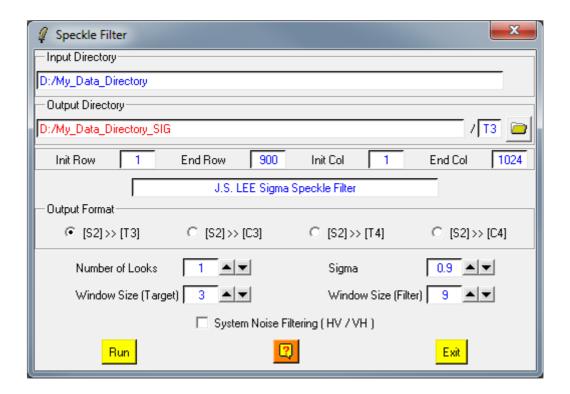


# J.S. Lee Sigma Speckle Filter



## **Description:**

This function is used to apply a Polarimetric Speckle filtering on polarimetric raw binary data.

The proposed polarimetric Speckle Filetr is:

• J.S. Lee Sigma filter.

According to the input data format, indicated in the widget, different compatible output data formats are proposed according the following table:

Input Data Format	<b>Output Data Format</b>
(2x2) Sinclair matrix [S2]	[T3], [T4], [C3], [C4]
(3x3) Coherency matrix [T3]	[T3]
(4x4) Coherency matrix [T4]	[T4]
(2x2) Covariance matrix [C2]	[C2]
(3x3) Covariance matrix [C3]	[C3]
(4x4) Covariance matrix [C4]	[C4]
Dual Polarimetric Elements (Sxx, Sxy)	[C2]

#### **Comments:**

Parameters written in Red can be modified directly by the user from the keyboard.

### **Input/Output Arguments:**

Indicates the location of the considered **Main Directory** (**MD**) Input

**Directory** containing the polarimetric data sets to be filtered.

Indicates the location of the filtered data output directory. **Output** 

**Directory** The default value is set automatically to:

Main Directory SIG / YY.

where YY is associated with the Output Data Format (C2, C3, C4,

**T3** or **T4**).

## **Output Image Number of Rows/Columns:**

The output image numbers of rows and columns are initialised to the input data set dimensions.

Users wishing to process a sub-part of the initial image can modify the **Init** and **End** values of the converted images rows and columns.

Note: init and end values have to remain within the range defined by the input image dimensions.

#### **Filtering Parameters:**

**Number of** Users have to set the Input data equivalent number of looks used to

Looks compute the a priori input speckle noise variance.

The default value of N is set to 1.

Window size Users have to set the size of the (N\*N) sliding window used to (target)

detect bright point target, compute the MMSE and fix the Sigma

The default value of N is set to 3.

**Sigma** Sigma value

Users have to set the size of the (N\*N) sliding window used to Window size

(filter) compute the local estimate of the average matrix.

The default value of N is set to 9.

#### Reference

<sup>&</sup>quot;Improved Sigma Filter for Speckle Filtering of SAR imagery", J.S. Lee, J.H Wen, T. Ainsworth, K.S Chen, A.J Chen, IEEE GRS Letters - 2008