits shown: ATLAS (2019) [1], ATLAS (2022) [2], BEBC(Barouki et al) [3], Belle [4], Borexino [5], CHARM [6], CMS (2022) [7], Cosmology [8], DELPHI (long) [9], DELPHI (short) [9], KENU [10], NA62 [11], PIENU (2017) [12], PIENU (Bryman et al) [10], PMNS Unitarity [**?** ], T2K [13], TRIUMF [14].

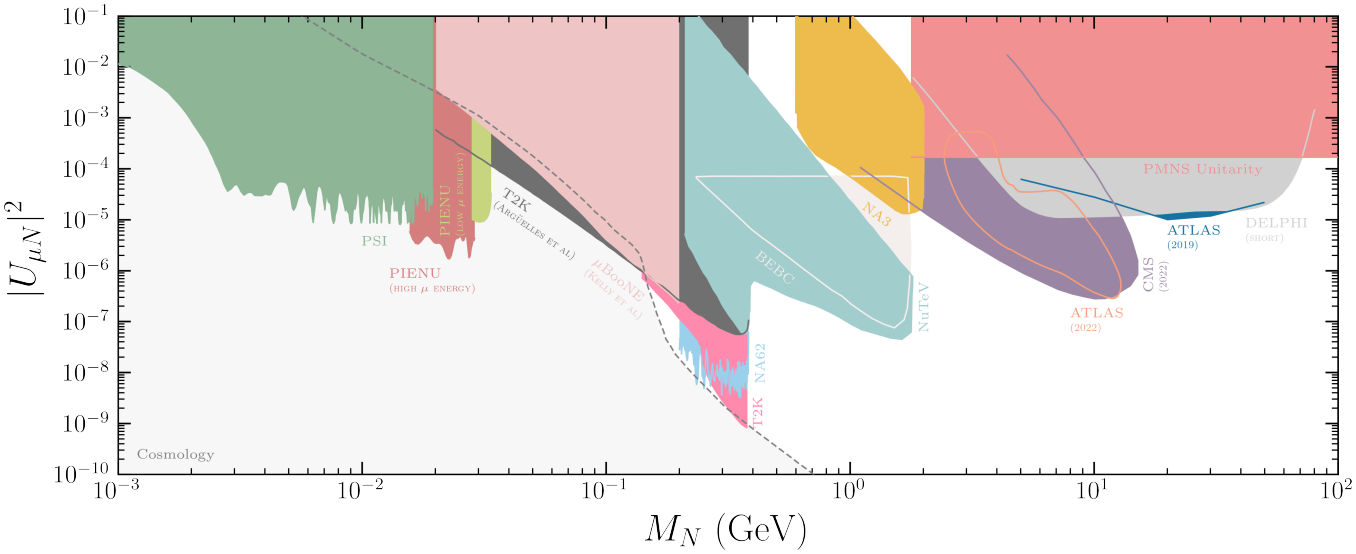


FIG. 2. Constraints on $|U_{\mu N}|^2$ as a function of the HNL mass $m_N$. Limits shown: $\mu$BooNE (Kelly et al) [15], ATLAS (2019) [1], ATLAS (2022) [2], BEBC [16], CMS (2022) [7], Cosmology [8], DELPHI (short) [9], NA3 [17], NA62 [18], NuTeV [19], PIENU(high $\mu$ energy) [20], PIENU(low $\mu$ energy) [20], PMNS Unitarity [**?** ], PSI [21], T2K [13], T2K (Argüelles et al) [22].


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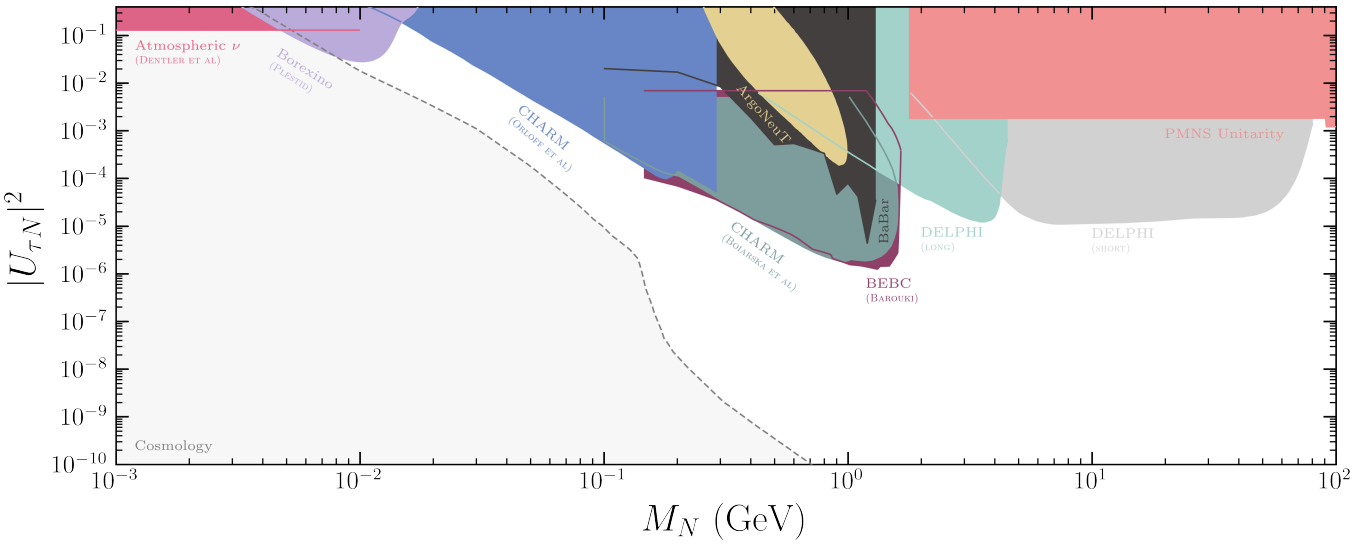


FIG. 3. Constraints on $|U_{\tau N}|^2$ as a function of the HNL mass $m_N$. Limits shown: ArgoNeuT [23], Atmospheric $\nu$ (Dentler et al) [24], BEBC(Barouki) [3], BaBar [25], Borexino (Plestid) [26], CHARM (Boiarska et al) [27], CHARM (Orloff et al) [28], Cosmology [8], DELPHI (long) [9], DELPHI (short) [9], PMNS Unitarity [**?** ].


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