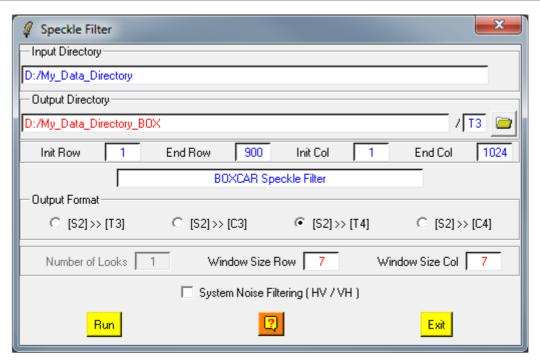


# **Speckle Filter**



### **Description:**

This function is used to apply a Polarimetric Speckle filtering on polarimetric raw binary data.

The different proposed polarimetric Speckle Filetrs are:

- Box Car filter
- Gauss filter
- IDAN filter
- Refined Lee filter.
- PWF (Polarimetric Whitening Filter)

According to the input data format, indicated in the widget, different compatible output data formats are proposed according the following table:

Input Data Format	<b>Output Data Format</b>
(2x2) Sinclair matrix [S2]	[T3], [T4], [C3], [C4]
(3x3) Coherency matrix [T3]	[T3]
(4x4) Coherency matrix [T4]	[T4]
(2x2) Covariance matrix [C2]	[C2]
(3x3) Covariance matrix [C3]	[C3]
(4x4) Covariance matrix [C4]	[C4]
Dual Polarimetric Elements (Sxx, Sxy)	[C2]
Intensities (Ixx, Ixy)	(Ixx, Ixy)

#### **Comments:**

Parameters written in Red can be modified directly by the user from the keyboard.

### **Input/Output Arguments:**

**Input** Indicates the location of the considered **Main Directory (MD)** 

**Directory** containing the polarimetric data sets to be filtered.

**Output** Indicates the location of the filtered data output directory.

**Directory** The default value is set automatically to :

Main Directory\_XXX / YY.

where **XXX** is associated with the selected Speckle Filter (BOX, GSS, IDAN LEE) and where **YY** is associated with the Output

Data Format (C2, C3, C4, T3 or T4).

# **Output Image Number of Rows/Columns:**

The output image numbers of rows and columns are initialised to the input data set dimensions.

Users wishing to process a sub-part of the initial image can modify the **Init** and **End** values of the converted images rows and columns.

Note: init and end values have to remain within the range defined by the input image dimensions.

## **Filtering Parameters:**

Window size Users have to set the size of the (N\*N) sliding window used to

compute the local estimate of the average matrix.

The default value of N is set to 7.

Number of

Looks

Users have to set the Input data equivalent number of looks used to

compute the a priori input speckle noise variance.

The default value of N is set to 1.

Note: The IDAN (Intensity Driven Adaptive Neighbourhood) speckle filter functionality is a contribution by G. Vasile and E. Trouve from LISTIC – Polytech' Savoie.