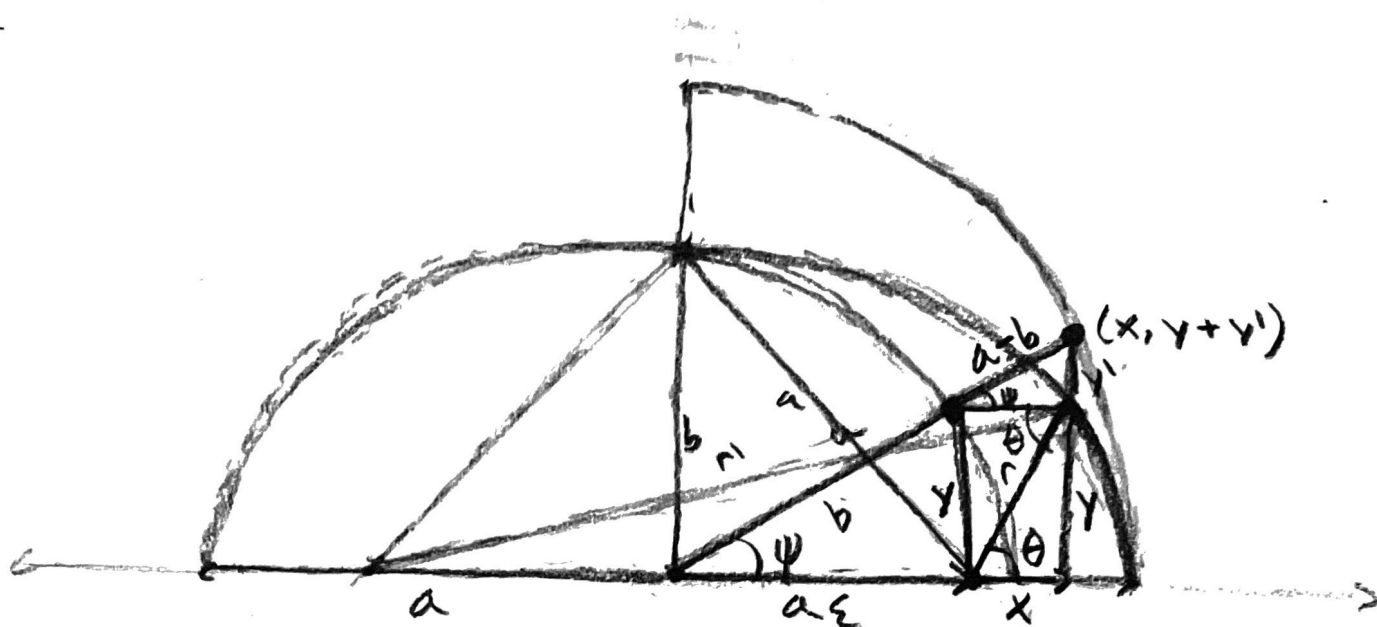


Kepler's Equation



The equation of an ellipse is:

$$\frac{(x + a\epsilon)^2}{a^2} + \frac{y^2}{b^2} = 1$$

additionally, by geometry, $\cos \psi = \frac{x + a\epsilon}{a}$;
therefore:

$$\cos^2 \psi + \frac{y^2}{b^2} = 1$$

By the Pythagorean Identity,
 $\cos^2 \psi + \sin^2 \psi = 1$, so $\sin \psi = \frac{y}{b}$.

2) We then have:

$$y = b \sin \psi \quad x = a(\cos \psi - \epsilon) \\ = a\sqrt{1-\epsilon^2} \sin \psi$$

As $r^2 = x^2 + y^2$:

$$r^2 = a^2((1-\epsilon^2)\sin^2 \psi + \cos^2 \psi - 2\epsilon \cos^2 \psi + \epsilon^2)$$

As $\sin^2 \psi = 1 - \cos^2 \psi$:

$$r^2 = a^2[1 - \cancel{\epsilon^2} - \cancel{\cos^2 \psi} + \epsilon^2 \cos^2 \psi + \cancel{\cos^2 \psi} - 2\epsilon \cos \psi + \cancel{\epsilon^2}]$$

$$= a^2[1 + \epsilon^2 \cos^2 \psi - 2\epsilon \cos \psi]$$

$$= a^2(1 - \epsilon \cos \psi), \text{ and so:}$$

$$r(\psi) = a(1 - \epsilon \cos \psi)$$

Thus far, our derivation has been purely geometrical. Now, we can use our shape equation $r(\theta)$ to relate θ and ψ . For a body under the an inverse square force, $\theta(r)$ is:

$$\theta(r) = \int \frac{(1/r^2) dr}{\sqrt{2\mu(E + \frac{K}{r} - \frac{L^2}{2\mu r^2})}}$$

$$\theta = \frac{l}{\sqrt{2m}} \int \frac{(1/r^2) dr}{\sqrt{E + \frac{k}{r} - \frac{l^2}{2mr^2}}}, \text{ letting } u = \frac{1}{r}, du = -\frac{1}{r^2} dr.$$

$$= -\frac{l}{\sqrt{2m}} \int \frac{du}{\sqrt{E + ku - \frac{l^2}{2m}u^2}}, \text{ where } a^2 = \frac{l^2}{2m} \quad \begin{aligned} & au - \frac{k}{2a} \\ & -\frac{k}{2a} - \frac{k}{2}u \quad \frac{k^2}{4a^2} \end{aligned}$$

$$\text{As } E + ku - \frac{l^2}{2m}u^2 = E + \frac{k^2}{4a^2} - \left(au - \frac{k}{2a}\right)^2$$

$$= -\frac{l}{\sqrt{2m}} \int \frac{du}{\sqrt{\left(E + \frac{k^2}{4a^2}\right) - \left(au - \frac{k}{2a}\right)^2}}, \text{ letting } b^2 = E + \frac{k^2}{4a^2},$$

$$w = au - \frac{k}{2a}, dw = a du.$$

$$= -\frac{l}{a\sqrt{2m}} \int \frac{dw}{\sqrt{b^2 - w^2}}, \text{ letting } w = b \cos \phi, dw = -b \sin \phi d\phi$$

$$= -\frac{l}{a\sqrt{2m}} \int \frac{\sin \phi d\phi}{\sqrt{1 - \cos^2 \phi}} = \frac{1}{a\sqrt{2m}} \phi = \frac{l}{a\sqrt{2m}} \arccos\left(\frac{w}{b}\right)$$

$$= \frac{l}{a\sqrt{2m}} \arccos\left(\frac{au - \frac{k}{2a}}{\sqrt{E + \frac{k^2}{4a^2}}}\right) = \frac{l}{a\sqrt{2m}} \arccos\left(\frac{\frac{l}{\sqrt{2m}r} - \frac{k/\sqrt{2m}}{l}}{\sqrt{E + \frac{2k^2m}{l^2}}}\right)$$

$$= \frac{l}{a\sqrt{2m}} \arccos\left(\frac{\frac{l^2}{2mr} - l}{\sqrt{\frac{El^2}{2k^2m} + l}}\right), \text{ otherwise correct}$$

Which is all equivalent to θ , so:

$$\cos(\theta) = \frac{\frac{\ell^2}{2\mu r} - 1}{\sqrt{\frac{E\ell^2}{2k^2\mu} + 1}}, \text{ where } \epsilon = \sqrt{\frac{E\ell^2}{2k^2\mu} + 1},$$
$$\alpha = \frac{\ell^2}{2\mu}, \text{ so:}$$

$$r = \frac{\alpha}{1 + \epsilon \cos \theta} \text{ (which describes a conic section with eccentricity } \epsilon)$$

$$r + \epsilon r \cos \theta = \alpha, \text{ (a) } \alpha = a(1 - \epsilon^2):$$

Because the orbit is an ellipse, we require that $\alpha = a(1 - \epsilon^2)$, so: $\frac{E\ell^2}{2k^2\mu} + 1 = 1 - \epsilon^2$

$$r + \epsilon r \cos \theta = a(1 - \epsilon^2) + 1 = \frac{aE}{2k^2} \left(\frac{\ell^2}{2\mu} \right)$$

$$\epsilon r \cos \theta = a(1 - \epsilon^2) - r, \text{ adding } \epsilon r \text{ to both sides!}$$

$$\epsilon r + \epsilon r \cos \theta = \epsilon r + a(1 - \epsilon^2) - r$$

$$\epsilon r(1 + \cos \theta) = r(\epsilon - 1) + a(1 - \epsilon)(1 + \epsilon)$$

$$\epsilon r(1 + \cos \theta) = (1 - \epsilon)[a(1 + \epsilon) - r]$$

Now, we recall that $r = a(1 - \epsilon \cos \psi)$; making this substitution, it follows that:

$$\cancel{\epsilon r}(1 + \epsilon \cos \theta) = (1 - \epsilon) [\cancel{a}(1 + \epsilon) - \cancel{a}(1 - \epsilon \cos \psi)]$$

$$\cancel{\epsilon r}(1 + \epsilon \cos \theta) = \cancel{a}(1 - \epsilon) [\cancel{1} + \cancel{\epsilon} \cos \psi]$$

$$r(1 + \epsilon \cos \theta) = a(1 - \epsilon)(1 + \epsilon \cos \psi)$$

Going back to the relation $r + \epsilon r \cos \theta = a(1 - \epsilon^2)$ we instead subtract ϵr , giving:

$$\epsilon r \cos \theta + \epsilon r = a(1 - \epsilon^2) - r - \epsilon r$$

$$\epsilon r(\cos \theta - 1) = a(1 - \epsilon)(1 + \epsilon) - r(1 + \epsilon)$$

$$\epsilon r(\cos \theta - 1) = (1 + \epsilon) [a(1 - \epsilon) - r]$$

$$\text{as } r = a(1 - \epsilon \cos \psi),$$

$$\epsilon r(\cos \theta - 1) = (1 + \epsilon) [a(1 - \epsilon) - a(1 - \epsilon \cos \psi)]$$

$$\epsilon r(\cos \theta - 1) = a(1 + \epsilon) [1 - \epsilon - 1 + \epsilon \cos \psi]$$

$$\cancel{\epsilon r}(\cos \theta - 1) = \cancel{a}(1 + \epsilon) [\cancel{-1} + \cancel{\epsilon} \cos \psi]$$

$$r(\cos\theta - 1) = a(1+e)(\cos\psi - 1)$$

Dividing the result from subtracting er by that from adding er gives:

$$\frac{\cos\theta + 1}{\cos\theta - 1} = \frac{1+e}{1-e} \cdot \frac{1 - \cos\psi}{1 + \cos\psi}$$

and as $\tan \frac{\theta}{2} = \sqrt{\frac{\cos\theta - 1}{\cos\theta + 1}}$:

$$\tan \frac{\theta}{2} = \sqrt{\frac{1+e}{1-e}} \tan \frac{\psi}{2}$$

~~$$\frac{\theta}{2} = \frac{\psi}{2} \arctan \left(\sqrt{\frac{1+e}{1-e}} \right)$$~~

~~$$\theta = \psi \arctan \left(\sqrt{\frac{1+e}{1-e}} \right)$$~~

Differentiating $\tan \frac{\theta}{2} = \sqrt{\frac{1+e}{1-e}} \tan \frac{\psi}{2}$ gives:

$$\frac{1}{2} \sec^2 \frac{\theta}{2} d\theta = \frac{1}{2} \sec^2 \frac{\psi}{2} \sqrt{\frac{1+e}{1-e}} d\psi$$

$$d\theta = \frac{\cos^2 \frac{\theta}{2}}{\cos^2 \frac{\psi}{2}} \sqrt{\frac{1+e}{1-e}} d\psi$$

as $r(1+\cos\theta) = a(1-e)(1+\cos\psi)$, $r = a(1-e) \frac{1+\cos\psi}{1+\cos\theta}$, however
 as $\cos^2 \frac{\theta}{2} = \frac{1}{2}(1+\cos\theta)$;

$$r = a(1-e) \frac{\cos^2(\frac{\psi}{2})}{\cos^2(\frac{\theta}{2})}$$

Kepler's second law states, that equal areas are swept out by the orbit in equal times. Say that the period of orbit is T ; the area of the ellipse is $A = \pi ab$. Because area is swept out at a constant rate, this rate must be the area swept out in an orbit in a time T , $\frac{\Delta A}{\Delta t} = \frac{A}{T} = \frac{\pi ab}{T}$; we can equate this with the area integral, $\int dA = \frac{1}{2} \int r^2 d\theta$ (analogous to $\Delta \theta r^2$ for a circle). That is:

$$\frac{\pi ab}{T} = \frac{1}{2} \int r^2 d\theta$$

$$\text{As } d\theta = \frac{\cos^2 \frac{\theta}{2} \sqrt{1+\epsilon}}{\cos^2 \frac{\psi}{2} \sqrt{1-\epsilon}} d\psi \quad r = a(1-\epsilon) \frac{\cos^2 \frac{\psi}{2}}{\cos^2 \frac{\theta}{2}};$$

$$\frac{\pi ab}{T} = \frac{1}{2} \int a(1-\epsilon) \sqrt{1+\epsilon} r d\psi$$

$$\text{Taking } r = a(1-\epsilon \cos \psi):$$

$$\frac{\pi ab}{T} = \frac{1}{2} a^2 \sqrt{1-\epsilon^2} \int 1 - \epsilon \cos \psi d\psi$$

$$\frac{\pi b}{T} = \frac{a}{2} \sqrt{1-\epsilon^2} (\psi - \epsilon \sin \psi)$$

$$\text{as } b = \epsilon \sqrt{1 - \epsilon^2};$$

$$\frac{R \cancel{\epsilon \sqrt{1 - \epsilon^2}}}{T} + = \frac{\cancel{\epsilon}}{2} \sqrt{1 - \cancel{\epsilon^2}} (\psi - \epsilon \sin \psi)$$

$$\frac{2R}{T} + = \psi - \epsilon \sin \psi$$

which must somehow be approximated,
for $\psi(t)$.