

The University of New Mexico
School of Engineering
Electrical and Computer Engineering Department

ECE 535 Satellite Communications

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Module # 3-2: 2.19, 2.23, 2.24, 2.25, 2.27

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2.19 Determine the Julian days for the following dates and times: midnight March 10, 1999; noon, February 23, 2000; 16:30h, March 1, 2003; 3 P.M., July 4, 2010.

Midnight March 10, 1999:

$$1999 = 2451178.5$$

$$\text{Day of Year} = 69$$

$$\text{UT} = 13 \text{ hours} = 0.5416667$$

$$\begin{aligned}\text{Julian Day} &= 2451178.5 + 69 + 0.5416667 \\ &= 2,451,248.5\end{aligned}$$

Noon, February 23, 2000:

$$2000 = 2451543.5$$

$$\text{Day of Year} = 54$$

$$\text{UT} = 12 \text{ hours} = 0.5$$

$$\begin{aligned}\text{Julian Day} &= 2451543.5 + 54 + 0.5 \\ &= 2,451,598\end{aligned}$$

16:30h, March 1, 2003:

$$2003 = 2452639.5$$

$$\text{Day of Year} = 60$$

$$\text{UT} = 16.5 = 0.6875$$

$$\begin{aligned}\text{Julian Day} &= 2452639.5 + 60 + 0.6875 \\ &= 2,452,700.1875\end{aligned}$$

3 P.M., July 4, 2010:

$$2010 = 2455196.5$$

$$\text{Day of Year} = 185$$

$$\text{UT} = 3 \text{ P.M.} = 15:00\text{h} = 0.625$$

$$\text{Julian Day} = 2,455,382.125$$

2.23 The Molnya 3-(25) satellite has the following parameters specified: perigee height 462 km; apogee height 40,850 km; period 736 min; inclination 62.8°. Using an average value of 6371 km for the earth's radius, calculate (a) the semi-major axis and (b) the eccentricity. (c) Calculate the nominal mean motion n_0 . (d) Calculate the mean motion. (e) Using the calculated value for a, calculate the anomalistic period and compare with the specified value. Calculate (f) the rate of regression of the nodes, and (g) the rate of rotation of the line of apsides.

$$h_p = 462 \text{ km}$$

$$\text{Inclination } (i) = 62.8^\circ (1.096 \text{ radian})$$

$$h_a = 40,850 \text{ km}$$

$$R = 6,371 \text{ km}$$

$$P = 736 \text{ min}$$

$$\mu = 3.986005 \times 10^{14} \text{ m}^3/\text{s}^2$$

(a) Semi-major axis

$$r_p = 40,850 \text{ km} + 6,371 \text{ km} = 47,221 \text{ km}$$

$$r_a = 462 \text{ km} + 6,371 \text{ km} = 6,833 \text{ km}$$

$$a = \frac{r_a + r_p}{2} \xrightarrow{\text{yields}} \frac{47,221 \text{ km} + 6,833 \text{ km}}{2} = 27,027 \text{ km}$$

(b) Eccentricity

$$e = \frac{r_a}{a} - 1 \xrightarrow{\text{yields}} \frac{47,221 \text{ km}}{27,027 \text{ km}} - 1 = 0.747$$

(c) Nominal Mean Motion n_0

$$n_0 = \sqrt{\frac{\mu}{a^3}} \xrightarrow{\text{yields}} \sqrt{\frac{3.986005 \times 10^{14} \text{ m}^3/\text{s}^2}{(27,027 \times 10^3 \text{ m})^3}} = \sqrt{\frac{3.986005 \times 10^{14} \text{ m}^3/\text{s}^2}{19.742 \times 10^{22} \text{ m}^3}} = \sqrt{2.019 \times 10^{-8}}$$

$$n_0 = 1.421 \times 10^{-4} \text{ rad/sec} * 86400 \text{ sec} = 12.277 \text{ rad/day}$$

(d) Mean Motion

$$n = \frac{2\pi}{P} \xrightarrow{\text{yields}} \frac{2\pi}{736 \text{ min} * 60 \text{ sec}} = \frac{2\pi}{44160} = 1.423 \times 10^{-4} \text{ rad/sec} * 86400 \text{ sec} = 12.276 \text{ rad/day}$$

(e) With a = 27,027km, find anomalistic period

$$P_A = \frac{2\pi}{n_0} \xrightarrow{\text{yields}} \frac{2\pi}{1.421 \times 10^{-4} \text{ rad/sec}} = 44,218.88 \text{ s} / 60 \text{ s} = 736.98 \text{ min}$$

(f) Rate of Regression of the Nodes

$$K = \frac{nK_1}{a^2(1-e^2)^2} \xrightarrow{\text{yields}} \frac{(1.423 \times 10^{-4} \text{ rad/sec})(66,063.1704 \text{ km}^2)}{(27,027 \text{ km})^2(1-(0.747)^2)^2} = \frac{9.394}{1.42705 \times 10^8} = 6.583 \times 10^{-8}$$

$$\frac{d\Omega}{dt} = -K \cos(i) \xrightarrow{\text{yields}} -6.583 \times 10^{-8} (\cos(1.096 \text{ rad})) = -3.0009 \times 10^{-8} \text{ rad/sec} = -0.149 \text{ deg/day}$$

(g) Rate of Rotation of the line of apsides

$$\frac{d\omega}{dt} = K(2 - 2.5 \sin^2(i)) \xrightarrow{\text{yields}} 6.583 \times 10^{-8} (2 - 2.5 \sin^2(1.096 \text{ rad})) = 1.448 \times 10^{-9} \text{ rad/sec} = 0.007 \text{ deg/day}$$

2.24 Repeat the Calculations in Problem 2.23 for tan inclination of 63.435

Inclination(i) = 63.435° or 1.107 rad

(a) Semi-major axis

$$r_p = 40,850\text{km} + 6,371\text{km} = 47,221\text{ km}$$

$$r_a = 462\text{km} + 6,371\text{ km} = 6,833\text{ km}$$

$$a = \frac{r_a + r_p}{2} \xrightarrow{\text{yields}} \frac{47,221\text{km} + 6,833\text{km}}{2} = 27,027\text{km}$$

(b) Eccentricity

$$e = \frac{r_a}{a} - 1 \xrightarrow{\text{yields}} \frac{47,221\text{km}}{27,027\text{km}} - 1 = 0.747$$

(c) Nominal Mean Motion n_0

$$n_0 = \sqrt{\frac{\mu}{a^3}} \xrightarrow{\text{yields}} \sqrt{\frac{3.986005 \times 10^{14} \text{m}^3/\text{s}^2}{(27,027 \times 10^3 \text{m})^3}} = \sqrt{\frac{3.986005 \times 10^{14} \text{m}^3/\text{s}^2}{19.742 \times 10^{22} \text{m}}} = \sqrt{2.019 \times 10^{-8}}$$

$$n_0 = 1.421 \times 10^{-4} \text{ rad/sec} * 86400 \text{ sec} = 12.277 \text{ rad/day}$$

(d) Mean Motion

$$n = \frac{2\pi}{P} \xrightarrow{\text{yields}} \frac{2\pi}{736 \text{min} * 60 \text{sec}} = \frac{2\pi}{44160} = 1.423 \times 10^{-4} \text{ rad/sec} * 86400 \text{ sec} = 12.276 \text{ rad/day}$$

(e) With $a = 27,027\text{km}$, find anomalistic period

$$P_A = \frac{2\pi}{n_0} \xrightarrow{\text{yields}} \frac{2\pi}{1.421 \times 10^{-4} \text{ rad/sec}} = 44,218.88 \text{s} / 60 \text{s} = 736.98 \text{min}$$

(f) Rate of Regression of the Nodes

$$K = \frac{nK_1}{a^2(1-e^2)^2} \xrightarrow{\text{yields}} \frac{(1.423 \times 10^{-4} \text{ rad/sec})(66,063.1704 \text{km}^2)}{(27,027 \text{km})^2(1-(0.747)^2)^2} = \frac{9.394}{1.42705 \times 10^8} = 6.583 \times 10^{-8}$$

$$\frac{d\Omega}{dt} = -K \cos(i) \xrightarrow{\text{yields}} -6.583 \times 10^{-8} (\cos(1.107 \text{rad})) = -2.944 \times 10^{-8} \text{ rad/sec} = -0.146 \text{ deg/day}$$

(g) Rate of Rotation of the line of apsides

$$\frac{d\omega}{dt} = K(2 - 2.5 \sin^2(i)) \xrightarrow{\text{yields}} 6.583 \times 10^{-8} (2 - 2.5 \sin^2(1.107 \text{rad})) = 2.3037 \times 10^{-9} \text{ rad/sec} = 0.0114 \text{ deg/day}$$

2.25 Determine the orbital condition necessary for the argument of perigee to remain stationary in the orbital plane. The orbit for a satellite under this condition has an eccentricity of 0.001 and a semi-major axis of 27,000km. At a given epoch the perigee is exactly on the line of Aries. Determine the satellite position relative to this line after a period of 30 days from epoch.

$$e = 0.001$$

$$a = 27,000\text{km}$$

For the satellite to remain stationary in the orbital plane, Inclination(i) = 63.435° or 1.0715 rad

Find mean motion:

$$n = \sqrt{\frac{\mu}{a^3}} \xrightarrow{\text{yields}} \sqrt{\frac{3.986005 \times 10^{14} \text{m}^3/\text{s}^2}{(27,000 \times 10^3 \text{m})^3}} = \sqrt{\frac{3.986005 \times 10^{14} \text{m}^3/\text{s}^2}{19.683 \times 10^{22} \text{m}}} = \sqrt{2.025 \times 10^{-8}}$$

$$n = 1.42306 \times 10^{-4} \text{ rad/sec}$$

Find K:

$$K = \frac{nK_1}{a^2(1-e^2)^2} \xrightarrow{\text{yields}} \frac{(1.42306 \times 10^{-4} \text{rad/sec})(66,063.1704 \text{km}^2)}{(27,000 \text{km})^2(1-(0.001)^2)^2} = 1.289604 \times 10^{-8} \text{rad/sec}$$

Satellite Position:

$$\frac{d\Omega}{dt} = -K \cos(i) \xrightarrow{\text{yields}} -1.289604 \times 10^{-8} (\cos(1.0715 \text{rad})) = -6.1748 \text{rad/sec}$$

$$\frac{d\Omega}{dt}_{30 \text{ days}} = -\frac{6.1748 \text{rad}}{\text{sec}} * 2592000 = -0.016 \text{rad/day}_{30} = -0.917^\circ$$

Note: In MS Excel, the position is exactly 0.856°W most likely due to significant digits/rounding.

2.27 A satellite has an inclination of 90° and an eccentricity of 0.1. At epoch, which corresponds to the time of perigee passage, the perigee height is 2,643.24km directly over the north pole. Determine (a) the satellite mean motion. For 1 day after epoch determine (b) the true anomaly, (c) the magnitude of the radius vector to the satellite, and (d) the latitude of the subsatellite point.

Inclination (i) = 90° (polar orbit)

$$\mu = 3.986005 \times 10^{14} \text{ m}^3/\text{s}^2$$

$$h_p = 2,643.24 \text{ km}$$

$$K_1 = 66,063.1704 \text{ km}^2$$

(a) Mean Motion

$$r_p = 2,643.24 \text{ km} + 6,371 \text{ km} = 9,014.24 \text{ km}$$

$$a = \frac{r_p}{1 - e} \xrightarrow{\text{yields}} \frac{9,014.24 \text{ km}}{1 - 0.1} = 10,015.82 \text{ km}$$

$$n = \sqrt{\frac{\mu}{a^3}} \xrightarrow{\text{yields}} \sqrt{\frac{3.986005 \times 10^{14} \text{ m}^3/\text{s}^2}{(10,015.82 \times 10^3 \text{ m})^3}} = 6.29 \times 10^{-4} \text{ s}^{-1} * 86400 = 54.42 \text{ rad/day}$$

(b) True Anomaly

Assume $M \approx E$ for small eccentricity of 0.1.

$$\tan\left(\frac{v}{2}\right) = \sqrt{\frac{1+e}{1-e}} \tan \frac{E}{2} \xrightarrow{\text{yields}} \sqrt{\frac{1.1}{0.9}} \tan\left(\frac{54.42}{2}\right) = 1.074 \tan(27.21 \text{ rad}) = -1.9361 \text{ rad}$$

$$\left(\frac{v}{2}\right) = \tan^{-1}(-1.9361 \text{ rad}) = -1.094 \text{ rad}$$

$$v = 2 * -1.094 = -2.1881 \text{ rad or } -125.363 \text{ deg}$$

(c) Magnitude of the Radius Vector

$$r = \frac{a(1 - e^2)}{1 + e \cos(v)} \xrightarrow{\text{yields}} \frac{(10,015.82 \text{ km})(1 - 0.1^2)}{1 + 0.1 \cos(-125.37)} = \frac{9,915.6618 \text{ km}}{0.94212} = 10,524.8395 \text{ km}$$

(d) Latitude of the Subsatellite Point

$$\phi = \sin^{-1}(\sin(v + \omega) \sin(i)) \xrightarrow{\text{yields}} \sin^{-1}(\sin(-125.37) \sin(90))$$

$$\phi = \sin^{-1}(-0.8154) * (1) \cong -54.63^\circ$$

$$\phi = 90^\circ - 54.63^\circ = 35.37^\circ \text{ S lat (due to south to north orbit, passing the apogee)}$$