大学物理笔记

多普勒效应

source: v_s , observer: v_O , distance L

issued at t=0, x=0, recieved at $t_1^\prime, L+v_Ot_1^\prime$

issued at $t=T, x=v_ST$, received at $t_2^\prime, L+v_Ot_2^\prime$

由运动学:

$$egin{aligned} v_P t_1' &= L + v_O t_1' \ v_P (t_2' - T) &= L + v_O t_2' - v_S T \end{aligned}$$

于是

$$t_1' = rac{L}{v_P - v_O}$$
 $t_2' = rac{L}{v_P - v_O} + rac{v_P - v_S}{v_P - v_O}T$
 $T' = t_2' - t_1' = rac{v_P - v_S}{v_P - v_O}T$
 $u' = rac{v_P - v_O}{v_P - v_S}
u$

Discussion:

 $v_O > v_S$ (including $v_S \leq 0$):separating, u' <
u, Doppler redshift

 $v_S>v_O$ (including $v_O\leq 0$):approaching, u'>
u , Doppler blueshift

靠近:频率变高;远离:频率变低

考虑波源的运动与波振面的运动(supersonic speed $v_S>v_p$)

马赫锥: $sin heta=rac{vP}{vS}$

驻波(stationary wave)

推导

由
$$kx-\omega t-\phi=const$$
,求导得 $v=rac{\omega}{k}$
$$u_1=Acos(kx-\omega t-\phi)$$

$$u_2=Acos(kx+\omega t)$$

$$=Acos(k(x+rac{\omega}{k}t))$$
 (1) 从而 $u_1+u_2=2Acos(kx-rac{\phi}{2})cos(wt+rac{\phi}{2})$

令
$$(u_1+u_2)_{x=0}=0$$
,得 $\phi=\pi$; 令 $(u_1+u_2)_{x=L}=0$,得 $L=\frac{n\pi}{k}=\frac{n\lambda}{2}$
注意 $\lambda=T\cdot v_P=\frac{2\pi}{\omega}\cdot\frac{\omega}{k}=\frac{2\pi}{k}$, $k=\frac{2\pi}{\lambda}$, $\omega=vk$,有
$$f=\frac{\omega}{2\pi}=\frac{kv}{2\pi}=n\frac{v}{2L}$$
 $=f_n$ $=nf_1$

波节(node),波腹(antinode)的定义

• 半波损失

fixed hard boundary; from high spped to low speed

由
$$(1)$$
式,令 $(u_1+u_2)_{x=0}=0$,得 $\phi=\pi$,即为在固定端的相差

第9章 Relativistic Mechanics 相对论力学

9.1 伽利略变换

- 观测者:一套矫正好的时钟,可记录 x, y, z, t
- ・ 结果 $t_A'-t_B'=t_A-t_B$, $x_A'-x_B'=x_A-x_B-u(t_A-t_B)$ 同时测量: $x_A'-x_B'=x_A-x_B$

速度直接叠加加速度不变

- Galileo's relativity: 力学定律在不同惯性参考系中相同
- 电学:不满足麦克斯韦方程组

9.2 洛伦兹变换 Lorentz Transformation

- Two basic principles of Special Theory of Relativity
 - 所有惯性参考系中物理定律相同(力学,电学等)
 - 。 光速不变
- 推导

假定变换为线性变换(考虑匀速直线运动);

待定系数:

$$x' = a_{11}x + a_{12}y + a_{13}z + a_{14}t \ y' = y \ z' = z$$

For arbitrary y,z if x'=0,x=ut;则

$$x' = a_{11}(u)(x - ut), a_{11}(0) = 1 (1.1)$$

By the priciple of relativity, we have

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$$x = a_{11}(-u)(x'+ut')$$
, 带入(1.1), 有
 $t' = f(x, x'(u, t)) = a_{44}t + a_{41}x$

从而

$$x' = a(x - ut)$$

$$t' = b(t - ex)$$
(1.2)

解得 $x=rac{1}{\Delta}(bx'+aut')$

与
$$x = a_{11}(-u)(x'+ut')$$
比较,得 $a = b$

考虑光速不变:假设 t=t'=0时在原点发出光信号,有

$$x^{2} + y^{2} + z^{2} = c^{2}t^{2}$$

 $x'^{2} + y'^{2} + z'^{2} = c^{2}t'^{2}$

代入 (1.2)式,有

$$e=rac{u}{c^2}$$
 $a=rac{1}{\sqrt{1-eta^2}}=\gamma=\sqrt{1-rac{v^2}{c^2}}$ 其中 $eta=rac{u}{c}$

完整的变换关系为

$$x' = \gamma(x - ut)$$

 $y' = y$
 $z' = z$
 $t' = \gamma(t - \frac{u}{c^2}x)$

考虑事件对 $(x'_1, t'_1), (x'_2, t'_2)$,洛伦兹变换是

$$\Delta x' = \gamma(\Delta x - u\Delta t)$$
 $\Delta y' = \Delta y$
 $\Delta z' = \Delta z$

$$\Delta t' = \gamma(\Delta t - \frac{u}{c^2}\Delta x)$$
(1.3)

- 讨论
 - \circ 逆变换: $u \rightarrow -u$ 直接根据相对性原理
 - 。 Invariance of spacetime interval 时空间隔不变: $(\Delta s)^2=(c\Delta t)^2-(\Delta x)^2-(\Delta y)^2-(\Delta z)^2=(\Delta s')^2$
 - 。 伽利略变换:时间归时间 空间归空间 时间空间分别不变
- 事件对
 - \circ 光信号联系: $(c\Delta t)^2=(\Delta x)^2, \Delta s=0$
 - 。 固有时间间隔与固有空间间隔

Simultaneous measurement $\Delta x'=\ell_0$ in $S'(\Delta t'=0)$ 形成一个事件 the same pair of events in S:

$$\Delta x = \gamma (\Delta x' + u \Delta t') = \gamma \Delta x' = \gamma \ell_0$$
 同时的相对性 $\Delta t = \gamma (\Delta t + \frac{u}{c^2} \Delta x') = 0$

也即
$$\ell=\ell_0\sqrt{1-rac{v^2}{c^2}}=rac{\ell_0}{\gamma}$$

• 用光速不变推出钟慢:

$$\Delta t'=rac{2L_0}{c} \ (rac{u\Delta t}{2})^2+L_0^2=(rac{c\Delta t}{2})^2($$
用到光速不变 $)$ $\Delta t=rac{2L_0}{\sqrt{c^2-u^2}}=rac{2L_0/c}{\sqrt{1-eta^2}}=\gamma\Delta t'$

• 用光速不变推出尺缩:

$$c\Delta t_1 = L + u\Delta t_1, c\Delta t_2 = L - u\Delta t_2$$
 $\Delta t = \Delta t_1 + \Delta t_2 = \frac{L}{c+u} + \frac{L}{c-u} = \frac{2L}{c} \frac{1}{1-\beta^2}$
 $\Delta t' = \frac{2L_0}{c} = \gamma \Delta t$
 $\frac{2L_0}{c\gamma} = \frac{2L}{c} \frac{1}{1-\beta^2}$
于是 $L = \frac{L_0}{\gamma}$

- 因果律及信号传播(Causality and signal speed)
 - \circ An event $P(x_P,t_P)$ causes an event $Q(x_Q,t_Q)$. 信号传输速度

$$v_S = rac{x_Q - x_P}{t_Q - t_P} = rac{\Delta x}{\Delta t}$$

在 S'(u) 系中时间间隔为

$$\Delta t' = \gamma \Delta t (1 - \frac{uv_S}{c^2})$$

为保证先后次序,必须有

$$1-rac{uv_S}{c^2} < 0$$
则 $v_S < c$

• 光多普勒效应

$$\circ$$
 $ext{th}(1.3), \Delta t = \sqrt{rac{1+eta}{1-eta}} \Delta t'$

$$\circ \;\; {
m redshift:} \,
u_0 = \sqrt{rac{1-eta}{1+eta}}
u_e$$

• 波动方程: 用洛伦兹变换恰好符合(洛伦兹不变)

9.3 时空图和孪生paradox

- Minkowski Space
 - 。 洛伦兹变换在时空图中的体现
- Twin paradox: 在宇宙飞船加速,减速的过程中经历了非惯性系
- Pole-barn Paradox: 火车进隧道;火车系:隧道收缩;地面系:火车收缩
 - 涉及同时的相对性:看能否在地面系同时关上前门和后门(火车系觉得没有同时)
 - 。 地面系的观测者觉得能关进去,在火车系的观测者看来后门早就关起来了
- Visual Apperance
 - 。 球不变 立方体转过一个角度

9.4 相对论运动学

• 速度变换公式(求导可得)

$$v'_{x} = rac{v_{x} - u}{1 - v_{x} rac{u}{c^{2}}} \ v'_{y} = rac{v_{y}}{\gamma(1 - v_{x} rac{u}{c^{2}})} \ v'_{z} = rac{v_{z}}{\gamma(1 - v_{x} rac{u}{c^{2}})}$$

9.5 相对论动力学(Relativistic Dynamics)

- 广义相对论: 引力场,非惯性系
- 讨论 m, p, F, E_k, a
- Discuss a completely inelastic collision(保持动量守恒)

$$\circ \ \ \operatorname{In} S'(-u): u, -u \to 0, 0$$

$$\circ$$
 In $S:v_x,0 o u,u;$ $v_x=rac{u+u}{1+urac{u}{c^2}}(9.5.1)$

$$\circ$$
 而 $mrac{u+u}{1+rac{uu}{2}}+m imes 0=2mu$,动量守恒不成立.

• 假定
$$m \equiv m(v), m_0 \equiv m(0)$$
,有

$$egin{aligned} m(v_x)+m_0&=m_{total}(u)\ m(v_x)\cdot v_x+m_0\cdot 0&=m_t(u)u\ &=[m(v_x)+m_0]u\ rac{x}{lpha m(v_x)}&=rac{m_0}{rac{v_x}{u}-1} \end{aligned}$$

由
$$(9.5.1)$$
, $\dfrac{v_x}{u}-1=\sqrt{1-eta^2}$,则 $m(v)=\gamma m_0$. 于是

$$egin{aligned} p &= \gamma m_0 v \ E_k &= \int F \cdot dx = \int rac{d}{dt}(mv) \cdot dx = \int v \cdot d(mv) \ orall d(mv) &= v dm + m dv \ orall E_k &= \int (v^2 dm + mv dv) \end{aligned}$$

考虑
$$m^2(1-rac{v^2}{c^2})=m_0^2, m^2(c^2-v^2)=m_0^2c^2$$
,两边微分得
$$c^22mdm-v^22mdm-m^22vdv=0, \\ c^2dm-v^2dm-mdv=0, c^2dm=v^2dm+mdv$$

于是

$$egin{aligned} E_k &= \int c^2 dm = mc^2 - m_0 c^2 \ &= m_0 c^2 \Big[rac{1}{\sqrt{1 - rac{v^2}{c^2}}} - 1 \Big] \ &= (\gamma - 1) m_0 c^2 \end{aligned}$$

- 讨论
 - 。 牛顿力学中

$$E_k = \int v \cdot d(mv) = \int mv \cdot dv = rac{1}{2} mv^2$$

- 相对论中,mass-energy $E=mc^2=E_k+m_0c^2$ 已知动能求速度: $\gamma=rac{E_k}{m_0c^2}+1$
- 相对论能量-动量关系

$$\circ$$
 $E=\sqrt{c^2p^2+m_0^2c^4}$

• 若
$$m_0 = 0, E = cp, p = \frac{E}{c}$$

• 若 $cp << m_0 c^2$,泰勒展开后得到牛顿力学.

期末考试大致内容

- 相对论
- 理想气体(或给定气体的状态方程) 等温膨胀,绝热等过程, 功和热等
- 理想气体的性质: 速度,动能等
- 过程中温度的变化,熵的变化
- 亥姆霍兹自由能, 求平衡态(可能有数学上求导等)
- 比热容, 温度.
- 相变 三相点.

第十章 温度

10.1 Equilibruim state

System	exchange
isolated	×
closed	energy
open	energy&matter

- Thermodynamic Equilibruim state: 宏观状态不随时间改变,除非外界条件发生变化.
- description of equilibruim state: state variables
 - extensive quantity: $F(n \ systems) = nF(1 \ system)$
 - \circ intensive quantity: $F(n \ systems) = F(1 \ system)$
- relaxation time *⊤*: 恢复平衡所用时间
- Quasi-static process 准静态过程: $t >> \tau$, 每一点都是平衡态, 可以在状态图上画出.
- No-dissipative, quasi-static process is reversible. (无耗散准静态过程是可逆的); All natural process(自发过程) is irreversible.
 - 。 系统和外界同时恢复初态

10.2 Thermal equilibruim and temperature

- The zeroth law of thermodynamics: 热平衡的传递性(热平衡的系统有共同性质)
- 温度: 衡量热平衡的一种物理性质
- 各种温标

10.4 物态方程

• 微分的常见写法
$$(\frac{\partial U}{\partial V})_T$$
 , $(\frac{\partial U}{\partial V})_P$. 保持右下标不变

•
$$P, V, T$$
不独立, $f(P, V, T) = 0 \rightarrow U(T, P) = U(T, V(P, T)) = U(T, V) = 0$

• 等压线膨胀系数
$$a_\ell=rac{1}{L}rac{\partial L}{\partial T}$$
; $lpha_V=rac{1}{V}ig(rac{\partial V}{\partial T}ig)_P=rac{1}{L^3}ig(rac{\partial L^3}{\partial T}ig)_P=3lpha_\ell$

- 等温线膨胀系数 $\kappa_T = -rac{1}{V}rac{\partial V}{\partial p_T}$
- 体积变化方程

$$egin{aligned} rac{dV}{V} &= rac{1}{V} ig(rac{\partial V}{\partial T}ig)_p dT + rac{1}{V} ig(rac{\partial V}{\partial p}ig)_T dp = lpha_V(T) dT - \kappa_T(p) dp \ ig(\hat{\Xi} ig) ig) \end{aligned}$$
 $egin{aligned} lnrac{V}{V_0} &= lpha_V \Delta T - \kappa_T \Delta p \ V &= V_0 (1 + lpha_V \Delta T - \kappa_T \Delta p) \end{aligned}$

Van der Waal's equation

$$\left[p+a(\frac{n}{V})^2\right](V-nb)=nRT$$

推导: 对1mol理想气体:

$$p = \frac{RT}{V_m} \leftarrow \frac{RT}{V_m - b}$$
 (相互作用,体积减小)
$$\leftarrow \frac{RT}{V_m - b} - \frac{a}{V^2}$$
 (相互作用,压强减小,体积越小作用力越大)

第十一章 热力学第一定律

11.1 功,内能和热力学第一定律

- 准静态过程压活塞 $ar{d}W=-par{d}V$, $W=-\int_{V_i}^{V_f}pdV$ (依赖于 P-V图上的路径,不是全微分)
- 绝热功: Internal energy function $\Delta U = U_B U_A \equiv W_{BA} \equiv -W_{AB}$ (用与路径无关的绝热功定义内能变化)
- 非绝热功: $\Delta U=W+Q$ 外界做的功和外界传给内部的热量; 微分 $dU=\bar{d}Q+\bar{d}W$ (功和热不是全微分,与路径有关)

11.3 热容和比热容(Heat capacity and specific heat capacity)

•
$$C = \lim_{\Delta T o 0} rac{\Delta Q}{\Delta T}$$
; specific: $C_m = rac{C}{n}$

- 。 C=0: 绝热(adiabatic)过程, $\Delta Q=0$
- 。 $C=\infty$: 等温(isothermal)过程, $\Delta T=0$

•
$$C_V = \lim_{\Delta T \to 0} (\frac{\Delta Q}{\Delta T})_V = \lim_{\Delta T \to 0} (\frac{\Delta U}{\Delta T})_V = (\frac{\partial U}{\partial T})_V$$

•
$$C_P = \lim_{\Delta T \to 0} (\frac{\Delta Q}{\Delta T})_p = \lim_{\Delta T \to 0} (\frac{\Delta U + p\Delta V}{\Delta T})_p = (\frac{\partial (U + pV)}{\partial T})_p \equiv (\frac{\partial H}{\partial T})_{p'} H = U + pV$$
: enthalpy, 焓

- ratio of specific heat: $\gamma = \frac{C_p}{C_V}$
- heat capacity of ideal gases

monatomic ideal gas
$$C_V=rac{3}{2}nR$$
 双原子 ideal gas $C_V=rac{5}{2}nR$ 三原子 ideal gas $C_V=rac{6}{2}nR$

11.4 Free expansion and internal energy of gas

• Case of free expansion

$$U = U(T, V)$$

$$dU = \left(\frac{\partial U}{\partial T}\right)_V dT + \left(\frac{\partial U}{\partial V}\right)_T dV$$

$$T = T(U, V)$$

$$dU = \left(\frac{\partial U}{\partial T}\right)_V \left[\left(\frac{\partial T}{\partial U}\right)_V dU + \left(\frac{\partial T}{\partial V}\right)_U dV\right] + \left(\frac{\partial U}{\partial V}\right)_T dV$$
对照左右,得
$$\left(\frac{\partial U}{\partial T}\right)_V \left(\frac{\partial T}{\partial U}\right)_V = 1$$

$$\left(\frac{\partial U}{\partial V}\right)_T = -\left(\frac{\partial U}{\partial T}\right)_V \left(\frac{\partial T}{\partial V}\right)_U = -C_V \left(\frac{\partial T}{\partial V}\right)_U \to$$
無耳系数 = 0(理想气体) 故 $U = U(T)$

- 焦耳的实验非常不严谨
- 理想气体:

$$dU(T) = rac{dU}{dT}dT = C_V dT$$
 $U = U_0 + C_V T$
 $dQ = dU - dW = C_V dT + p dV$
理想气体: $d(pV) = d(nRT)$
 $p dV + V dp = nR dT, p dV = nR dT - V dp$
 $dQ = (C_V + nR) dT - V dp$
 $C_p = C_V + nR o C_V = rac{nR}{\gamma - 1}, C_p = rac{\gamma nR}{\gamma - 1}$

11.5 Adiabatic equation

$$egin{aligned} dQ &= dU - dW, \ &= C_V dT + p dV \ C_V dT + p dV &= 0 \ orall V dp + p dV &= nRdT = (\gamma - 1)C_V dT \end{aligned}$$

消去 dT,得

$$egin{aligned} Vdp + \gamma pdV &= 0 \ rac{dp}{p} + \gamma rac{dV}{V} &= 0 \ lnp + \gamma lnV &= C \ pV^{\gamma} &= C \end{aligned}$$

由 $\gamma > 1$, 绝热线比等温线陡峭

· adiabatic work

$$egin{align} W_S &= -\int_{V_1}^{V_2} p dV = -C \int_{V_1}^{V_2} rac{dV}{V^\gamma} = rac{C}{\gamma-1} (rac{V_2}{V_2^\gamma} - rac{V_1}{V_1^\gamma}) \ &= rac{1}{\gamma-1} (p_2 V_2 - p_1 V_1) \ &= rac{nR}{\gamma-1} (T_2 - T_1) \ &= C_V (T_2 - T_1) = \Delta U \ \end{cases}$$

• 声音的传播: 绝热

11.6 卡诺循环

- 高温热源 T_1 等温膨胀 \rightarrow 绝热膨胀 \rightarrow 低温热源 T_2 等温压缩 \rightarrow 绝热压缩
- $T_1, U = U(T_1) = const$ 热机吸热,做功
- T_2 热机放热
- 整个循环: $\Delta U=0, \Delta Q_{12}+\Delta Q_{34}-\Delta W=0$

效率
$$\eta = \frac{\Delta W}{\Delta Q_{12}} = \frac{\Delta Q_{12} + \Delta Q_{34}}{\Delta Q_{12}} = \frac{Q_1 - Q_2}{Q_1}$$
 (考察做功过程中的放热 Q_2 , 放热越多效率越低) 等温过程: $Q_1 = \Delta W_{12} = \int_{V_1}^{V_2} p dV = \int_{V_1}^{V_2} \frac{nRT_1}{V} dV = nRT_1 ln \frac{V_2}{V_1}$ $Q_2 = -\Delta Q_{34} = nRT_2 ln \frac{V_3}{V_4}$ $\eta = 1 - \frac{T_2 ln \frac{V_3}{V_4}}{T_1 ln \frac{V_2}{V_1}}$ $2 \rightarrow 3: T_1 V_2^{\gamma-1} = T_2 V_3^{\gamma-1}$ $4 \rightarrow 1: T_1 V_1^{\gamma-1} = T_2 V_4^{\gamma-1}$ $\eta = 1 - \frac{T_2}{T_1}$

要学会将此推导推广到其他热机上

- 可逆过程: 外界不可逆
- 逆卡诺循环: 从低温热源吸热 到高温热源被做功,放热

第十二章 热力学第二定律

12.1 内容

- No process is possible whose sole result is the absorption of heat from a reservoir and the conversion of this heat into work.
- No process is possible whose sole result is the transfer of heat from a cooler to a hotter body.
- 自发过程具有方向性

12.2 卡诺定理

• 任何热机效率不超过可逆热机

• 证明: 任意热机做功驱动可逆热机的逆过程, 反证法与热二矛盾

12.3 熵与熵原理

Q₂ < 0放热:

效率不大于卡诺热机:
$$\eta_A=1+\frac{Q_2}{Q_1}\leq 1-\frac{T_2}{T_1}$$

$$\sum_{i=1}^2\frac{Q_i}{T_i}\leq 0$$
 一般过程: $\sum_{i=1}^n\frac{Q_i}{T_i}\leq 0$ (系统与很多热源接触)
$$\oint \frac{dQ}{T}\leq 0,$$
 取等:可逆
$$写作全微分: $\frac{dQ}{T}\equiv dS$
$$S_B-S_A=\int_A^B(\frac{dQ}{T})_R$$$$

克劳修斯不等式的说明: 考虑多热源与单一热源 T_0 由多个卡诺机联合