# Maxent Example - Non-interacting Case

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#### Abstract

This document is a tutorial on the use of Maxent, a program for doing analytical continuation using the maximum entropy method. It will explain how to provide the program with the proper parameter file, data format for particle-hole symmetric data (in both  $\tau$  and  $i\omega_n$ ), and understand the output. This program uses the ALPSCore libraries[1, 2]

# Contents

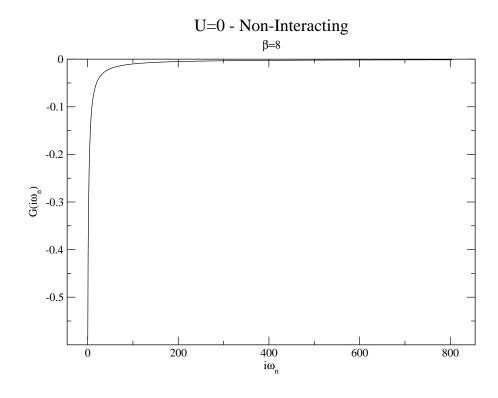
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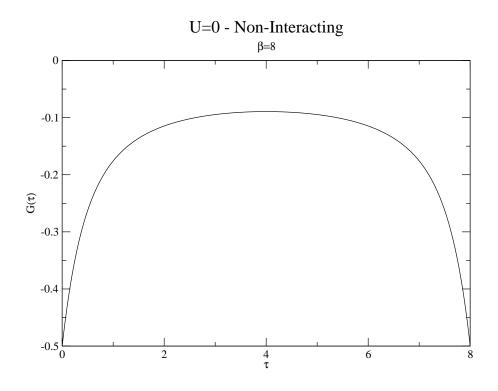
# 1 Introduction

Using DMFT we can set U=0 and generate the non-interacting Hubbard model

$$H = -t \sum_{\langle ij \rangle \sigma} c^{\dagger}_{j\sigma} c_{i\sigma}$$

This produces a Green's function in Matsubara space and Time space:





### 2 File Structure

We'e included several files that will be used to generate the remainder of this document:

```
Filenames and Descriptions
• G_{-}im = \text{Im}[G(i\omega_n)], also input format for maxent – column format: i\omega_n \text{ Im}[G(i\omega_n)] \sigma_{I,n}
                                              Input file - G im
  0.39269908169872 - 0.58900239090596 \ 1 \text{e} - 05
  1.1780972450962 - 0.40986302909581 \ 1e-05
  1.9634954084936 -0.32440959736089 1e-05
• G re = \operatorname{Re}[G(i\omega_n)]
     - column format: i\omega_n \operatorname{Re}[G(i\omega_n)] \sigma_{R,n}
                                              Input file - G re
  #Freq 0 1e-05
  0.39269908169872 4.1987536135299e-16 1e-05
  1.1780972450962 2.1665780666625e-16 1e-05
• Gtau = G(\tau), also input format for maxent
     - column format: \tau G(\tau) \sigma_{\tau}
                                              Input file - Gtau
  0 -0.5 1e-05
  0.0078125 -0.49377185389756 1e-05
  0.015625 -0.48766306852771 1e-05
```

## 3 Using Maxent

These files are easily used with Maxent. Here is the frequency space input file:

#### Param File in.param

```
BETA=8 #inverse temperature NDAT=1024 #num of data points NFREQ=500 #num of output frequencies DATASPACE=frequency #G(i\omega) KERNEL=fermionic #fermionic/bosonic values PARTICLE_HOLE_SYMMETRY=1 #0/1 DATA="G_im" #location of data file
```

Maxent then produces the following output:

#### Maxent output

```
./maxent in.param
Using flat default model
using kernel fermionic in domain frequency with ph symmetry
Kernel is set up
# 0
        6.27911e+06
        1.16635e+06
# 1
# 2
        258981
        68739
# 3
# 4
        17040.5
# 5
        3761.79
# 6
        744.024
# 7
        132.167
minimal chi2: 2.55387e-05
alpha it: 0
                WARNING: iteration reached max_it without converging, your
   minimizer is having problems. Please be careful!
Q = 0.5 chi^2 - \alpha * entropy: 987272
                                         norm: 1.034
                WARNING: iteration reached max_it without converging, your
alpha it: 1
   minimizer is having problems. Please be careful!
Q = 0.5 chi^2-\alpha*entropy: 155.23
                                         norm: 1.00046
alpha it: 2
                Q = 0.5 chi^2-\alpha*entropy: 93.5
                                                           norm: 1.00035
                Q = 0.5 chi^2-\alpha*entropy: 82.1164
alpha it: 3
                                                           norm: 1.00033
alpha it: 4
                Q = 0.5 chi^2-\alpha + entropy: 72.0799
                                                           norm: 1.0003
. . . .
alpha it: 55
                Q = 0.5 chi^2 - \alpha * entropy: 0.0395693 norm: 1.00001
                                                           norm: 1.00001
alpha it: 56
                Q = 0.5 \text{chi}^2-\alpha*entropy: 0.0337866
alpha it: 57
                Q = 0.5 chi^2-\alpha*entropy: 0.028796
                                                           norm: 1
alpha it: 58
                Q = 0.5 chi^2-\alpha*entropy: 0.0245097
alpha it: 59
                Q = 0.5 chi^2-\alpha*entropy: 0.0208447
                                                           norm: 1
Ng: 7.3691
chi2 max: 3.66901
posterior probability of the default model: 1.59791e-24
```

#### 3.1 Time input file

BETA=8

The input file for  $G(\tau)$  can be slightly more complex. You can either supply the tau grid points through the parameter file or input file, with the param file used over input. An example of such a param file is included:

```
Param File in_tau.param
```

4

#inverse temperature

Using your own  $\tau$  input generates a message confirming the choice of  $\tau$  grid:

```
Using flat default model using kernel fermionic in domain time with ph symmetry
Using param input tau points
Kernel is set up
```

When  $\tau$  grid data is solely used from an input file, Maxent reads Using data file tau points

### 3.2 Output Guide

```
Using flat default model using kernel fermionic in domain frequency with ph symmetry Kernel is set up
```

These are the setup messages, confirming your input choices.

```
# 0 6.27911e+06

# 1 1.16635e+06

# 2 258981

# 3 68739

# 4 17040.5

# 5 3761.79

# 6 744.024

# 7 132.167

minimal chi2: 2.55387e-05
```

These represent the eigenvalues that are above precision after the single value decomposition (SVD). The last line represents the smallest  $\chi^2$  value the program thinks it will achieve. If this is  $\gg 1$  there may be something wrong with your input

```
alpha it: 2 Q = 0.5chi^2-\alpha*entropy: 93.5 norm: 1.00035
alpha it: 3 Q = 0.5chi^2-\alpha*entropy: 82.1164 norm: 1.00033
...
```

The root finding procedure will print the iterations through  $\alpha$ values in the range given by the parameters (default: 60 values  $\in$  [0.01, 20]) If the first two or three do not minimize properly that is ok, as long as the rest continue normally. Notice that the norm stays  $\approx$ 1 for all iterations

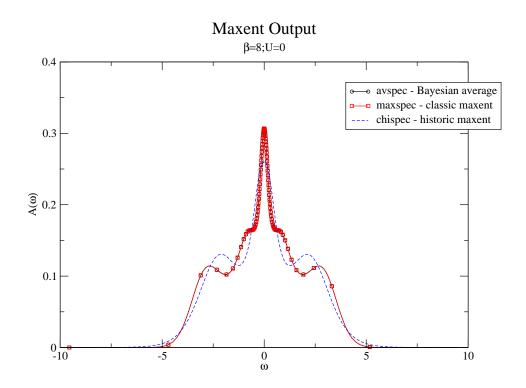
```
Ng: 7.3691
chi2 max: 3.66901
posterior probability of the default model: 1.59791e-24
```

This is posted after completing all  $\alpha$  values and root finding. Ng represents the number of "good input points," chi2 max is the maximum value of  $\chi^2$  in the  $\alpha$  iterations, and the last line is the probability that the default model is the correct representation of the spectral function. Note that that posterior probability has no normalization.

If text output is on, Maxent produces 8 files:

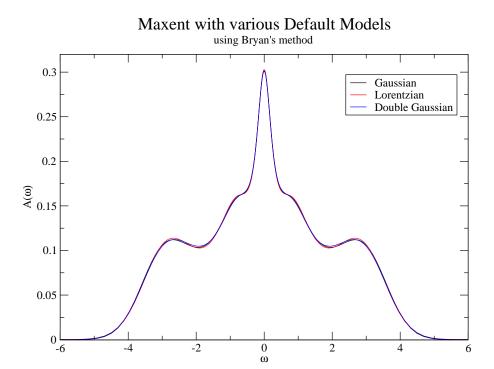
name.out.avspec.dat	Spectral function using Bayesian Averaging - Bryan's method
name.out. <u>chi2</u> .dat	Estimated $\chi^2$ for each $\alpha$ value solution
name.out. <u>chispec</u> .dat	Spectral function satisfying the best $\chi^2$ - classic Maxent
name.out.fits.dat	Fits of each $\alpha$ value, see comments in file
name.out.maxspec.dat	Spectral function with the highest probability - historic Maxent
name.out.out.h5	All output data in the hdf5 format
name.out.prob.dat	The posterior probability of each $\alpha$ value
name.out. <u>spex</u> .dat	All spectral functions produced; one for each $\alpha$

In our example here are the spectral outputs:



# 4 Fine-Tuning Output

Different default models shouldn't change the results much, but sometimes end up doing so. One must be wary not to use a default model whose entropy is too 'strong' so that Maxent gets stuck in that local minimum. Here are a variety of models from the above example:



With  $\sigma = 1, \Gamma = 0.5, \mu(\text{shift}) = 2$ . This represents well behaved data within Maxent.

### References

- [1] Alexander Gaenko, Emanuel Gull, Andrey E. Antipov, Lukas Gamper, and Gabriele Carcassi. ALPSCore: Version 0.4.5. May 2015. doi:10.5281/zenodo.17398.
- [2] B Bauer et al. The ALPS project release 2.0: open source software for strongly correlated systems. *Journal of Statistical Mechanics: Theory and Experiment*, 2011(05):P05001, 2011. doi:10.1088/1742-5468/2011/05/P05001.