User Guide for Corrfunc

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1 Introduction

Correlation functions are a statistical measure of a density field and are widely used in large-scale structure formation. Generally, the measurements are done $once^1$ on survey data and compared with model predictions in a Monte-Carlo Markov Chain. As such, the correlation functions have to be measured repeatedly during an MCMC. The codes presented here are meant to cover the typical scenarios of measuring correlation functions in theoryland. The primary consideration in writing these codes is speed²— the codes presented here should outperform any other CPU based correlation functions codes by a wide-margin.

 $^{^{1}}$ which is why I have not bothered with releasing the codes to measure correlation functions on data

²The secondary consideration was maintainability and ease of use for others. I have versions of these codes that are even faster but are potentially gibberish for anyone other than me!

2 Installation

The only requirements for the code to install is a valid C compiler, with OpenMP support. The AVX instruction set can only be used for CPU's later than 2011 (Intel Sandy Bridge/ AMD Bulldozer or later).

2.1 Getting the Source

You can obtain the source in two ways: i) Clone the mercurial repo (hg clone https://bitbucket.org/manodeep/corrfunc/) or ii) Download the tar archive (corrfunc.\$MAJOR.0.\$MINOR.tar.gz) and unpack it in the directory where you wish to keep the files (tar xvzf corrfunc.\$MAJOR.0.\$MINOR.tar.gz). Here, \$MAJOR and \$MINOR refer to the major and minor release versions (current \$MAJOR=1, \$MINOR=0). I will only change the \$MAJOR version if the API breaks.

The directory structure for the code looks like this:

```
corrfunc
  _paper
 \_ xi_theory
     benchmarks......IDL scripts to run benchmarks.
    _bin......will be created to copy executable
                        files when you run 'make install'.
    \_ examples\ldots\ldotsSource files for example C bindings
                        using the static libraries.
    _include......Header files for static libraries.
    _io.....Source files for reading in data.
     _{\scriptscriptstyle -}lib_{\scriptscriptstyle -}tib_{\scriptscriptstyle -}to copy static
                        libraries and python library after
                        you run 'make'.
    _python_bindings...Source files to generate python
                        bindings.
     tests......Correct outputs for tests.
       _data.....Mock galaxy catalogs for tests.
    _utils......Source files for creating 3-D grid
                        and helper routines.
    \_ wp\ldots\ldots Source files for w_p(r_p) .
    \_ xi\_of\_r\ldots\ldotsSource files for \xi(r).
    \_ xi_rp_pi\ldots . Source files for \xi(r_p,\pi) .
```

Option Type	Option Name	Default State	Requires	Notes	
	PERIODIC	Enabled	None	Enables periodic boundary conditions.	
Science	OUTPUT_RPAVG	Disabled	DOUBLE_PREC	Outputs the average pair-separation in each bin. $\xi(r)$ and $w_p(r_p)$ can be slower by more than $2\times$, $\xi(r_p, \pi)$ is less affected.	
	DOUBLE_PREC	Disabled	None	Computations are done using double precision. Slower and requires more RAM.	
Code	USE_AVX	Enabled	CPU and compiler with AVX support	CPUs later than 2011 have AVX support. Code will run much faster with this option.	
	USE_OMP	Enabled	OpenMP capable compiler	Since clang does not support OpenMP yet, common.mk will stop compilation with clang when this flag is enabled.	

Table 1: List of compilations options, what the options mean and their dependencies for the codes.

2.2 Code Options

There are a few code options that control both the Science case and the code compilation. All of these options are located in 'common.mk' in the base directory ('corrfunc'). Edit the first few lines to set these options (see Table. 1 for details):

- Science options PERIODIC, OUTPUT_RPAVG
- Code options DOUBLE_PREC, USE_AVX and USE_OMP

Depending on your Science use-case and the cpu/compiler, you will want to set the different options. Once you set those options, you should set the C compiler, CC (available options are icc, gcc, clang). Once you have set the compiler, installing should be as simple as typing 'make' and 'make install' in the xi_theory directory. All the libraries are intentionally chosen to be static libraries just to avoid any path conflicts. However, on MAC OSX, you may have to do more to get the library to work – so I have outlined some of the scenarios in Section 2.4.

2.3 Linux

If the installation went well, you should have an executable called run_correlations in the examples directory. Type ./run_correlations in the examples directory and you should see the code in action. The C source file run_correlations.c also serves as an example to use the $\xi(r)$, $\xi(r_p,\pi)$ and $w_p(r_p)$ libraries in C.

2.4 Mac OSX

There can be two issues on MACs. One is that the default gcc assembler supplied by XCode or macports is too old and does not support AVX instructions even when the CPU does. One way to get around this is by using the clang assembler even when compiling with gcc. The easiest way to do it is by replacing the default assembler with the as script in the paper directory (taken from this url). Copy this as script to the appropriate directory (/opt/local/bin/ for me since I use macports gcc on my laptop). Another problem might come with running the python example codes in the python_bindings directory. If you get an error message:

• Fatal Python error: PyThreadState_Get: no current thread

when you run python call_correlation_functions.py, then the following steps might fix the problem (these are also noted in the FAQ). This error occurs when the python library used at compile time is not the same as the runtime python library. In all cases that I have seen, this error occurs when using the conda package manager for python³.

- Change the relative path for the shared python library _countpairs.so. You can change the relative path by issuing the command:

 install_name_tool -change libpython2.7.dylib `python-config --prefix`/lib/libpython2.7.dylib _countpairs.so
- Add to the fallback library path environment variable.

 export DYLD_FALLBACK_LIBRARY_PATH=`python-config --prefix`/lib:\DYLD_FALLBACK_LIBRARY_PATH
- If both of the above methods fail, then create a symbolic link

 ln -s `python-config --prefix`/lib/libpython2.7.dylib

If all went well, then you should be able to run the run_correlations code in the examples directory as well as execute python call_correlation_functions.py in the python_bindings directory. In all of the above examples, I have assumed that the relevant python library is libpython2.7.dylib (the default under conda) – you may have to replace it with your python library version.

³This behaviour is by design according to conda

3 Running the Codes

3.1 Input Files

The codes currently can handle these types of input data files:

- ascii White-space separated columns, format code is 'a'.
- csv Comma-separated values, format code is 'c'.
- fast-food Fast-food, fortran binary format, format code is 'f'. The fast-food file format is described in detail in Section 3.1.1.

For the ascii and csv files, the code reads in the first three columns as the co-moving X/Y/Z arrays. Note, that more columns can be present but the code will ignore those columns.

3.1.1 The fast-food file format

The fast-food format is a fortran binary format – all fields are surrounded with 4 bytes padding. These value of these padding bytes is the number of bytes of data contained in between the padding bytes. For example, to write out 20 bytes of data in a fast-food file format would require a total of 4 + 20 + 4 = 28 bytes. The first and last 4 bytes of the file will contain the value 20 – showing that 20 bytes of real data are contained in between the two paddings.

The fast-food file consists of a header:

```
int idat[5];
float fdat[9];
float znow;
```

For the purposes of these correlation function codes, the only useful quantity is idat[1] which contains N – the number of particles in the data file.

()	1	2	3	4	5	6	7	8	9 10 11	12 13 14 15 16 17 18 19 20 21 22 23	24 25 26 27
		2	0			Iı	nt			N	3 Integers	20

The next 56 bytes contain two other fields (10 floats + 4 padding bytes of 4 bytes each) and their corresponding paddings. Since these bytes do not

contain any data that are useful in the context of these codes we will skip over the contents of these bytes. In io.c, I actually fseek over these bytes.

After this header, the actual X/Y/Z values are stored. The first 4 bytes after the header contains N*4, for float precision or N*8, for double precision where N=idat[1], is the number of particles in the file. After all the X values there will be another 4 bytes containing N*4 or N*8. Note, that even when the X/Y/Z arrays are written out in double-precision, the padding is still 4 bytes.

Byte-structure of the X/Y/Z arrays in a fast-food file

floats	doubles		
0 1 2 3	0 1 2 3 4 5 6 7		
N*4	N*8		
X[0]	X[0]		
X[1]	X[1]		
:	:		
X[N-1]	X[N-1]		
N*4	N*8		

3.2 Specifying the radial bins

The codes were intentionally designed to read in a set of (somewhat) arbitrary⁴ set of bin specifications from a file. This way, you can specify disjoint bin-edges as well as use 0.0 as a bin edge (which would be impossible if log bins are assumed). The bins are to be specified in a (white-space separated) text file in this manner:

```
r_low[0] r_high[0]
r_low[1] r_high[1]
.
.
.
.
r_low[nbins-1] r_high[nbins-1]
```

⁴I assume bins are non-overlapping

where, r_low[i] and r_high[i] are the left and right edges of the i'th bin respectively. The text files should contains as many lines as the number of bins desired. The logbins executable can be used to create such a text file containing log bins. The syntax for running logbins is:

./logbins <rmin> <rmax> <nbins> > <text filename>
An example of such a file with radial bins is the file bins in the tests

Note, all of the three codes print the correlation function to **stdout** – so be sure to redirect **stdout** to an output file.

3.3 Running $\xi(r)$

directory.

To run the correlation function in 3-D, you will need to run the executable DD from either the bin or xi_of_r directory. The inputs to DD are:

- file1 the file name for the first file.
- format1 the file format for the first file.
- file2 the file name for the second file.
- format2 the file format for the second file.
- binfile the file name for the file containing the bins (see Section 3.2)
- nthreads the number of OpenMP threads to use (only required when the Makefile option USE_OMP is enabled)

Thus, the code can compute an auto-correlation (when file1,format1 and file2,format2 are identical) or a cross-correlation (when file1 and file2 are different). Out of the box, some sample DD call directory) can be:

- auto-correlation ./DD ../tests/data/gals_Mr19.ff f ../tests/data/gals_Mr19.ff f ../tests/bins 4 > Mr19_output.DD
- Cross-correlation ./DD ../tests/data/cmassmock_Zspace.ff f
 ../tests/data/random_Zspace.ff f ../tests/bins 4 > cmass_output.DR

The output (printed to stdout) has nbins columns <Npairs> <rpavg> <rmin> <rmax>, where the <rpavg> column contains 0.0 unless the Makefile option OUTPUT_RPAVG has been enabled. Here, each line of the output represents the i'th radial bin.

3.4 Running $\xi(r_p, \pi)$

To run the correlation function in 2-D for $\xi(r_p, \pi)$, you will need to run the executable DDrppi from either the bin or xi_rp_pi directory. The inputs to DDrppi are:

- file1 the file name for the first file.
- format1 the file format for the first file.
- file2 the file name for the second file.
- format2 the file format for the second file.
- binfile the file name for the file containing the bins (see Section 3.2)
- pimax the maximum distance to consider in the π direction. The code uses Z axis as the π (line-of-sight) direction.
- nthreads the number of OpenMP threads to use (only required when the Makefile option USE_OMP is enabled)

Thus, the code can compute an auto-correlation (when file1,format1 and file2,format2 are identical) or a cross-correlation (when file1 and file2 are different). Out of the box, some sample DDrppi call directory) can be:

- auto-correlation ./DDrppi ../tests/data/gals_Mr19.ff f
 ../tests/data/gals_Mr19.ff f ../tests/bins 40.0 4 > Mr19_output_rppi.DD
- Cross-correlation ./DDrppi ../tests/data/cmassmock_Zspace.ff f ../tests/data/random_Zspace.ff f ../tests/bins 80.0 4 > cmass_output_rppi.DR

The output (printed to stdout) has nbins columns <Npairs> <rpavg> <log(rmin)> <pi_upper>, where the <rpavg> column contains 0.0 unless the Makefile option OUTPUT_RPAVG has been enabled. The code bins in 1 Mpc/h in the π direction by default. Therefore, the total number of bins in the output file will be $nbins \times \pi_{max}$, where each radial bin is further broken into π_{max} bins along the π direction.

There is a code wprp in the xi_rp_pi directory that can combine the output of DDrppi for DD, DR and RR counts and use Landy-Szalay estimator to produce a projected correlation function.

3.5 Running $w_p(r_p)$

To run the projected correlation function, you will need to run the executable wp from either the bin or wp directory. The inputs to wp are:

- boxsize the boxsize for the periodic box.
- file file name.
- format file format.
- binfile the file name for the file containing the bins (see Section 3.2)
- pimax the maximum distance to consider in the π direction. The code uses Z axis as the π (line-of-sight) direction.
- nthreads the number of OpenMP threads to use (only required when the Makefile option USE_OMP is enabled)

The wp code only computes an auto-correlation with PERIODIC boundary conditions, irrespective of the options set in common.mk.

• ./wp 420.0 ../tests/data/gals_Mr19.ff f ../tests/bins 40.0 4 > Mr19_output.wp

The output (printed to stdout) has nbins columns <wp> <rpavg> <rmin> <rmax> <Npairs>, where the <rpavg> column contains 0.0 unless the Makefile option OUTPUT_RPAVG has been enabled.

4 Code Design

The following sections are taken directly from the associated paper. The fundamental idea behind the code can be broken down into the following steps:

- Given the max. separation, r_{max} , for the correlation function, grid the entire the particle distribution with cell width r_{max} . In case of $\xi(r_p, \pi)$ and $w_p(r_p)$, the X/Y bins are r_{max} while the Z bins is π_{max} .
- In each 3-D cell, store all the particles that are located in that cell in contiguous X/Y/Z arrays. This is contained in a structure defined in 4.1.
- Loop over all particles in every cell (this loop uses OpenMP parallelization if USE_OMP is enabled). This is the index1 loop in countpairs.c, countpairs_rp_pi.c and countpairs_wp.c. The loop variable for the loop over each particle is i.
- Now, for every particle in the input list, we only need to look at particles in the neighbouring cells (see Fig. 1).
- Loop over neighbouring cells in all 3 dimensions. Corresponding variables are iiix, iiiy, iiiz. These variables are constructed from the loop variables iix, iiy, iiz to account for periodic boundary conditions.
- Once we have a triplet of iiix, iiiy, iiiz, we can construct the index for the neighbouring cell. Variable index2 gives the index for the neighbouring cell.
- Compute the distances between target particle and some particle bunch (where a chunk is 4 double or 8 floats) using AVX intrinsics. The slower, non-AVX version of the code computes the distances serially and does not take advantage of the CPU architecture. This is the loop with loop variable j in countpairs.c, countpairs_rp_pi.c, countpairs_wp.c.

⁵I apologize for this atrocious variable naming scheme.

- Check if any of the distances are less than r_{max} . If not, continue processing the data. If yes, then update the histograms using AVX bit-masks and popcount.
- Continue until all particles in neighbouring cell are done.

4.1 How to Maintain Cache Locality within the Grid

For all pairs around a given target galaxy, we need to compute distances to all points within all neighbouring 3-d cells. We ensure that the particle locations are contiguous by moving them into the following C struct in the order in which they arrive.

```
typedef struct{
  DOUBLE *x;
  DOUBLE *y;
  DOUBLE *z;
  int64_t nelements;
} cellarray;
```

Listing 4.1: Definition of the cellarray structure. This structure contains the X/Y/Z positions of all the particles that are in one 3-D cell.

5 Calling the C Libraries

All of the correlation function codes create a corresponding static library rather than a dynamic/shared library. This was a design decision intended to minimize path-issues for the end-user. After the libraries have been created, it is fairly straightforward to use them in an external C/python code.

5.1 C bindings

The examples contains the files run_correlations.c that shows how to use the three types of correlation function libraries from C. Essentially, the process consists of including the appropriate header file and passing the appropriate arrays in the functions.

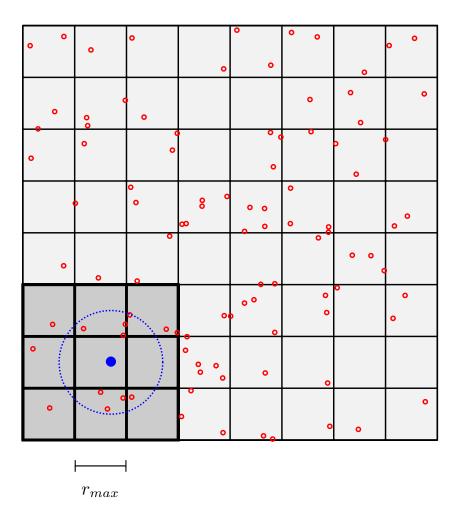


Figure 1: A 2-D grid showing the bin-lattice partitioning scheme. The bigger square show the entire domain, the red circles show a random distribution of 100 particles. Let's say we want to compute all pairs for the target blue point, then we would only have to consider red points that are within one cell (the dark shaded region). A circle with radius r_{max} is also drawn to shown the actual pairs that will eventually count in the correlation function.

5.1.1 API for $\xi(r)$

```
typedef struct{
  uint64_t *npairs;
  DOUBLE *rupp;
  DOUBLE *rpavg;
  int nbin;
} results_countpairs;
results_countpairs * countpairs(
const int64_t ND1, const DOUBLE * const X1, const DOUBLE * const Y1, const
    DOUBLE * const Z1,
const int64_t ND2, const DOUBLE * const X2, const DOUBLE * const Y2, const
    DOUBLE * const Z2,
#ifdef USE_OMP
const int numthreads,
#endif
const int autocorr,
const char *binfile) __attribute__((warn_unused_result));
void free_results(results_countpairs **results);
5.1.2 API for \xi(r_p, \pi)
typedef struct{
  uint64_t *npairs;
  DOUBLE *rupp;
  DOUBLE *rpavg;
  DOUBLE\ pimax;
  int nbin;
  int npibin;
} results_countpairs_rp_pi;
results_countpairs_rp_pi * countpairs_rp_pi(
const int64_t ND1, const DOUBLE *X1, const DOUBLE *Y1, const DOUBLE *Z1,
const int64_t ND2, const DOUBLE *X2, const DOUBLE *Y2, const DOUBLE *Z2,
#ifdef USE_OMP
const int numthreads,
#endif
const int autocorr,
const char *binfile ,
                      __attribute__ ((warn_unused_result));
const double pimax)
void free_results_rp_pi(results_countpairs_rp_pi **results);
5.1.3
        API for w_p(r_p)
typedef struct{
  uint64_t *npairs;
  DOUBLE * wp;
  DOUBLE *rupp;
  DOUBLE *rpavg;
  DOUBLE pimax;
  int nbin;
} results_countpairs_wp;
```

```
results_countpairs_wp *countpairs_wp(
const int64_t ND1, DOUBLE * restrict X1, DOUBLE * restrict Y1, DOUBLE *
    restrict Z1,
const double boxsize,
#ifdef USE_OMP
const int numthreads,
#endif
const char *binfile,
const double pimax) __attribute__((warn_unused_result));
void free_results_wp(results_countpairs_wp **results);
```

5.2 Python Bindings

The python_bindings directory contains python bindings for python 2.x. Note that python3 is not supported out of the box⁶. If all went well, then typing python call_correlation_functions.py should run the example python code. If you get an error (and you are on a MAC), then refer to Section 2.4 or the FAQ.

If you edit the common.mk file and compile for double precision arithmetic, then be sure to change the line:

```
dtype=np.float32
to
dtype=np.float64.
Otherwise, you will get a TypeError at runtime.
```

6 Extending the Code

6.1 Different Type of Input Data File

All of the codes use io.c in the io sub-directory to read-in the data. If you want to specify a different file format, the easiest way would be to edit io.c. Decide on the file format code and add another strncmp case in io.c. Remember that the x/y/z are declared as void pointers, so you can not directly reference the x/y/z pointers. If you do add support for a different file-type, please submit a pull request and I will be happy to merge it into the code-base.

⁶In the future, I might switch to cython to cover both python2 and python3

6.2 Computing a different type of correlation function

Let's say, you want to compute a marked correlation function. Now, you will need to read-in/create the marks for each individual point. And then you will have to add an appropriate field to the cellarray structure (see 4.1).

6.3 Using SSE instead of AVX

If your CPU is too old and does not support AVX, then you can still use SSE instrinsics to compute the correlation functions. However, this will require replacing all of the AVX sections with corresponding SSE intrinsics. Email me and I will guide you through the conversion process.