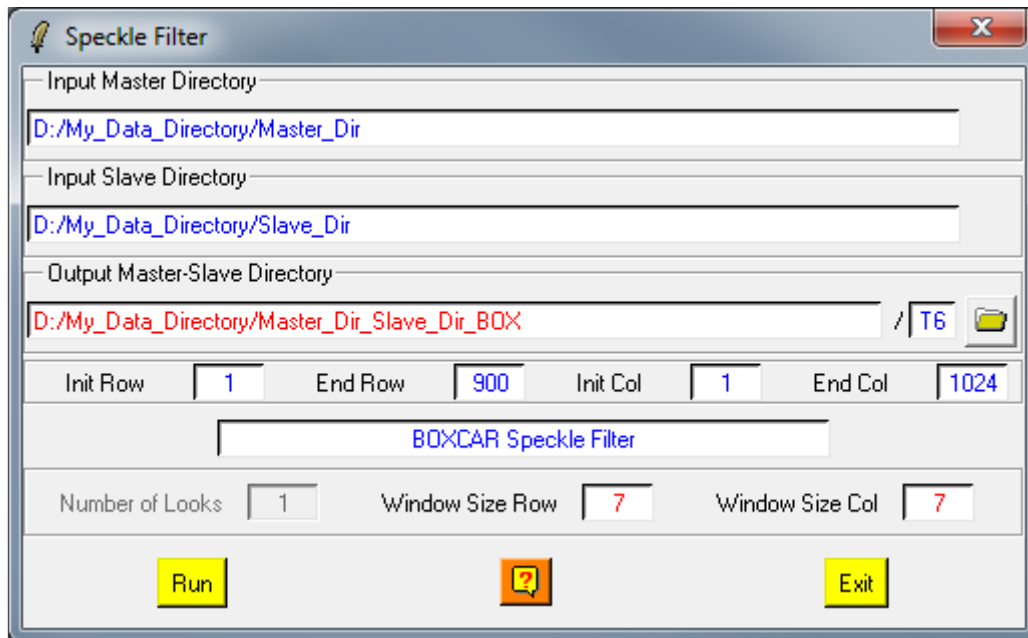


Speckle Filter



The screenshot shows the 'Speckle Filter' window with the following fields and controls:

- Input Master Directory:** D:/My_Data_Directory/Master_Dir
- Input Slave Directory:** D:/My_Data_Directory/Slave_Dir
- Output Master-Slave Directory:** D:/My_Data_Directory/Master_Dir_Slave_Dir_BOX / T6
- Init Row:** 1, **End Row:** 900, **Init Col:** 1, **End Col:** 1024
- BOXCAR Speckle Filter:** (Selected)
- Number of Looks:** 1, **Window Size Row:** 7, **Window Size Col:** 7
- Buttons:** Run, [?] (Help), Exit

Description:

This function is used to apply a Polarimetric Speckle filtering on polarimetric interferometric raw binary data.

The different proposed polarimetric Speckle Filters are :

- Box Car filter
- Gauss filter
- Refined Lee filter.

According to the input data format, indicated in the widget, different compatible output data formats are proposed according the following table:

Input Data Format	Output Data Format
2 x (2x2) Sinclair matrix [S2]	[T6]
(6x6) Coherency matrix [T6]	[T6]
2 x Dual Polarimetric Elements (Sxx, Sxy)	[T4]
(4x4) Coherency matrix [T4]	[T4]

Comments:

Parameters written in Red can be modified directly by the user from the keyboard.

Input/Output Arguments:

Input Master Directory	Indicates the location of the considered Master Main Directory (Master-MD) containing the polarimetric data sets to be filtered.
Input Slave Directory	Indicates the location of the considered Slave Main Directory (Slave-MD) containing the polarimetric data sets to be filtered.
Output Master-Slave Directory	Indicates the location of the filtered data output directory. The default value is set automatically to : Master-MD_Slave-MD_XXX / YY . where XXX is associated with the selected Speckle Filter (BOX, GSS, LEE) and where YY is associated with the Output Data Format (T6 or T4).

Output Image Number of Rows/Columns:

The output image numbers of rows and columns are initialised to the input data set dimensions.

Users wishing to process a sub-part of the initial image can modify the **Init** and **End** values of the converted images rows and columns.

Note: init and end values have to remain within the range defined by the input image dimensions.

Filtering Parameters:

Window size	Users have to set the size of the (N*N) sliding window used to compute the local estimate of the average matrix. The default value of N is set to 7 .
Number of Looks	Users have to set the Input data equivalent number of looks used to compute the a priori input speckle noise variance. The default value of N is set to 1 .
