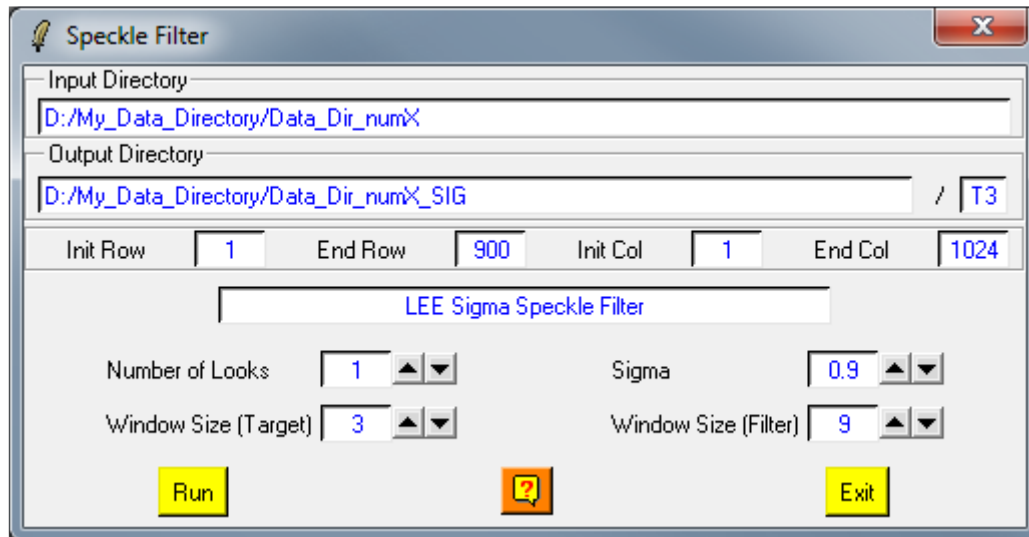


## J.S. Lee Sigma Speckle Filter



### Description:

This function is used to apply a Polarimetric Speckle filtering on polarimetric raw binary data.

The proposed polarimetric Speckle Filter is :

- J.S. Lee Sigma filter.

According to the input data format, indicated in the widget, different compatible output data formats are proposed according the following table:

| Input Data Format                     | Output Data Format |
|---------------------------------------|--------------------|
| (2x2) Sinclair matrix [S2]            | [T3]               |
| (3x3) Coherency matrix [T3]           | [T3]               |
| Dual Polarimetric Elements (Sxx, Sxy) | [C2]               |
| (2x2) Covariance matrix [C2]          | [C2]               |

## Comments:

Parameters written in Red can be modified directly by the user from the keyboard.

## Input/Output Arguments:

|                         |  |
|-------------------------|--|
| <b>Input Directory</b>  | Indicates the location of the considered <b>Main Directory (MD)</b> containing the polarimetric data sets to be filtered.  |
| <b>Output Directory</b> | Indicates the location of the filtered data output directory.<br>The default value is set automatically to :<br><b>Main Directory_SIG / YY.</b><br>where <b>YY</b> is associated with the Output Data Format ( <b>C2</b> or <b>T3</b> ). |

## Output Image Number of Rows/Columns:

The output image numbers of rows and columns are initialised to the input data set dimensions.

Users wishing to process a sub-part of the initial image can modify the **Init** and **End** values of the converted images rows and columns.

Note: init and end values have to remain within the range defined by the input image dimensions.

## Filtering Parameters:

|                             |  |
|-----------------------------|--|
| <b>Number of Looks</b>      | Users have to set the Input data equivalent number of looks used to compute the a priori input speckle noise variance.<br>The default value of N is set to <b>1</b> .              |
| <b>Window size (target)</b> | Users have to set the size of the (N*N) sliding window used to detect bright point target, compute the MMSE and fix the Sigma Range<br>The default value of N is set to <b>3</b> . |
| <b>Sigma</b>                | Sigma value  |
| <b>Window size (filter)</b> | Users have to set the size of the (N*N) sliding window used to compute the local estimate of the average matrix.<br>The default value of N is set to <b>9</b> .                    |

---

## Reference

"Improved Sigma Filter for Speckle Filtering of SAR imagery", J.S. Lee, J.H Wen, T. Ainsworth, K.S Chen, A.J Chen, IEEE GRS Letters - 2008