

Space Flight Modelling and Atmospheric Reentry

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1. Newton's Gravity Equation and the Two Body Problem

2. Rescaling

3. Implementation

4. Preview on Plans and Future Talks

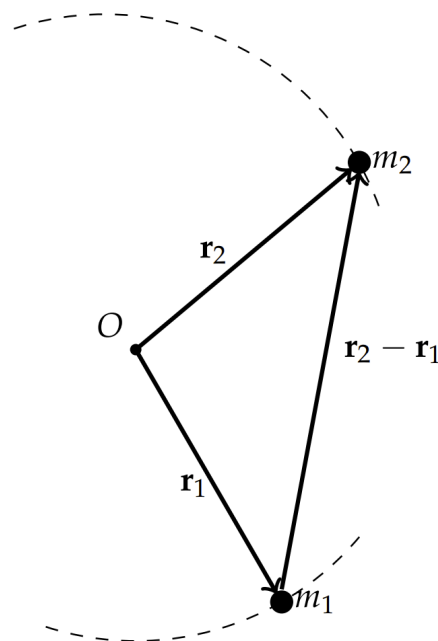


Figure: Scheme of the two body problem ¹

2 Body Problem

- Describes the problem where the future motion of two masses needs to be determined

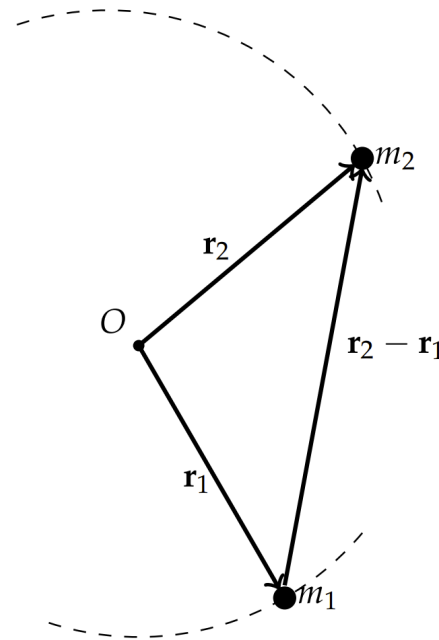


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2 Body Problem

- Describes the problem where the future motion of two masses needs to be determined
- \exists plane E s.t. both masses $m_i \in E \Rightarrow$ we can treat the space as 2-dimensional for now

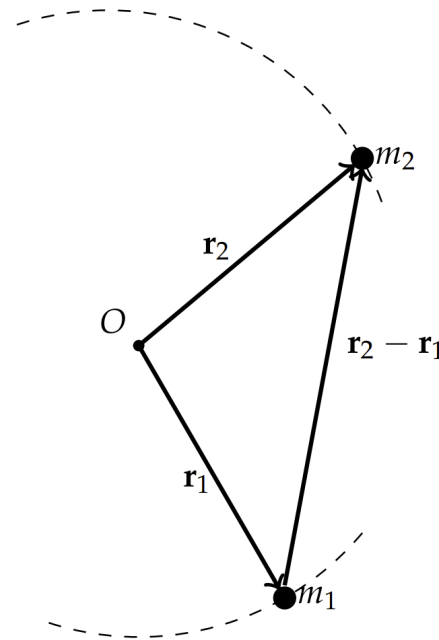


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$$m_1 \cdot a_1(t) = F_{21} = G \cdot \frac{m_1 \cdot m_2}{\|r_2(t) - r_1(t)\|^2} \cdot \frac{r_2(t) - r_1(t)}{\|r_2(t) - r_1(t)\|}$$
$$m_2 \cdot a_2(t) = F_{12} = G \cdot \frac{m_2 \cdot m_1}{\|r_1(t) - r_2(t)\|^2} \cdot \frac{r_1(t) - r_2(t)}{\|r_1(t) - r_2(t)\|}$$

$$\ddot{r}_1(t) = \frac{G \cdot m_2}{\|r_2(t) - r_1(t)\|^3} \cdot (r_2(t) - r_1(t))$$
$$\ddot{r}_2(t) = \frac{G \cdot m_1}{\|r_1(t) - r_2(t)\|^3} \cdot (r_1(t) - r_2(t))$$

For a two-dimensional model with coordinates $r_1(t) = (x_1(t), y_1(t))^T$ and velocities $\dot{r}_1(t) = (u_1(t), v_1(t))^T$ and $\alpha(t) = \frac{G}{\sqrt{(x_1 - x_2)^2 + (y_1 - y_2)^2}^3}$, we get an ODE-system of 2nd order:

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$$\frac{d}{dt} \begin{pmatrix} x_1 \\ y_1 \\ x_2 \\ y_2 \\ u_1 \\ v_1 \\ u_2 \\ v_2 \end{pmatrix} = \begin{pmatrix} u_1 \\ v_1 \\ u_2 \\ v_2 \\ \alpha m_2 (x_2 - x_1) \\ \alpha m_2 (y_2 - y_1) \\ \alpha m_2 (x_1 - x_2) \\ \alpha m_2 (y_1 - y_2) \end{pmatrix} \quad (1)$$

Representation with one Fixed Mass



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$$\frac{d}{dt} \begin{pmatrix} x \\ y \\ u \\ v \end{pmatrix} = \begin{pmatrix} u \\ v \\ -\mu \frac{x}{\|\mathbf{r}\|^3} \\ -\mu \frac{y}{\|\mathbf{r}\|^3} \end{pmatrix}$$

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Rescaling of Base Quantities



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Quantity	SI-Unit	Rescaled unit
Mass	1 kg	$1 m_E = 5.9772 \cdot 10^{24} \text{ kg}$
Time	1 s	1 day = 86400 s
Distance	1 m	1 AU = 149 597 870 700 m

Dependent Quantities

Quantity	SI-Unit	Rescaled unit
velocity	$1 \frac{m}{s}$	$1 \frac{AU}{day} \hat{=} \frac{1.49 \cdot 10^{11} \frac{m}{AU}}{8.64 \cdot 10^4 \frac{s}{day}} \cdot 1 \frac{AU}{day} \approx 1.72 \cdot 10^6 \frac{m}{s}$
μ	$1 \frac{m^3}{s^2}$	$1 \frac{AU^3}{day^2} \hat{=} \frac{\left(1.49 \cdot 10^{11} \frac{m}{AU}\right)^3}{\left(8.64 \cdot 10^4 \frac{s}{day}\right)^2} \cdot 1 \frac{AU^3}{day^2} \approx 4.43 \cdot 10^{23} \frac{m^3}{s^2}$

Constants

Quantity	SI-Value	Rescaled value
G	$6.67 \cdot 10^{-11} \frac{m^3}{kg \cdot s^2}$	$\tilde{G} = 6.67 \cdot 10^{-11} \frac{m^3}{kg \cdot s^2} \cdot \frac{5.98 \cdot 10^{24} \frac{kg}{m_E} \cdot \left(8.64 \cdot 10^4 \frac{s}{day}\right)^2}{\left(1.49 \cdot 10^{11} \frac{m}{AU}\right)^3} \approx 9.38 \cdot 10^{-10} \frac{AU^3}{m_E \cdot s^2}$
m_{sun}	$2 \cdot 10^{30} kg$	$\tilde{m}_{sun} \hat{=} \frac{2 \cdot 10^{30} kg}{5.98 \cdot 10^{24} \frac{kg}{m_E}} \approx 334\,448 m_E$

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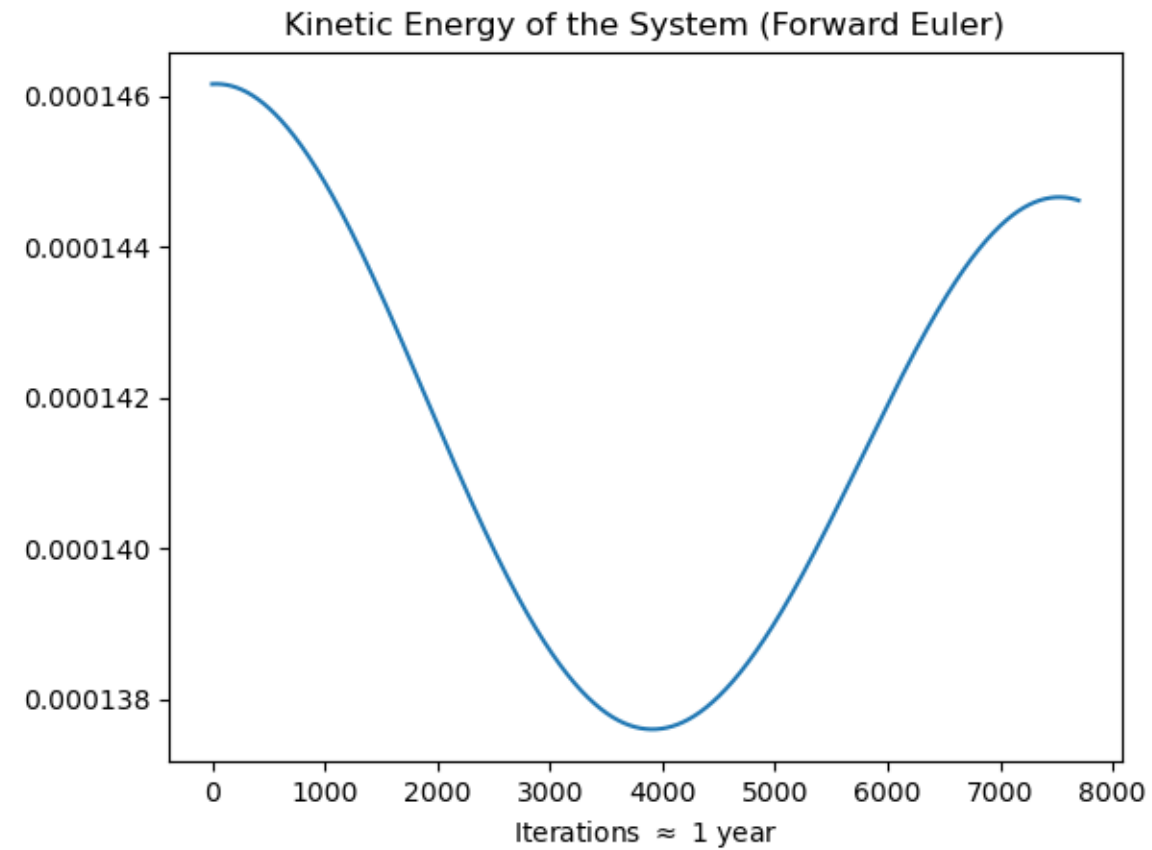
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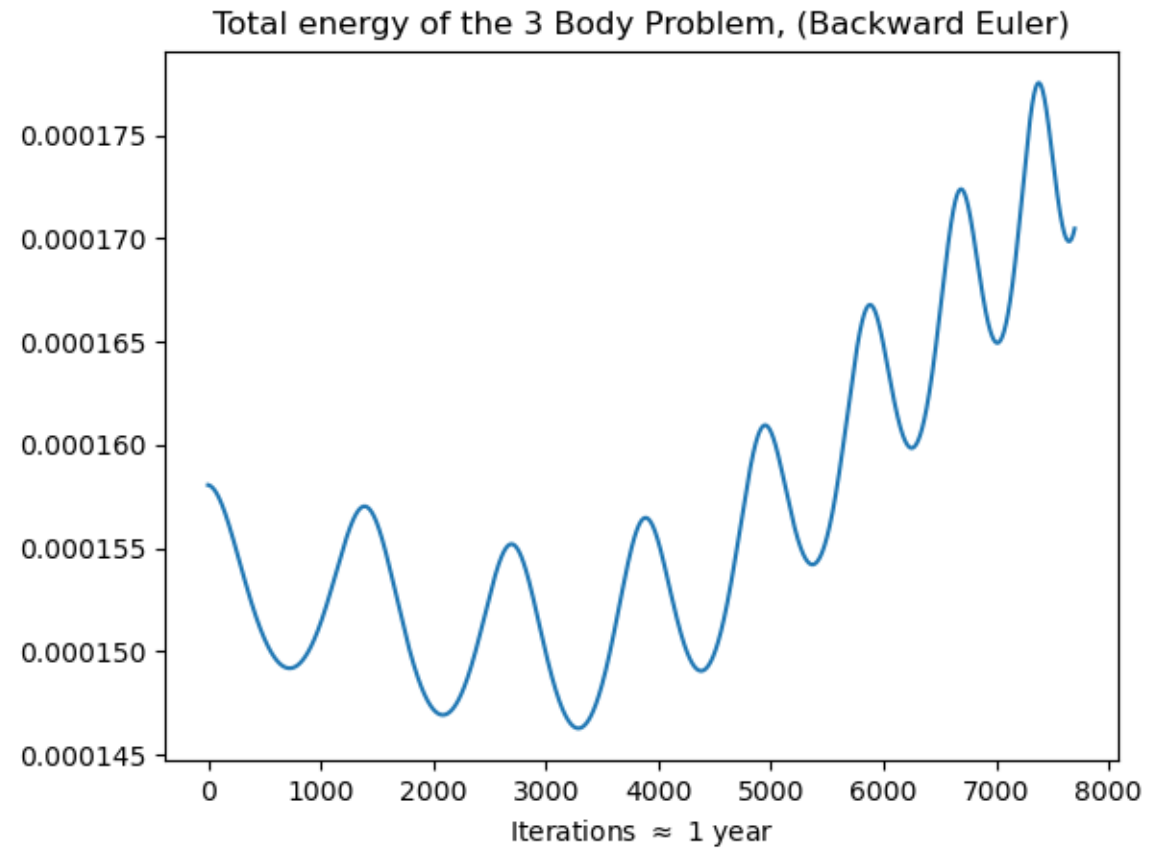
Energy

$$\begin{aligned} E_{tot}(t) &= E_{kin_1}(r_1(t)) + E_{kin_2}(r_2(t)) + E_{pot}(\mathbf{r}(t)) \\ &= \frac{1}{2}m_1\dot{r}_1^2 + \frac{1}{2}m_2\dot{r}_2^2 - \frac{Gm_1m_2}{||\mathbf{r}||} \\ &= \frac{1}{2} (m_1(u_1^2 + v_1^2) + m_2(u_2^2 + v_2^2)) - \frac{Gm_1m_2}{||\mathbf{r}||} \stackrel{!}{=} const. \end{aligned}$$

$$\begin{aligned}\alpha_n &= \frac{\mu}{\sqrt{x_{n-1}^2 + y_{n-1}^2}^3} \\ x_n &= x_{n-1} + \Delta t \cdot u_{n-1} \\ y_n &= y_{n-1} + \Delta t \cdot v_{n-1} \\ u_n &= u_{n-1} - \Delta t \cdot \alpha_n x_{n-1} \\ v_n &= v_{n-1} - \Delta t \cdot \alpha_n y_{n-1} \\ t_n &= t_{n-1} + \Delta t\end{aligned}$$



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Expansion for n bodies

$$\ddot{r}_i(t) = G \cdot \sum_{j \neq i} \frac{m_j}{\|r_j(t) - r_i(t)\|^3} (r_j(t) - r_i(t)) \quad i = 1, \dots, n$$

System Energy

$$E_{tot}(t) = \sum_{i=1}^N E_{kin_i}(t) + E_{pot}(t)$$

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- Discuss leaving entering the atmosphere

