

Practicalities of Muon Data Analysis

Francis Pratt (ISIS)

ISIS Muon Training Course

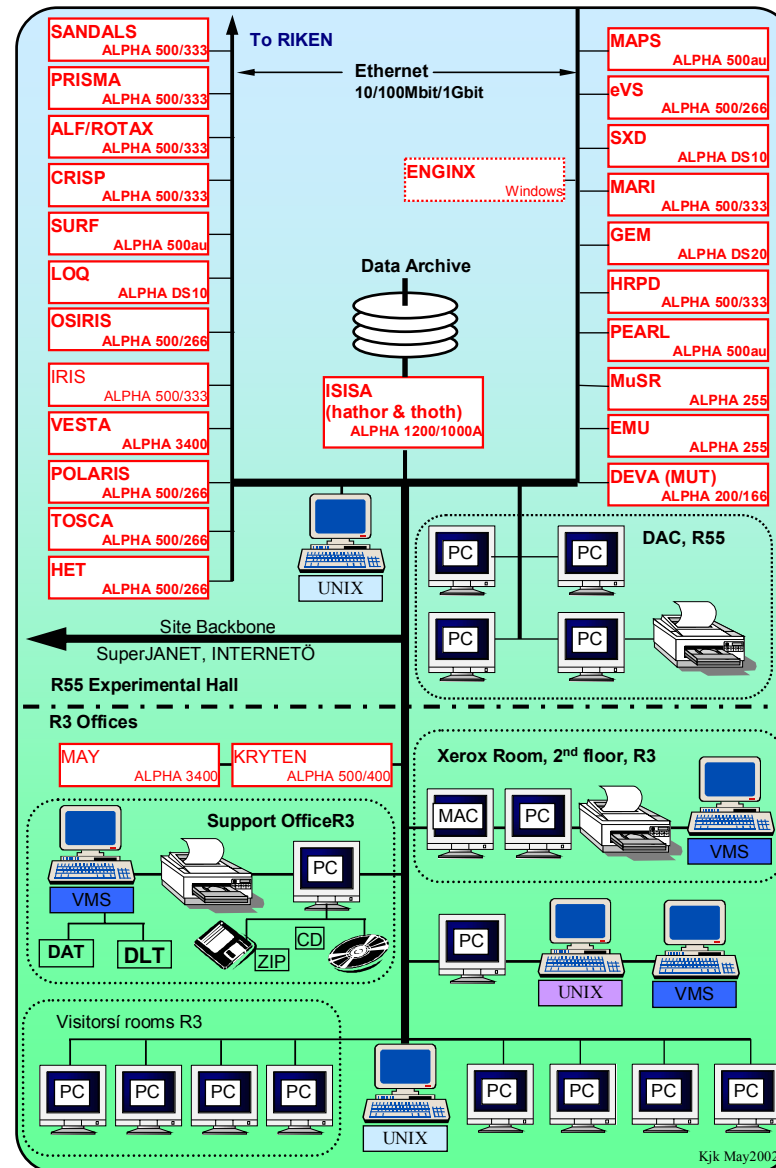
February 2005

Outline of Talk

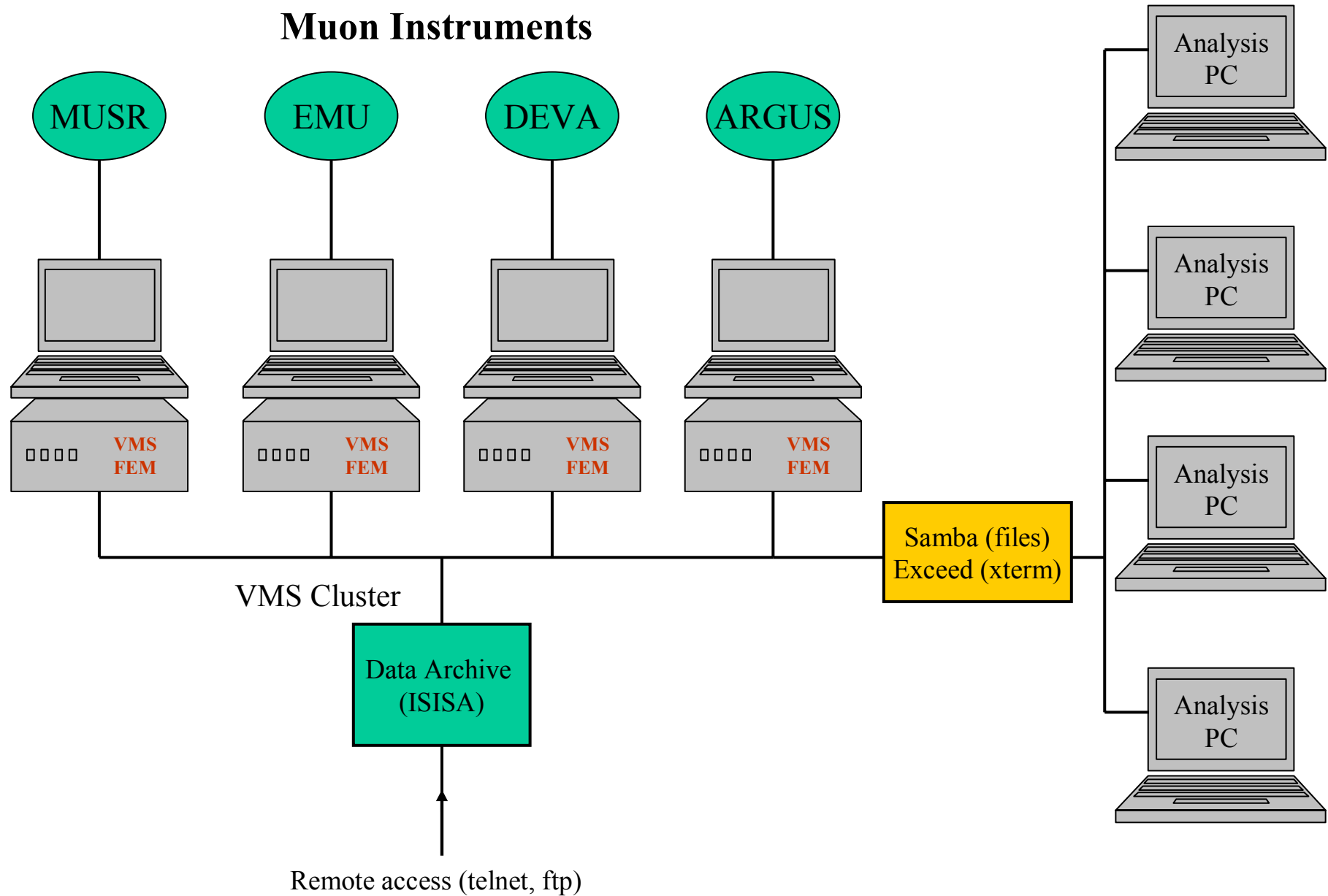
Computing infrastructure at ISIS

- Network organisation
- Front-end and data analysis machines
- Instrument control programs
- Data formats and data access

ISIS Computing (2002)

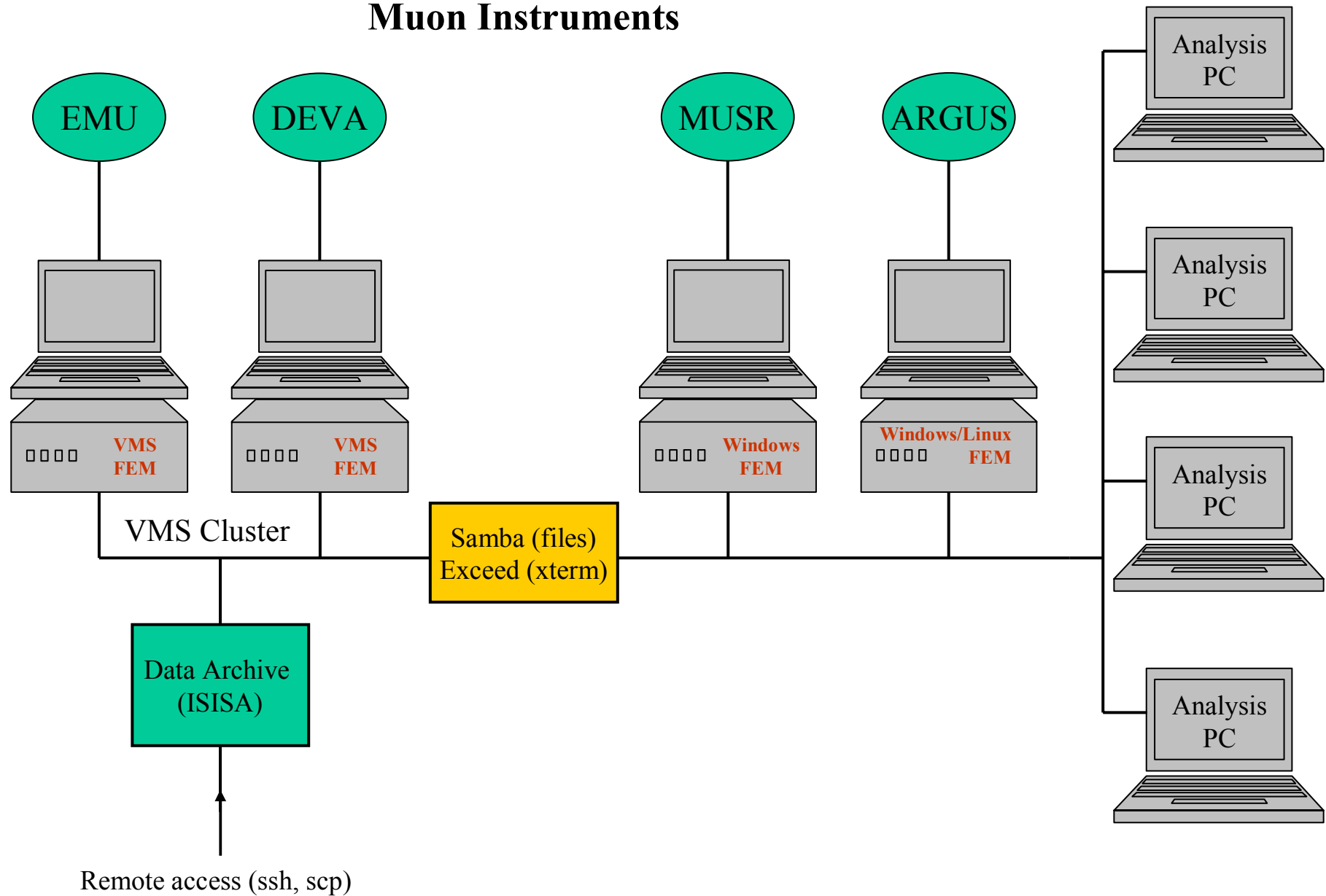


Original ISIS μ SR Computing Layout



Current ISIS μ SR Computing Layout

Muon Instruments



Using the ISIS Computers

- VMS alpha cluster

Login to ISISA from an x-terminal window, use the account details available at each instrument

- Individual PCs

Login as user: isiscnc\muontc (isismuonstc)

ISIS Computing Support can be contacted by emailing
support@isise.rl.ac.uk or ISISsupport@rl.ac.uk
or by phoning extension 1763

Using Personal Laptops at ISIS

Connecting Laptops at ISIS

Visitors to ISIS are welcome to connect their laptop to our network in the R70 hostel as well as the R55 experimental hall and R3 Offices. The ISIS internet connection is behind the CCLRC site firewall, which allows most connections going out, but connections coming in are restricted.

Physical Connections

Network sockets are available in all instrument cabins and public areas. Wireless access is also available in many parts of the R55 experimental hall (not in the MUSR or ARGUS cabins though).

IP address

IP addresses are allocated automatically using DHCP. Please set your network settings to use DHCP and reboot (if you have fixed settings for your home institution it may be worth recording their values)

Mail

Receiving mail should work immediately. Sending SMTP mail needs to go out through **outbox.rl.ac.uk**. Access to web based email will work once a proxy server has been configured.

Web

Web access off site needs to go through a proxy server. Set your browser to use the automatic configuration script **<http://wwwcache.rl.ac.uk/proxy.pac>**

Printers

Access to certain printers is enabled. Please ask your local contact for information on connecting to your nearest printer.

Remote Access

- All external access is via ISIS (isis.rl.ac.uk)
(use ssh for terminal login
and scp for file copying)
- A unified Data Portal covering all of the CLRC Facility data is under development

Data Acquisition Control Systems

Currently we have three different systems:

MCS

The original VMS-based system
(EMU, DEVA)

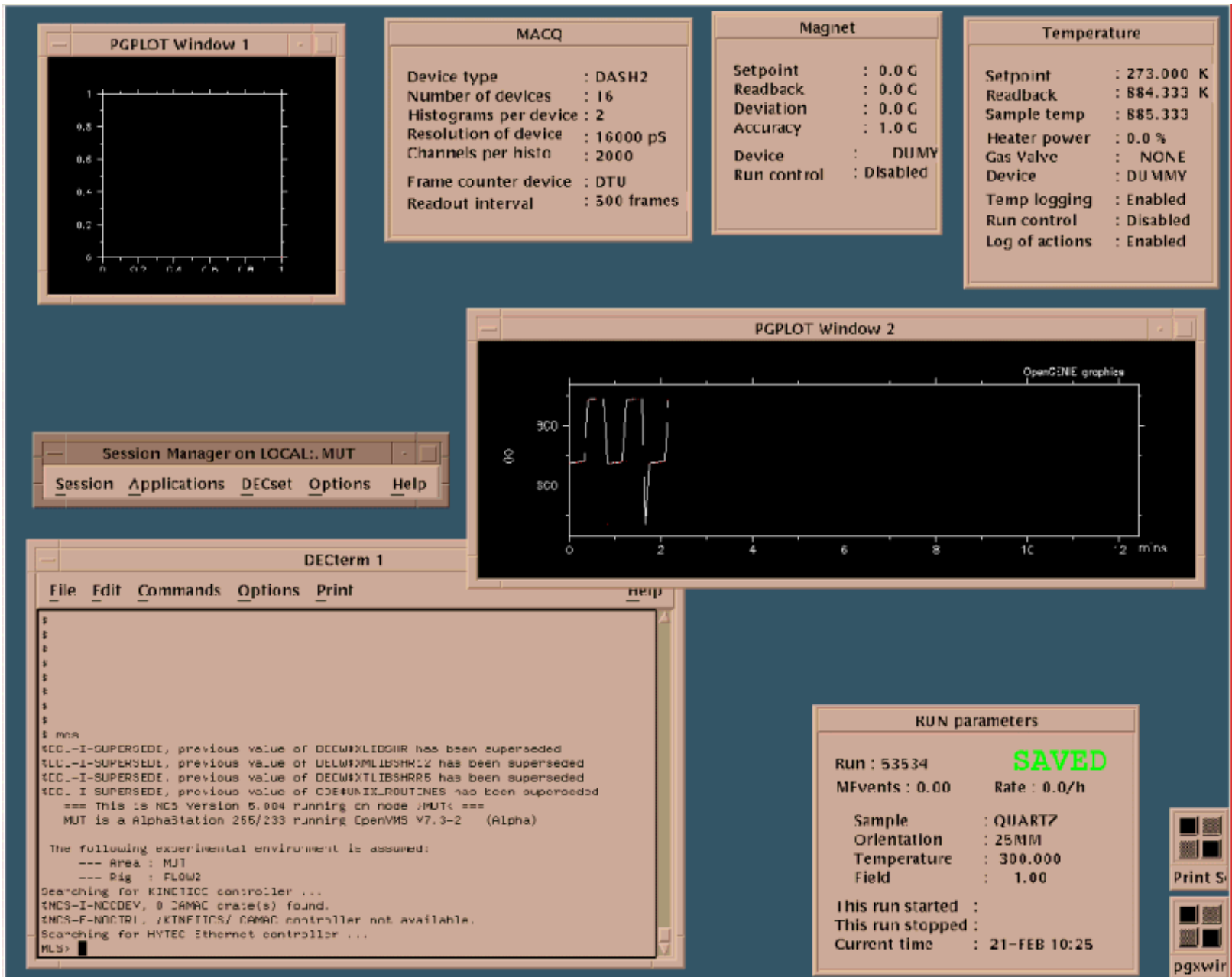
MACS

PC-based system (ARGUS, 2002 on)

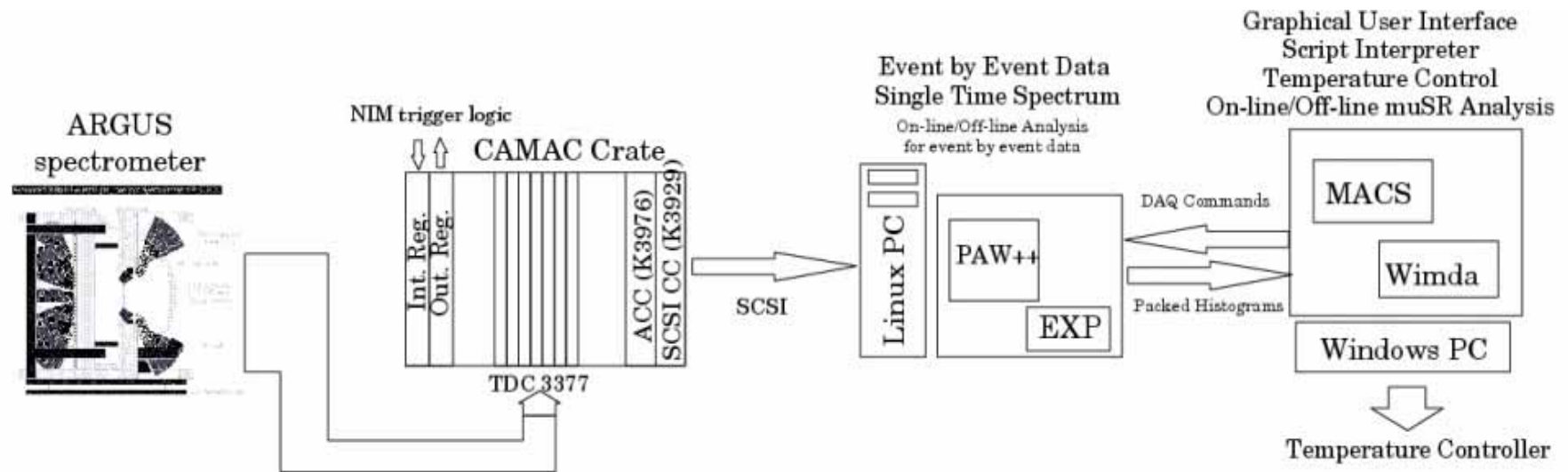
SECI

PC-based system (MUSR, 2004 on)

MCS

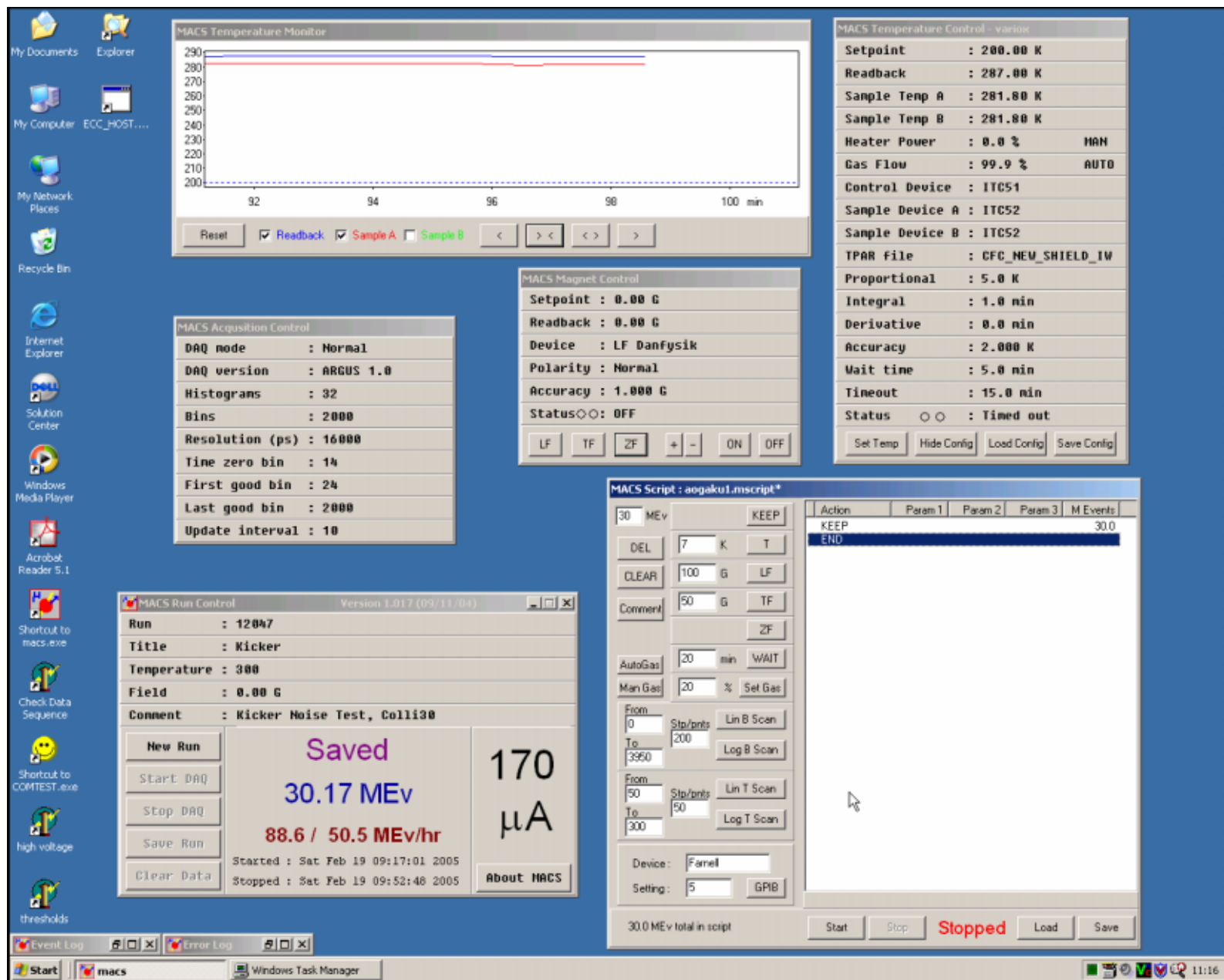


MACS



Hybrid Linux/Windows system

MACS



SECI

ISIS Sample Environment Control Interface, Version: SEC12 - 1.0.1852.20305, SECICOM - 1.0.1852.20310

File Configuration Options Help

Display page: Dashboard Config Name: blue_its503 User: must_mgr

MUSR IS RUNNING, RUN NUMBER: 00001174, TIME: 19/02/2005 11:29:12
TITLE: na1CO T=10.0 F=0.0

RBNumber: 15426
Phone: 6789
User: FB, AO, GL, PM
Total Good: 88769
Total Raw: 88787
This good: 88769
This raw: 88787

Start: Sat 19-Feb-2005 10:57:58
RunTime This: 1871
RunTime Tot: 1871
Beam Curr/Tot: 0/0
Total Counts: 15.47
CountRate: 32.20
Mon Spectrum: 1

Mon From/Ta: 10/200
Mon Counts: 164092
Channels: 4000
Spectra: 64
DAETimeSec: ISIS
Bytes Used: 0
Pds Curr/Tot: 1/1

a_Selected_Magnet: Active ZF
Field_Danlysis: 0 G
Field_T20: 0 G
Field_ZF_Magnitude: 0.00013 G
Field_ZF_Mode: 1
Steer_HSM: -0.066 A
Steer_VSM: -0.37 A
Temp_Cryostat: 10.145 K
Temp_Gas: 49.9 %
Temp_GasAuto: 0
Temp_HeaterAuto: True
Temp_Power: 0.7 %

Instrument
Magnets
Temperature
Beamline
Steering Magnets
Data Collection
DAE
Scripting/Logging
Dashboard
MkScript
Open Genie
Help and Bugzilla
Bugzilla
Machine Status
Beam Status
MCR News

Temperature

Temp_Cryostat
setpoint
Temp_Sam

PAUSE
CLEAR

Magnets

Field_T20: 0 G
Field_T20 Setpoint: 0 G
Field_T20 run control: Off
Field_ZF_Magnitude: 0.00013 G
Field_ZF_Magnitude run control: Off
Field_ZF_Mode: 1
Field_ZF_Mode Setpoint: 1
Field_ZF_Mode run control: Off
Field_Danlysis: 0 G
Field_Danlysis Setpoint: 0 G
Field_Danlysis run control: Off

Steering Magnets

Steer_HSM: -0.066 A
Steer_HSM Setpoint: -0.066 A
Steer_HSM run control: Off
Steer_VSM: -0.37 A
Steer_VSM Setpoint: -0.37 A
Steer_VSM run control: Off

Temperature

Temp_Cryostat: 10.145 K
Temp_Cryostat Setpoint: 10 K
Temp_Cryostat run control: Off
Temp_Sample: 9.959 K
Temp_Sample run control: Off
Temp_Power: 0.7 %
Temp_Power Setpoint: 0.7 %
Temp_Power run control: Off
Temp_HeaterAuto: True
Temp_HeaterAuto Setpoint: 0
Temp_HeaterAuto run control: Off
Temp_Gas: 49.9 %
Temp_Gas Setpoint: 49.9 %
Temp_Gas run control: Off
Temp_GasAuto: 0
Temp_GasAuto Setpoint: 0
Temp_GasAuto run control: Off

Open GENIE

```

Converting Integer parameter to Real
In procedure "SETTEMP", Converting supplied "TEMP=" parameter into a "Real"
>> begin
Sample [na1CO]?
Orientation [powder]?
Temperature [10.0]?
Field [0.0]?
Geometry (L or T) [1]?
RBNo [15426]?
Experimental Team [FB, AO, GL, PM]?
Comment [ ]?
    Informational printing is ON
>> end
Sample = na1CO
Orientation = powder
Temp = 10.0
Field = 0.0
Geometry = 1
RB Number = 15426
Experimental Team = FB, AO, GL, PM
Comment =
Is the run information correct (y/n)? y
  
```

Start DataSocket Server C:\WINDOWS\System32... C:\WINDOWS\System32... LabVIEW Form1 ISIS Sample Environmenk... Open GENIE MkScript3

Address C:\Labview modules\Eurotherms\Source Code Go 11:29 Saturday

Data Formats for Muon Data

- MCS □ VMS Fortran binary (mangled by VMS ftp!)
- MACS □ standard binary
- SECI □ NeXus: hierarchical, extendable format

Typical run file is 200-600 kb in size

Data compression reduces the size by up to a factor of 7
(bzip2 is the most efficient zip algorithm for muon data)

NeXus

The NeXus hierarchical data format has three components:

A set of subroutines

to make it easy to read and write NeXus files

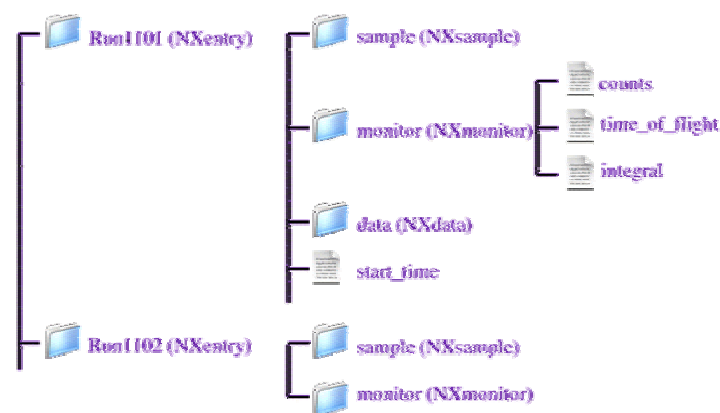
A set of design principles

to help people understand what is in them

A set of instrument definitions

to allow the development of more portable analysis software

Example part of data structure:



NeXus webpage:

www.nexus.anl.gov

ISIS muons NeXus webpage:

www.isis.rl.ac.uk/muons/data_analysis/nexus/intro.htm

μSR Data Formats in Use Worldwide

- □ PSI (Switzerland)
 - .dat (VMS binary)
 - .bin (standard binary)
- □ TRIUMF (Canada)
 - .tri (VMS binary)
 - .mud (hierarchical)
- □ KEK/JPARC (Japan)
 - .kek (VMS binary)

Finding the Muon Data

ARGUS	From the ISIS network:	<code>\\mirch\macs\argus0005634.ral</code>
	From ISISA:	not available
DEVA	From the ISIS network:	<code>\\ndavms\mutdata\r20000.ral</code>
	From ISISA:	<code>mut\$disk0:[data.mut]r20000.ral</code>
EMU	From the ISIS network:	<code>\\ndavms\emudata\r20000.ral</code>
	From ISISA:	<code>emu\$disk0:[data.emu]r20000.ral</code>
MUSR	From the ISIS network:	<code>\\ndavms\musrdata\musr00001025.nxs</code>
	From ISISA:	<code>musr\$disk0:[data.musr]musr00001025.nxs</code>

Special instrument (only for the muon training course practical !)

MUONTC	From the ISIS network:	<code>\\ndavms\muontc_data\r20000.ral</code>
	From ISISA:	<code>muontc\$disk0:[data.muontc]r20000.ral</code>

Temperature Logs

- ☐ Usually stored in a tlog subdirectory within the data directory.

Same name as the data file but with extension .tlog

- ☐ NeXus files store the tlog data internally

Taking Data Away

- Copy to own laptop
- Burn CD using PC in instrument cabin
- □ Send to remote system back home ñ pushing is easier than pulling!
- Fetch using remote login from home

Retrieving Older Data

Restoring files from the ISIS Data Archive

Archiver Restore Times

Restores for last cycle within 10 minutes. Restores for relatively recent data within 1 hour.

Older runs are restored from tape starting at 10:00 am for requests placed during the previous 24hrs (excluding weekends).

Restoring RAW Files

In order to restore ISIS RAW files, simply type **RESTISIS** from the DCL command prompt.

usage:

RESTISIS [options] inst_name low_run high_run

options:

- l is used to retrieve (T)LOG files
- s is used to retrieve SAV files.
- v will produce a verbose (diagnostic) output.
- d can be used to specify a restore directory.

e.g.:

RESTISIS -l -d SYS\$SCRATCH MUSR 12345 12350

Muon Data Analysis Software Used at ISIS

UDA (VMS, General Purpose)

RUMDA (VMS, General Purpose)

MESA (VMS, Maximum Entropy for TF Studies)

WiMDA (Windows, General Purpose)

many user groups have also developed their own programs

Muon Data Analysis Software Used at ISIS

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RUMDA (VMS, General Purpose)

MESA (VMS, Maximum Entropy for TF Studies)

WiMDA (Windows, General Purpose) - practical session

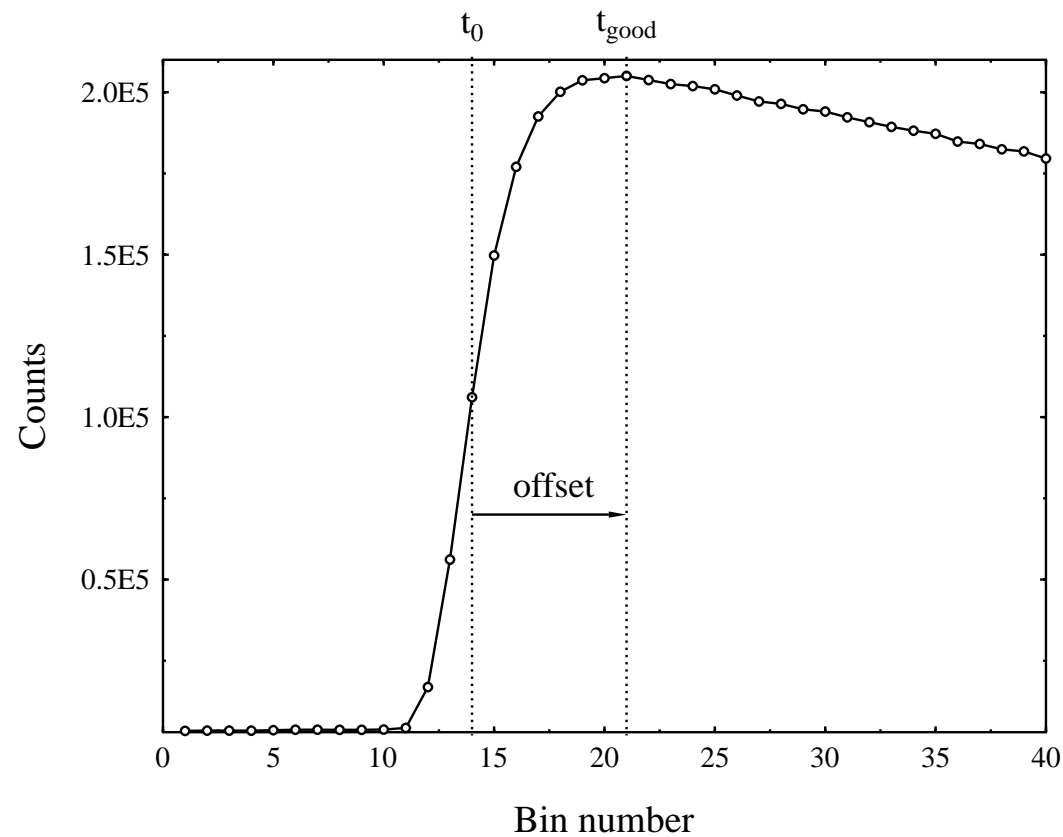
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Main Stages in Muon Data Analysis

1. Preparing the data to be analysed, 'setting up'
2. Fitting the measured asymmetry to a chosen relaxation function; 'analysing'
3. Assessing the fitted relaxation parameters, which may involve a further stage of fitting these parameters to an appropriate model; 'modeling'
4. Preparing plots of the results of analysing and modelling the data, 'plotting'

1. Setting Up the Data

- a) Checking the time origin t_0 and the time of the first and last good data points



1. Setting Up the Data

b) Defining the detector grouping

e.g.

for LF/ZF :	Forward group	1-16
	Backward group	17-32

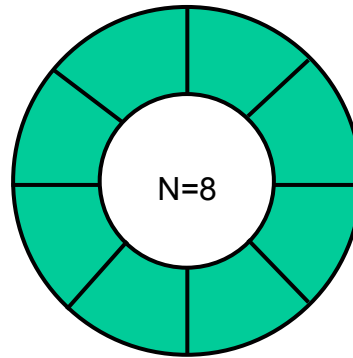
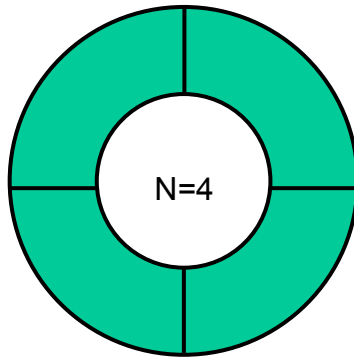
for TF :	Group1	1-4,	17-20
	Group2	5-8,	21-24
	Group3	9-12,	25-28
	Group4	13-16,	29-32

Notes:

ARGUS has 192 detectors (usually pregroupped to 32 histograms in the data file)

MUSR has 64 detectors (conveniently pregroupped to 32 histograms in the data analysis software)

Note that a dephasing effect will reduce the asymmetry of TF data if not enough groups are used:



$$\text{Dephasing factor} = \sin(\pi/N) / (\pi/N)$$

TF Groups

Dephasing Factor

16

99 %

8

98 %

4

90 %

2

64 %

i.e. 8 TF groups are sufficient for most purposes

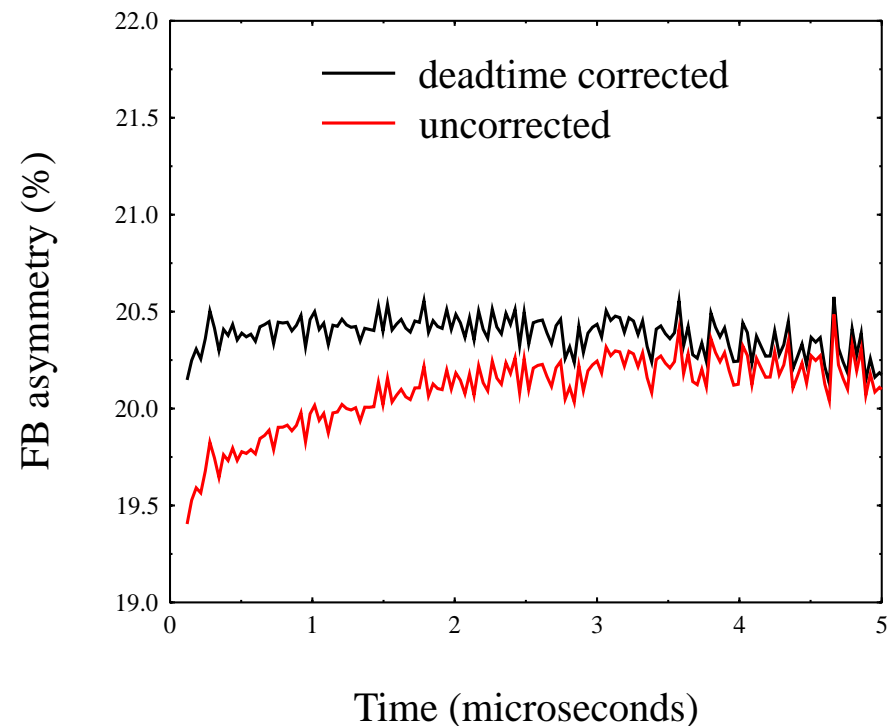
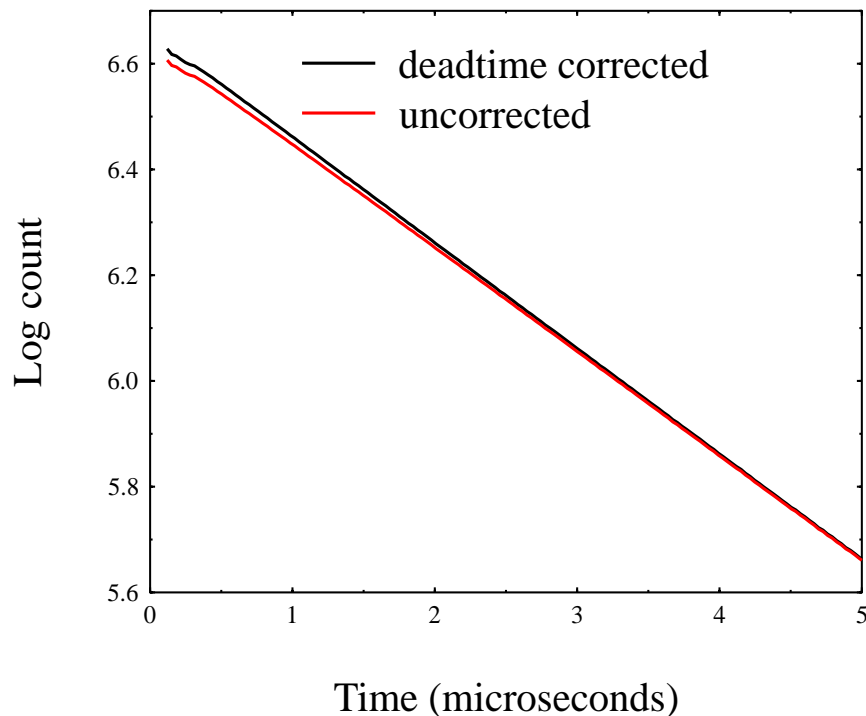
c) Correction for counting loss due to counter deadtime

characterised by a deadtime τ for each detector channel, typically $\tau \sim 10$ ns

deadtimes for particular instruments are obtained from high statistics calibration runs using Ag

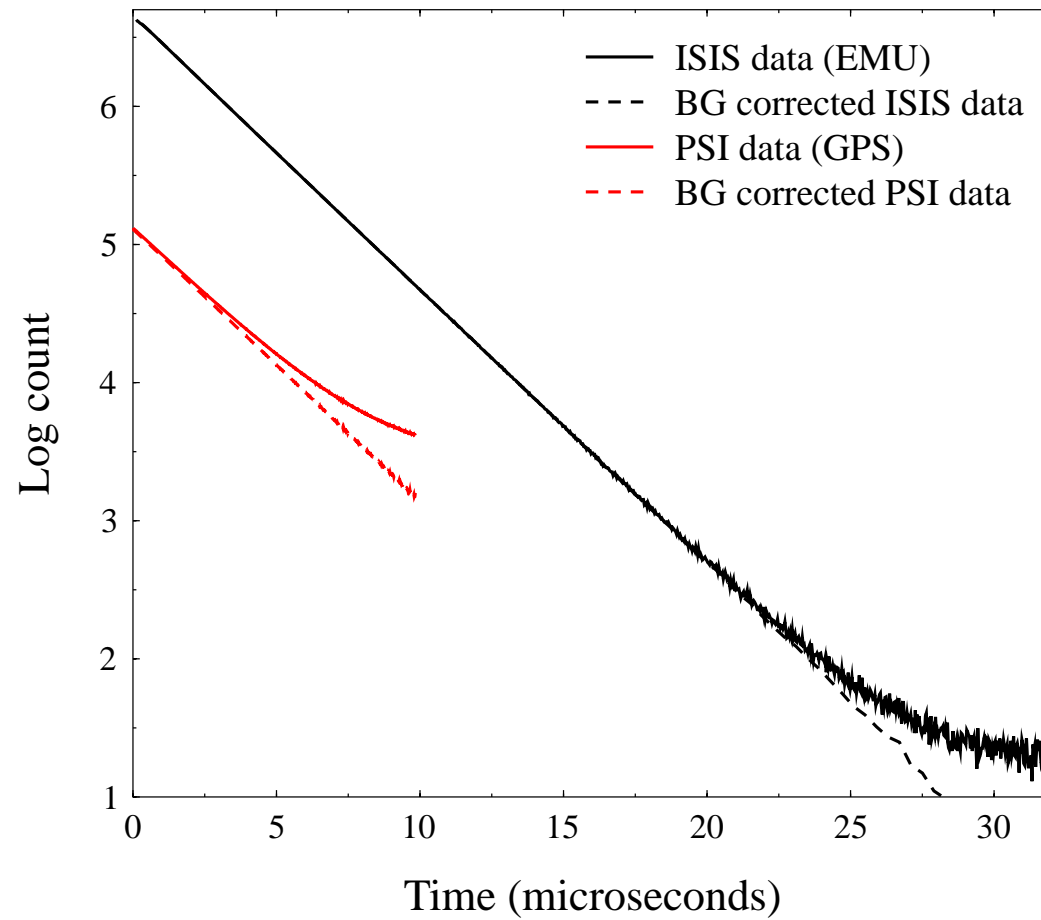
data rate correction to the observed rate r_{ob} is applied to give the true rate r

the simplest form of correction is $r = r_{\text{ob}} / (1 - r_{\text{ob}}\tau)$



d) Correction for steady background count rate

can be included as part of the fitting procedure



e) Choice of binning:

The standard raw time bin for ISIS data is 16 ns. It is often useful to choose to increase the bin size for data analysis

bin width : tradeoff between number of points
and fitting speed

allows separate focus on fast and slow
parts of the relaxation

fixed/variable: variable binning compensates for the
deteriorating signal-to-noise at longer times

best to keep to fixed binning for weakly
damped oscillations, e.g. TF studies

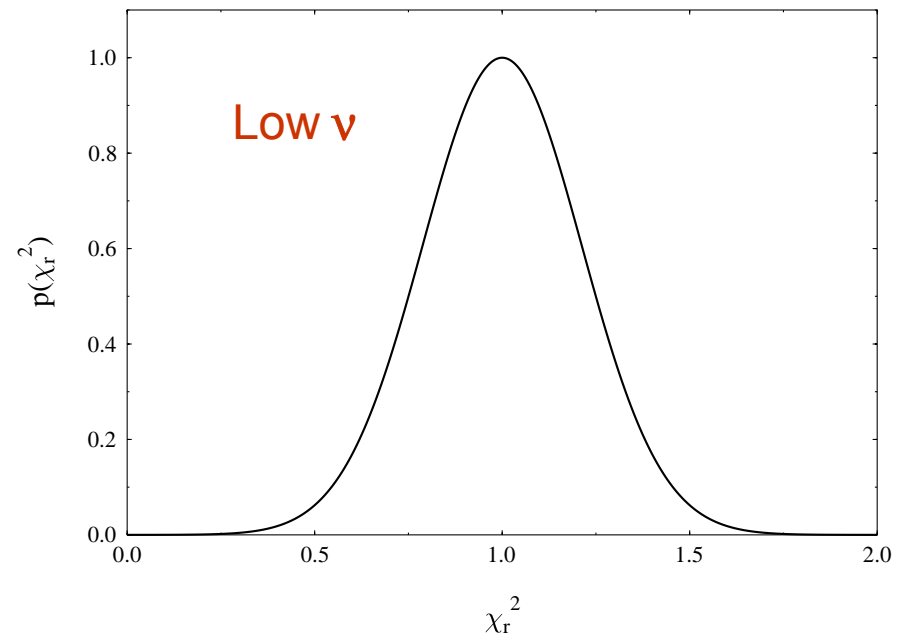
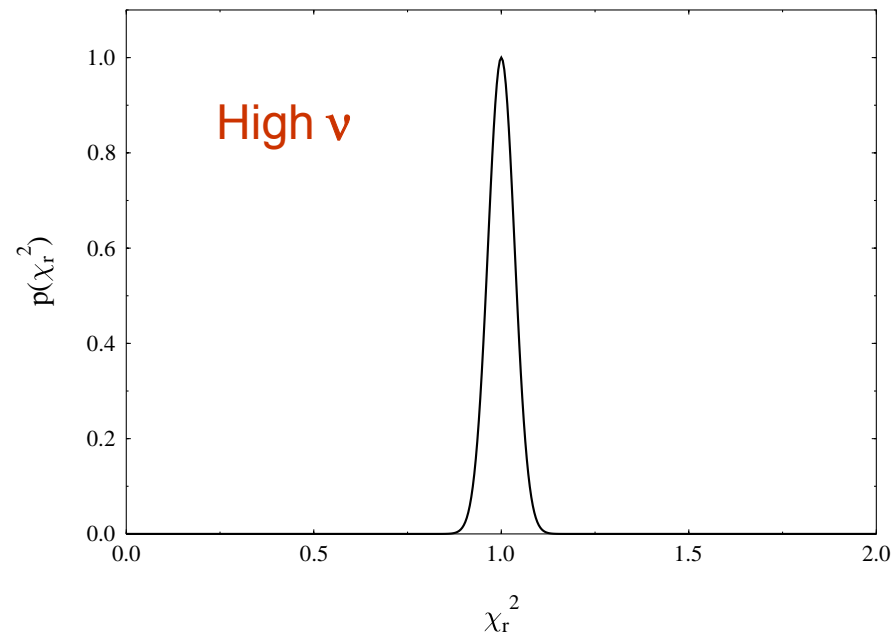
2. Analysing the Relaxation

- Try fitting to possible alternative relaxation functions
- ☐ Look for systematic deviations of the fit from the data \tilde{n} are additional relaxation components needed?
- Use the reduced chi-squared χ_r^2 to judge the quality of the fit and appropriateness of the model

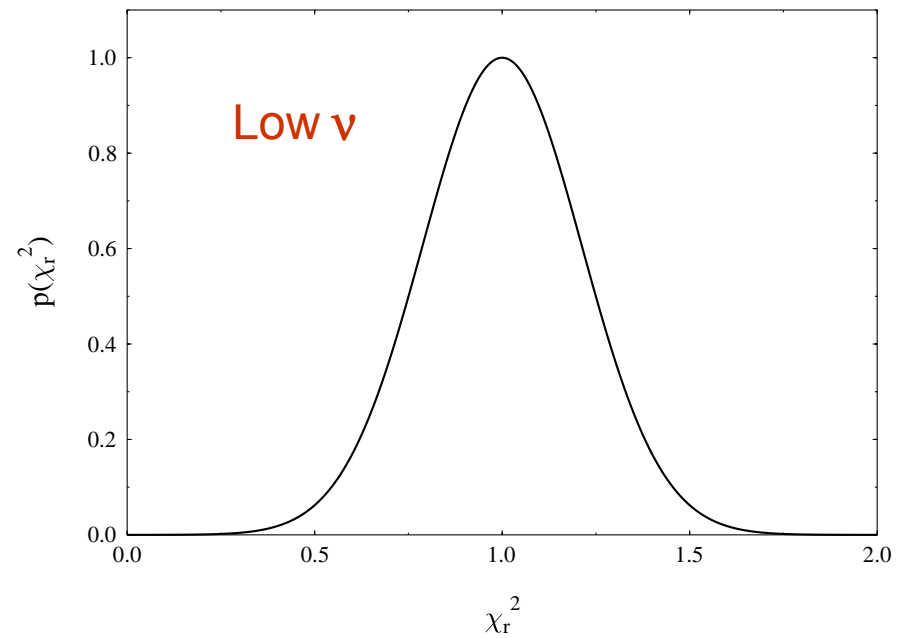
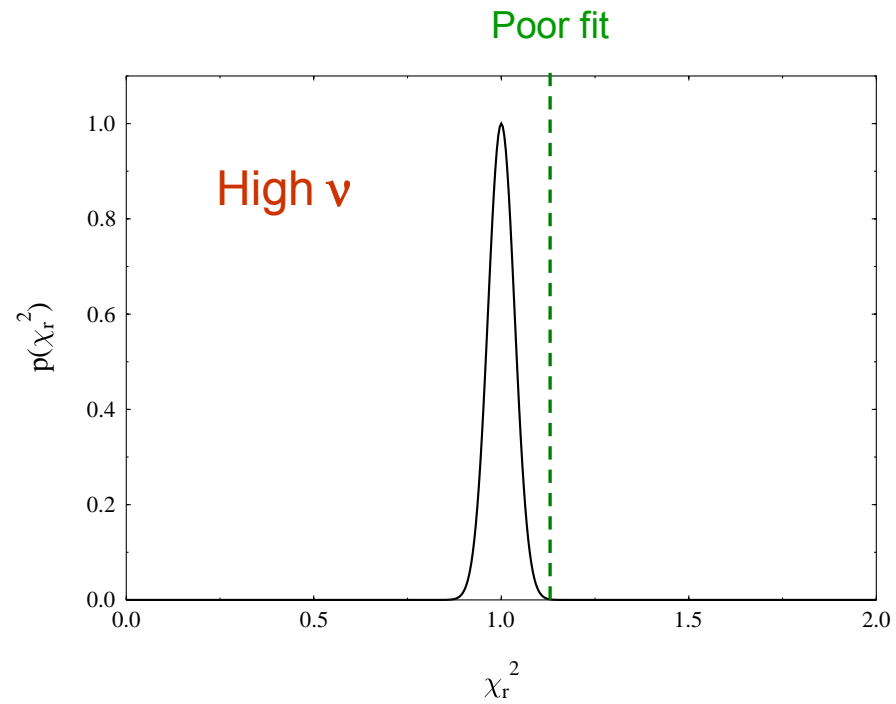
$$\chi_r^2 = \sum_{i=1}^N \left(\frac{y_i - y(x_i; p_1, p_2 \dots p_m)}{\sigma_i} \right)^2 / (N - m)$$

N-m = v is the number of degrees of freedom

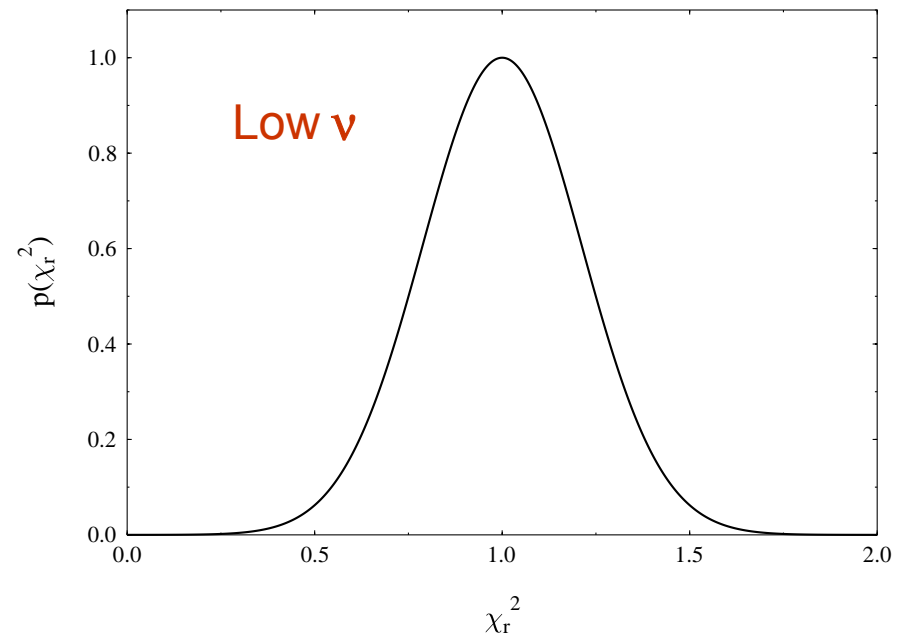
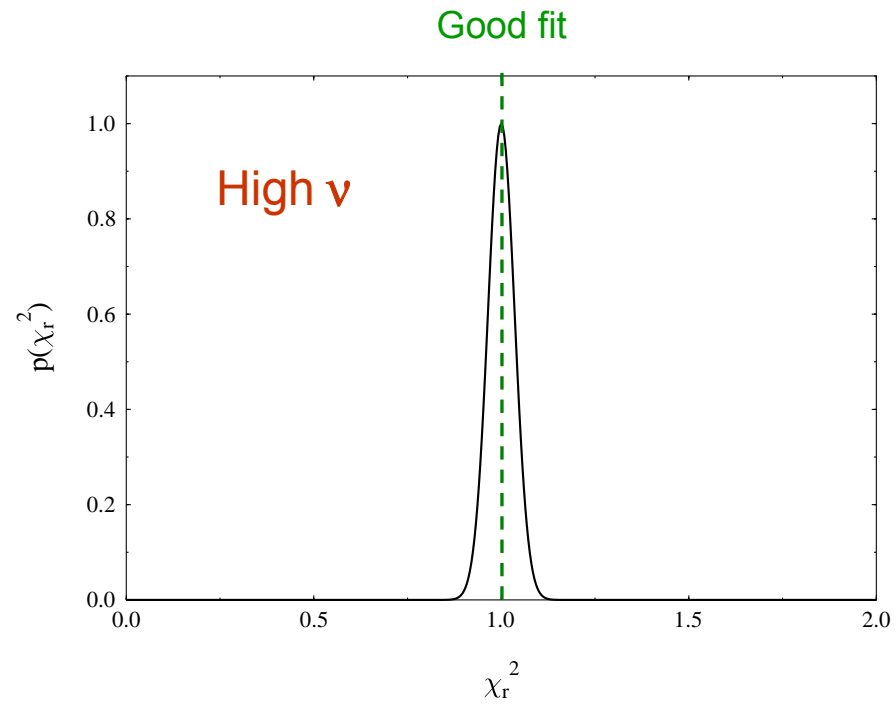
Expected standard error of χ_r^2 is $(2/\nu)^{1/2}$



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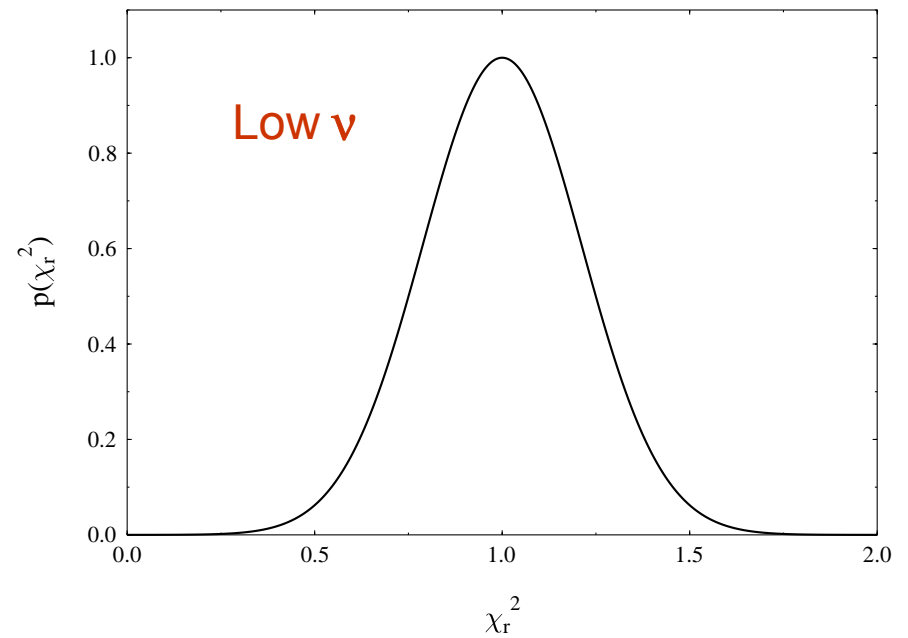
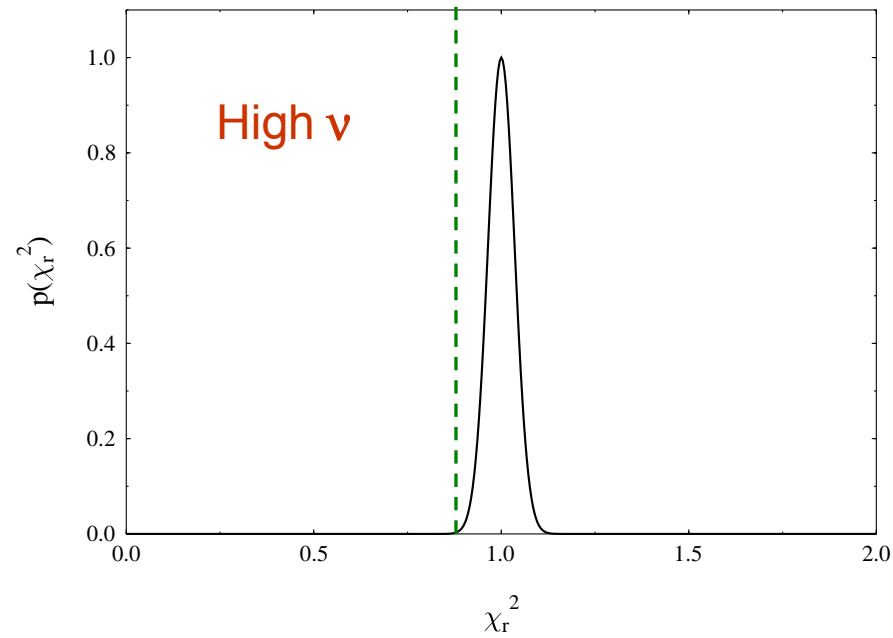


Expected standard error of χ_r^2 is $(2/\nu)^{1/2}$

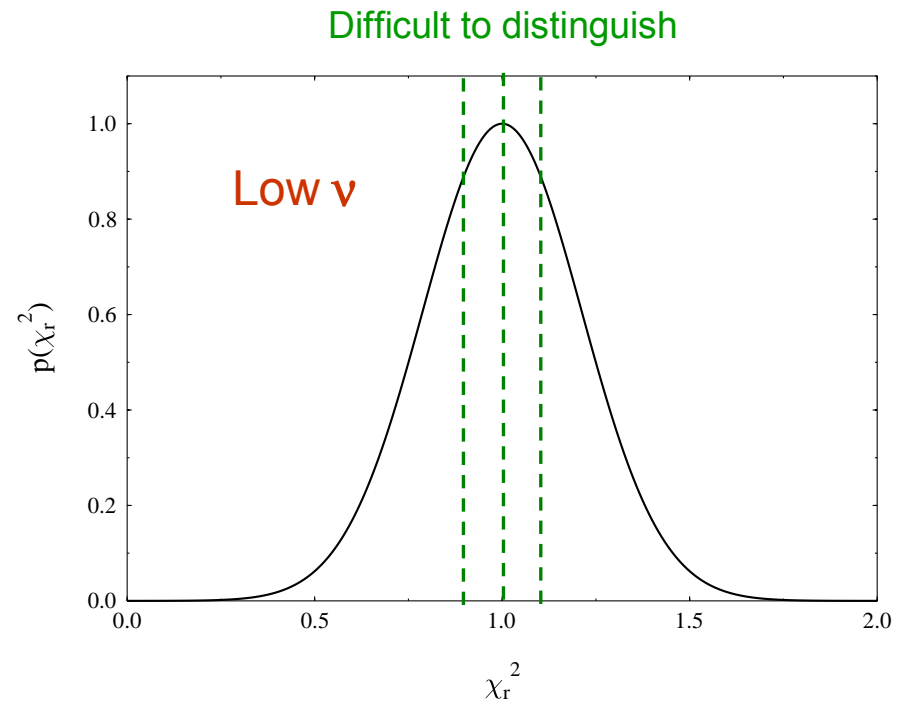
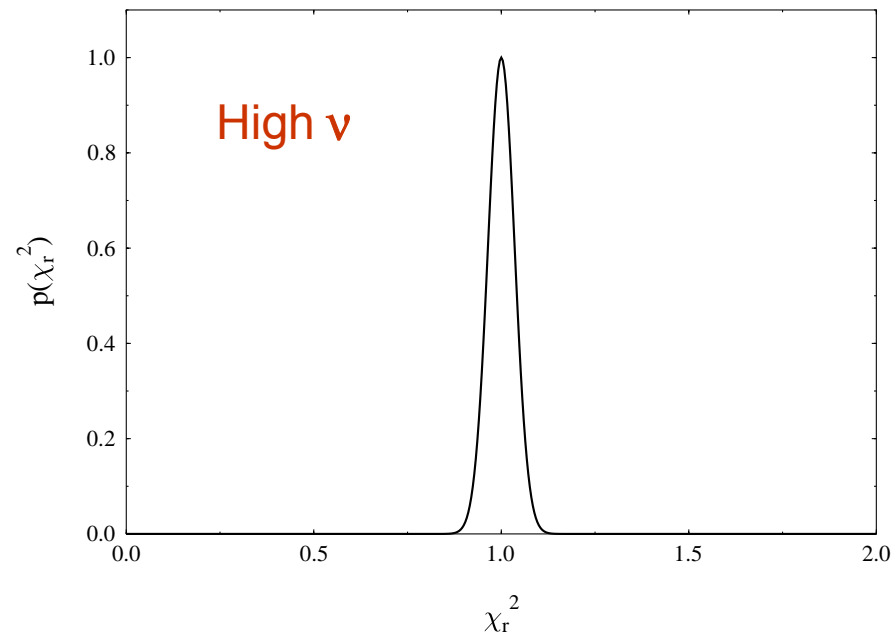


Expected standard error of χ_r^2 is $(2/\nu)^{1/2}$

Too good fit!



Expected standard error of χ_r^2 is $(2/v)^{1/2}$



3. Modelling Fitted Parameter Sets

- A further stage of fitting involves modelling the fitted parameters for a related set of runs:
 - e.g. following the temperature dependence of a precession frequency within a magnetic phase
 - or fitting the field dependence of a relaxation rate to an appropriate model

4. Plotting the Results

Close integration with the fitting process is desirable for rapid feedback on:

- the data quality

- the state of the analysis

- the progress of the experiment

WiMDA/GLE used in the data analysis workshop

Origin popular general-purpose Windows
plotting package

Analysis of Complex Rotation Spectra

1. Fourier and All Poles transforms
2. Maximum Entropy spectral analysis
3. Time domain analysis versus frequency domain analysis

Fourier and All-Poles Transforms

FFT (Fast Fourier Transform) is the standard way to convert from time domain to frequency domain.

FFT assumes frequency spectrum is well represented by array of evenly spaced points, which works well for spectra containing broad spectral features.

However, if the spectrum contains very narrow features, other types of frequency transform can work better.

The **All-Poles (maxent)** method is one such method, which makes an expansion of the data in terms of a series of sharp frequencies

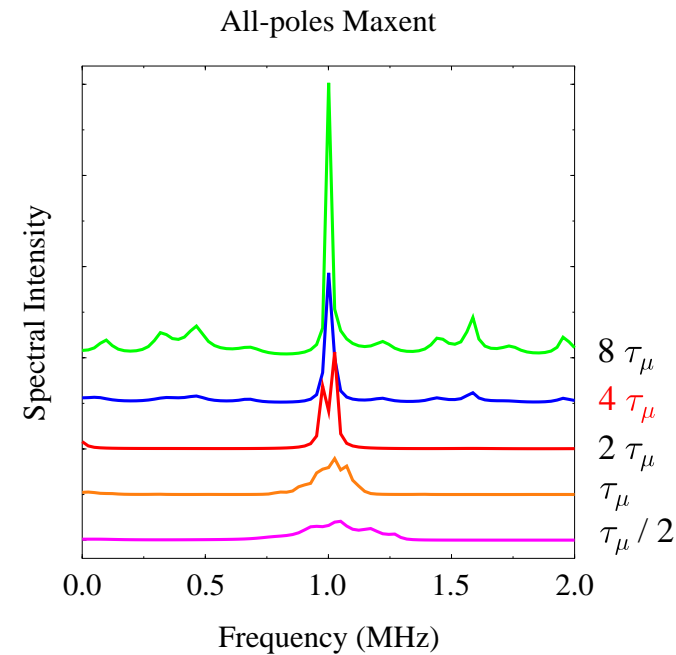
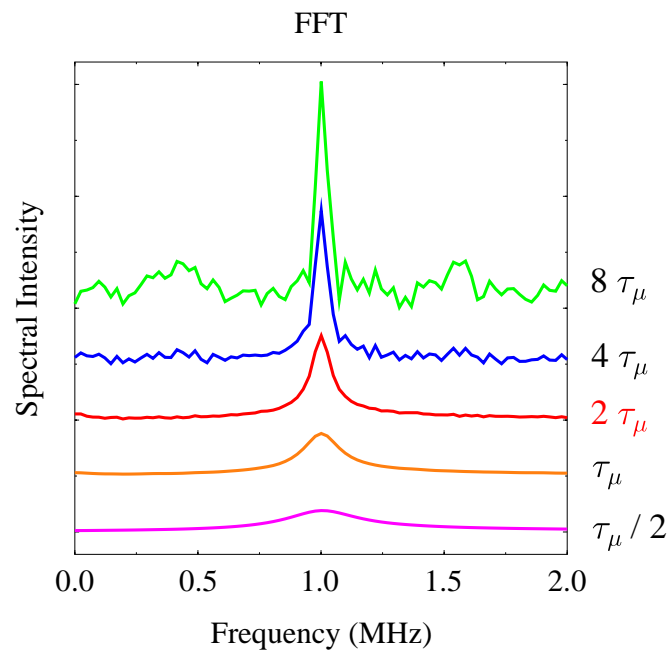
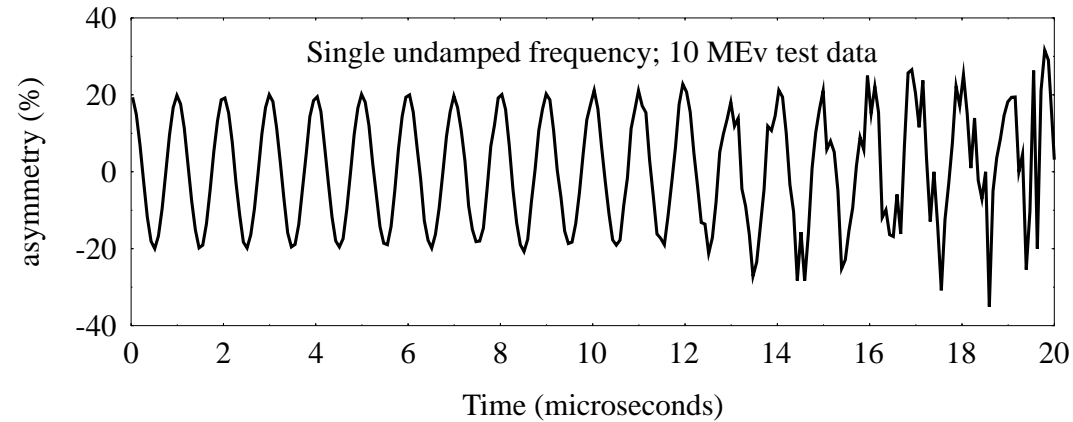
See Press et al, Numerical Recipes, CUP for further details of the All-Poles transform

All transform methods assume that the data error is independent of time, which is clearly not the case for μ SR data.

Data filtering (apodization) is essential before transforming.

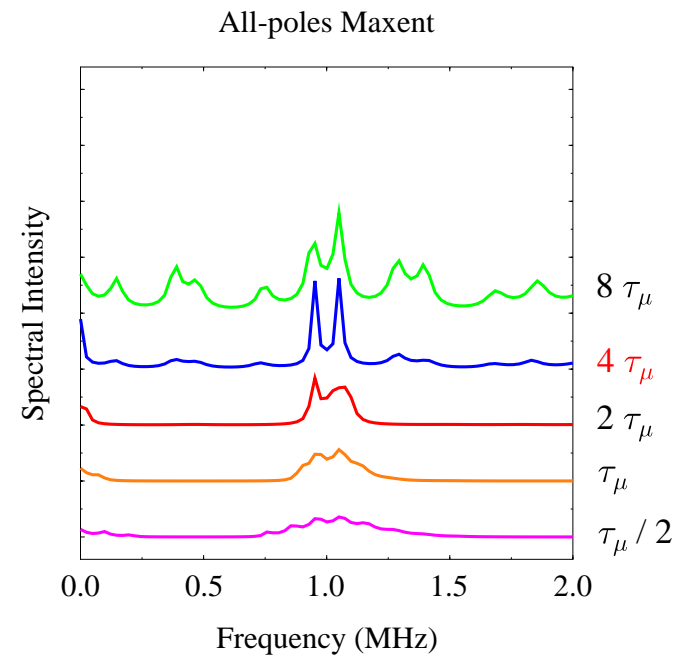
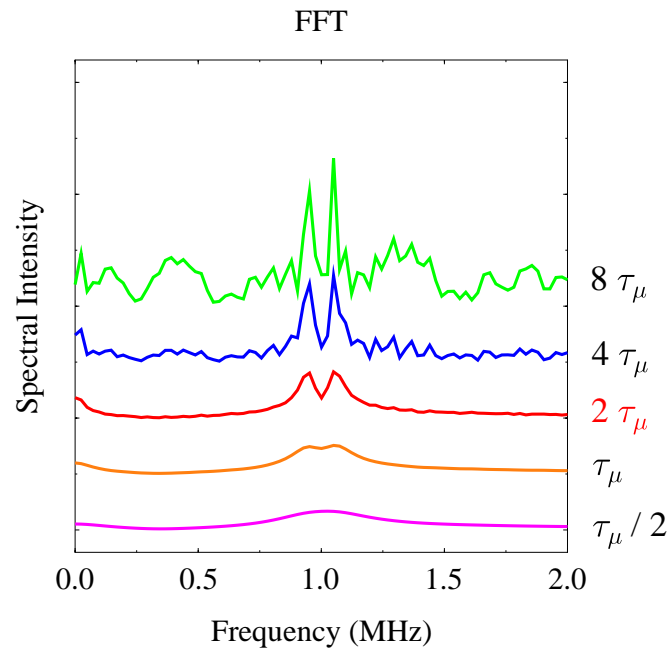
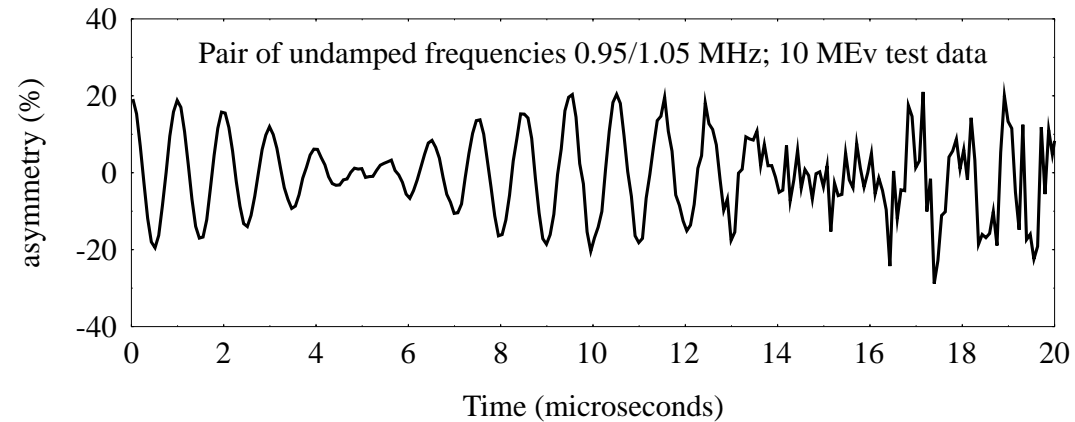
Fourier and All-Poles Transforms

Optimal filtering time constant for a single undamped test frequency



Fourier and All-Poles Transforms

A close pair of undamped test frequencies



The Maximum Entropy Method

Avoids the noise problem and need for filtering; takes data errors fully into account

Iterative procedure for constructing the frequency spectrum with the minimum structure (i.e. maximum entropy) that is consistent with the measured data

Entropy here is determined from the frequency spectrum p_k

$$S = -\sum_k \frac{p_k}{b} \log \frac{p_k}{b}$$

The procedure involves maximising $S - \lambda \chi^2$, where λ is a Lagrange multiplier

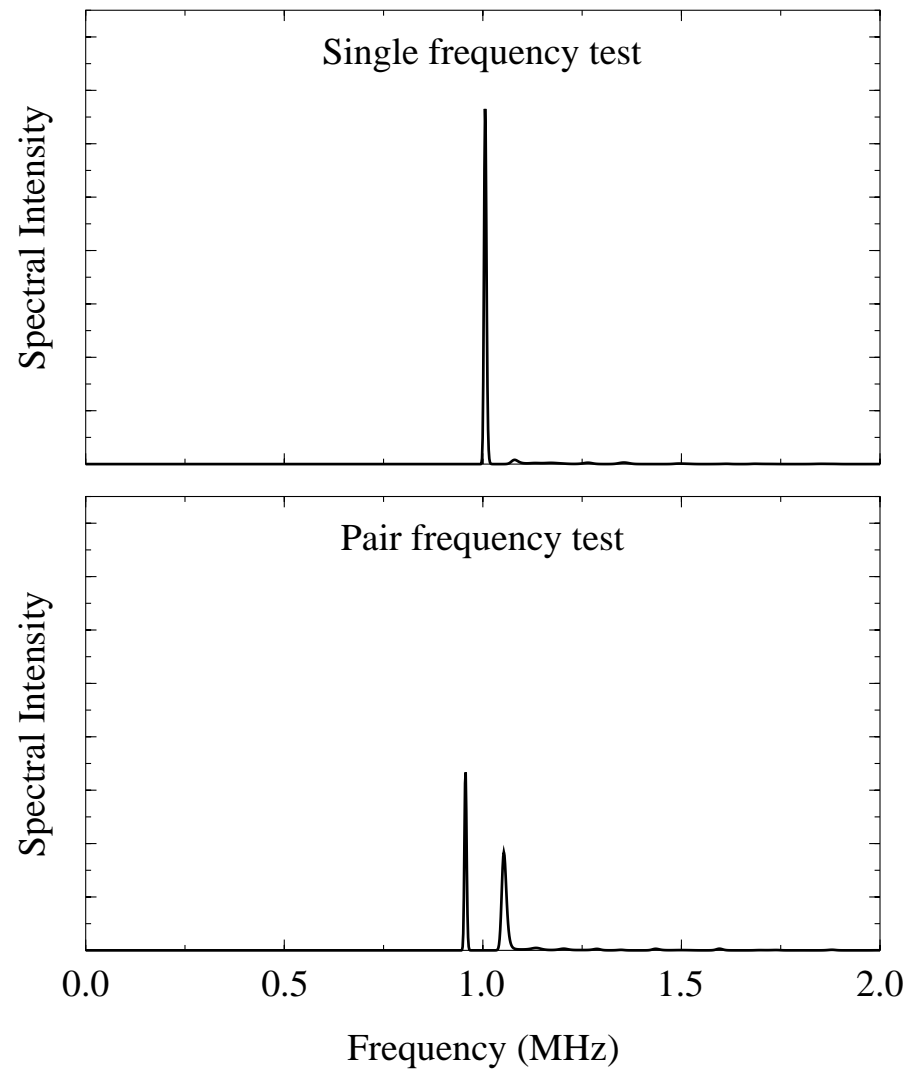
See Rainford and Daniell, Hyperfine Interactions 87, 1129 (1994)
for a detailed discussion of using Maximum Entropy in μ SR

for a general reference see:

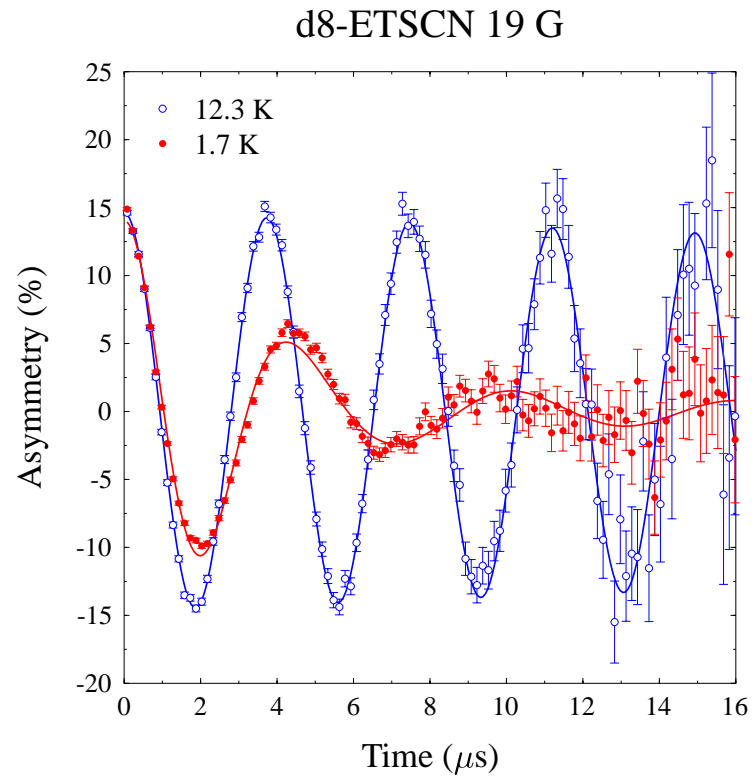
Maximum Entropy in Action, Buck and Macaulay, OUP (1991)

The Maximum Entropy Method

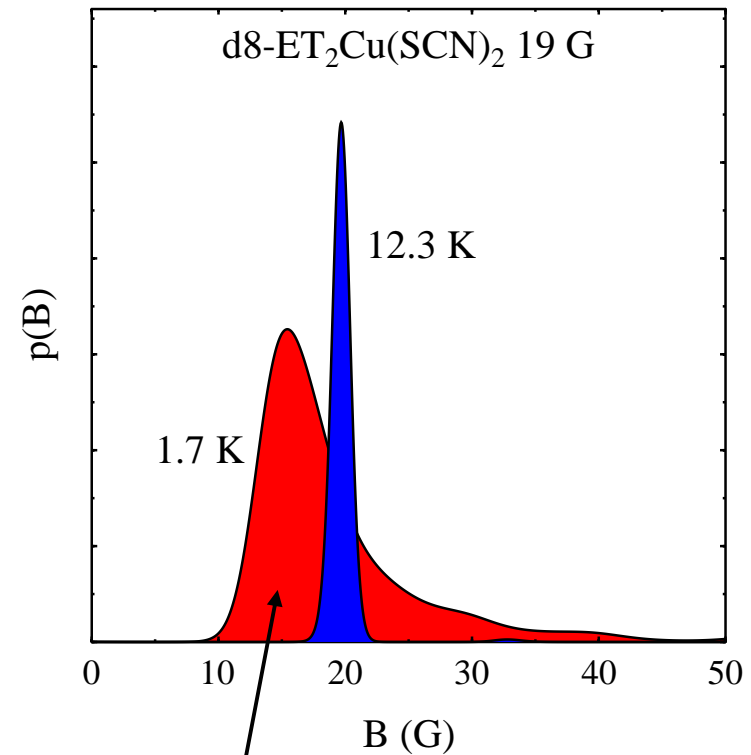
Demonstration using the test data for the transforms



Organic Superconductor Example

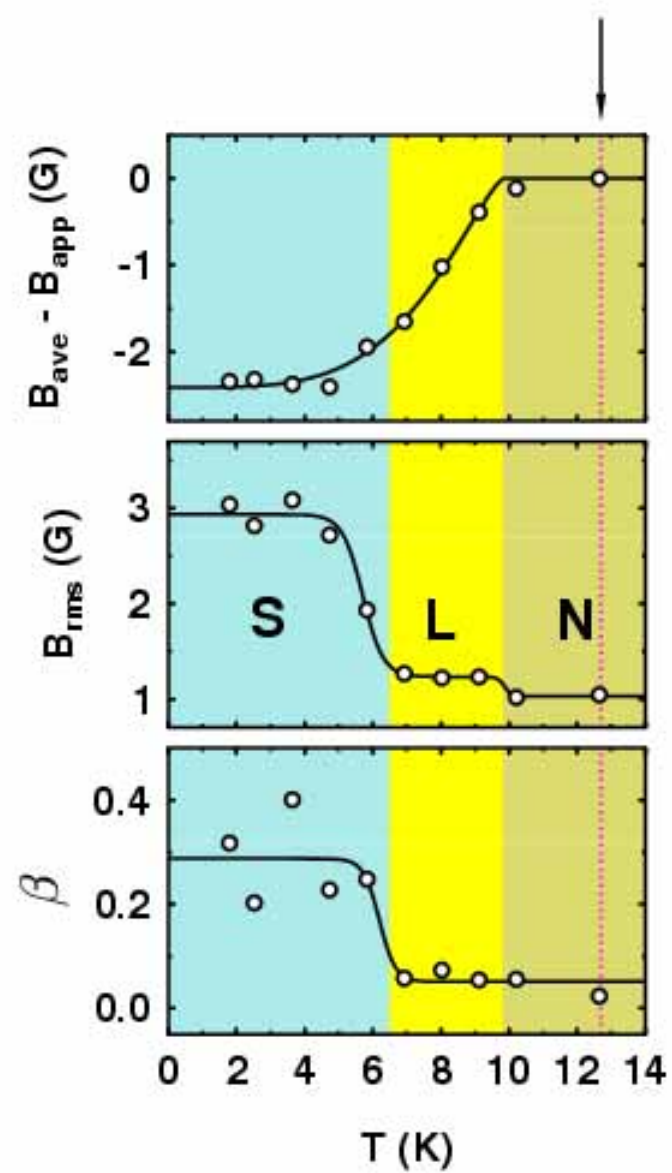
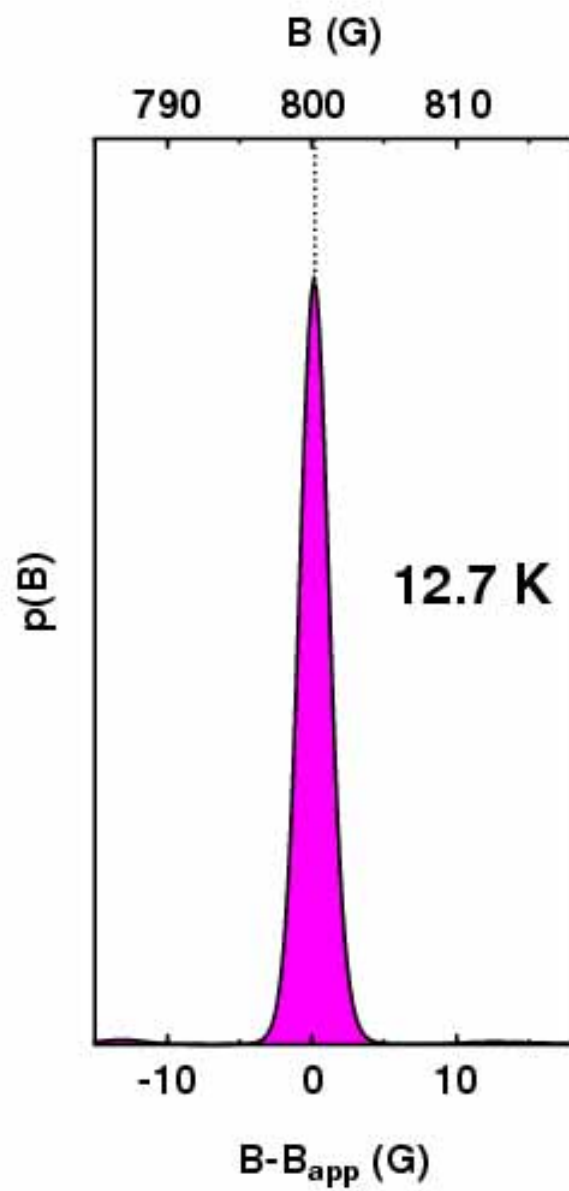


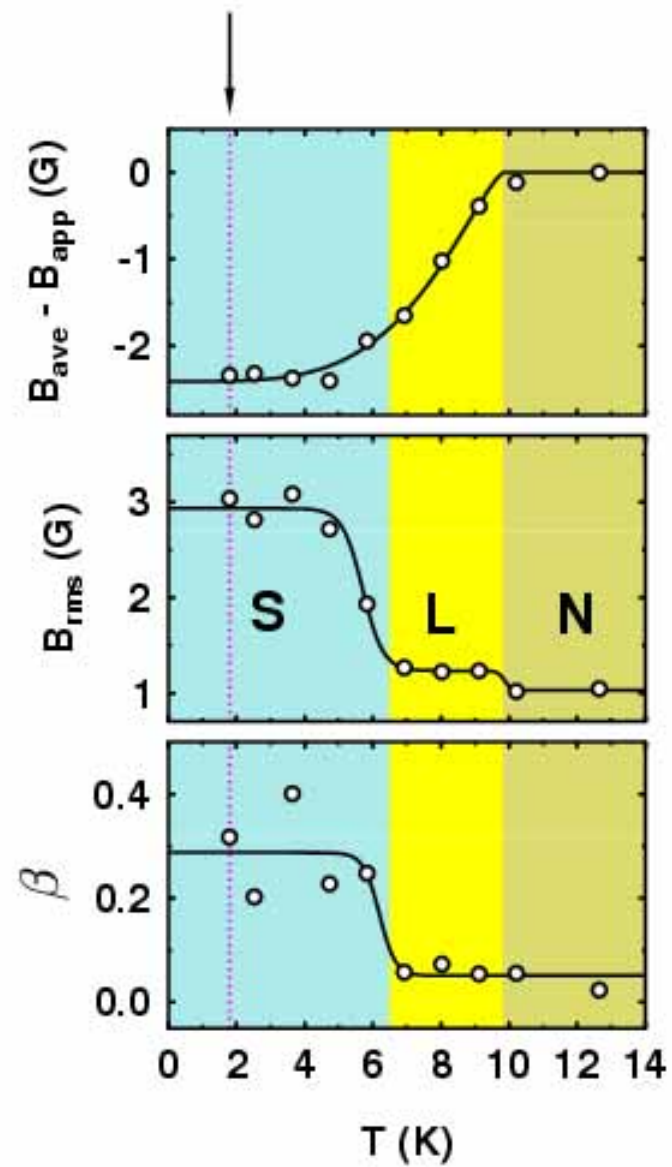
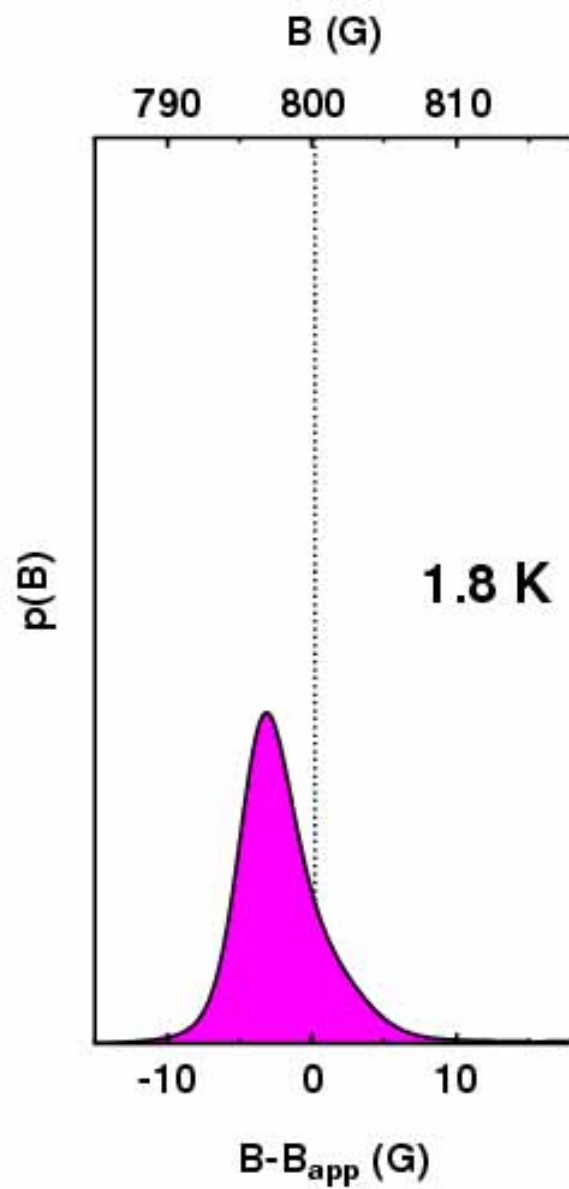
Maximum Entropy Spectra

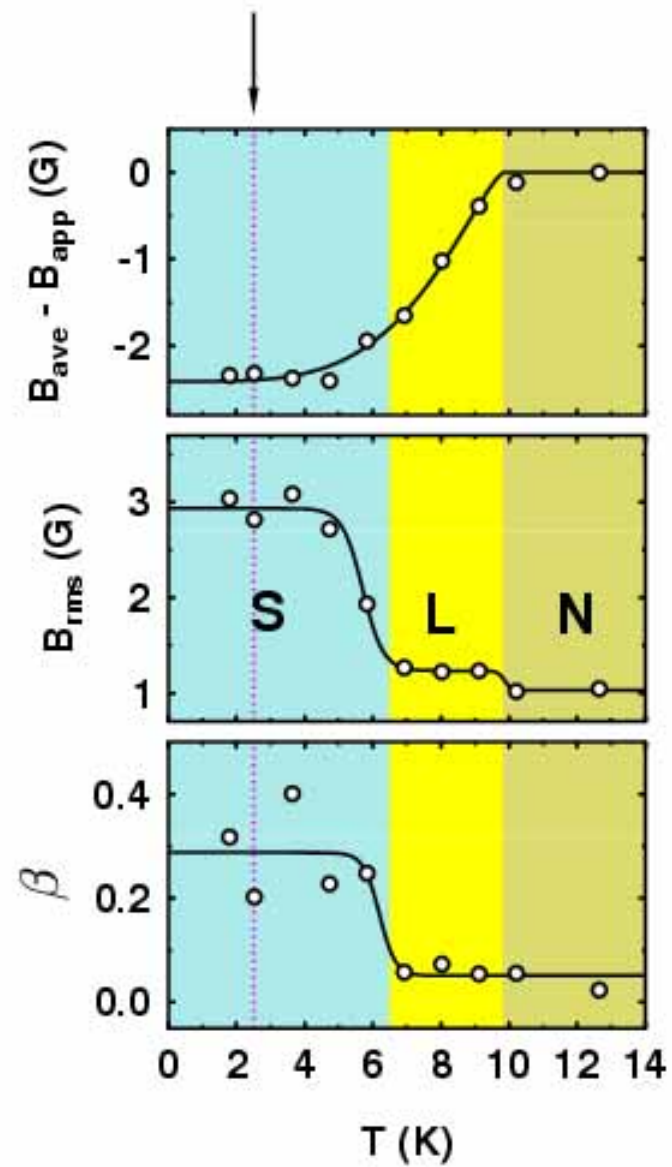
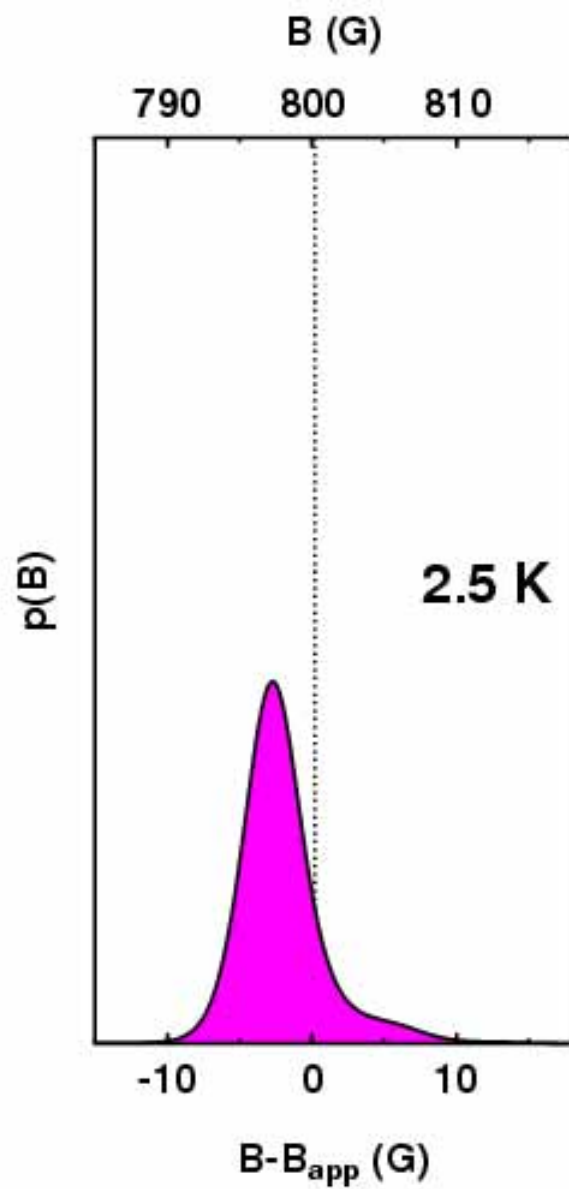


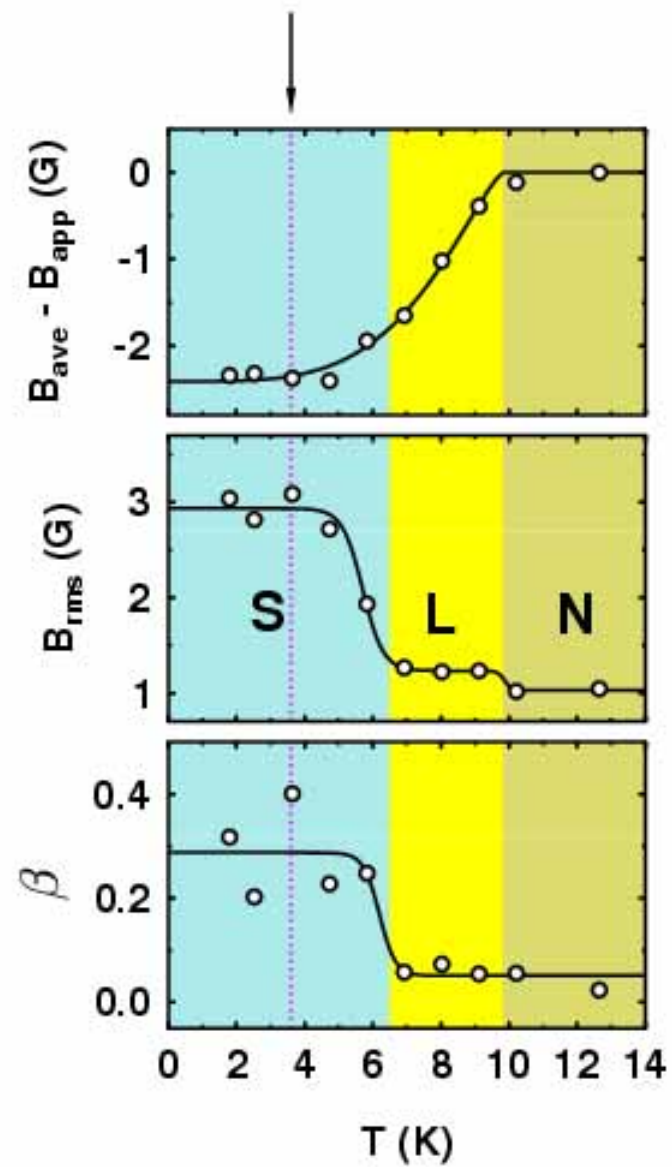
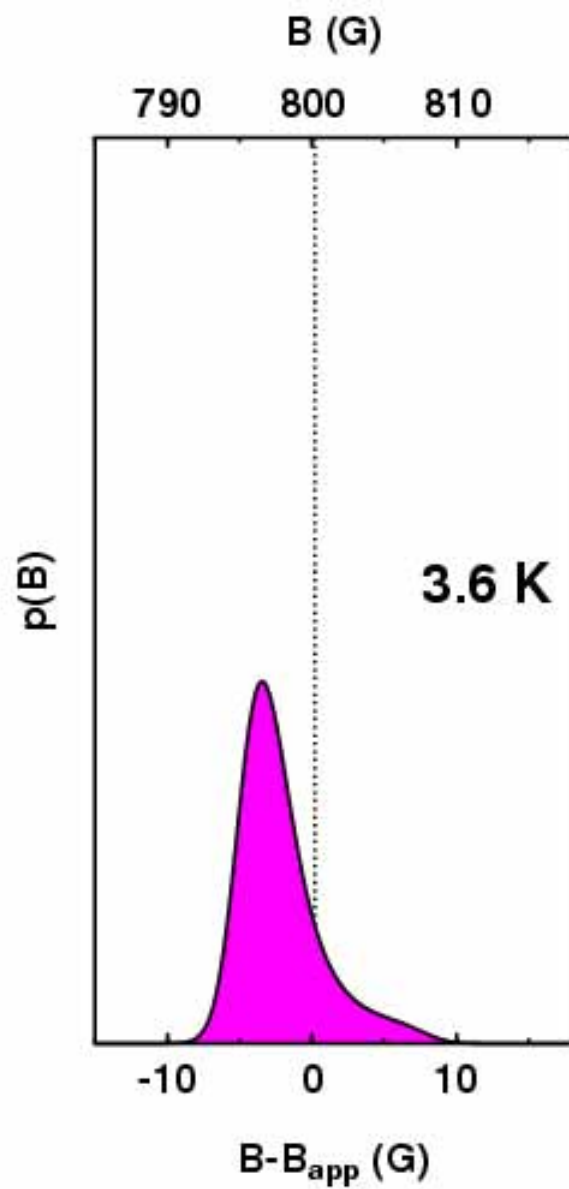
Characteristic field distribution
due to vortex lattice

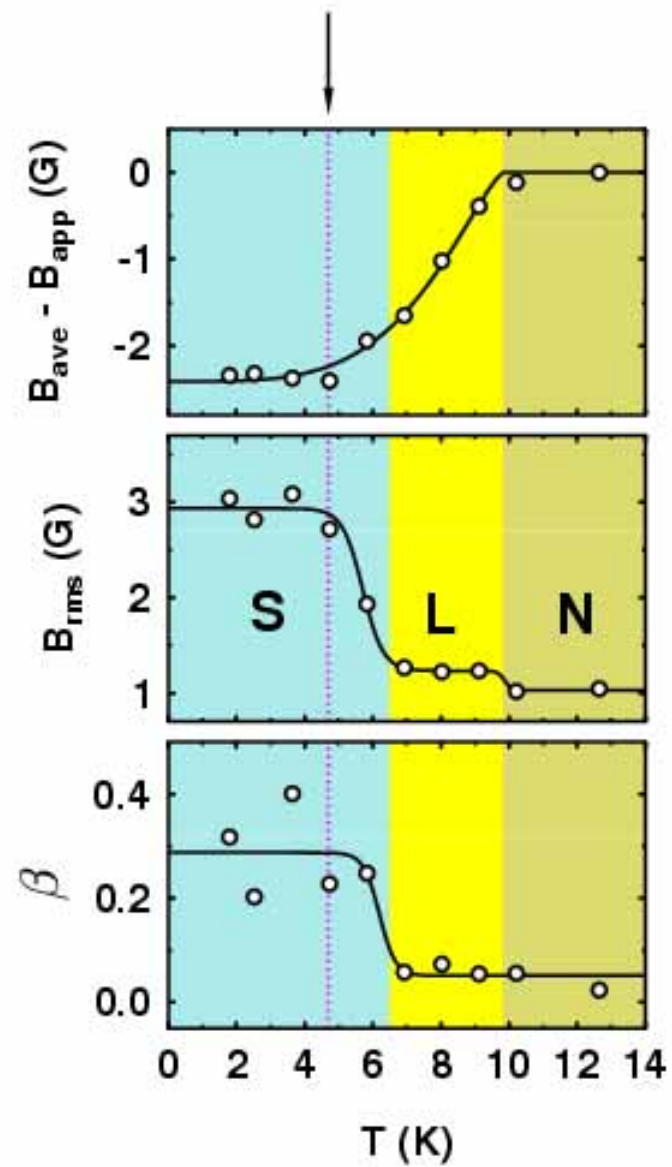
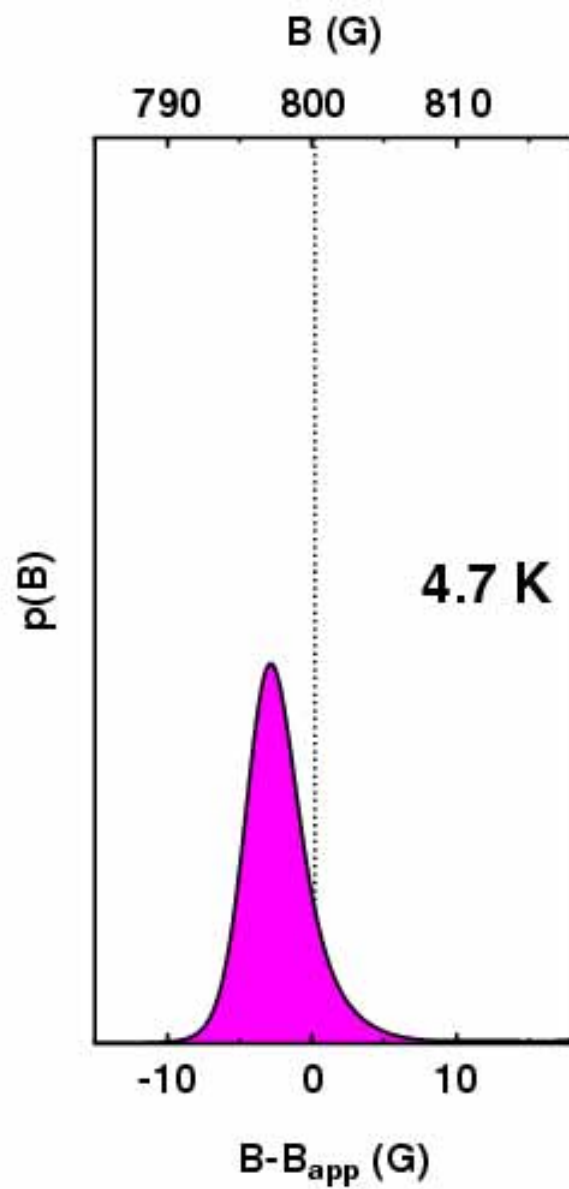
Melting of the Vortex Lattice

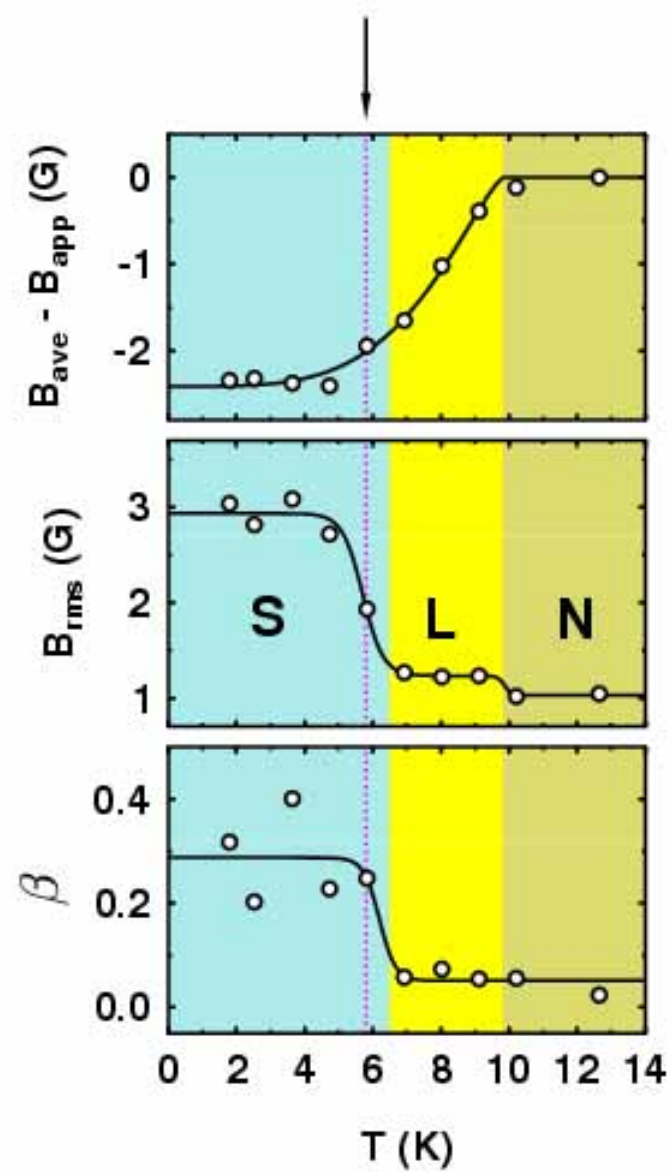
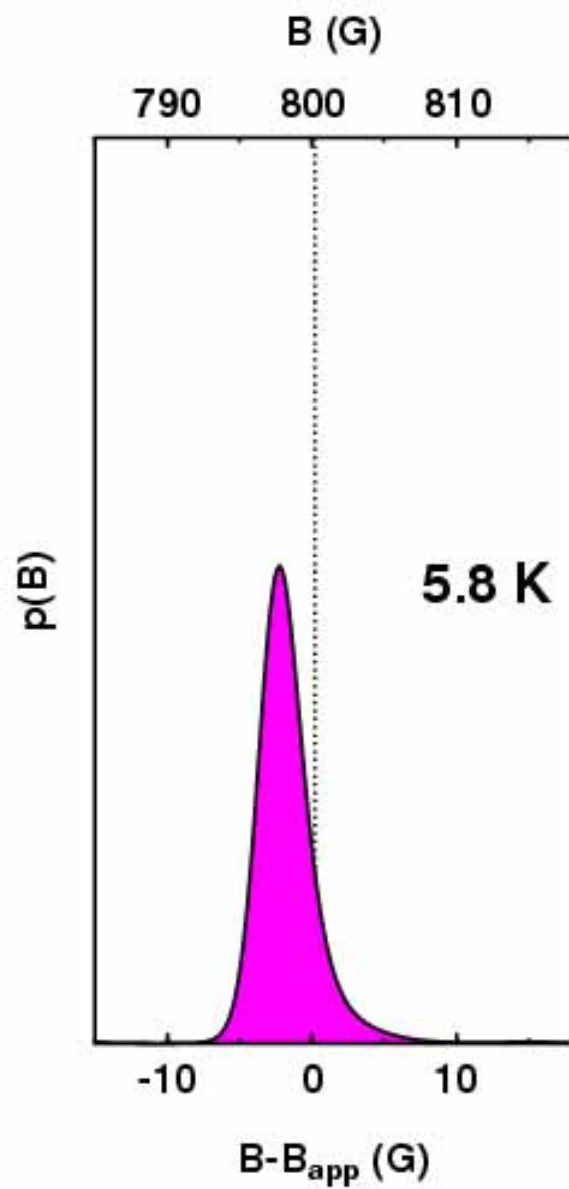


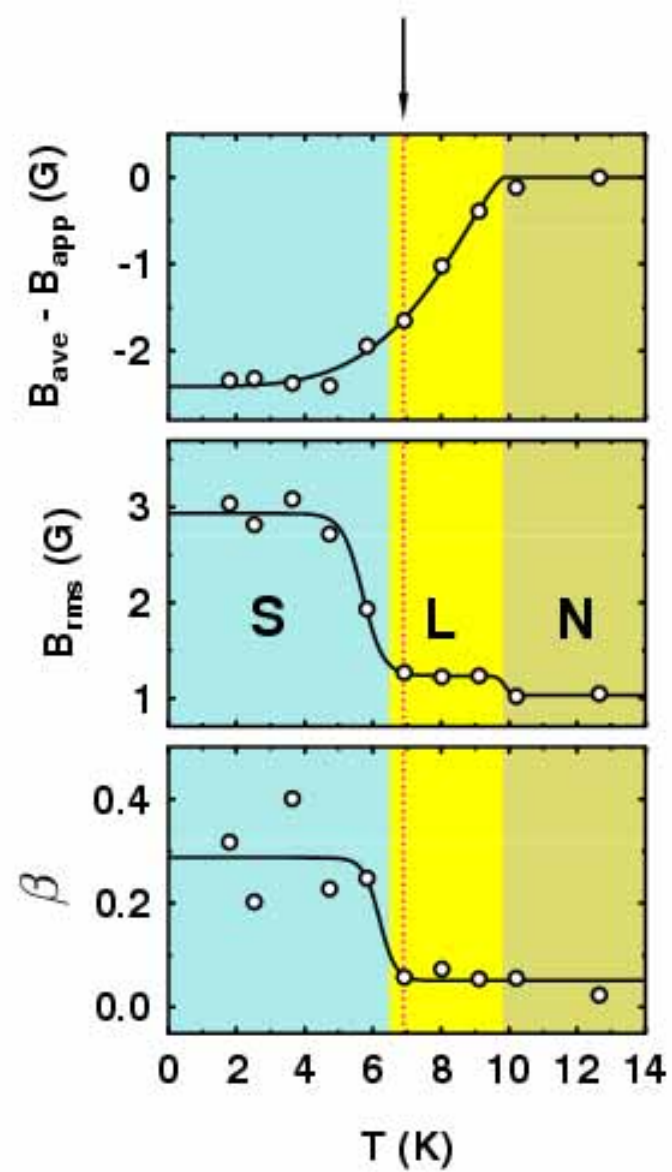
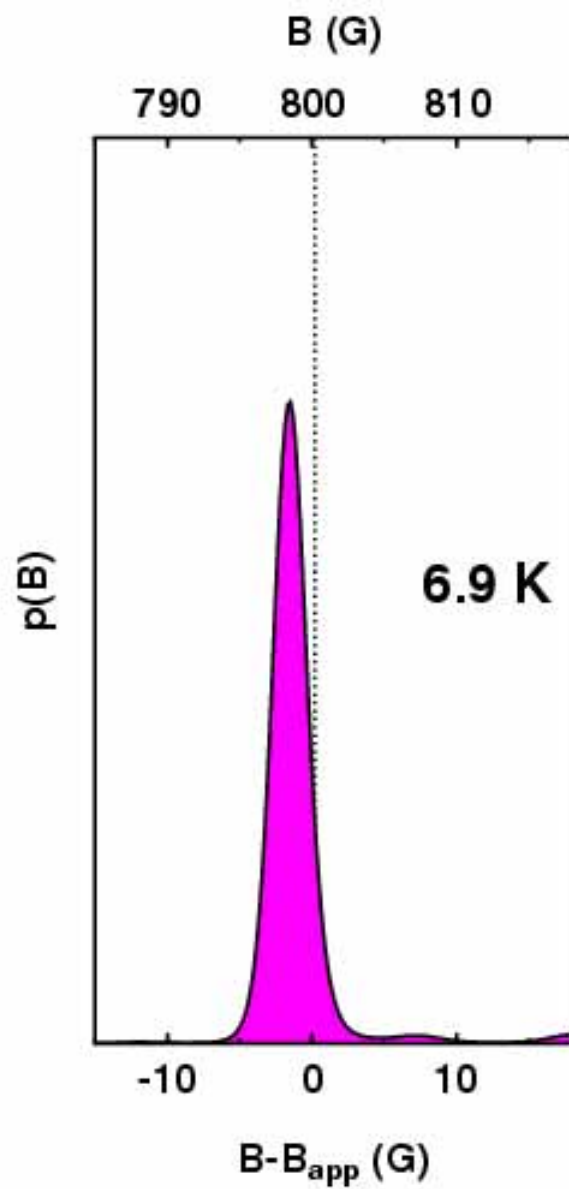


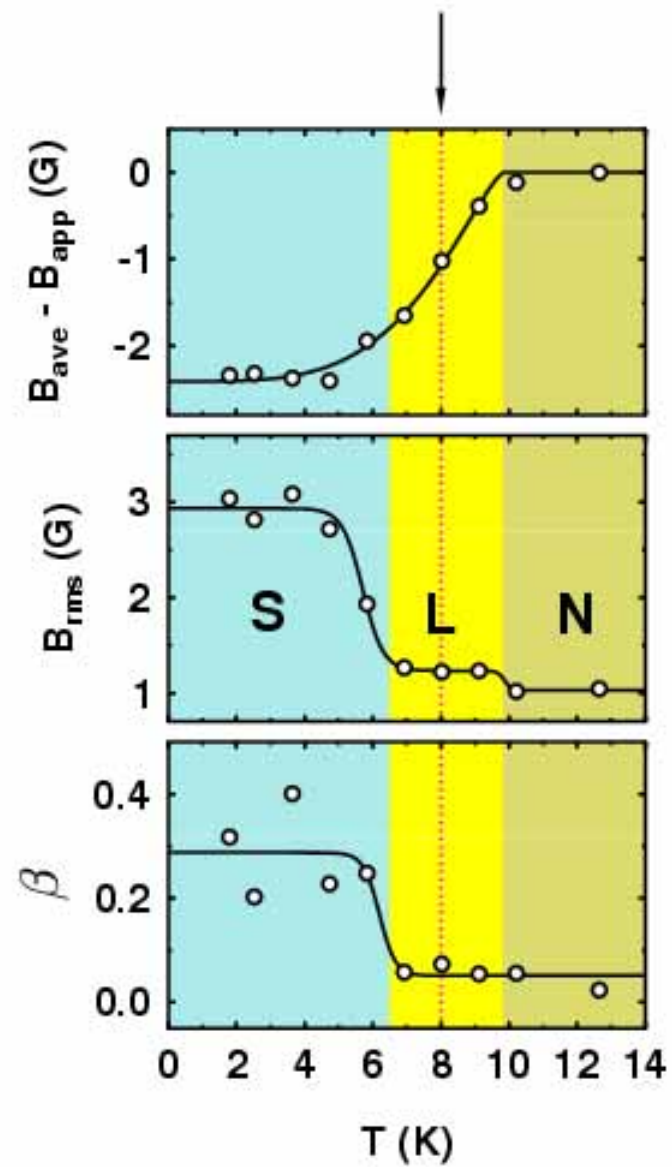
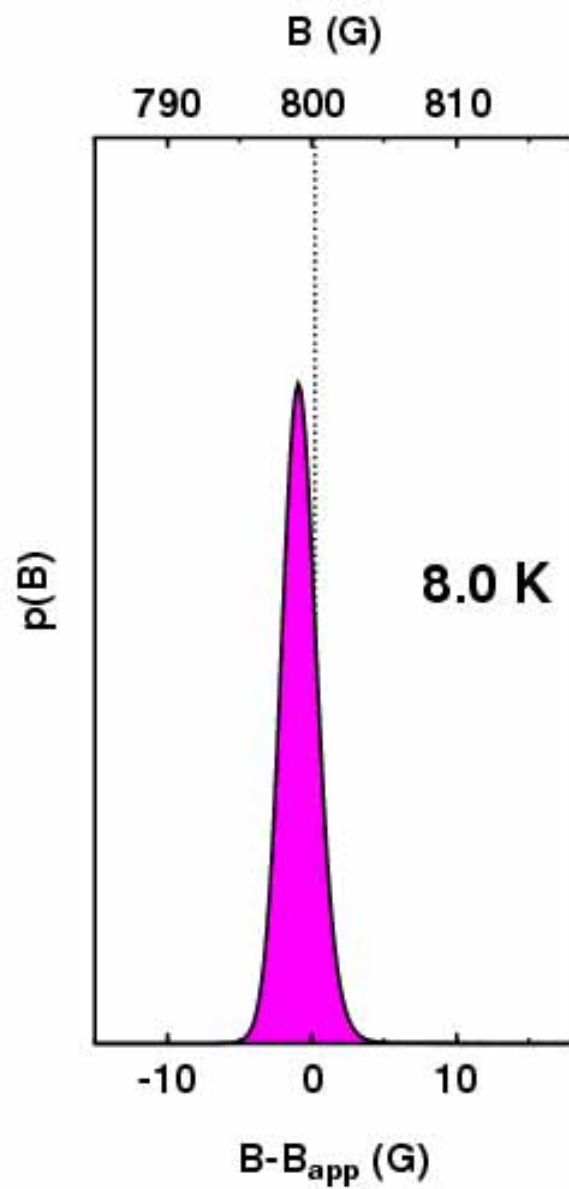


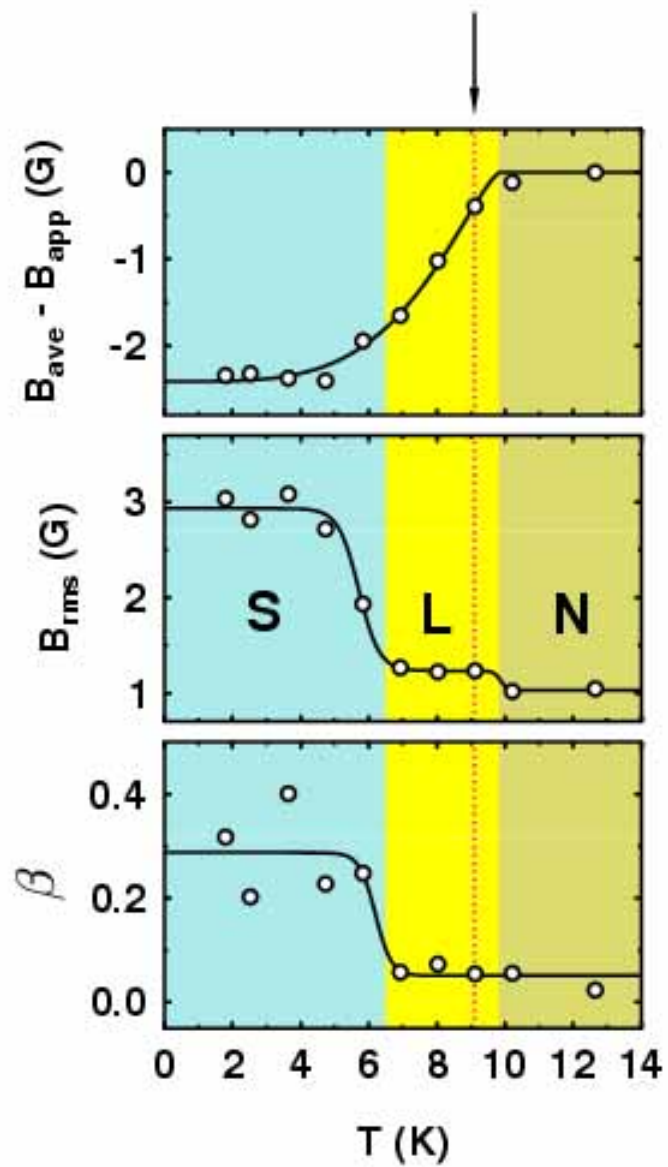
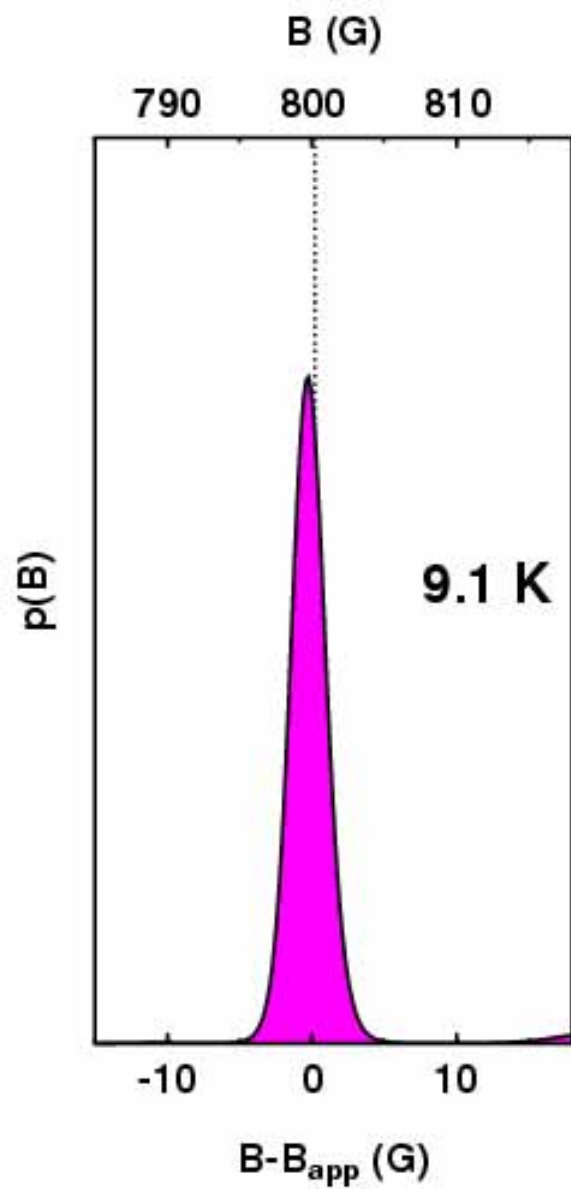


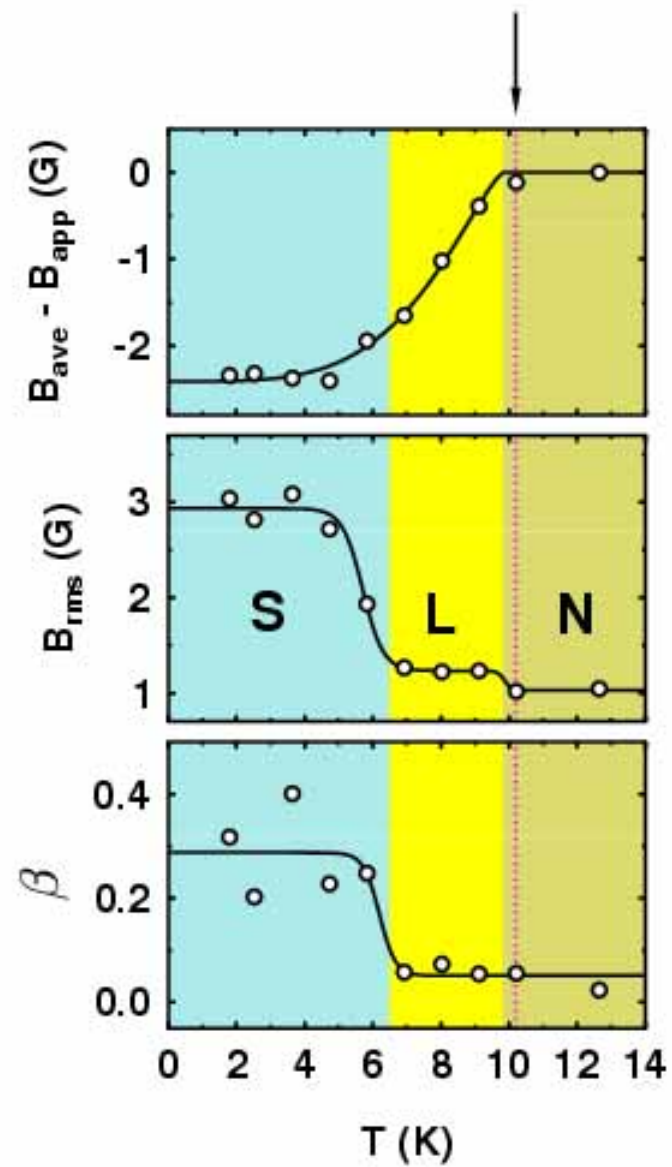
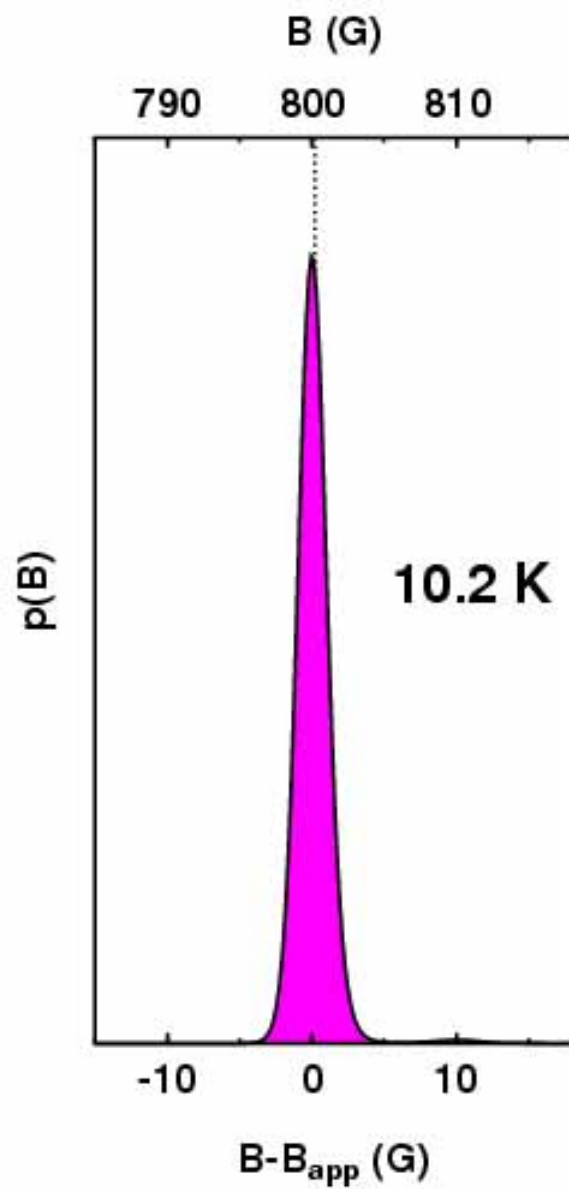


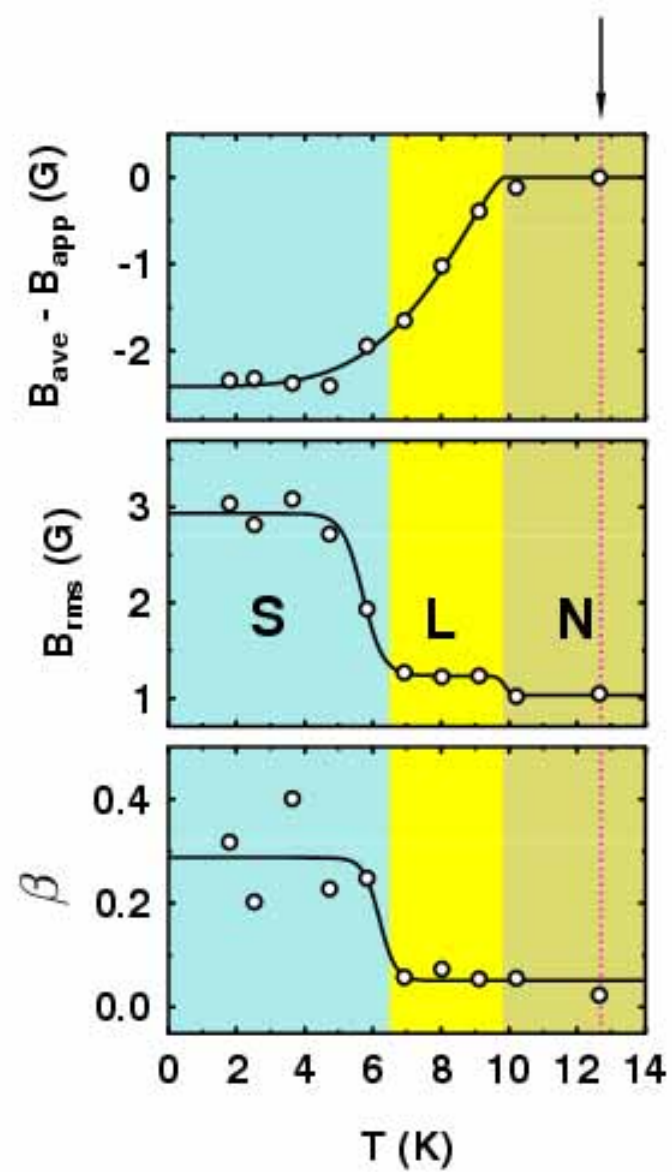
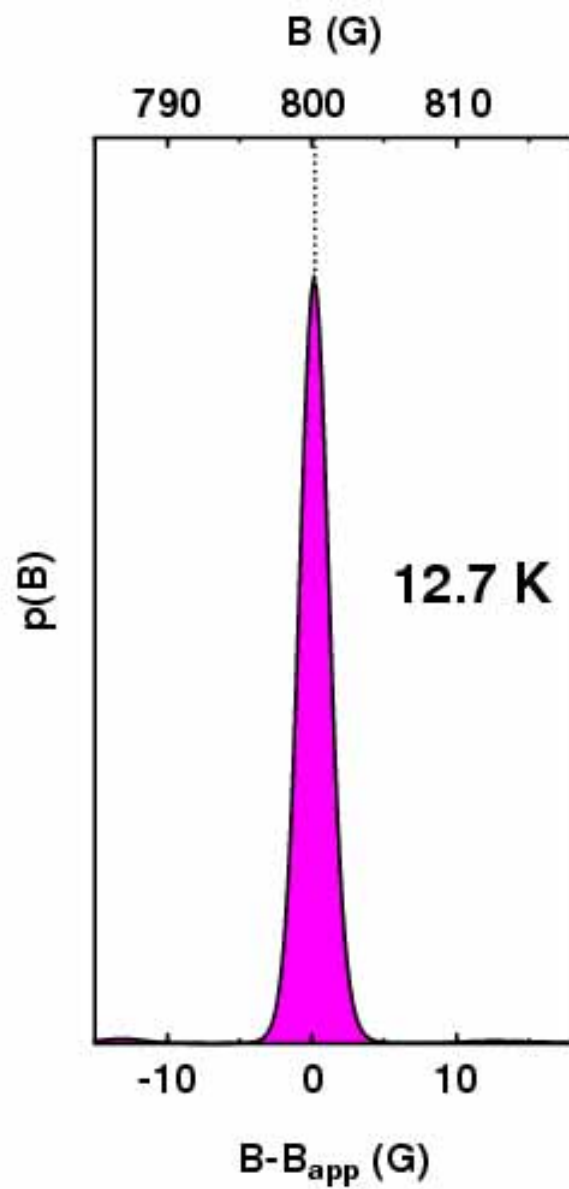


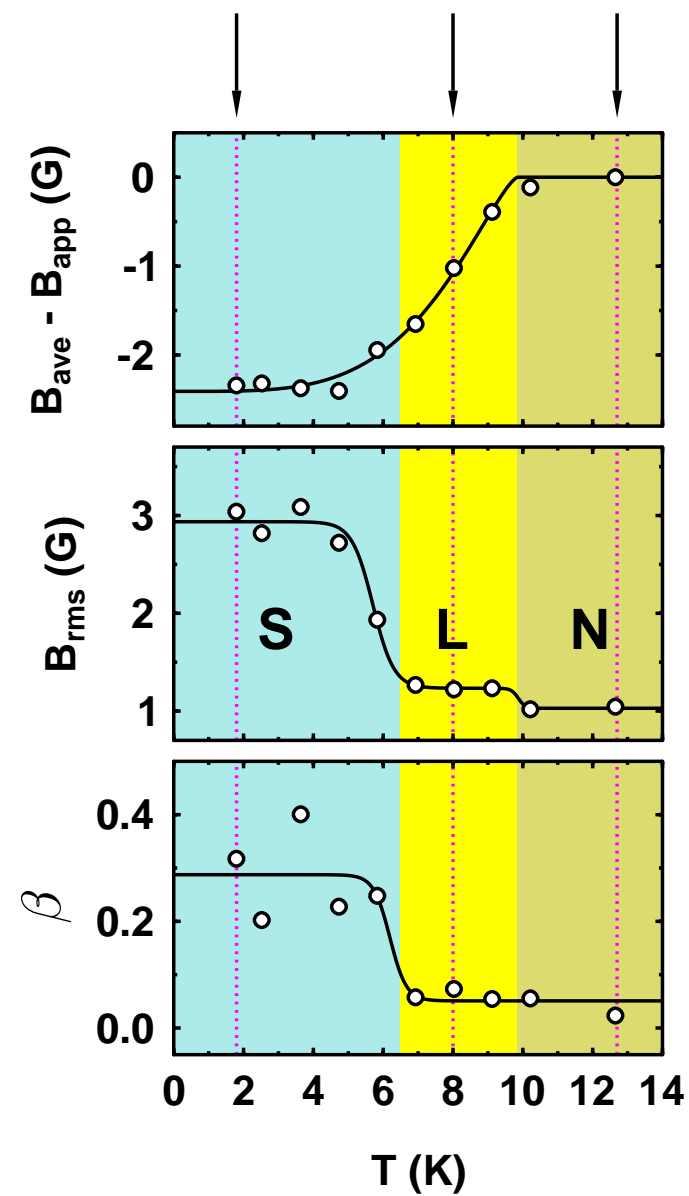
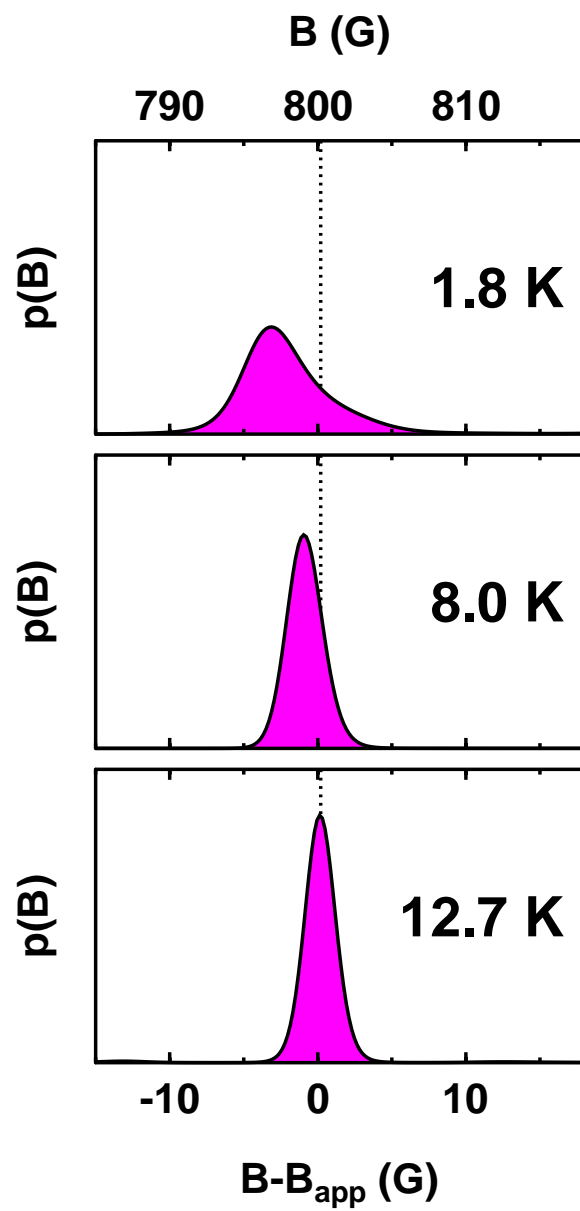












Time Domain Analysis versus Frequency Domain Analysis

Single Frequency

	Freq (MHz)	Width (MHz)
Test Data	1.0000	0.000
Time domain fit	0.9998(1)	0.001(1)
Maximum Entropy	1.006	0.003

Pair of Frequencies

	Freq (MHz)		Width (MHz)	
Test Data	0.9500,	1.0500	0.000,	0.000
Time domain fit	0.9493(1)	1.0499(3)	0.003(3)	0.004(3)
Maximum Entropy	0.956	1.054	0.002	0.005

Time Domain Analysis versus Frequency Domain Analysis

Summary

Transforms are good for determining a qualitative picture of data:

FFT best for spectra containing relatively **broad** features

All-poles transform best for spectra composed of **sharp** features

Maximum Entropy gives an unbiased view of the data but **Time Domain Fitting** gives best ultimate accuracy, provided the correct model is being used.

Combination of **Frequency Domain** and **Time Domain** analysis works best

Next:

Practical Data Analysis Workshop

back in R78