

# Some computational results for generalized pressure Schur complement eigenvalues of the surface Stokes problem

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## 1 Bilinear forms and matrices

We set  $n_{\mathbf{A}}$  to be the number of velocity d.o.f. and  $n_{\mathbf{S}}$  to be the number of pressure d.o.f. Vector stiffness, divergence, pressure mass, normal stabilization, and full stabilization matrices resulting from TraceFEM discretization of the surface Stokes problem [1] are defined via

$$\begin{aligned} \langle \mathbf{A} \bar{\mathbf{u}}, \bar{\mathbf{v}} \rangle &= \int_{\Gamma} (E_s(\mathbf{u}) : E_s(\mathbf{v}) + \mathbf{u} \cdot \mathbf{v} + \tau u_N v_N) ds + \rho_u \int_{\Omega_h^{\Gamma}} ([\nabla \mathbf{u}] \hat{\mathbf{n}}) \cdot ([\nabla \mathbf{v}] \hat{\mathbf{n}}) d\mathbf{x}, \quad \mathbf{A} \in \mathbb{R}^{n_{\mathbf{A}} \times n_{\mathbf{A}}}, \\ \langle \mathbf{B} \bar{\mathbf{u}}, \bar{q} \rangle &= - \int_{\Gamma} q \operatorname{div}_{\Gamma} \mathbf{u} ds, \quad \mathbf{B} \in \mathbb{R}^{n_{\mathbf{S}} \times n_{\mathbf{A}}}, \\ \langle \mathbf{M}_0 \bar{\mathbf{p}}, \bar{q} \rangle &= \int_{\Gamma} p q ds, \quad \mathbf{M}_0 \in \mathbb{R}^{n_{\mathbf{S}} \times n_{\mathbf{S}}}, \\ \langle \mathbf{C}_n \bar{\mathbf{p}}, \bar{q} \rangle &= \rho_p \int_{\Omega_h^{\Gamma}} \frac{\partial p}{\partial \hat{\mathbf{n}}} \frac{\partial q}{\partial \hat{\mathbf{n}}} d\mathbf{x}, \quad \mathbf{C}_n \in \mathbb{R}^{n_{\mathbf{S}} \times n_{\mathbf{S}}}, \\ \langle \mathbf{C}_{\text{full}} \bar{\mathbf{p}}, \bar{q} \rangle &= \rho_p \int_{\Omega_h^{\Gamma}} \nabla p \cdot \nabla q d\mathbf{x}, \quad \mathbf{C}_{\text{full}} \in \mathbb{R}^{n_{\mathbf{S}} \times n_{\mathbf{S}}}, \end{aligned} \tag{1}$$

respectively. We use notations as in [1], in particular,  $\Omega_h^{\Gamma}$  is the domain consisting of tetrahedra cut by  $\Gamma$ . Here  $\bar{\mathbf{u}}$  denotes a vector of d.o.f. corresponding to a FE interpolant  $\mathbf{u}$  (analogously for  $\bar{\mathbf{p}}$  and  $p$ ). Mesh-dependent parameters are set as

$$\tau = h^{-2}, \quad \rho_u = \rho_p = h, \tag{2}$$

and  $h$  is the typical mesh size for tetrahedra from  $\Omega_h^{\Gamma}$ .  $\Gamma$  is chosen either as the unit sphere or torus,  $\Gamma = \Gamma_{\text{sph}}$  or  $\Gamma = \Gamma_{\text{tor}}$  (see Figure 1).

We also define matrices

$$\begin{aligned} \mathbf{C}_0 &:= \mathbf{0}, \\ \mathbf{M}_n &:= \mathbf{M}_0 + \mathbf{C}_n, \\ \mathbf{M}_{\text{full}} &:= \mathbf{M}_0 + \mathbf{C}_{\text{full}}. \end{aligned} \tag{3}$$

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We are interested in (generalized) extreme eigenvalues of the pressure Schur complement matrices

$$\begin{aligned}\mathbf{S}_0 &:= \mathbf{B} \mathbf{A}^{-1} \mathbf{B}^T, \\ \mathbf{S}_n &:= \mathbf{S}_0 + \mathbf{C}_n, \\ \mathbf{S}_{\text{full}} &:= \mathbf{S}_0 + \mathbf{C}_{\text{full}},\end{aligned}\tag{4}$$

i.e. in solving

$$\mathbf{S}_\star \mathbf{x} = \lambda \mathbf{M}_\star \mathbf{x},\tag{5}$$

where “ $\star$ ” stands for “0,” “ $n$ ,” or “full.” We denote by  $0 = \lambda_1 < \lambda_2 \leq \dots \leq \lambda_{n_s} = O(1)$  the spectrum of (5).

## 2 Solution description

Computing  $\mathbf{A}^{-1}$  in (4) becomes troublesome already for  $h = 5.21 \times 10^{-2}$  ( $n_{\mathbf{A}} = 32736$  for  $\mathbf{u} \in \mathbf{P}_1$  FE space): although  $\mathbf{A}$  is sparse,  $\mathbf{A}^{-1}$  is dense and consumes 8.5+ GB in double-precision arithmetic. A quick research [showed](#) that **Mathematica** has no built-in matrix-free eigenvalue routines. Intel MKL’s FEAST algorithm for computing (generalized) eigenvalues in an interval [is suitable for matrix-free implementations](#); however, it requires some expensive operations to be implemented (e.g. matrix-matrix multiplications  $\mathbf{Y} \leftarrow \mathbf{S}_\star \mathbf{X}$ ,  $\mathbf{Y} \leftarrow \mathbf{M}_\star \mathbf{X}$  and approximating the action of inverses in the form  $\mathbf{y} \leftarrow (\sigma \mathbf{M}_\star - \mathbf{S}_\star)^{-1} \mathbf{x}$ ).

Taking this into account, instead of (5) we consider a perturbed<sup>1</sup> problem

$$\underbrace{\begin{bmatrix} \mathbf{A} & \mathbf{B}^T \\ \mathbf{B} & -\mathbf{C}_\star \end{bmatrix}}_{\mathcal{A}_\star :=} \begin{bmatrix} \mathbf{x} \\ \mathbf{y} \end{bmatrix} = \mu \underbrace{\begin{bmatrix} \epsilon \mathbf{A} & \\ & \mathbf{M}_\star \end{bmatrix}}_{\mathcal{M}_\star^\epsilon :=} \begin{bmatrix} \mathbf{x} \\ \mathbf{y} \end{bmatrix}\tag{6}$$

with  $0 < \epsilon \ll 1$ . For  $\mathcal{A}_0$  and  $\mathcal{M}_0^\epsilon$  we have

$$\mu = -\lambda + o(1) \quad \text{or} \quad \epsilon^{-1} + \lambda + o(1), \quad \epsilon \rightarrow 0.\tag{7}$$

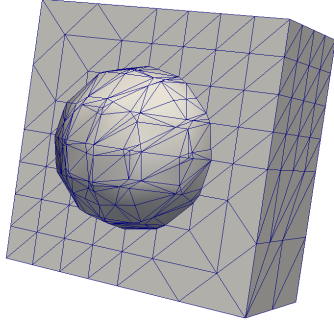
This makes it easy to pick only “correct” eigenvalues. To ease the computation further we replace the  $(1, 1)$ -block of  $\mathcal{M}_\star^\epsilon$  with  $\epsilon \mathbf{I}$ .

To make sure that results are consistent we solve (6) for  $\epsilon = 10^{-5}$  and  $\epsilon = 10^{-6}$ ; for the coarse mesh levels we also check that the direct solution of (5) and the iterative one for (6) coincide.

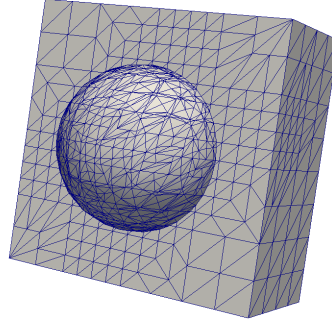
## 3 Numerical results: dependency of the spectrum on the mesh size

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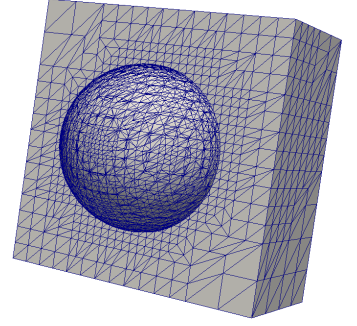
<sup>1</sup>The majority of generalized eigenvalue solvers require left-hand-side matrix to be Hermitian and right-hand-side matrix to be Hermitian **positive definite**; that’s why we need to introduce  $\epsilon > 0$ .



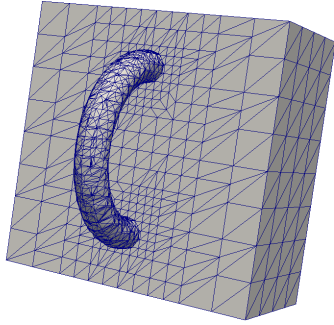
(a)  $h = 8.33 \times 10^{-1}$



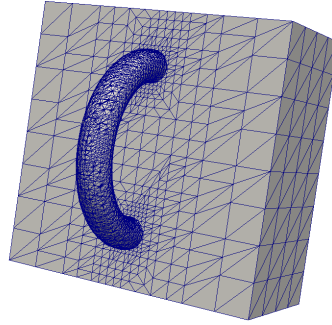
(b)  $h = 4.17 \times 10^{-1}$



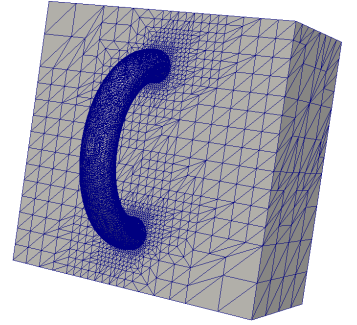
(c)  $h = 2.08 \times 10^{-1}$



(d)  $h = 2.08 \times 10^{-1}$



(e)  $h = 1.04 \times 10^{-1}$



(f)  $h = 5.21 \times 10^{-2}$

Figure 1: First three mesh levels for  $\Gamma_{\text{sph}}$  (top) and  $\Gamma_{\text{tor}}$  (bottom)

Table 1: Spectrum of (5) for  $\mathbf{P}_1 - P_1$ 

 (a)  $\Gamma = \Gamma_{\text{sph}}$ 

$h$	$n_{\mathbf{A}}$	$n_{\mathbf{S}}$	$\mathbf{S}_0$		$\mathbf{S}_n$		$\mathbf{S}_{\text{full}}$	
			$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$
$8.33 \times 10^{-1}$	153	51	$1.32 \times 10^{-2}$	1.42	$7.48 \times 10^{-1}$	1.13	$9.58 \times 10^{-1}$	1.06
$4.17 \times 10^{-1}$	570	190	$5.12 \times 10^{-3}$	1.04	$5.77 \times 10^{-1}$	1.	$8.54 \times 10^{-1}$	1.
$2.08 \times 10^{-1}$	1992	664	$4.4 \times 10^{-3}$	$7.93 \times 10^{-1}$	$3.87 \times 10^{-1}$	1.	$6.71 \times 10^{-1}$	1.
$1.04 \times 10^{-1}$	8292	2764	$2.01 \times 10^{-3}$	$7.79 \times 10^{-1}$	$2.19 \times 10^{-1}$	1.	$5.82 \times 10^{-1}$	1.
$5.21 \times 10^{-2}$	32736	10912	$6.04 \times 10^{-5}$	$9.81 \times 10^{-1}$	$1.17 \times 10^{-1}$	1.	$5.37 \times 10^{-1}$	1.
$2.6 \times 10^{-2}$	131592	43864	$3.53 \times 10^{-5}$	$8.67 \times 10^{-1}$	$5.72 \times 10^{-2}$	1.	$5.16 \times 10^{-1}$	1.
$1.3 \times 10^{-2}$	525864	175288	$2.16 \times 10^{-6}$	$7.34 \times 10^{-1}$	$2.84 \times 10^{-2}$	1.	$5.04 \times 10^{-1}$	1.

 (b)  $\Gamma = \Gamma_{\text{tor}}$ 

$h$	$n_{\mathbf{A}}$	$n_{\mathbf{S}}$	$\mathbf{S}_0$		$\mathbf{S}_n$		$\mathbf{S}_{\text{full}}$	
			$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$
$2.08 \times 10^{-1}$	972	324	$5.04 \times 10^{-2}$	4.93	$2.84 \times 10^{-1}$	1.35	$3.64 \times 10^{-1}$	1.19
$1.04 \times 10^{-1}$	4740	1580	$2.99 \times 10^{-3}$	3.83	$1.58 \times 10^{-1}$	1.02	$3.35 \times 10^{-1}$	1.01
$5.21 \times 10^{-2}$	19704	6568	$1.11 \times 10^{-3}$	5.45	$7.73 \times 10^{-2}$	1.01	$3.25 \times 10^{-1}$	1.
$2.6 \times 10^{-2}$	80808	26936	$1.2 \times 10^{-4}$	5.42	$3.07 \times 10^{-2}$	1.01	$3.21 \times 10^{-1}$	1.
$1.3 \times 10^{-2}$	327036	109012	$1.77 \times 10^{-5}$	5.23	$1.18 \times 10^{-2}$	1.01	$3.16 \times 10^{-1}$	1.

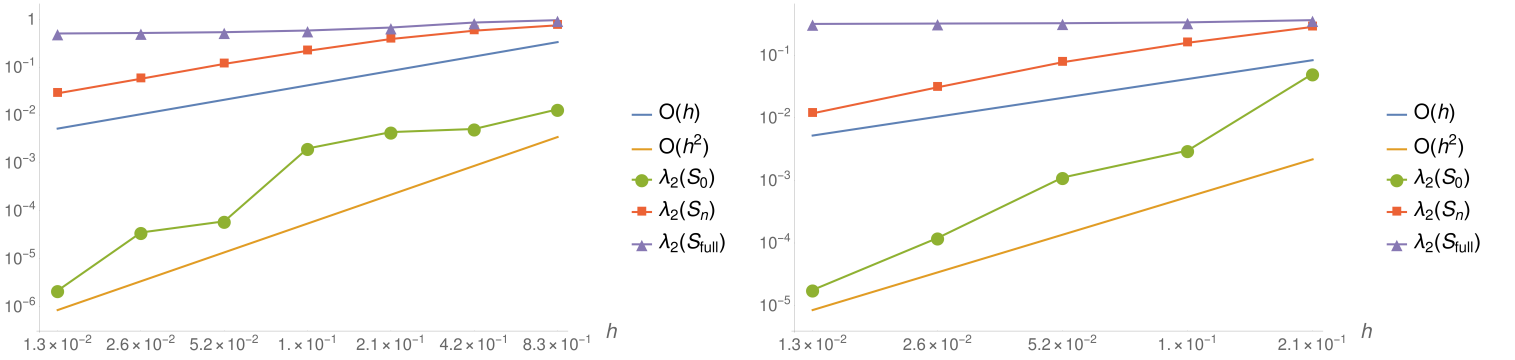

 Figure 2: Log-log plot of  $\lambda_2$  for Tables 1a (left) and 1b (right)

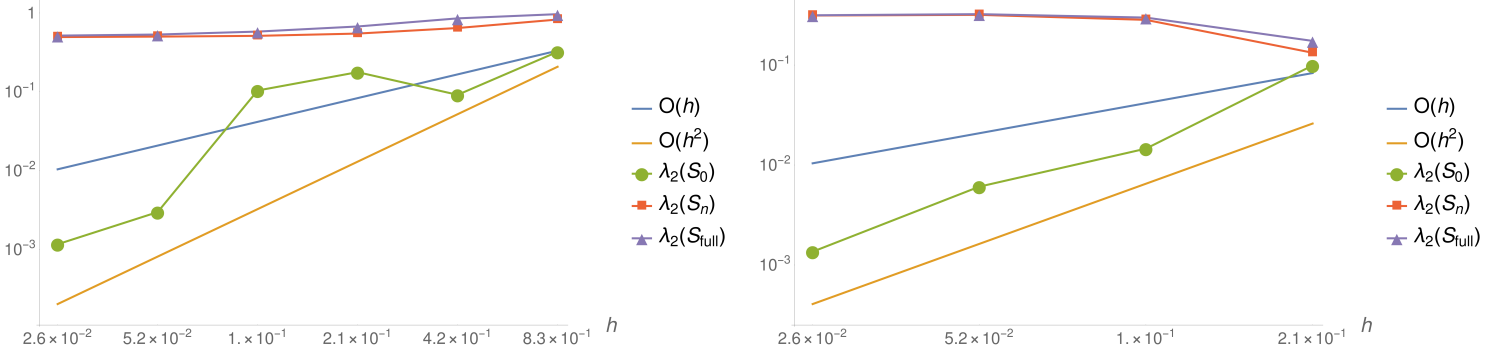
Table 2: Spectrum of (5) for  $\mathbf{P}_2 - P_1$ 

 (a)  $\Gamma = \Gamma_{\text{sph}}$ 

$h$	$n_{\mathbf{A}}$	$n_{\mathbf{S}}$	$\mathbf{S}_0$		$\mathbf{S}_n$		$\mathbf{S}_{\text{full}}$	
			$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$
$8.33 \times 10^{-1}$	789	51	$3.22 \times 10^{-1}$	1.73	$8.27 \times 10^{-1}$	1.17	$9.68 \times 10^{-1}$	1.07
$4.17 \times 10^{-1}$	3240	190	$9.17 \times 10^{-2}$	1.08	$6.45 \times 10^{-1}$	1.	$8.56 \times 10^{-1}$	1.
$2.08 \times 10^{-1}$	11718	664	$1.78 \times 10^{-1}$	$8.31 \times 10^{-1}$	$5.49 \times 10^{-1}$	1.	$6.75 \times 10^{-1}$	1.
$1.04 \times 10^{-1}$	48762	2764	$1.04 \times 10^{-1}$	$8.35 \times 10^{-1}$	$5.14 \times 10^{-1}$	1.	$5.82 \times 10^{-1}$	1.
$5.21 \times 10^{-2}$	193014	10912	$2.99 \times 10^{-3}$	$9.89 \times 10^{-1}$	$5.02 \times 10^{-1}$	1.	$5.34 \times 10^{-1}$	1.
$2.6 \times 10^{-2}$	775998	43864	$1.17 \times 10^{-3}$	$7.9 \times 10^{-1}$	$4.96 \times 10^{-1}$	1.	$5.17 \times 10^{-1}$	1.

 (b)  $\Gamma = \Gamma_{\text{tor}}$ 

$h$	$n_{\mathbf{A}}$	$n_{\mathbf{S}}$	$\mathbf{S}_0$		$\mathbf{S}_n$		$\mathbf{S}_{\text{full}}$	
			$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$	$\lambda_2$	$\lambda_{n_{\mathbf{S}}}$
$2.08 \times 10^{-1}$	5184	324	$9.92 \times 10^{-2}$	3.89	$1.33 \times 10^{-1}$	1.37	$1.75 \times 10^{-1}$	1.19
$1.04 \times 10^{-1}$	27906	1580	$1.46 \times 10^{-2}$	4.35	$2.84 \times 10^{-1}$	1.04	$2.99 \times 10^{-1}$	1.02
$5.21 \times 10^{-2}$	116568	6568	$6.08 \times 10^{-3}$	4.85	$3.19 \times 10^{-1}$	1.01	$3.24 \times 10^{-1}$	1.01
$2.6 \times 10^{-2}$	477660	26936	$1.36 \times 10^{-3}$	4.92	$3.14 \times 10^{-1}$	1.01	$3.16 \times 10^{-1}$	1.


 Figure 3: Log-log plot of  $\lambda_2$  for Tables 2a (left) and 2b (right)

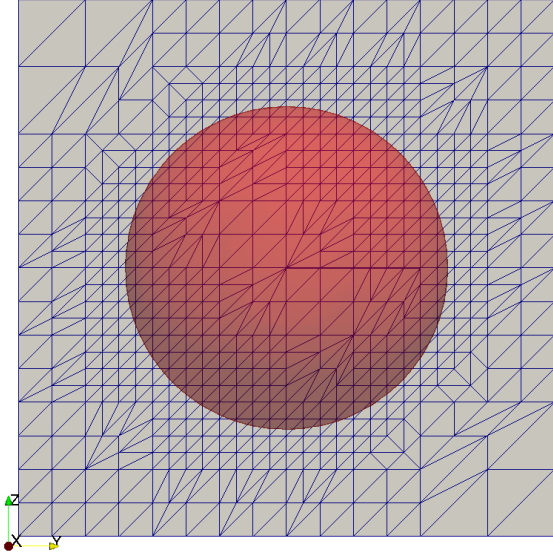
## 4 Numerical results: sensitivity of the spectrum to levelset shifts

In this section we investigate the sensitivity of the spectrum to levelset shifts

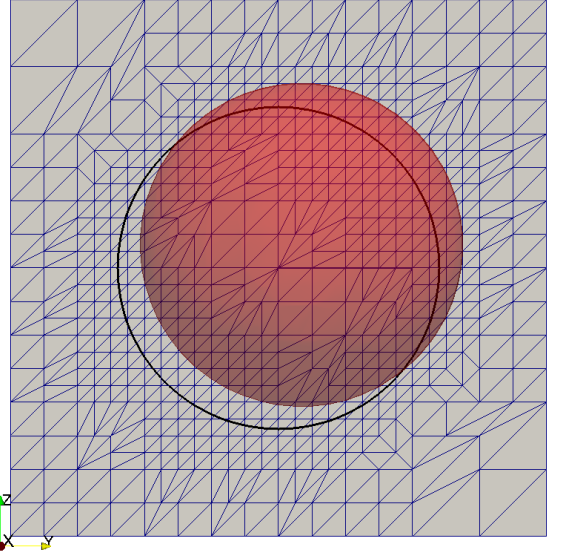
$$\Gamma \mapsto \Gamma + \alpha \mathbf{s}, \quad (8)$$

for some  $\alpha \in \mathbb{R}$  and  $\mathbf{s} \in \mathbb{R}^3$ ,  $\|\mathbf{s}\| = 1$ .

We construct the bulk mesh  $\Omega_h^\Gamma$  and then perform the assembly of matrices (1) using the shifted levelset (8). That is, the refinement of  $\Omega_h^\Gamma$  is performed using  $\Gamma$ , not  $\Gamma + \alpha \mathbf{s}$ , and  $\Omega_h^{\Gamma + \alpha \mathbf{s}}$  is never constructed. We choose  $\alpha \in [0, h]$  to guarantee the appearance of “small cuts” in  $\Omega_h^\Gamma$ .



(a)  $\Gamma_{\text{sph}}$



(b)  $\Gamma_{\text{sph}} + \alpha \mathbf{s}$

Figure 4: The unit sphere (left) and the shifted unit sphere (right). Here  $\mathbf{s} = (0, 1, 1)^T / \sqrt{2}$ ,  $\alpha = 0.2$ , and  $h = 2.08 \times 10^{-1}$ . The bulk mesh  $\Omega_h^\Gamma$  is computed for  $\Gamma_{\text{sph}}$  and then used for  $\Gamma_{\text{sph}} + \alpha \mathbf{s}$

Table 3: Spectrum of (5) for perturbed levelset  $\Gamma_{\text{sph}} + \alpha \mathbf{s}$ . Here  $\mathbf{s} = (1, 1, 1)^T / \sqrt{3}$ ,  $h = 1.04 \times 10^{-1}$

(a)  $\mathbf{P}_1 - P_1$

Surface	$\mathbf{S}_0$		$\mathbf{S}_n$		$\mathbf{S}_{\text{full}}$	
	$\lambda_2$	$\lambda_{n\mathbf{S}}$	$\lambda_2$	$\lambda_{n\mathbf{S}}$	$\lambda_2$	$\lambda_{n\mathbf{S}}$
$\Gamma_{\text{sph}}$	$2.006 \times 10^{-3}$	$7.79 \times 10^{-1}$	$2.19 \times 10^{-1}$	1.	$5.818 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.1 h \mathbf{s}$	$4.832 \times 10^{-4}$	$8.01 \times 10^{-1}$	$2.195 \times 10^{-1}$	1.	$5.818 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.3 h \mathbf{s}$	$7.278 \times 10^{-4}$	$8.17 \times 10^{-1}$	$2.203 \times 10^{-1}$	1.	$5.818 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.5 h \mathbf{s}$	$3.121 \times 10^{-4}$	$8.67 \times 10^{-1}$	$2.221 \times 10^{-1}$	1.	$5.82 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.7 h \mathbf{s}$	$1.438 \times 10^{-3}$	1.51	$2.254 \times 10^{-1}$	1.	$5.82 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + h \mathbf{s}$	$1.79 \times 10^{-3}$	2.07	$2.332 \times 10^{-1}$	1.	$5.827 \times 10^{-1}$	1.

(b)  $\mathbf{P}_2 - P_1$

Surface	$\mathbf{S}_0$		$\mathbf{S}_n$		$\mathbf{S}_{\text{full}}$	
	$\lambda_2$	$\lambda_{n\mathbf{S}}$	$\lambda_2$	$\lambda_{n\mathbf{S}}$	$\lambda_2$	$\lambda_{n\mathbf{S}}$
$\Gamma_{\text{sph}}$	$1.041 \times 10^{-1}$	$8.35 \times 10^{-1}$	$5.138 \times 10^{-1}$	1.	$5.841 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.1 h \mathbf{s}$	$1.705 \times 10^{-3}$	$8.58 \times 10^{-1}$	$5.138 \times 10^{-1}$	1.	$5.841 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.3 h \mathbf{s}$	$2.293 \times 10^{-3}$	$8.81 \times 10^{-1}$	$5.137 \times 10^{-1}$	1.	$5.841 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.5 h \mathbf{s}$	$5.63 \times 10^{-3}$	$9.35 \times 10^{-1}$	$5.138 \times 10^{-1}$	1.	$5.844 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + 0.7 h \mathbf{s}$	$8.188 \times 10^{-3}$	1.77	$5.138 \times 10^{-1}$	1.	$5.843 \times 10^{-1}$	1.
$\Gamma_{\text{sph}} + h \mathbf{s}$	$1.93 \times 10^{-2}$	2.21	$5.142 \times 10^{-1}$	1.	$5.852 \times 10^{-1}$	1.

## References

- [1] M. Olshanskii, A. Quaini, A. Reusken, and V. Yushutin. A finite element method for the surface stokes problem. *SIAM Journal on Scientific Computing*, 40(4):A2492–A2518, 2018.