# 太极图形课

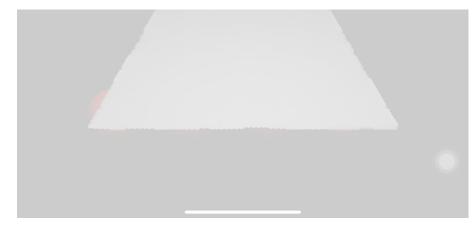
第09讲 Fluid Simulation 01: The Particle-based Methods



#### Where are we?



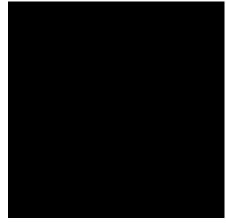
**Procedural Animation** 



Deformable Simulation

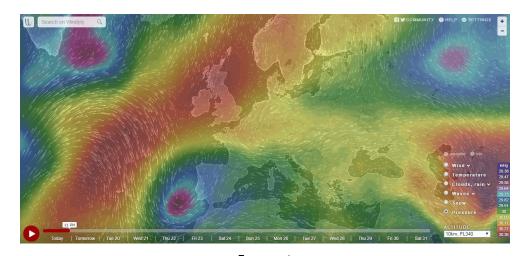


Rendering



Fluid Simulation

#### Fluid simulation

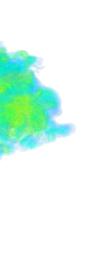




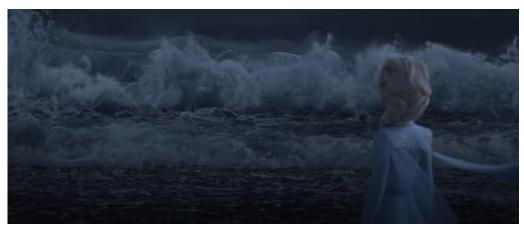


Forecast

Design



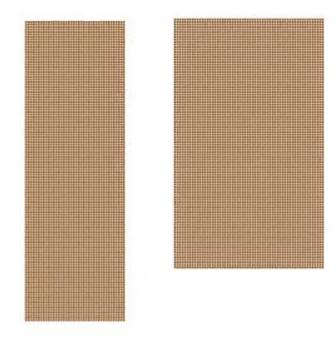
VFX Game



Animation

# Code of the day

- Code: <a href="https://github.com/taichiCourse01/taichi\_sph">https://github.com/taichiCourse01/taichi\_sph</a>
- Code courtesy of @erizmr

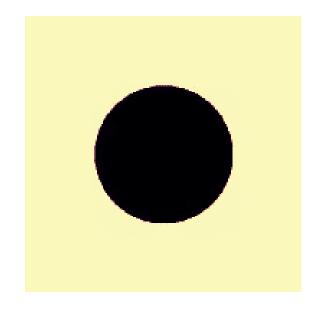


#### Outline today

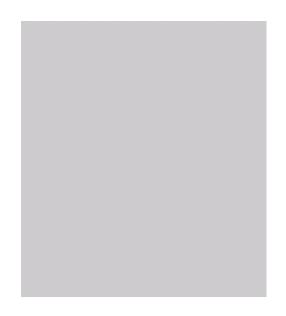
- Incompressible fluid dynamics
  - Incompressible Navier–Stokes equations
- Time discretization
  - Operator splitting
  - Integration with the weakly compressible (WC) assumption
- Spatial discretization
  - Smoothed particle hydrodynamics (SPH)
- Implementation details (WCSPH)
  - Simulation Pipeline
  - Boundary conditions
  - Neighbor search

Incompressible fluid dynamics

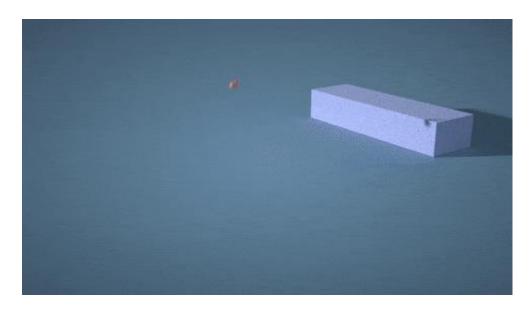
### Compressible and incompressible fluids



Compressible shock wave

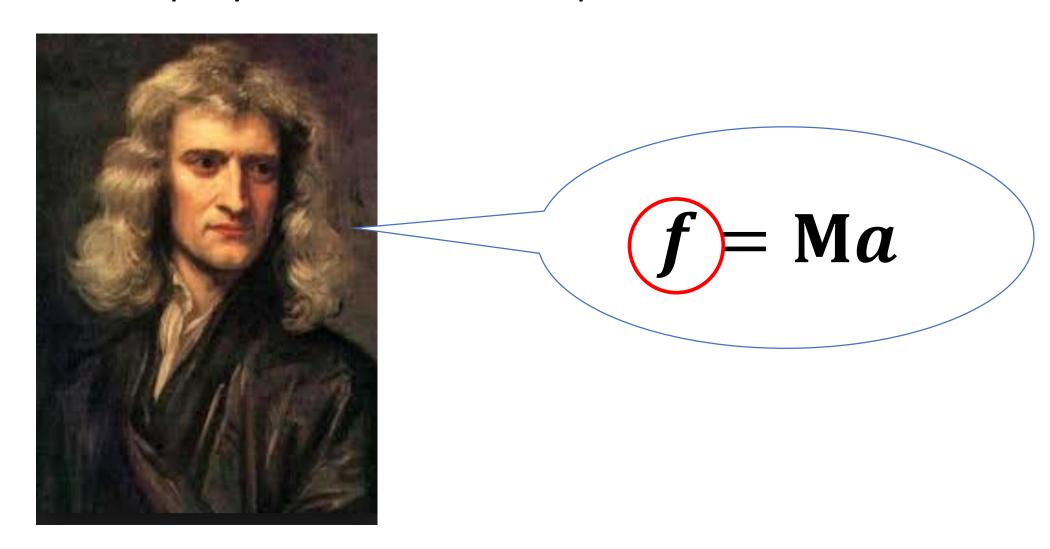


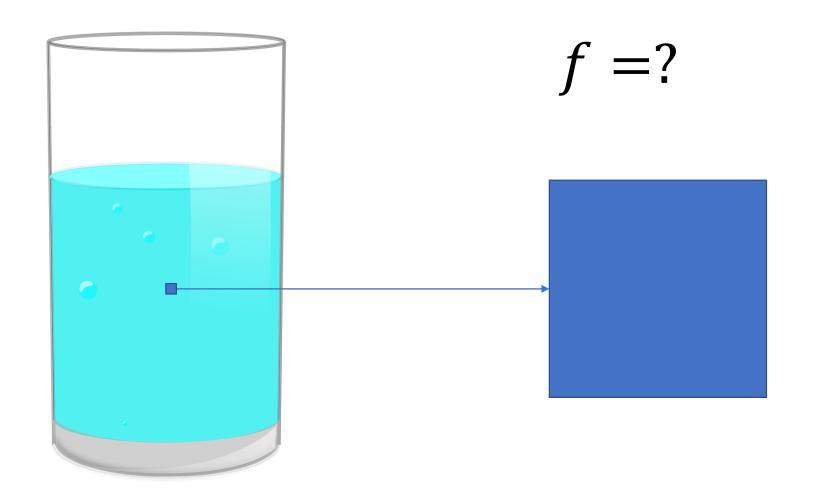
Incompressible smoke

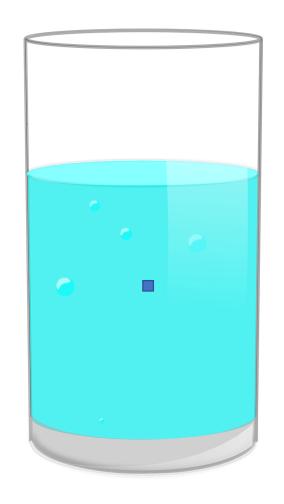


Incompressible liquid

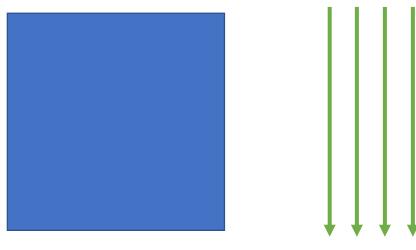
### Laws of physics for incompressible fluids

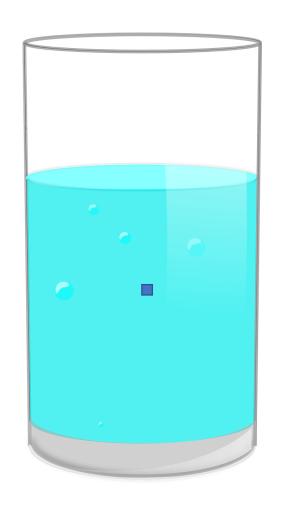


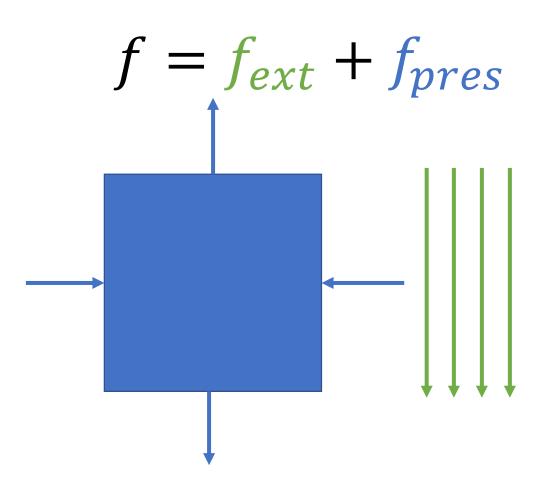


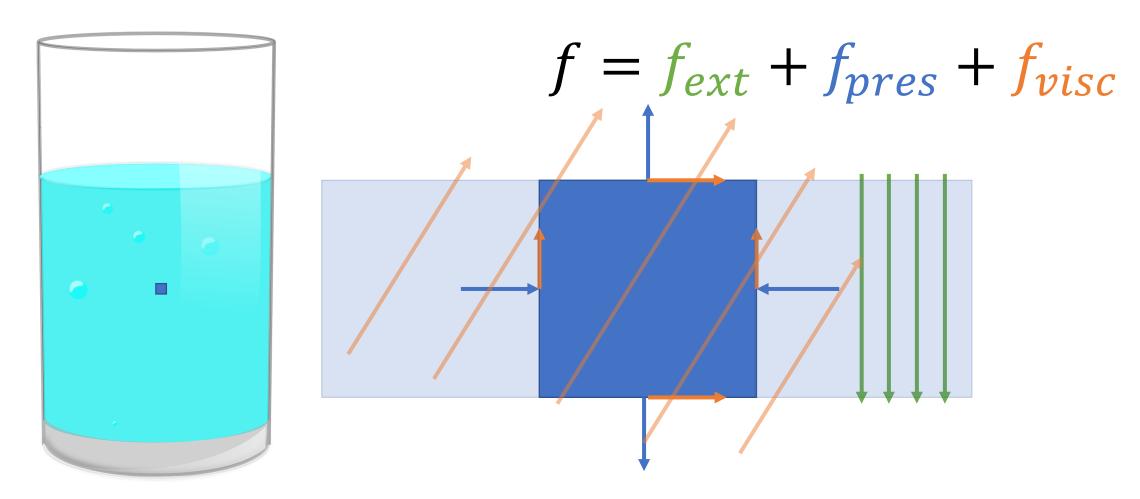


$$f = f_{ext}$$

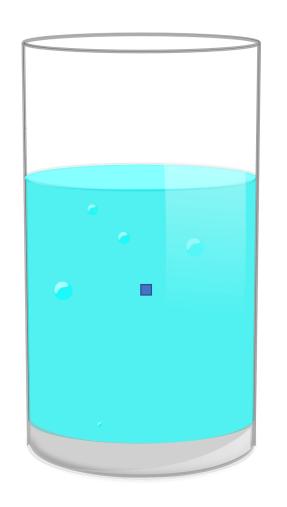


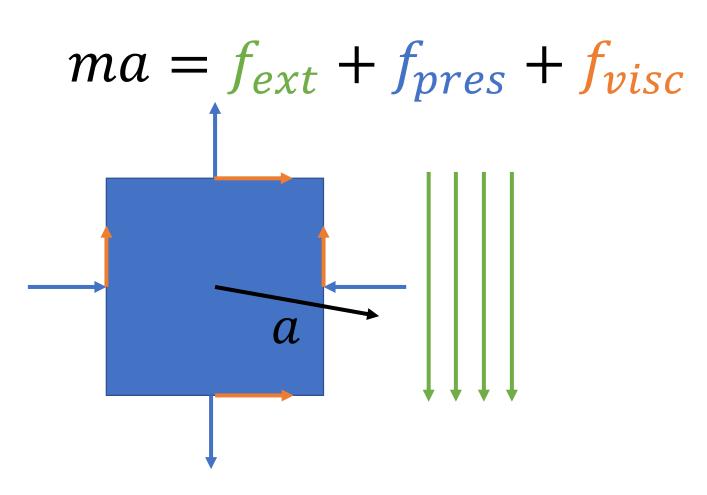




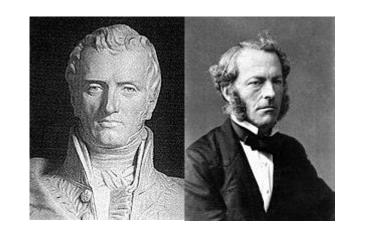


#### Laws of motion for incompressible fluids





#### Incompressible Navier-Stokes equation



$$ma = f_{ext} + f_{pres} + f_{visc}$$

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

# The spatial derivative operators in 3D

$$\nabla = \begin{bmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \\ \frac{\partial}{\partial z} \end{bmatrix}$$

- Gradient  $\nabla : \mathbb{R}^1 \to \mathbb{R}^3$ 
  - $grad s = \nabla s = \left[\frac{\partial s}{\partial x}, \frac{\partial s}{\partial y}, \frac{\partial s}{\partial z}\right]^T$
- Divergence  $\nabla : \mathbb{R}^3 \to \mathbb{R}^1$

• 
$$div \ v = \nabla \cdot v = \frac{\partial v_x}{\partial x} + \frac{\partial v_y}{\partial y} + \frac{\partial v_z}{\partial z}$$

- Curl  $\nabla \times : \mathbb{R}^3 \to \mathbb{R}^3$ 
  - $curl\ v = \nabla \times v = \left[\frac{\partial v_z}{\partial y} \frac{\partial v_y}{\partial z}, \frac{\partial v_z}{\partial x} \frac{\partial v_x}{\partial z}, \frac{\partial v_y}{\partial x} \frac{\partial v_x}{\partial y}\right]^T$
- Laplace  $\Delta = \nabla^2 = \nabla \cdot \nabla : \mathbb{R}^n \to \mathbb{R}^n$ 
  - $laplace \ s = div \ (grad \ s) = \nabla^2 s = \frac{\partial^2 s}{\partial x^2} + \frac{\partial^2 s}{\partial y^2} + \frac{\partial^2 s}{\partial z^2}$

#### Incompressible Navier-Stokes equation

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

 $\rho$ : mass-density

 $\frac{D(\cdot)}{Dt}$ : material derivative

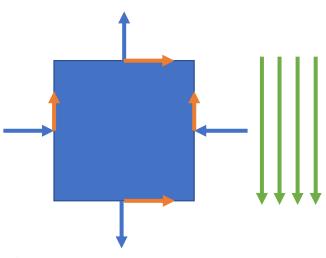
g: gravity  $(g = [0, -9.8, 0]^T \text{ in 3D})$ 

p: pressure,  $p = k(\rho - \rho_0)$ 

 $\mu$ : shear modulus (dynamics visc.),  $\nu = \frac{\mu}{\rho_0}$ : kinematic viscosity

 $\nabla$ : vector differential operator ( $\nabla = \left[\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}\right]^T$  in 3D)

$$abla^2 = \Delta$$
: Laplace operator,  $\Delta f = \nabla \cdot \nabla f \ (\Delta f = \frac{\partial^2}{\partial x^2} f + \frac{\partial^2}{\partial y^2} f + \frac{\partial^2}{\partial z^2} f \text{ in 3D})$ 



# The ma in f = ma

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

This is simply "mass" times "acceleration" divided by "volume"

#### External force term

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

Gravitational force divided by "volume"

#### Viscosity term

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

$$\nabla = \begin{bmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \\ \frac{\partial}{\partial z} \end{bmatrix}$$

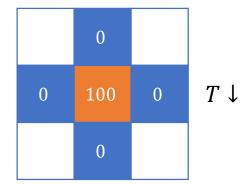
$$\nabla^2 = \Delta: \text{ Laplace operator (or diffusion operator)}$$

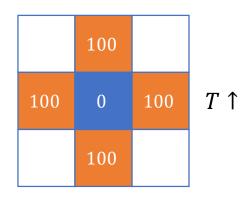
$$\nabla^2: \text{ takes a scalar/vector, returns a scalar/vector}$$

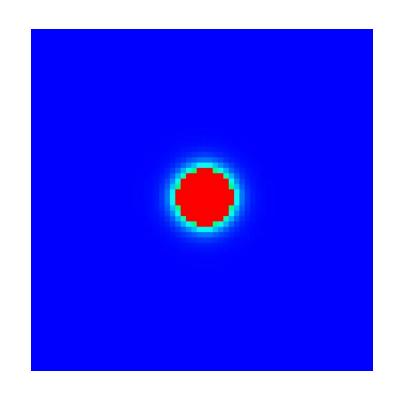
$$\nabla^2 f = \frac{\partial^2}{\partial x^2} f + \frac{\partial^2}{\partial y^2} f + \frac{\partial^2}{\partial z^2} f \text{ in 3D}$$

### Still remember the diffusion problem?

$$\bullet \frac{\partial T}{\partial t} = \kappa \nabla^2 T$$



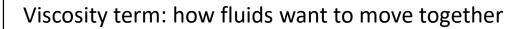


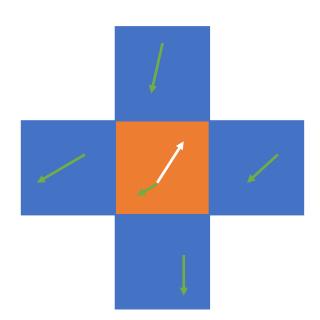


#### Viscosity term

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$





#### Viscosity term

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

$$\mu$$

 $\mu$ : some fluids are more viscous than others

#### Pressure term

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

$$\nabla \cdot gradient operator$$

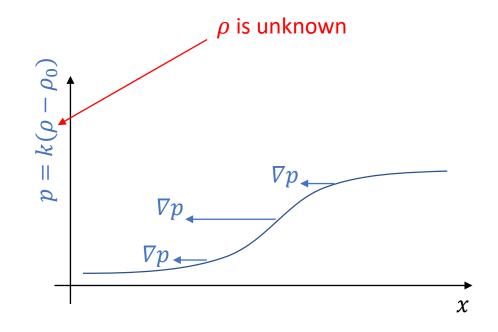
$$\nabla \cdot takes a scalar, returns a vector$$

$$\nabla = \begin{bmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \\ \frac{\partial}{\partial z} \end{bmatrix}$$

Pressure term: fluids do not want to change volume

#### Pressure term

$$\begin{array}{c}
Dv \\
\rho \stackrel{\frown}{=} pg - \nabla p + \mu \nabla^2 v \\
\nabla \cdot v = 0 \\
\nabla : \text{ gradient operator} \\
\nabla : \text{ takes a scalar, returns a vector}
\end{array}$$



Pressure term: fluids do not want to change volume

#### Divergence free condition

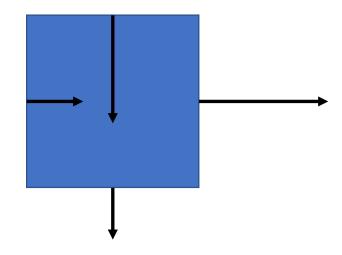
$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

$$\nabla = \begin{bmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \\ \frac{\partial}{\partial z} \end{bmatrix}$$

$$\nabla \cdot : \text{ divergence operator } \nabla \cdot : \text{ takes a vector, returns a scalar } \nabla \cdot : \text{ takes a vector, returns a vector.}$$

$$\nabla \cdot v = \frac{\partial v_x}{\partial x} + \frac{\partial v_y}{\partial y} + \frac{\partial v_z}{\partial z} = 0$$



#### Divergence free condition

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

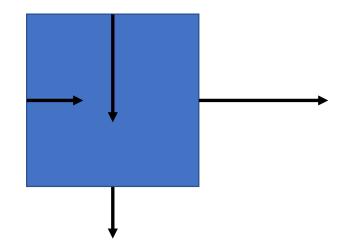
$$\nabla \cdot v = 0 \iff \frac{D\rho}{Dt} = (v \cdot \nabla)\rho = 0$$

Divergence  $(\nabla \cdot v)$  free: outbound flow equals to inbound flow The mass conserving condition

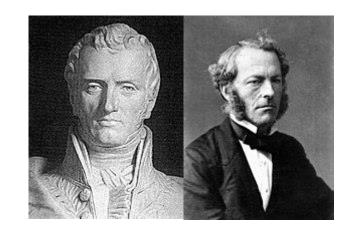
#### Divergence free condition

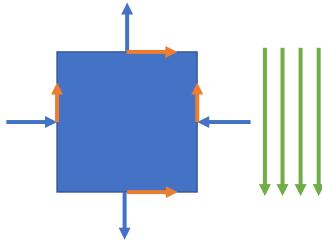
$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$
 
$$\nabla \cdot v = 0 \quad \leftarrow \text{The incompressibility assumption goes here!}$$

Divergence ( $\nabla \cdot v$ ) free: outbound flow equals to inbound flow The mass conserving condition



#### Incompressible Navier-Stokes equation

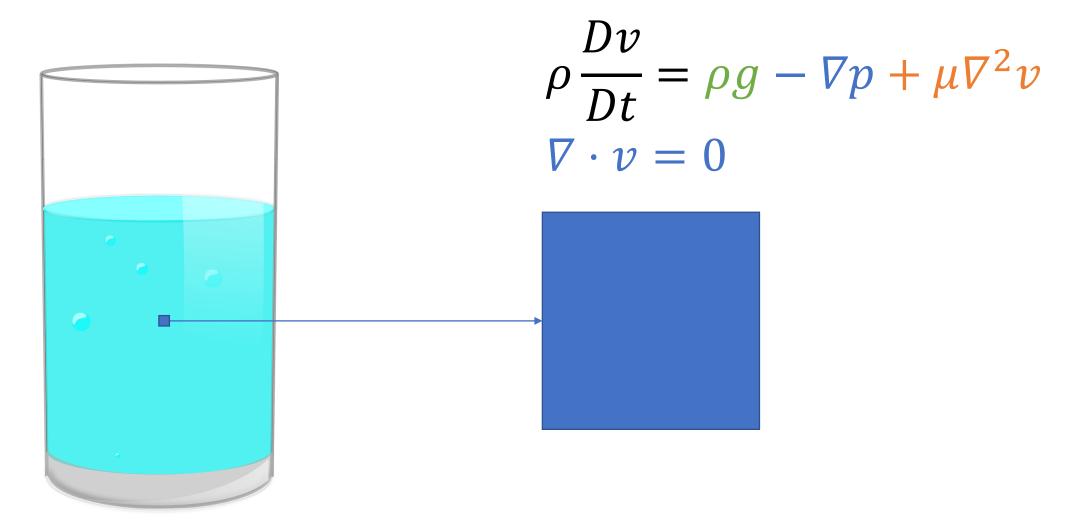




$$ho \, rac{D v}{D t} = 
ho g - 
abla p + \mu 
abla^2 v ext{ } ext{ ← The momentum equation}$$
  $abla v = 0 ext{ } ext{ ← The mass conserving cond}$ 

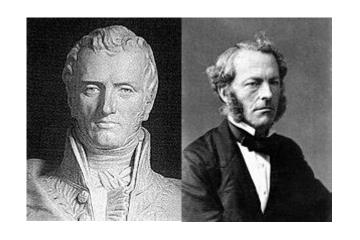
←The mass conserving condition

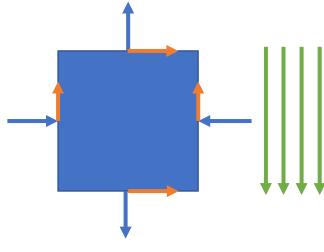
#### Integrate incompressible Navier-Stokes equation?



Temporal discretization

#### Incompressible Navier-Stokes equation





$$ho \, rac{D v}{D t} = 
ho g - rac{1}{V p} + \mu V^2 v \;$$
  $\leftarrow$  The momentum equation

$$\nabla \cdot \nu = 0$$

←The mass conserving condition

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

• The divergence free condition  $\nabla \cdot v = 0$  comes to the rescue:

• 
$$\nabla \cdot v = 0 \iff \frac{D\rho}{Dt} = 0$$

- Integrate the incompressible Navier-Stokes equation in steps:
  - Step 1: input  $v_n$ , output  $v_{n+0.5}$

• 
$$\rho \frac{Dv}{Dt} = \rho g + \mu \nabla^2 v$$

- Step 2: input  $v_{n+0.5}$ , output  $v_{n+1}$ 
  - $\rho \frac{Dv}{Dt} = -\nabla p$
  - $\nabla \cdot v = 0$

This integration method is sometime referred as "Operator splitting" or "Advection-Projection" in different contexts

#### Full time integration

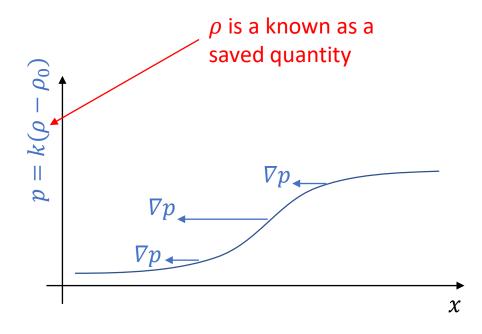
- Given  $x_n$ ,  $v_n$ :
  - Step 1: Advection / external and viscosity force integration
    - Solve:  $dv = g + \nu \nabla^2 v_n$
    - Update:  $v_{n+0.5} = v_n + \Delta t \ dv$
  - Step 2: Projection / pressure solver
    - Solve:  $\frac{dv}{dv} = -\frac{1}{\rho} \nabla \left( k(\rho \rho_0) \right)$  and  $\frac{D\rho}{Dt} = \nabla \cdot (v_{n+0.5} + dv) = 0$
    - Update:  $v_{n+1} = v_{n+0.5} + \Delta t \, dv$
  - Step 3: Update position
    - Update:  $x_{n+1} = x_n + \Delta t \ v_{n+1}$
  - Return  $x_{n+1}$ ,  $v_{n+1}$

#### The weakly compressible assumption

• Storing the density  $\rho$  as an individual variable that advect with the velocity field:

$$\frac{Dv}{Dt} = g - \frac{1}{\rho}\nabla p + v\nabla^2 v$$

$$\frac{\nabla \cdot v = 0}{\nabla v}$$



#### Integrate with the weakly compressible assumption

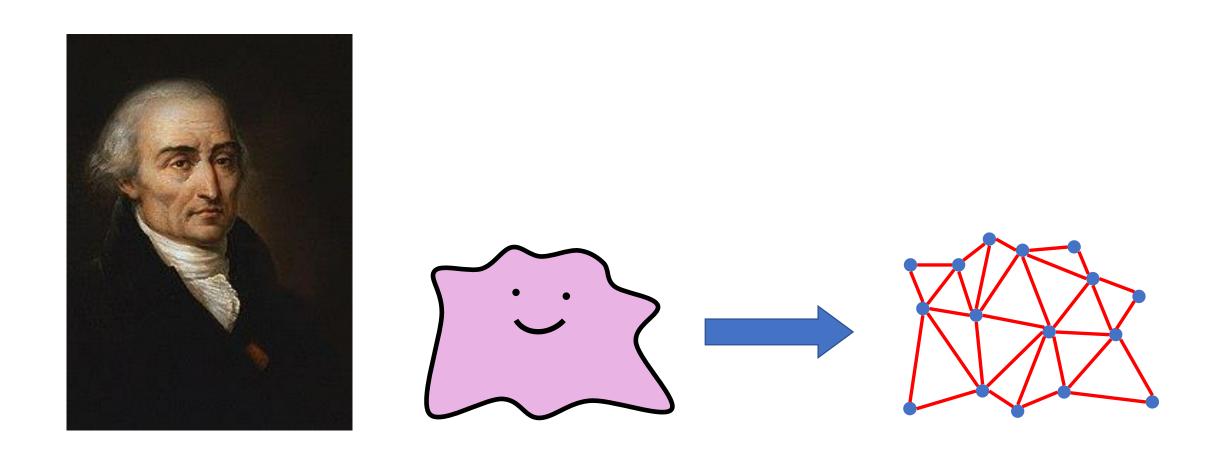
$$\frac{Dv}{Dt} = g - \frac{1}{\rho}\nabla p + v\nabla^2 v$$

- Given  $x_n$ ,  $v_n$ :
  - Step 1: Advection / external and viscosity force integration
    - Solve:  $dv = g + \nu \nabla^2 v_n$
    - Update:  $v_{n+0.5} = v_n + \Delta t \ dv$
  - Step 2: Projection / pressure solver
    - Solve:  $\frac{dv}{dv} = -\frac{1}{\rho} \nabla (k(\rho \rho_0))$
    - Update:  $v_{n+1} = v_{n+0.5} + \Delta t \, dv$
  - Step 3: Update position
    - Update:  $x_{n+1} = x_n + \Delta t \ v_{n+1}$
  - Return  $x_{n+1}$ ,  $v_{n+1}$

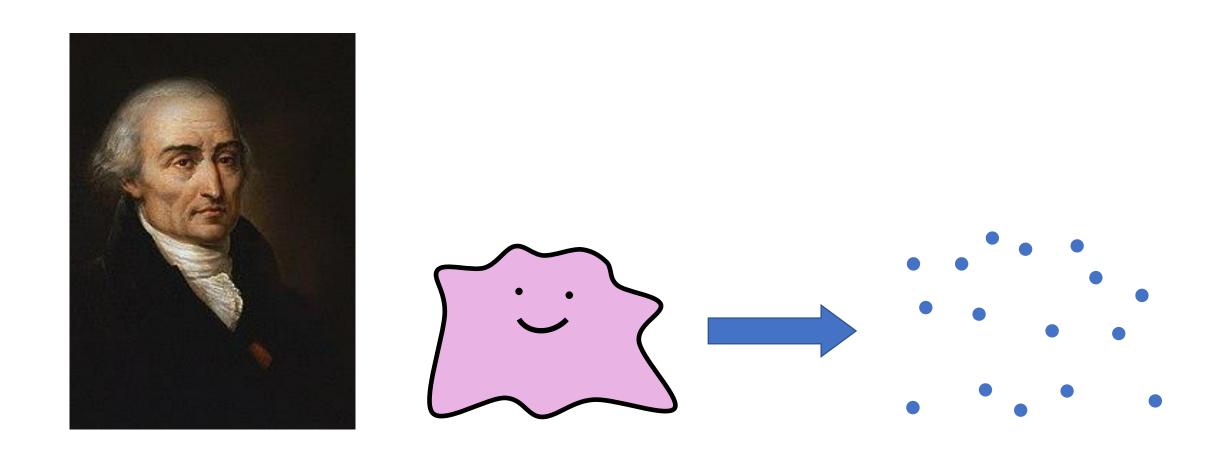
This is nothing but *Symplectic Euler* integration.

Spatial discretization (Lagrangian view)

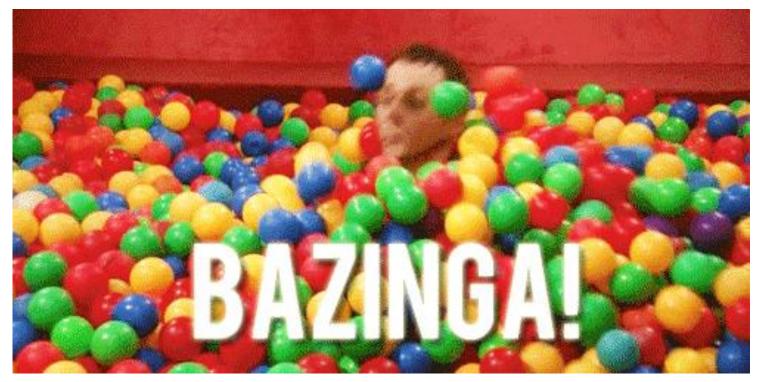
## Previously in this course (mesh-based simulation)



## Today in this course (mesh-free simulation)

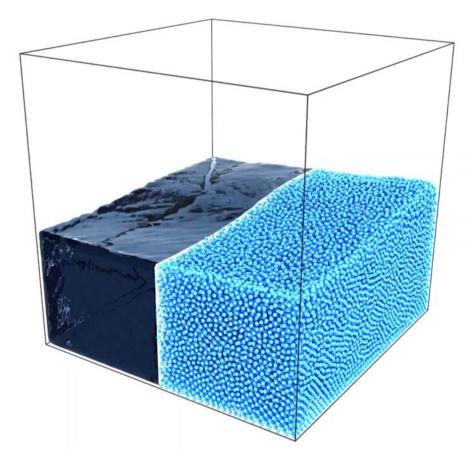


## Fluid can be discretized using particles as well



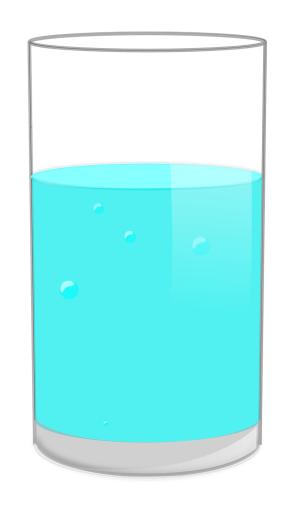
[The Big Bang Theory]

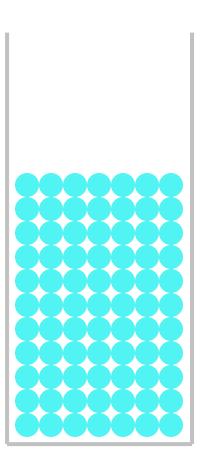
# Fluid can be discretized using particles as well



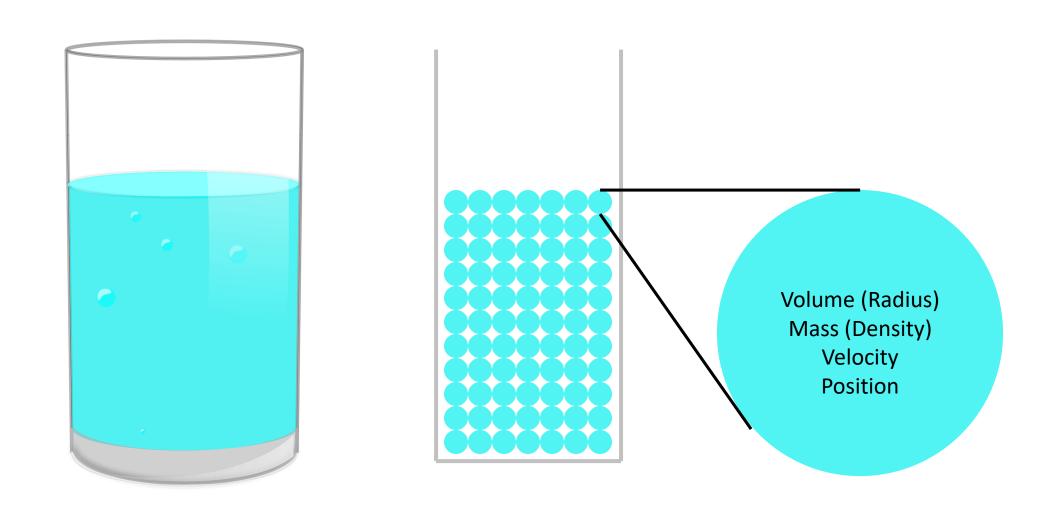
[www.dive-solutions.de]

## Fluid -> Particles

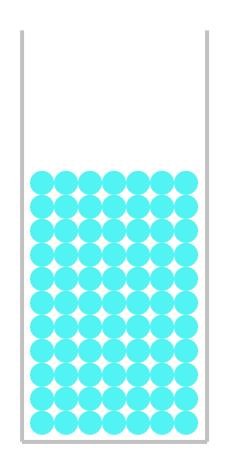




# What do we store in a particle?



#### Fluid dynamics with particles (weakly compressible)



Continuous view:

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

Discrete view (using particles):

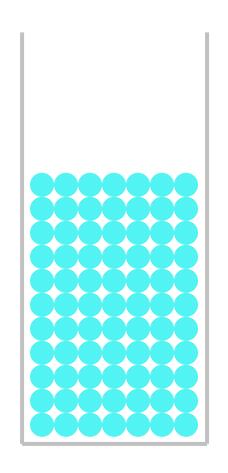
$$\frac{dv_i}{dt} = g - \frac{1}{\rho} \nabla p(x_i) + \nu \nabla^2 v(x_i) \quad , \text{ where } \nu = \frac{\mu}{\rho_0}$$

Time integration (Symplectic Euler):

 $a_i$ 

- for i in particles:
  - $v_i = v_i + \Delta t \ a_i$
- for i in particles:
  - $x_i = x_i + \Delta t \ v_i$

#### Fluid dynamics with particles (weakly compressible)



Continuous view:

$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

Discrete view (using particles):

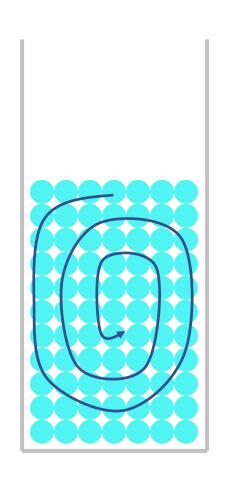
$$\frac{dv_i}{dt} = g - \frac{1}{\rho} \nabla p(x_i) + \nu \nabla^2 v(x_i) \quad , \text{ where } v = \frac{\mu}{\rho_0}$$

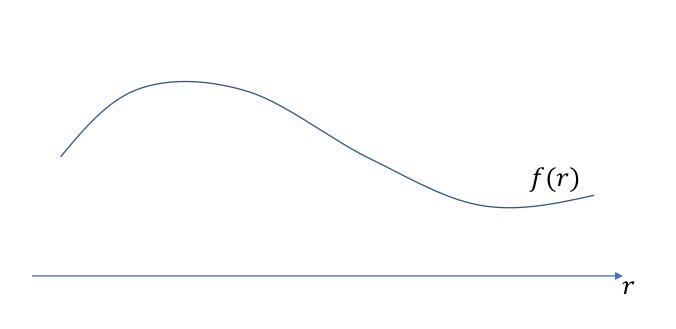
 $a_i$ 

Time integration (Symplectic Euler):

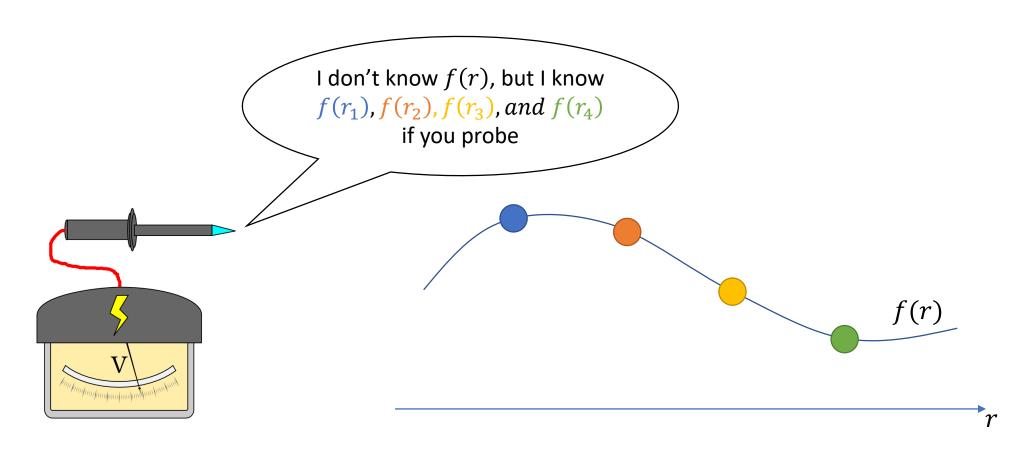
- for i in particles:
  - $v_i = v_i + \Delta t \ a_i$
- for i in particles:
  - $x_i = x_i + \Delta t \ v_i$

# How to evaluate a function? $\rho(x_i) \nabla p(x_i) \nabla^2 v(x_i)$





# Imaging a magic probe...

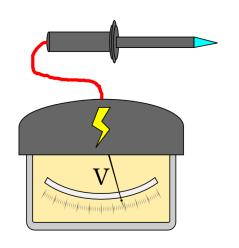


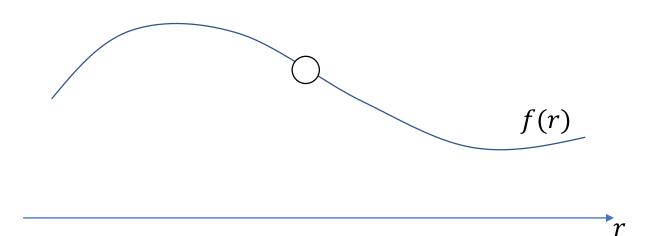
## A trivial identity with Dirac delta

$$f(r) = \int_{-\infty}^{\infty} f(r')\delta(r - r')dr'$$

$$f(r) = \int_{-\infty}^{\infty} f(r')\delta(r - r')dr'$$

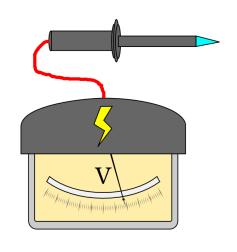
$$\delta(r) = \begin{cases} +\infty, & \text{if } r = 0 \\ 0, & \text{otherwise} \end{cases} \text{ and } \int_{-\infty}^{\infty} \delta(r)dr = 1$$

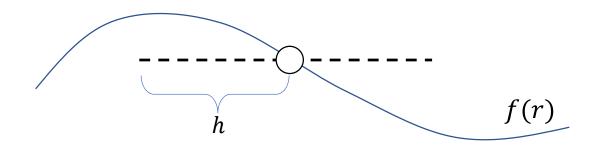




#### Let us widen the Dirac delta

$$f(r) pprox \int f(r')W(r-r',h)dr'$$
 , where  $\lim_{h\to 0} W(r,h) = \delta(r)$ 



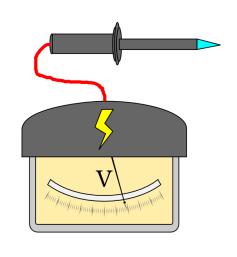


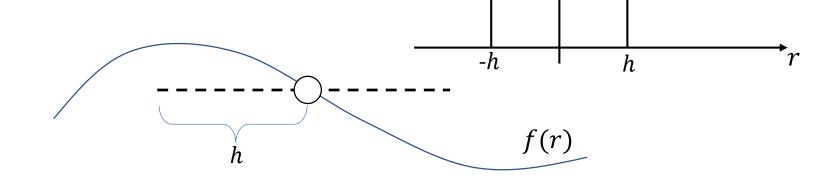
W(r,h): kernel function

- Sum to unity:  $\int W(r,h)dr = 1$
- Symmetric: W(r,h) = W(-r,h)
- Compact support: W(r,h) = 0 if |r| > h

#### A trivial kernel function

$$f(r) pprox \int f(r') W(r-r',h) dr'$$
 , where  $\lim_{h o 0} W(r,h) = \delta(r)$ 

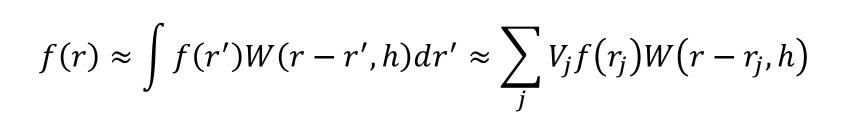


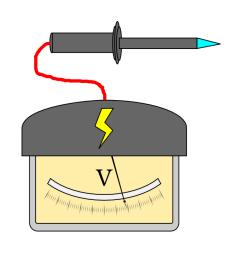


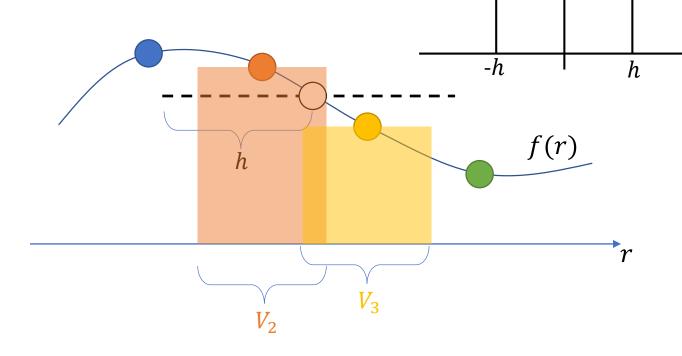
$$W(r,h) = \begin{cases} \frac{1}{2h}, & \text{if } |r| < h \\ 0, & \text{otherwise} \end{cases}$$

W(r,h)

## Finite probes: from integration to summation



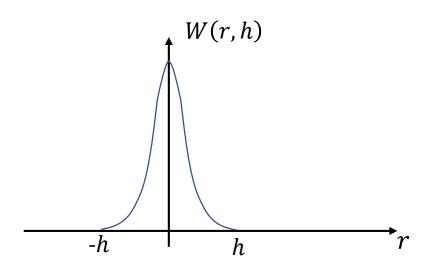


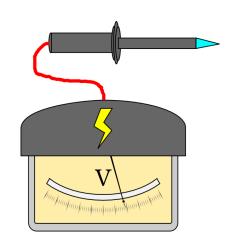


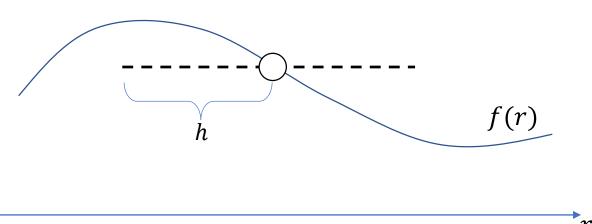
W(r,h)

#### A smoother kernel function

$$f(r) \approx \sum_{j} V_{j} f(r_{j}) W(r - r_{j}, h)$$



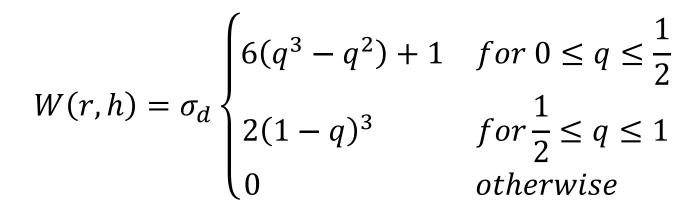




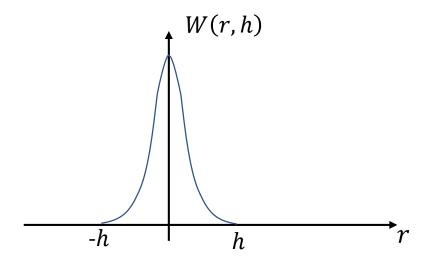
Smooth W(r, h): "we trust the closer probes better"

#### A smoother kernel function

$$f(r) \approx \sum_{j} V_{j} f(r_{j}) W(r - r_{j}, h)$$



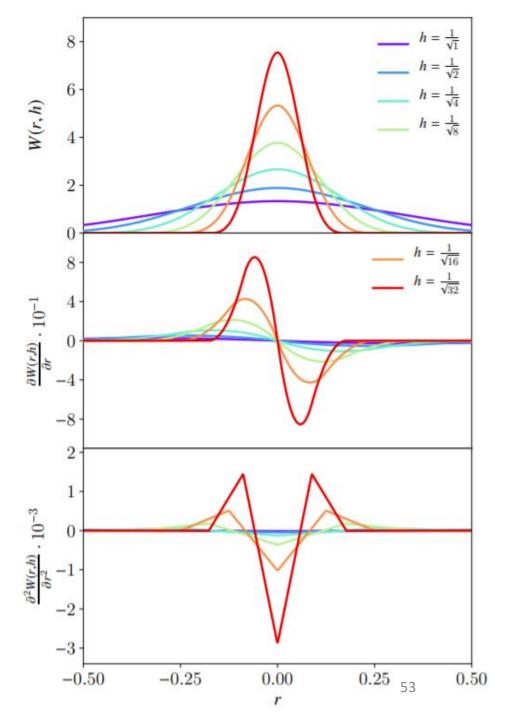
with 
$$q = \frac{1}{h} ||r||$$
,  $\sigma_1 = \frac{4}{3h}$ ,  $\sigma_2 = \frac{40}{7\pi h^2}$ ,  $\sigma_3 = \frac{8}{\pi h^3}$ 



## A smoother kernel function

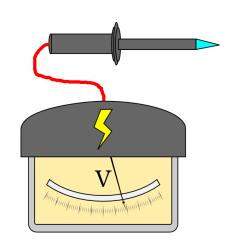
$$W(r,h) = \sigma_d \begin{cases} 6(q^3 - q^2) + 1 & for \ 0 \le q \le \frac{1}{2} \\ 2(1 - q)^3 & for \frac{1}{2} \le q \le 1 \\ 0 & otherwise \end{cases}$$

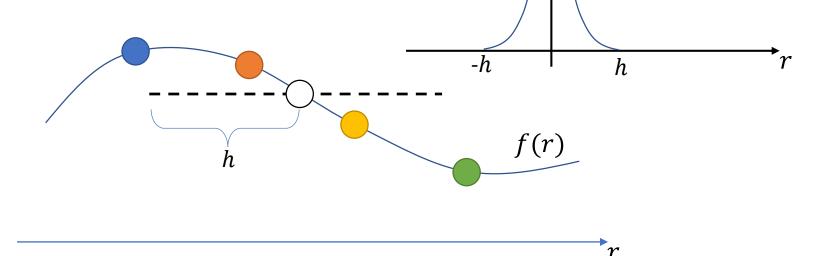
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## A Smoothed particle

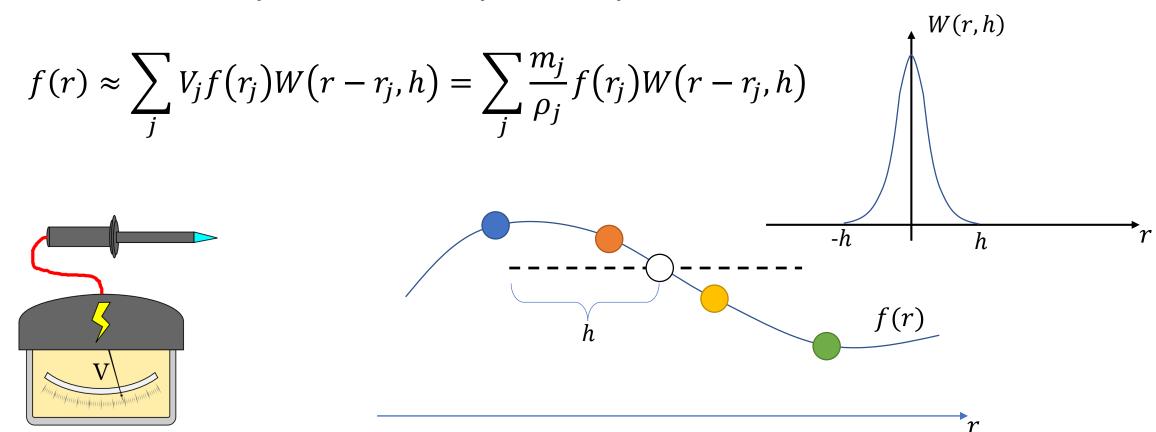
$$f(r) \approx \sum_{j} V_{j} f(r_{j}) W(r - r_{j}, h)$$





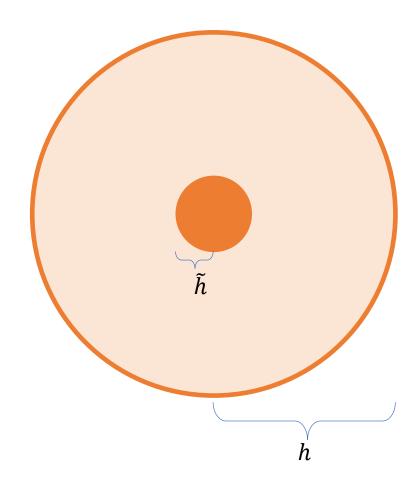
W(r,h)

# Smoothed particle hydrodynamics (SPH)



Smoothed particle hydrodynamics: theory and application to non-spherical stars [Gingold and Monaghan 1977][Link]

## A smoothed particle in 2D



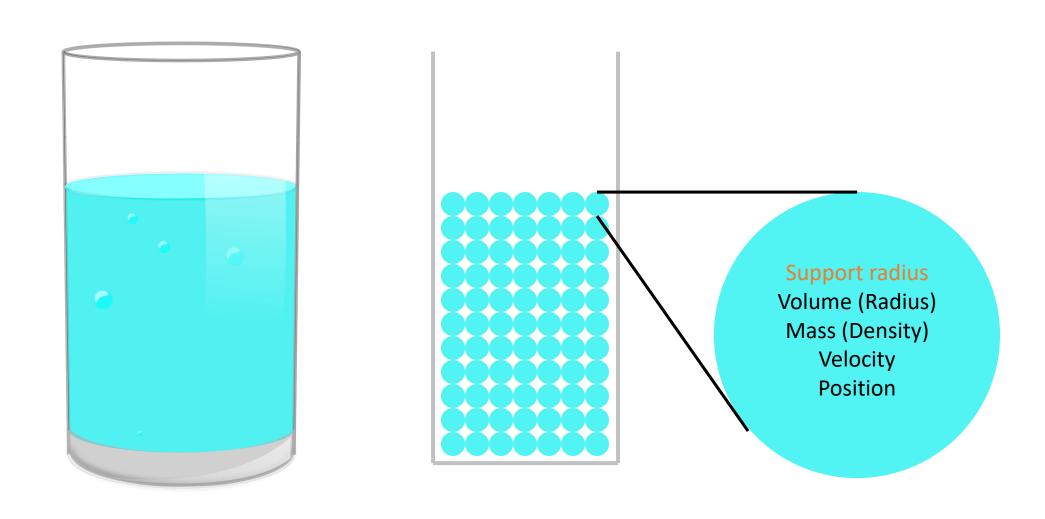
#### Intrinsic quantities:

- *h*: support radius
- $\tilde{h}$ : particle radius -> V: particle volume

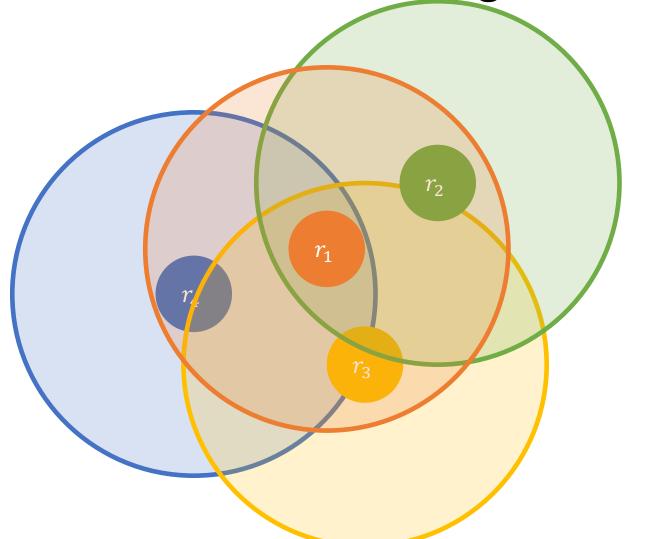
#### Time varying quantities:

- $\rho$ : density
- *v*: velocity
- *x*: position

## What do we store in a smoothed particle?



Evaluate 2D fields using the smoothed particles



$$f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) W(r - r_{j}, h)$$

$$f(r_1)$$

$$\approx \frac{m_2}{\rho_2} f(r_2) W(r_1 - r_2, h)$$

$$+ \frac{m_3}{\rho_3} f(r_3) W(r_1 - r_3, h)$$

$$+ \frac{m_4}{\rho_4} f(r_4) W(r_1 - r_4, h)$$

## Smoothed particle hydrodynamics (SPH)

#### SPH discretization:

• 
$$f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) W(r - r_{j}, h)$$

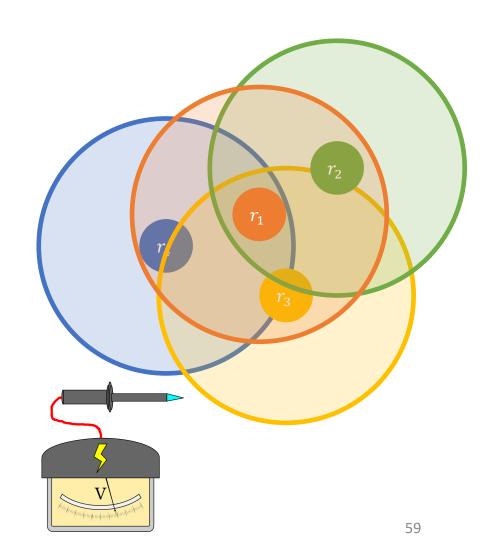
#### • SPH spatial derivatives:

• 
$$\nabla f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \nabla W(r - r_{j}, h)$$

• 
$$\nabla \cdot \mathbf{F}(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} \mathbf{F}(r_{j}) \cdot \nabla W(r - r_{j}, h)$$

• 
$$\nabla \times \mathbf{F}(r) \approx -\sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \times \nabla W(r - r_{j}, h)$$

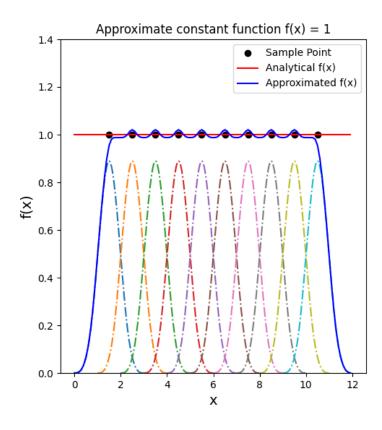
• 
$$\nabla^2 f(r) \approx \sum_j \frac{m_j}{\rho_j} f(r_j) \nabla^2 W(r - r_j, h)$$



• 
$$f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) W(r - r_{j}, h)$$

• 
$$\nabla f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \nabla W(r - r_{j}, h)$$

- Let  $f(r) \equiv 1$ , we have:
  - $1 \approx \sum_{j} \frac{m_{j}}{\rho_{i}} W(r r_{j}, h)$
  - $0 \approx \sum_{j} \frac{m_{j}}{\rho_{j}} \nabla W(r r_{j}, h)$



- Since  $f(r) \equiv f(r) * 1$ , we have:
  - $\nabla f(r) = \nabla f(r) * 1 + f(r) * \nabla 1$
  - Or equivalently:  $\nabla f(r) = \nabla f(r) f(r) * \nabla 1$

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  - Or equivalently:  $\nabla f(r) = \nabla f(r) f(r) * \nabla 1$

• 
$$\nabla f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \nabla W(r - r_{j}, h) - f(r) \sum_{j} \frac{m_{j}}{\rho_{j}} \nabla W(r - r_{j}, h)$$

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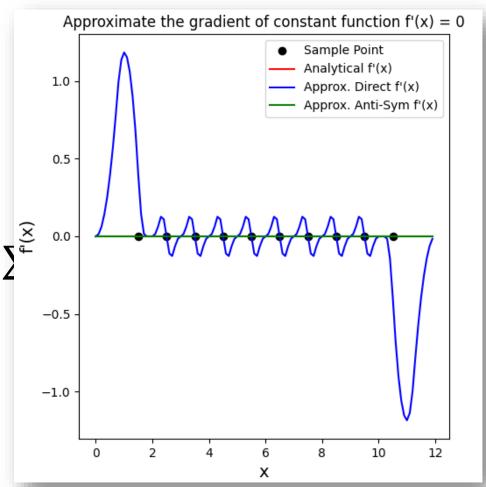
• 
$$\nabla f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \nabla W(r - r_{j}, h) - f(r) \sum_{j} \frac{m_{j}}{\rho_{j}} \nabla W(r - r_{j}, h)$$

• 
$$\nabla f(r) \approx \sum_{j} m_{j} \frac{f(r_{j}) - f(r)}{\rho_{j}} \nabla W(r - r_{j}, h)$$

The anti-symmetric form

- Since  $f(r) \equiv f(r) * 1$ , we have:
  - $\nabla f(r) = \nabla f(r) * 1 + f(r) * \nabla 1$
  - Or equivalently:  $\nabla f(r) = \nabla f(r) f(r) * \nabla 1$
- $\nabla f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \nabla W(r r_{j}, h) f(r) \sum_{k}^{3}$
- $\nabla f(r) \approx \sum_{j} m_{j} \frac{f(r_{j}) f(r)}{\rho_{j}} \nabla W(r r_{j}, h)$

The anti-symmetric form



A more general case:

• 
$$\nabla (f(r)\rho^n) = \nabla f(r) * \rho^n + f(r) * n\rho^{n-1}\nabla \rho$$

• 
$$\Rightarrow \nabla f(r) = \frac{1}{\rho^n} (\nabla (f(r) * \rho^n) - f(r) * n * \rho^{n-1} \nabla \rho)$$

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$$\nabla f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) \nabla W(r - r_{j}, h)$$

#### A more general case:

• 
$$\nabla (f(r)\rho^n) = \nabla f(r) * \rho^n + f(r) * n\rho^{n-1}\nabla \rho$$

• 
$$\Rightarrow \nabla f(r) = \frac{1}{\rho^n} (\nabla (f(r) * \rho^n) - f(r) * n * \rho^{n-1} \nabla \rho)$$

• 
$$\nabla f(r) \approx \frac{1}{\rho^n} \left( \sum_j \frac{m_j}{\rho_j} f(r_j) \rho_j^n \nabla W(r - r_j, h) - n \rho^{n-1} f(r) \sum_j \frac{m_j}{\rho_j} \rho_j \nabla W(r - r_j, h) \right)$$

#### • A more general case:

• 
$$\nabla (f(r)\rho^n) = \nabla f(r) * \rho^n + f(r) * n\rho^{n-1}\nabla \rho$$

• 
$$\Rightarrow \nabla f(r) = \frac{1}{\rho^n} (\nabla (f(r) * \rho^n) - f(r) * n * \rho^{n-1} \nabla \rho)$$

• 
$$\nabla f(r) \approx \frac{1}{\rho^n} \left( \sum_j \frac{m_j}{\rho_j} f(r_j) \rho_j^n \nabla W(r - r_j, h) - n \rho^{n-1} f(r) \sum_j \frac{m_j}{\rho_j} \rho_j \nabla W(r - r_j, h) \right)$$

• 
$$\nabla f(r) \approx \sum_{j} m_{j} \left( \frac{f(r_{j})\rho_{j}^{n-1}}{\rho^{n}} - \frac{nf(r)}{\rho} \right) \nabla W(r - r_{j}, h)$$

• 
$$\nabla f(r) \approx \sum_{j} m_{j} \left( \frac{f(r_{j})\rho_{j}^{n-1}}{\rho^{n}} - \frac{nf(r)}{\rho} \right) \nabla W(r - r_{j}, h)$$

- When n=-1
  - Type equation here.

• 
$$\nabla f(r) \approx \rho \sum_{j} m_{j} \left( \frac{f(r_{j})}{\rho_{j}^{2}} + \frac{f(r)}{\rho^{2}} \right) \nabla W(r - r_{j}, h)$$

The symmetric form

## Smoothed particle hydrodynamics (SPH)

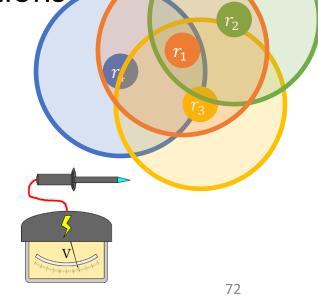
$$f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) W(r - r_{j}, h)$$

• Approximate a function f(r) using finite probes  $f(r_j)$ 

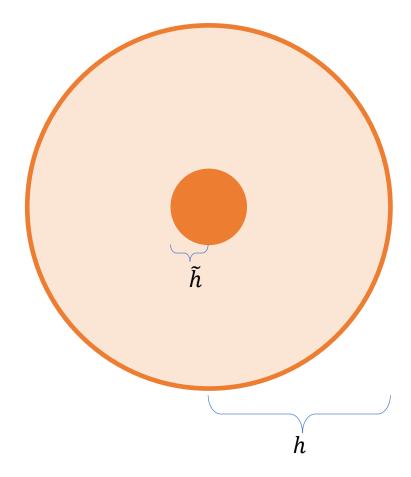
• The degree of freedom (r) goes inside the kernel functions

• Anti-sym:  $\nabla f(r) \approx \sum_{j} m_{j} \frac{f(r_{j}) - f(r)}{\rho_{j}} \nabla W(r - r_{j}, h)$ 

• Sym:  $\nabla f(r) \approx \rho \sum_{j} m_{j} \left( \frac{f(r_{j})}{\rho_{j}^{2}} + \frac{f(r)}{\rho^{2}} \right) \nabla W(r - r_{j}, h)$ 



## Quiz: which one is the smoothed particle in SPH?



The bigger circle?

or

The smaller circle?

#### Intrinsic quantities:

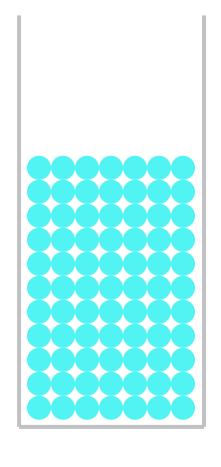
- *h*: support radius
- $ilde{h}$ : particle radius -> V: particle volume

#### Time varying quantities:

- $\rho$ : density
- v: velocity
- *x*: position

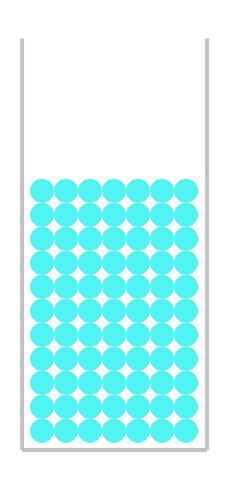
Implementation details (WCSPH)

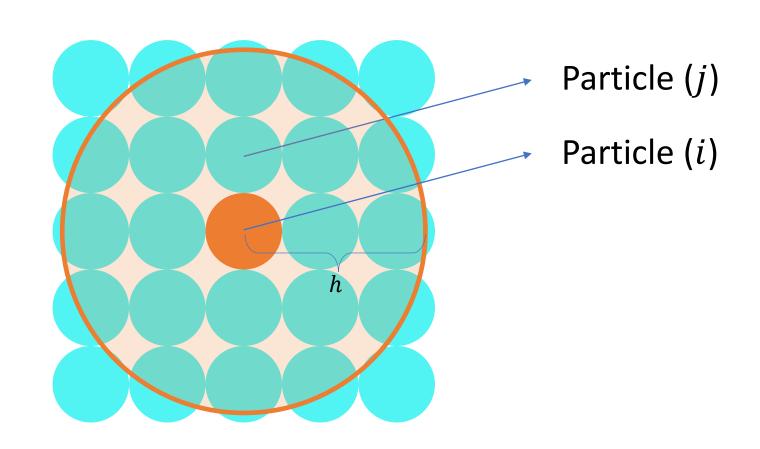
### Simulation pipeline



- for i in particles:
  - Search for neighbors j
- for i in particles:
  - Sample the velocity/density field using SPH
  - Compute force/acceleration using Navier-Stokes equation
- for i in particles:
  - Update velocity using acceleration
  - Update position using velocity

# Find a particle of interest (i) and its neighbors (j) within its support radius h





### Compute the acceleration for particle (i)

- for i in particles:
  - Step 1: Evaluate density

• 
$$\rho_i = \sum_j \frac{m_j}{\rho_j} \rho_j W(r_i - r_j, h) = \sum_j m_j W_{ij}$$

Step 2: Evaluate viscosity

• 
$$\nu \nabla^2 v_i = \nu \sum_j m_j \frac{v_j - v_i}{\rho_j} \nabla^2 W_{ij}$$

• Step 3: Evaluate pressure gradient

• 
$$-\frac{1}{\rho_i}\nabla p_i = -\frac{\rho_i}{\rho_i}\sum_j m_j\left(\frac{p_j}{\rho_j^2} + \frac{p_i}{\rho_i^2}\right)\nabla W_{ij}$$
, where  $p = k(\rho_j - \rho_0)$ 

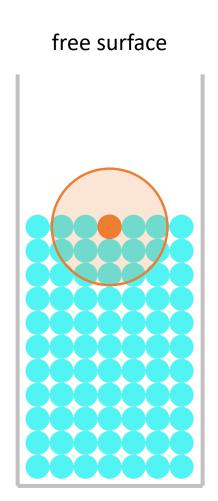
• 
$$\frac{dv_i}{dt} = g - \frac{1}{\rho_i} \nabla p_i + \nu \nabla^2 v_i$$

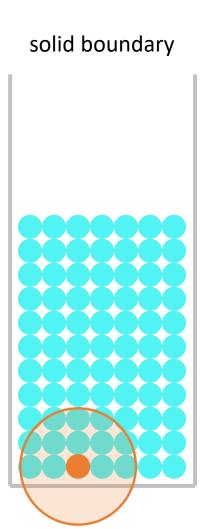
### Time integration (Symplectic Euler)

- for i in particles:
  - $v_i = v_i + \Delta t * \frac{dv_i}{dt}$
  - $x_i = x_i + \Delta t * v_i$

# Boundary conditions

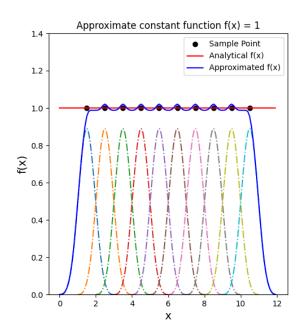
fluid

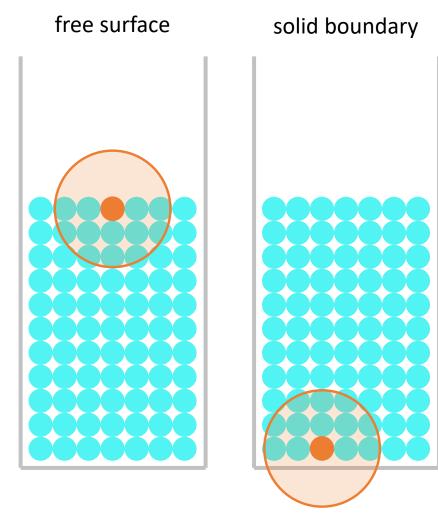




#### Problems of boundaries

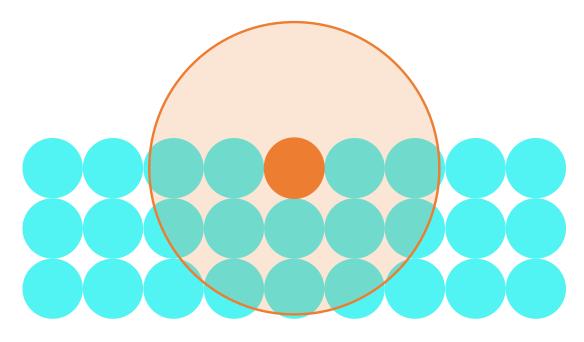
- Not enough samples within the supporting radius
  - Density: ↓
    - $\rho_i = \sum_j m_j W_{ij}$





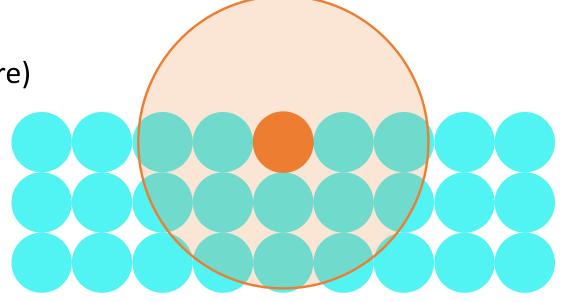
#### Free surface

- Problem:
  - Density ↓ Pressure ↓
  - Generate outward pressure



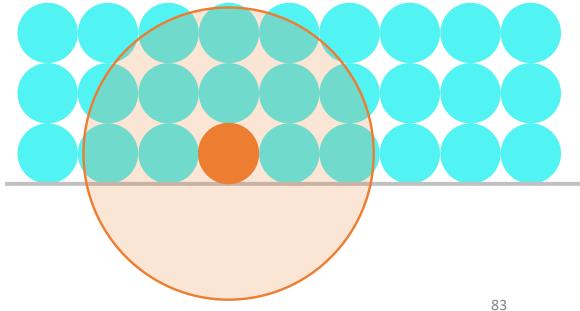
#### Free surface

- Problem:
  - Density ↓ Pressure ↓
  - Generate outward pressure
- Solution:
  - Clamp the negative pressure (everywhere)
  - $p = \max(0, k(\rho \rho_0))$



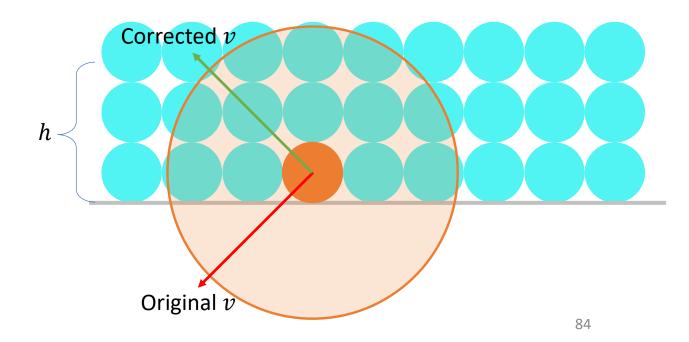
### Solid boundary

- Problem:
  - Density ↓ Pressure ↓
  - Fluid leakage (due to outbound velocity)



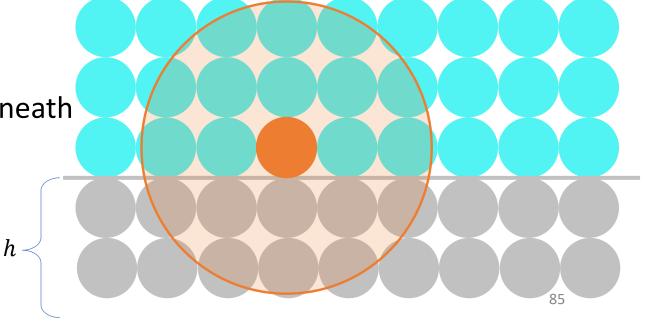
### Solid boundary

- Problem:
  - Density ↓ Pressure ↓
  - Fluid leakage (due to outbound velocity)
- $p = \max(0, k(\rho \rho_0))$
- Solution 1 for leakage:
  - Reflect the outbound velocity when close to boundary



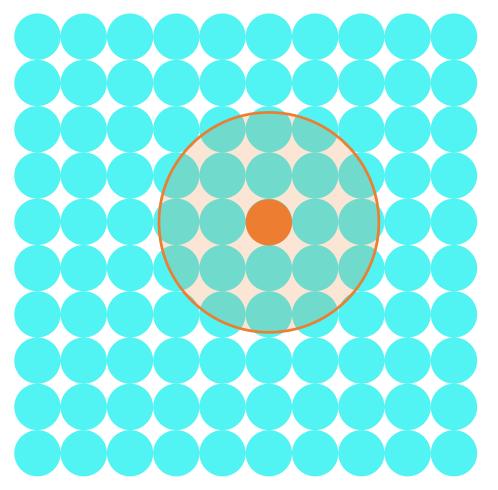
### Solid boundary

- Problem:
  - Density ↓ Pressure ↓
  - Fluid leakage (due to outbound velocity)
- $p = \max(0, k(\rho \rho_0))$
- Solution 2 for leakage:
  - Pad a layer of solid particles underneath the boundaries
    - $\rho_{solid} = \rho_0$
    - $v_{solid} = 0$



### Neighbor search

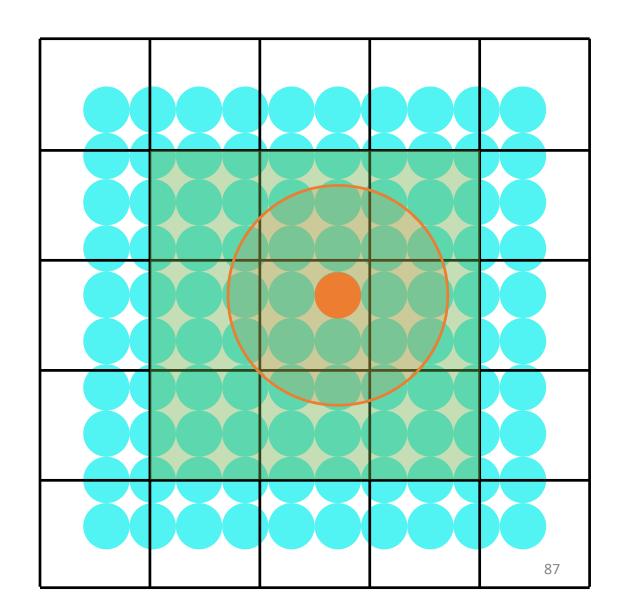
• Naïve search methods takes  $\mathcal{O}(n^2)$  time



### Neighbor search

• Naïve search methods takes  $\mathcal{O}(n^2)$  time

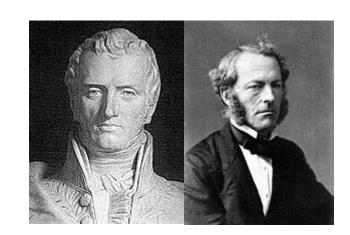
- A background grid can help
  - Common grid size = h (the support radius in SPH)
  - Each neighbor search takes 9 grids in 2D and 27 grids in 3D



- Incompressible fluid dynamics
  - Incompressible Navier–Stokes equations
- Time discretization
  - Operator splitting
  - Integration with the weakly compressible assumption



- Smoothed particle hydrodynamics (SPH)
- Implementation details (WCSPH)
  - Simulation Pipeline
  - Boundary conditions
  - Neighbor search



$$\rho \frac{Dv}{Dt} = \rho g - \nabla p + \mu \nabla^2 v$$

$$\nabla \cdot v = 0$$

- Incompressible fluid dynamics
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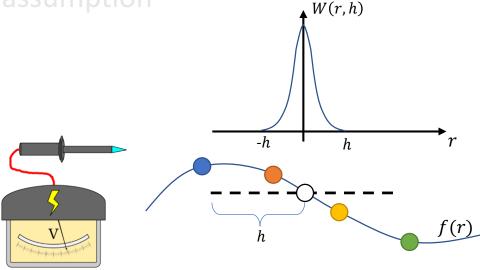
$$\rho \frac{Dv}{Dt} = \rho g + \mu \nabla^2 v$$

$$\rho \frac{Dv}{Dt} = -\nabla p$$

$$\nabla \cdot v = 0$$

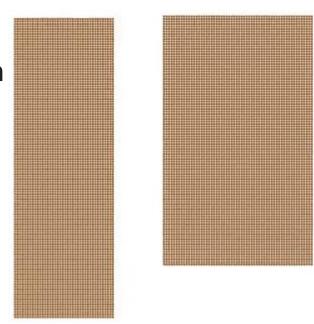
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$$f(r) \approx \sum_{j} \frac{m_{j}}{\rho_{j}} f(r_{j}) W(r - r_{j}, h)$$



- Incompressible fluid dynamics
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### Further readings

- Smoothed Particle Hydrodynamics [Monaghan 2005][Link]
- Smoothed particle hydrodynamics and magnetohydrodynamics [Price 2012][Link][Preprint]
- Smoothed Particle Hydrodynamics Techniques for the Physics Based Simulation of Fluids and Solids [Eurographics Tutorial, 2019][Link][Code]

### Next week





Homework

### Homework Today

- Download the repo (taichi\_sph):
  - https://github.com/taichiCourse01/taichi\_sph
- Try:
  - Designing your own scene
  - Implementing a particle based boundary handling
  - Changing the dense grid in the codebase to a sparse grid [03讲]

### Candidate projects for your final

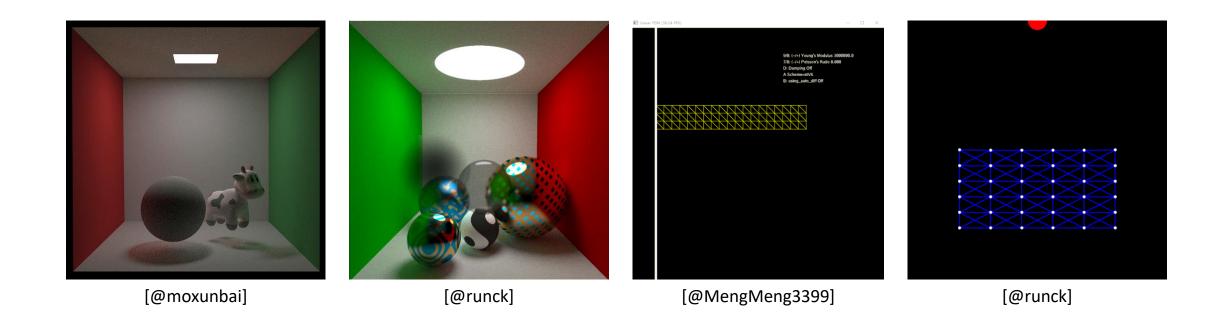
- Candidate topics:
  - Try advanced pressure (Poisson) solvers: IISPH/PCISPH/DFSPH [Chapter 4]
  - Put a statue/kinematically-controlled fan into your pond (One way coupling)
  - Throw a rubber duck into your pond (Two way coupling)
  - Render your fluid using your own path tracer (You may want a marching cube/square to construct the water surface)

- Both 2D and 3D projects are great!
  - As long as your pictures look great ☺

### Final project

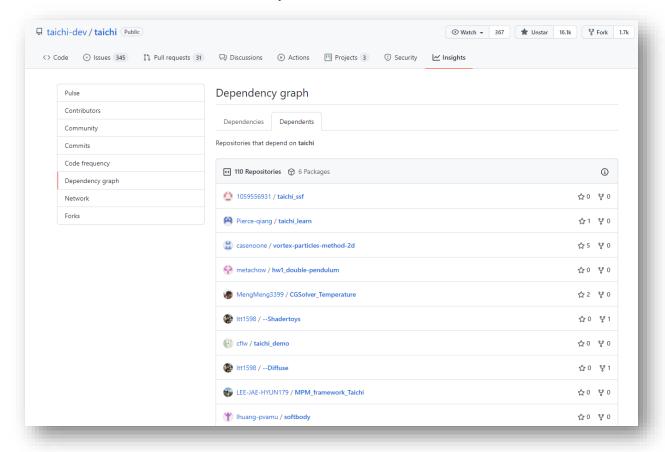
- 死线: 2022年1月3日
- 要求:
  - 使用作业模板
  - 需要有设计文档, 如果有参照代码也需要标明
- 题材:
  - 任何使用Taichi完成的内容(图形学更佳)
  - 可以参考每节图形课后给出的大作业选题灵感 [参考第07,09,10讲]
  - 鼓励实现任意图形学论文/图形学课程内容
- 形式:
  - 使用 GitHub/Gitee提交并邀请tgc01@taichi.graphics加入你的代码仓
  - 允许三人以下合作,记得管理多人合作的git commits
- 奖励:
  - 太极图形课第一季结业证书一份+神秘Taichi礼物一份

### Excellent homework assignments



## Gifts for the gifted

- Use **Template** for your homework
- Next check Dec. 14, 2021















# Questions?

本次答疑: 12/02 ←作业分享也在这里

下次直播: 12/07

直播回放: Bilibili 搜索「太极图形」

主页&课件: <a href="https://github.com/taichiCourse01">https://github.com/taichiCourse01</a>

主页&课件(backup): <a href="https://docs.taichi.graphics/tgc01">https://docs.taichi.graphics/tgc01</a>