

## Photonics aa 2021/2022

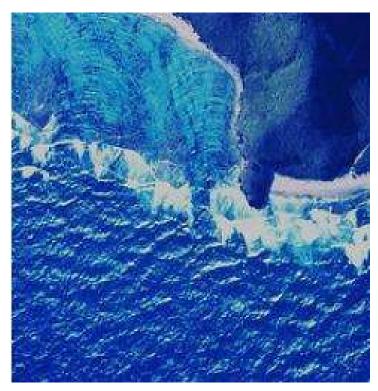
Prof. Maria Antonietta Vincenti Università degli Studi di Brescia

## Electromagnetic Optics: Reflection, Refraction and Scattering



### Refraction

### **Water Waves**



Waves refract where the water is shallower

### **E&M Waves**

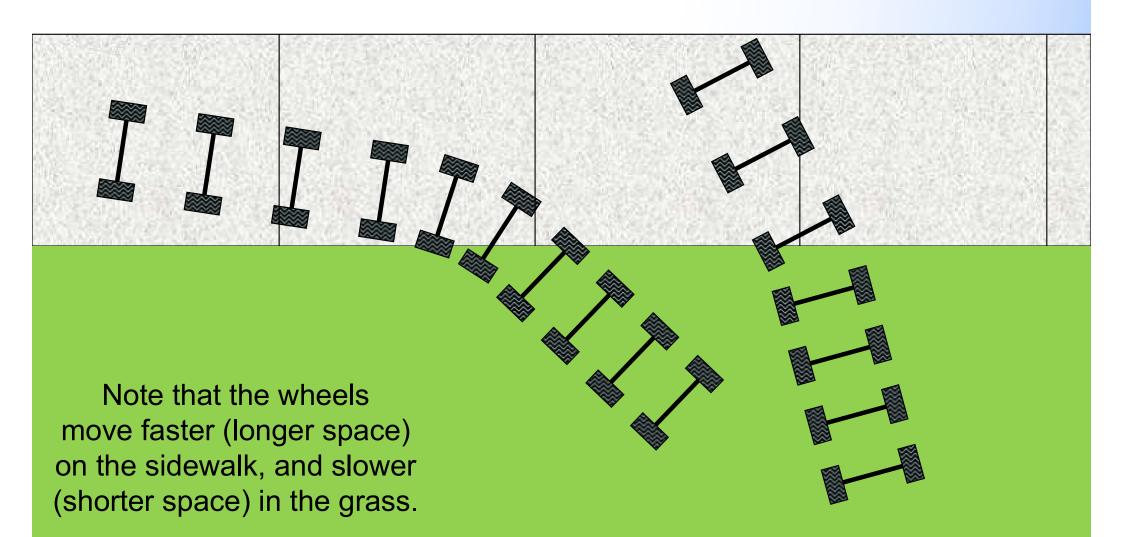


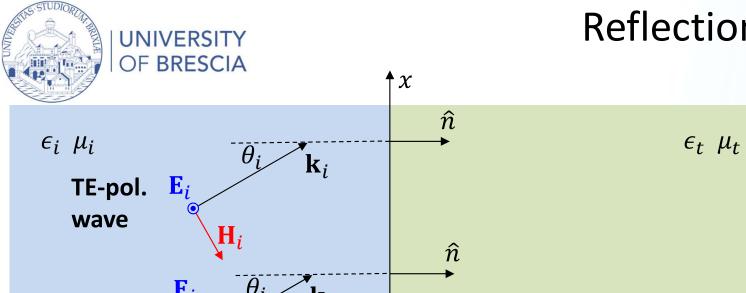
Refraction involves a change in the direction of wave propagation due to a change in propagation speed. It involves the oblique incidence of waves on media boundaries, and hence wave propagation in at least two dimensions.



### Refraction

Think of refraction as a pair of wheels on an axle going from a sidewalk onto grass. The wheel in the grass moves slower, so the direction of the wheel pair changes.





TM-pol.

wave

## Reflection and Refraction at an interface

Incidence and transmission media are assumed isotropic.

**Plane of Incidence**: The plane containing the incident wavevector  $\mathbf{k}_i$  and a vector that is normal to the interface is called the plane of incidence (in the figure above, the normal to the interface is  $\hat{z}$  and the x-z plane is the plane of incidence)

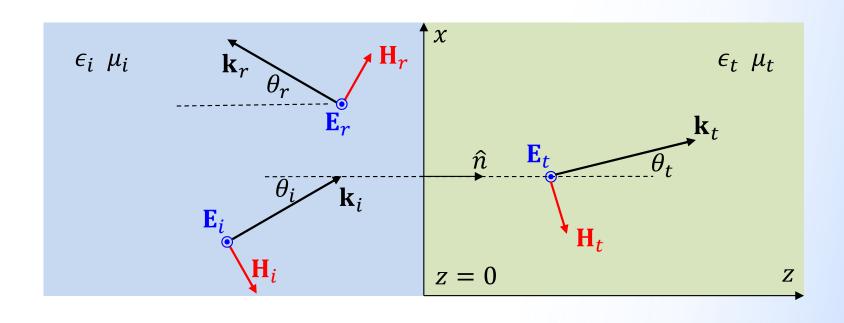
z=0

**TE Wave**: If the E-field of the wave is perpendicular to the plane of incidence then the wave is called a TE-wave (transverse electric polarization, also known as s-polarization)

**TM Wave**: If the H-field of the wave is perpendicular to the plane of incidence (i.e., if the E-field lies in the plane of incidence) then the wave is called a TM-wave (transverse magnetic polarization, also called p-polarization)



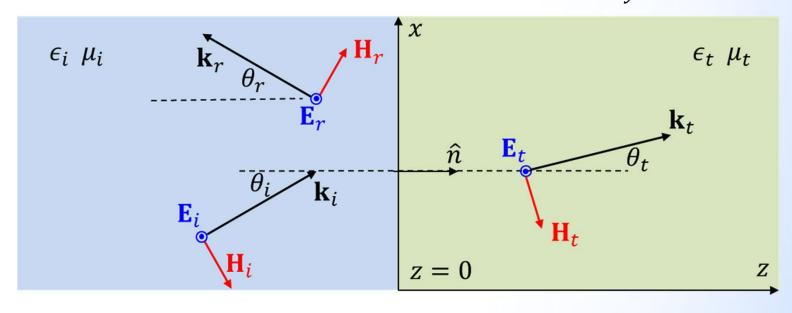
### **TE Wave**



$$\begin{aligned} \mathbf{k}_{i} &= k_{ix}\hat{x} + k_{iz}\hat{z} = k_{i} \Big[ \sin\left(\theta_{i}\right)\hat{x} + \cos\left(\theta_{i}\right)\hat{z} \Big], \quad \left| \mathbf{k}_{i} \right| = \omega\sqrt{\mu_{i}\varepsilon_{i}} = (\text{if nonmagnetic}) = \omega\frac{n_{i}}{c} \\ \mathbf{k}_{r} &= k_{rx}\hat{x} + k_{rz}\hat{z} = k_{r} \Big[ \sin\left(\theta_{r}\right)\hat{x} - \cos\left(\theta_{r}\right)\hat{z} \Big], \quad \left| \mathbf{k}_{r} \right| = \left| \mathbf{k}_{i} \right| = \omega\sqrt{\mu_{i}\varepsilon_{i}} = (\text{if nonmagnetic}) = \omega\frac{n_{i}}{c} \\ \mathbf{k}_{t} &= k_{tx}\hat{x} + k_{tz}\hat{z} = k_{t} \Big[ \sin\left(\theta_{t}\right)\hat{x} + \cos\left(\theta_{t}\right)\hat{z} \Big], \quad \left| \mathbf{k}_{t} \right| = \omega\sqrt{\mu_{t}\varepsilon_{t}} = (\text{if nonmagnetic}) = \omega\frac{n_{t}}{c} \end{aligned}$$



### **TE Wave – First Boundary Condition:** continuity of $E_{\gamma}$ at z=0



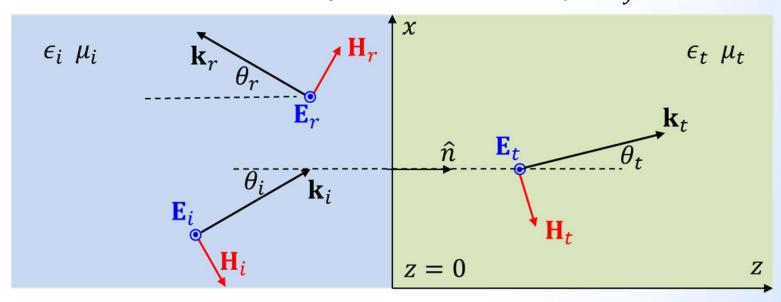
$$\begin{aligned} \mathbf{E}(\mathbf{r})|_{z<0} &= \hat{y}E_{i}e^{-j\mathbf{k}_{i}\cdot\mathbf{r}} + \hat{y}E_{r}e^{-j\mathbf{k}_{r}\cdot\mathbf{r}} \\ \mathbf{E}(\mathbf{r})|_{z>0} &= \hat{y}E_{t}e^{-j\mathbf{k}_{t}\cdot\mathbf{r}} + \hat{y}E_{r}e^{-j\mathbf{k}_{r}\cdot\mathbf{r}} \end{aligned} \qquad \begin{aligned} \mathbf{k}_{i} &= k_{i}\left[\sin\left(\theta_{i}\right)\hat{x} + \cos\left(\theta_{i}\right)\hat{z}\right] \\ \mathbf{k}_{r} &= k_{r}\left[\sin\left(\theta_{r}\right)\hat{x} - \cos\left(\theta_{r}\right)\hat{z}\right] \\ \mathbf{k}_{t} &= k_{t}\left[\sin\left(\theta_{t}\right)\hat{x} + \cos\left(\theta_{t}\right)\hat{z}\right] \end{aligned}$$

(1) At z = 0 the E-field parallel to the interface must be continuous across the interface for all x:

$$E_i e^{-jk_i \sin(\theta_i)x} + E_r e^{-jk_r \sin(\theta_r)x} = E_t e^{-jk_i \sin(\theta_i)x}$$



**TE Wave – First Boundary Condition:** continuity of  $E_{\nu}$  at z=0



The only way the above boundary condition can be satisfied for all x, and at all times t, is if all the x dependent phase factors are the same. This condition is called "phase matching condition", or conservation of the transverse wavevector, or conservation of the transverse momentum.

$$k_i \sin(\theta_i) = k_r \sin(\theta_r) = k_t \sin(\theta_t)$$

$$k_{ix} = k_{rx} = k_{tx}$$

This also implies that:

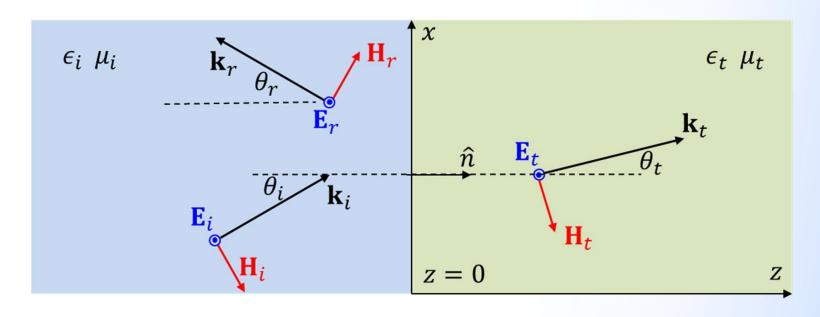
$$\sin(\theta_i) = \sin(\theta_r)$$

$$\theta_i = \theta_r$$

equals the angle of reflection



### TE Wave – Snell's Law



$$k_i \sin(\theta_i) = k_t \sin(\theta_t)$$

$$\omega \frac{n_i}{c} \sin(\theta_i) = \omega \frac{n_t}{c} \sin(\theta_t)$$

$$n_i \sin(\theta_i) = n_t \sin(\theta_t)$$
 Snell's law

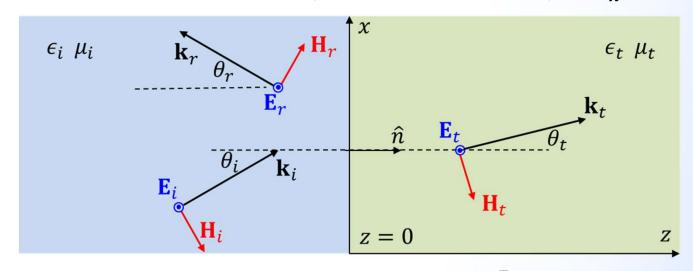
### Moreover:

$$E_i e^{-jk_i \sin(\theta_i)x} + E_r e^{-jk_r \sin(\theta_r)x} = E_t e^{-jk_i \sin(\theta_i)x}$$

$$(1)$$



### **TE Wave – Second Boundary Condition:** continuity of $H_x$ at z=0



$$\begin{aligned} \mathbf{H}(\mathbf{r})|_{z<0} &= \left(\hat{k}_i \times \hat{y}\right) \frac{E_i}{\eta_i} e^{-j\mathbf{k}_i \cdot \mathbf{r}} + \left(\hat{k}_r \times \hat{y}\right) \frac{E_r}{\eta_i} e^{-j\mathbf{k}_r \cdot \mathbf{r}} & \mathbf{k}_i &= k_i \left[\sin\left(\theta_i\right) \hat{x} + \cos\left(\theta_i\right) \hat{z}\right] \\ \mathbf{K}_r &= k_r \left[\sin\left(\theta_r\right) \hat{x} - \cos\left(\theta_r\right) \hat{z}\right] \\ \mathbf{H}(\mathbf{r})|_{z>0} &= \left(\hat{k}_t \times \hat{y}\right) \frac{E_t}{\eta_t} e^{-j\mathbf{k}_t \cdot \mathbf{r}} & \mathbf{k}_t &= k_t \left[\sin\left(\theta_t\right) \hat{x} + \cos\left(\theta_t\right) \hat{z}\right] \end{aligned}$$

(2) At z = 0 the H-field component parallel to the interface must be continuous for all x

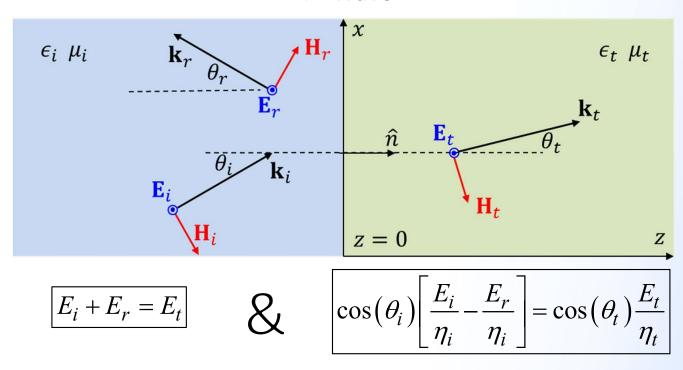
$$-\hat{x}\cos(\theta_{i})\frac{E_{i}}{\eta_{i}}e^{-jk_{i}\sin(\theta_{i})x} + \hat{x}\cos(\theta_{r})\frac{E_{r}}{\eta_{i}}e^{-jk_{r}\sin(\theta_{r})x} = -\hat{x}\cos(\theta_{t})\frac{E_{t}}{\eta_{t}}e^{-jk_{t}\sin(\theta_{t})x}$$

$$\cos(\theta_{i})\left[\frac{E_{i}}{\eta_{i}} - \frac{E_{r}}{\eta_{i}}\right] = \cos(\theta_{t})\frac{E_{t}}{\eta_{t}}$$
(2)



### Reflection and Transmission Coefficients

#### **TE Wave**



By solving the equation on the right using the equation on the left we get:

$$\left| t_{\text{TE}} = \frac{E_t}{E_i} = 2 \frac{\eta_t \cos(\theta_i)}{\eta_i \cos(\theta_t) + \eta_t \cos(\theta_i)} \right| \qquad \left| r_{\text{TE}} = \frac{E_r}{E_i} = \frac{\eta_t \cos(\theta_i) - \eta_i \cos(\theta_t)}{\eta_t \cos(\theta_i) + \eta_i \cos(\theta_t)} \right|$$

$$r_{\text{TE}} = \frac{E_r}{E_i} = \frac{\eta_t \cos(\theta_i) - \eta_i \cos(\theta_t)}{\eta_t \cos(\theta_i) + \eta_i \cos(\theta_t)}$$

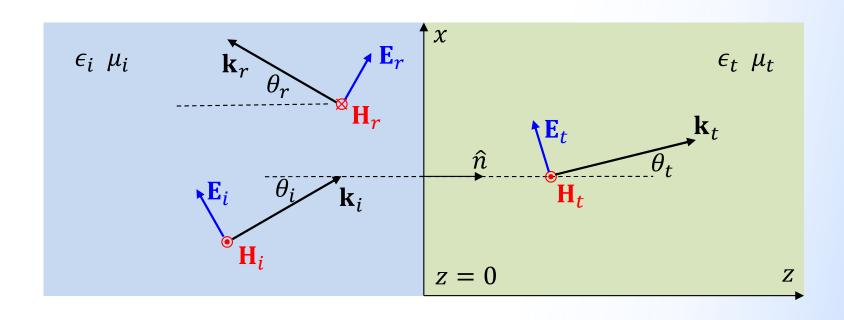
The two equations above can be re-written when considering  $\mu = \mu_0 = 1$  as:

$$\left| t_{\text{TE}} = \frac{E_t}{E_i} = 2 \frac{n_i \cos(\theta_i)}{n_t \cos(\theta_t) + n_i \cos(\theta_i)} \right|$$

$$r_{\text{TE}} = \frac{E_r}{E_i} = \frac{n_i \cos(\theta_i) - n_t \cos(\theta_t)}{n_i \cos(\theta_i) + n_t \cos(\theta_t)}$$



### **TM Wave**



The definitions of the wavevectors are identical for both polarizations:

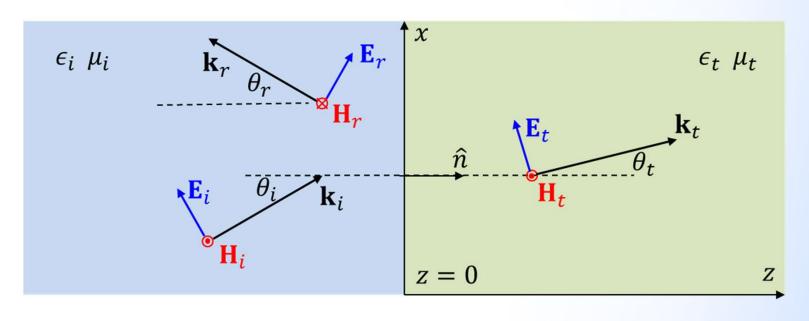
$$\mathbf{k}_{i} = k_{ix}\hat{x} + k_{iz}\hat{z} = k_{i}\left[\sin\left(\theta_{i}\right)\hat{x} + \cos\left(\theta_{i}\right)\hat{z}\right], \quad \left|\mathbf{k}_{i}\right| = \omega\sqrt{\mu_{i}\varepsilon_{i}} = (\text{if nonmagnetic}) = \omega\frac{n_{i}}{c}$$

$$\mathbf{k}_r = k_{rx}\hat{x} + k_{rz}\hat{z} = k_r \left[ \sin(\theta_r) \hat{x} - \cos(\theta_r) \hat{z} \right], \quad \left| \mathbf{k}_r \right| = \left| \mathbf{k}_i \right| = \omega \sqrt{\mu_i \varepsilon_i} = (\text{if nonmagnetic}) = \omega \frac{n_i}{c}$$

$$\mathbf{k}_{t} = k_{tx}\hat{x} + k_{tz}\hat{z} = k_{t}\left[\sin\left(\theta_{t}\right)\hat{x} + \cos\left(\theta_{t}\right)\hat{z}\right], \quad \left|\mathbf{k}_{t}\right| = \omega\sqrt{\mu_{t}\varepsilon_{t}} = (\text{if nonmagnetic}) = \omega\frac{n_{t}}{c}$$



### **TM Wave – First Boundary Condition:** continuity of $H_{\nu}$ at z=0



$$\mathbf{H}(\mathbf{r})|_{z<0} = \hat{y} \frac{E_i}{\eta_i} e^{-j\mathbf{k}_i \cdot \mathbf{r}} - \hat{y} \frac{E_r}{\eta_i} e^{-j\mathbf{k}_r \cdot \mathbf{r}}$$

$$\mathbf{H}(\mathbf{r})|_{z>0} = \hat{y} \frac{E_t}{\eta_t} e^{-j\mathbf{k}_t \cdot \mathbf{r}}$$

$$\mathbf{K}_i = k_i \left[ \sin(\theta_i) \hat{x} + \cos(\theta_i) \hat{z} \right]$$

$$\mathbf{k}_r = k_r \left[ \sin(\theta_r) \hat{x} - \cos(\theta_r) \hat{z} \right]$$

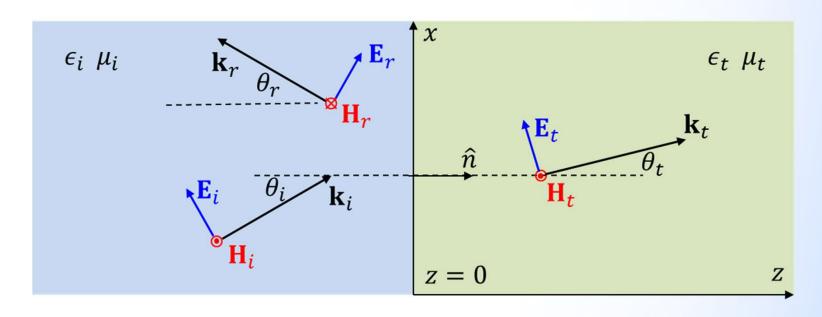
$$\mathbf{k}_t = k_t \left[ \sin(\theta_t) \hat{x} + \cos(\theta_t) \hat{z} \right]$$

(1) At z = 0 the H-field parallel to the interface must be continuous across the interface for all x:

$$\frac{E_i}{\eta_i} e^{-jk_i \sin(\theta_i)x} - \frac{E_r}{\eta_i} e^{-jk_r \sin(\theta_r)x} = \frac{E_t}{\eta_t} e^{-jk_t \sin(\theta_t)x}$$



### **TM Wave – First Boundary Condition**



The only way the above boundary condition can be satisfied for all x is if all the x dependent phase factors are the same. This condition is called "phase matching".

$$k_i \sin(\theta_i) = k_r \sin(\theta_r) = k_t \sin(\theta_t)$$

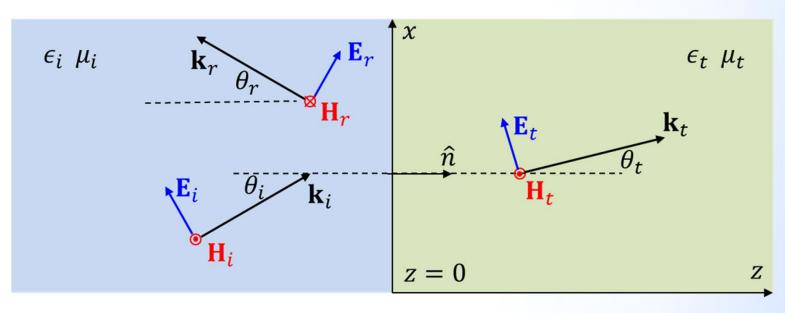
$$k_{ix} = k_{rx} = k_{tx}$$

This also implies that:

equals the angle of reflection



### TM Wave - Snell's Law



$$k_i \sin(\theta_i) = k_t \sin(\theta_t)$$

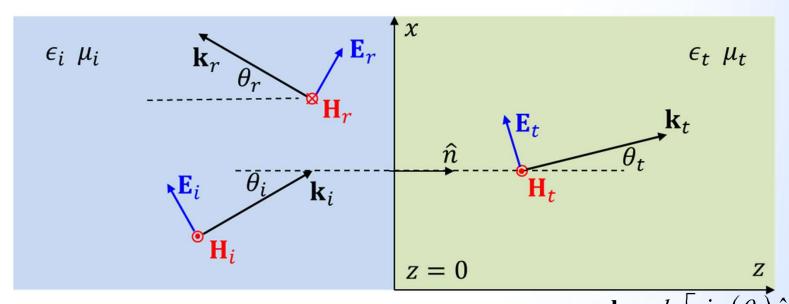
$$\omega \frac{n_i}{c} \sin(\theta_i) = \omega \frac{n_t}{c} \sin(\theta_t)$$

$$n_i \sin(\theta_i) = n_t \sin(\theta_t)$$
 Snell's law

### Moreover:



### **TM Wave – Second Boundary Condition**



$$\mathbf{E}(\mathbf{r})|_{z<0} = -(\hat{k}_i \times \hat{y}) E_i e^{-j\mathbf{k}_i \cdot \mathbf{r}} + (\hat{k}_r \times \hat{y}) E_r e^{-j\mathbf{k}_r \cdot \mathbf{r}}$$

$$\mathbf{k}_i = k_i \left[ \sin(\theta_i) \hat{x} + \cos(\theta_i) \hat{z} \right]$$

$$\mathbf{k}_r = k_r \left[ \sin(\theta_r) \hat{x} - \cos(\theta_r) \hat{z} \right]$$

$$\mathbf{E}(\mathbf{r})|_{z>0} = -(\hat{k}_t \times \hat{y}) E_t e^{-j\mathbf{k}_t \cdot \mathbf{r}}$$

$$\mathbf{k}_t = k_t \left[ \sin(\theta_t) \hat{x} + \cos(\theta_t) \hat{z} \right]$$

(2) At z = 0 the E-field component parallel to the interface must be continuous for all x

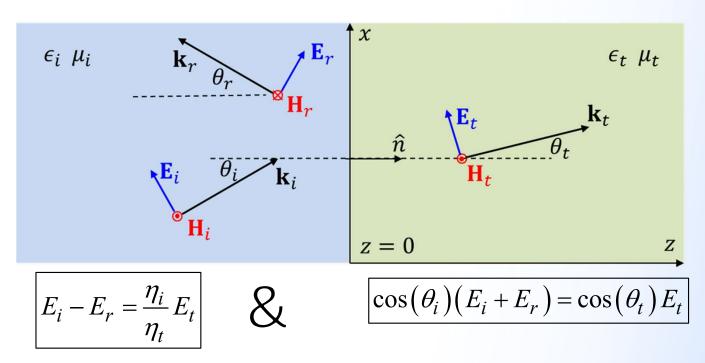
$$-\hat{x}\cos(\theta_i)E_ie^{-jk_i\sin(\theta_i)x} - \hat{x}\cos(\theta_r)E_re^{-jk_r\sin(\theta_r)x} = -\hat{x}\cos(\theta_t)E_te^{-jk_t\sin(\theta_t)x}$$

$$\cos(\theta_i)(E_i + E_r) = \cos(\theta_t)E_t$$
 (2)



### Reflection and Transmission Coefficients

### TM Wave



By solving the equation on the right using the equation on the left we get:

$$\left| t_{\text{TM}} = \frac{E_t}{E_i} = 2 \frac{\eta_t \cos(\theta_i)}{\eta_t \cos(\theta_t) + \eta_i \cos(\theta_i)} \right| \qquad \left| r_{\text{TM}} = \frac{E_r}{E_i} = \frac{\eta_t \cos(\theta_t) - \eta_i \cos(\theta_i)}{\eta_t \cos(\theta_t) + \eta_i \cos(\theta_i)} \right|$$

$$r_{\text{TM}} = \frac{E_r}{E_i} = \frac{\eta_t \cos(\theta_t) - \eta_i \cos(\theta_i)}{\eta_t \cos(\theta_t) + \eta_i \cos(\theta_i)}$$

The two equations above can be re-written when considering  $\mu = \mu_0 = 1$ , as:

$$t_{\text{TM}} = \frac{E_t}{E_i} = 2 \frac{n_i \cos(\theta_i)}{n_i \cos(\theta_t) + n_t \cos(\theta_i)}$$

$$r_{\text{TM}} = \frac{E_r}{E_i} = \frac{n_i \cos(\theta_t) - n_t \cos(\theta_i)}{n_i \cos(\theta_t) + n_t \cos(\theta_i)}$$



## Reflection and Transmission Coefficients

### For non magnetic materials

$$\mu = \mu_0$$

$$r_{\text{TE}} = \frac{n_i \cos(\theta_i) - n_t \cos(\theta_t)}{n_i \cos(\theta_i) + n_t \cos(\theta_t)}$$

$$t_{\text{TE}} = 2 \frac{n_i \cos(\theta_i)}{n_t \cos(\theta_t) + n_i \cos(\theta_i)}$$

$$r_{\text{TM}} = \frac{n_i \cos(\theta_t) - n_t \cos(\theta_i)}{n_i \cos(\theta_t) + n_t \cos(\theta_i)}$$

$$t_{\text{TM}} = 2 \frac{n_i \cos(\theta_i)}{n_i \cos(\theta_t) + n_t \cos(\theta_i)}$$

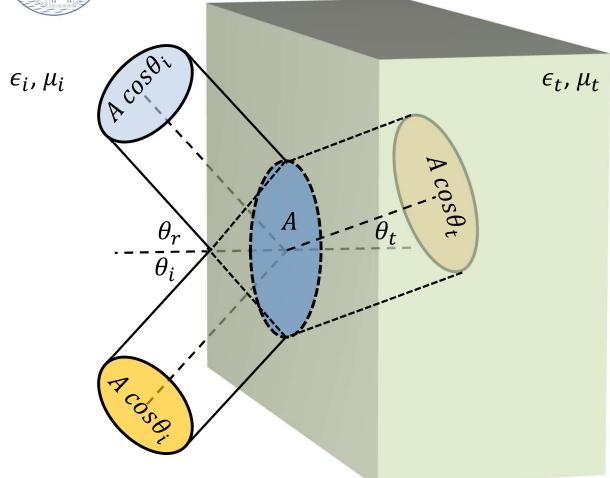
**TM Wave** 

### TE Wave

# incident wave interface incident wave interface E-field vectors are red. k vectors are black.

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### Reflectance and Transmittance



Consider a finite beam of uniform intensity, equal to the intensity of the input plane wave,  $I_i = |Re(S_i)| = \frac{|E_i|^2}{2\eta_i}$ Here we assume the input medium be lossless.

**Reflectance or Reflectivity** 

$$R = P_r/P_i$$

**Transmittance or transmittivity** 

$$T = P_t/P_i$$

Power of the input beam

$$P_i = I_i A_i = \frac{|E_i|^2}{2\eta_i} A \cos\theta_i$$

Power of the reflected beam

$$P_i = I_i A_i = \frac{|E_i|^2}{2\eta_i} A \cos \theta_i \qquad P_r = I_r A_r = \frac{|E_r|^2}{2\eta_i} A \cos \theta_i \qquad P_t = I_i A_i = \frac{|E_t|^2}{2\eta_t} A \cos \theta_t$$

Power of the transmitted beam

$$P_t = I_i A_i = \frac{|E_t|^2}{2\eta_t} A \cos \theta_t$$

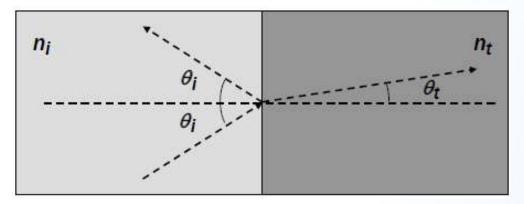
$$R = \frac{|E_r|^2}{|E_i|^2} = |r|^2$$

$$T = 1 - R = |t|^2 \frac{\eta_i \cos \theta_t}{\eta_t \cos \theta_i}$$

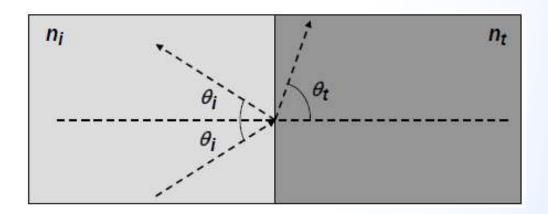


### Snell's Law

$$n_i \sin(\theta_i) = n_t \sin(\theta_t)$$



If  $n_i < n_t$  then  $\theta_t < \theta_i$  and the transmitted wave bends towards the normal

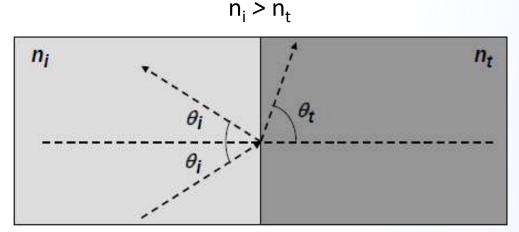


If  $n_i > n_t$  then  $\theta_t > \theta_i$  and the transmitted wave bends away the normal

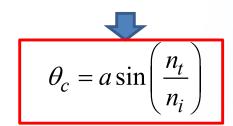


## Total Internal Reflection & Critical Angle

If  $\theta_i$  is increased, then  $\theta_t$  will eventually become 90°. The value of  $\theta_i$  for which  $\theta_t$  is 90° is called the critical angle  $\theta_c$ 



$$n_i \sin\left(\theta_c\right) = n_t \sin\left(\frac{\pi}{2}\right)$$



If  $\theta_i$  is increased beyond  $\theta_c$  the wave is not transmitted but is completely (100%) reflected at the interface back into the medium of incidence.

This phenomenon is called TOTAL INTERNAL REFLECTION and it happens for both TE and TM waves



### **Total Internal Reflection** Phase Matching

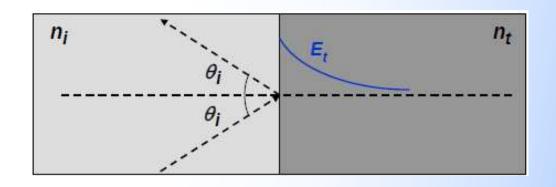
 $n_i > n_t$  and  $\theta_i > \theta_c$ 

The phase matching condition gives:

$$k_{ix} = k_{rx} = k_{tx}$$

$$k_{tx} = k_{ix} = k_{i} \sin(\theta_{i})$$





$$k_{tx}^2 = k_i^2 \sin^2(\theta_i) = \frac{\omega^2}{c^2} n_i^2 \sin^2(\theta_i)$$

The dispersion relation in the medium "t" is: 
$$k_t^2 = \frac{\omega^2}{c^2} n_t^2$$
  $\Rightarrow$   $k_{tx}^2 + k_{tz}^2 = \frac{\omega^2}{c^2} n_t^2$ 

$$k_{tz}^2 = \frac{\omega^2}{c^2} n_t^2 - k_{tx}^2 = \frac{\omega^2}{c^2} \left[ n_t^2 - n_i^2 \sin^2(\theta_i) \right] \qquad \Longrightarrow \qquad k_{tz} = \frac{\omega}{c} \left( n_t^2 - n_i^2 \sin^2(\theta_i) \right)$$
 This value is **NEGATIVE** when  $\theta_l > \theta_c$ 

$$k_{tz} = \frac{\omega}{c} \left( n_t^2 - n_i^2 \sin^2(\theta_i) \right)$$

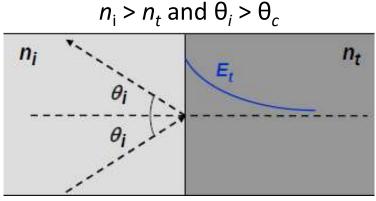
In other words the z-component of the wavevector has become completely imaginary, therefore we can write the E-field (assuming TE polarization) as:

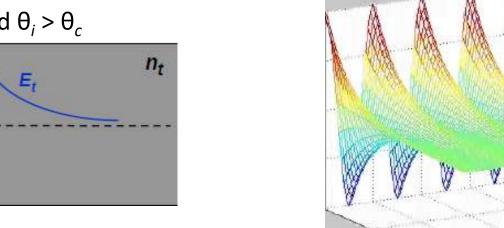
$$k_{tz} = -j\frac{\omega}{c}\sqrt{n_i^2\sin^2(\theta_i) - n_t^2} = -jk_{tz}$$

$$\mathbf{E}(\mathbf{r})\big|_{z>0} = \hat{y}E_t e^{-jk_{tx}x} e^{-k_{tz}^*z} \qquad \text{EVA}$$



### **Total Internal Reflection Electric Field Profile**



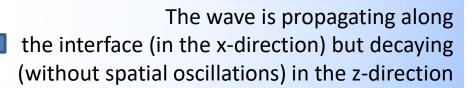


the E-field in the medium "t" (assuming TE polarization) is:

$$\mathbf{E}(\mathbf{r})\big|_{z>0} = \hat{y}E_t e^{-jk_{tx}x} e^{-k_{tz}^{"}z} = \hat{y}|E_t|e^{-j\varphi} e^{-jk_{tx}x} e^{-k_{tz}^{"}z}$$



$$\left| \mathbf{E}(\mathbf{r}) \right|_{z>0} = \hat{y} \left| E_t \right| e^{-k_{tz}^2 z} \cos \left( \omega t - k_{tx} x - \varphi \right) \right|$$



If  $\theta_i > \theta_c$  the wave is completely reflected back into the medium of incidence and we have:

$$k_{tz} = -j\frac{\omega}{c}\sqrt{n_i^2\sin^2(\theta_i) - n_t^2} = -jk_{tz}$$

$$|\Gamma| = 1$$
,  $\varphi$  Goos-Hanschen phase-shift

The reflection coefficient for the E-field (assuming TE wave) is: 
$$\Gamma = \frac{E_r}{E_i} = \frac{k_{iz} - k_{tz}}{k_{iz} + k_{tz}} = \frac{k_{iz} + jk_{tz}^{"}}{k_{iz} - jk_{tz}^{"}} = e^{i\varphi}$$



### Brewster's Angle



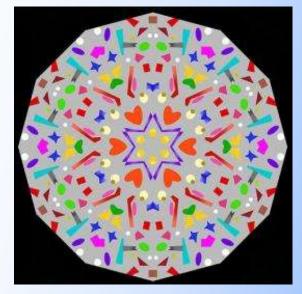
### **Today's Culture Moment**

### Sir David Brewster

- Scottish scientist
- Studied at University of Edinburgh at age 12
- Independently discovered Fresnel lens
- Editor of *Edinburgh Encyclopedia and contributor to Encyclopedia Britannica (7<sup>th</sup> and 8<sup>th</sup> editions)*
- Inventor of the Kaleidoscope
- Nominated (1849) to the National Institute of France.



1781 –1868



Kaleidoscope

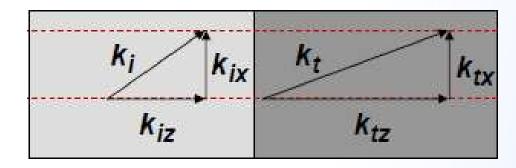


### Brewster's Angle

Question: Can one ever get the reflection coefficient to go to zero (very desirable to get rid of unwanted reflections in optics)?

$$\left| r_{\text{TE}} = \frac{E_r}{E_i} = \frac{n_i \cos(\theta_i) - n_t \cos(\theta_t)}{n_i \cos(\theta_i) + n_t \cos(\theta_t)} \right|$$

$$r_{\text{TM}} = \frac{H_r}{H_i} = \frac{n_i \cos(\theta_t) - n_t \cos(\theta_i)}{n_i \cos(\theta_t) + n_t \cos(\theta_i)}$$



For a TE wave reflection is zero if :  $n_i \cos(\theta_i) = n_t \cos(\theta_t)$   $\Longrightarrow k_i \cos(\theta_i) = k_t \cos(\theta_t)$ 

If:  $n_i \neq n_t$  then  $k_{iz} \neq k_{tz}$  since  $k_{ix} = k_{tx}$  for the phase-matching condition.

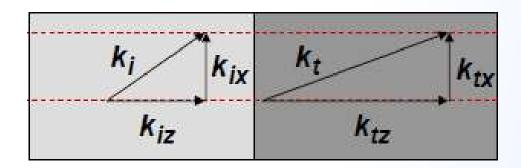


### Brewster's Angle

<u>Question</u>: Can one ever get the reflection coefficient to go to zero (very desirable to get rid of unwanted reflections in optics)?

$$\left| r_{\text{TE}} = \frac{E_r}{E_i} = \frac{n_i \cos(\theta_i) - n_t \cos(\theta_t)}{n_i \cos(\theta_i) + n_t \cos(\theta_t)} \right|$$

$$\left| r_{\text{TM}} = \frac{H_r}{H_i} = \frac{n_i \cos(\theta_t) - n_t \cos(\theta_i)}{n_i \cos(\theta_t) + n_t \cos(\theta_i)} \right|$$



For a TM wave reflection is zero if :  $n_i \cos(\theta_t) = n_t \cos(\theta_t)$ 

For Snell's law:  $n_i \sin(\theta_i) = n_t \sin(\theta_t)$ 

By squaring both equations, solving for the terms in  $\theta_i$  and then adding each corresponding term one can obtain:

## Practical Example: The Lens Flare!

If  $\theta_i$ =0 reflectance and transmittance are:

$$R = |r|^2 = \left(\frac{n_i - n_t}{n_i + n_t}\right)^2$$

$$T = \frac{n_t}{n_i} |t|^2 = 4 \left(\frac{n_i}{n_t + n_i}\right)^2 \frac{n_t}{n_i}$$

For a camera lens reflection/transmission occurs at the interface between glass and air. If  $n_i = 1$  and  $n_t = 1.5$ :

$$R = 4\%$$
 and  $T = 96\%$ 

Reflectance is "only" 4% but has big implication for photography





## Practical Example: Fresnel Equations in action!



Windows look like mirrors at night (when you're in a brightly lit room).



One-way mirrors (used by police to interrogate bad guys) are just partial reflectors (actually, with a very thin aluminum coating).



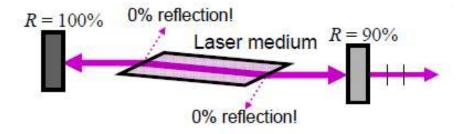
Smooth surfaces can produce pretty good mirror-like reflections, even though they are not made of metal.



## Practical Example: Fresnel Equations in action!



Optical fibers only work because of total internal reflection.



Many lasers use Brewster's angle components to avoid reflective losses



### Scattering

Scattering can be broadly defined as the redirection of radiation out of the original direction of propagation usually due to interactions with molecules and particles



Reflection, refraction, diffraction etc. are actually all just forms of scattering. The superposition of incident and scattered waves is what is actually observed.



### When is scattering important?

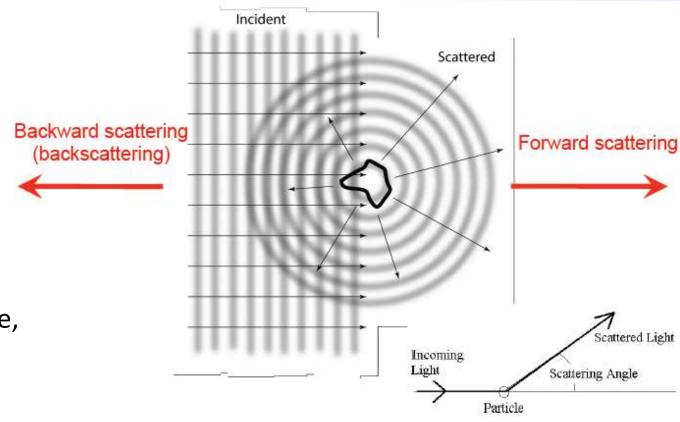
Scattering is negligible whenever gains in intensity due to scattering are small

compared to:

Losses due to extinction

 Gains due to thermal emission

 When considering direct radiation from a point source, such as the sun



In the UV, visible and near-IR bands, scattering is the dominant source of radiation along any line of sight, other than looking directly at the sun.



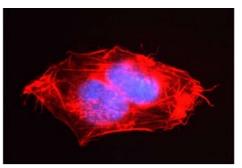
### Types of Scattering

Elastic scattering – the wavelength (frequency) of the scattered light is the same as the incident light (Rayleigh and Mie scattering)



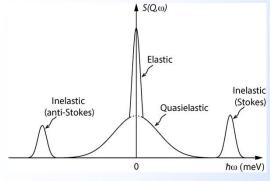
Inelastic scattering – the emitted radiation has a wavelength different from that of the

incident radiation (Raman scattering, fluorescence)



Quasi-elastic scattering – the wavelength (frequency) of the scattered light shifts (e.g.,

in moving matter due to Doppler effects)



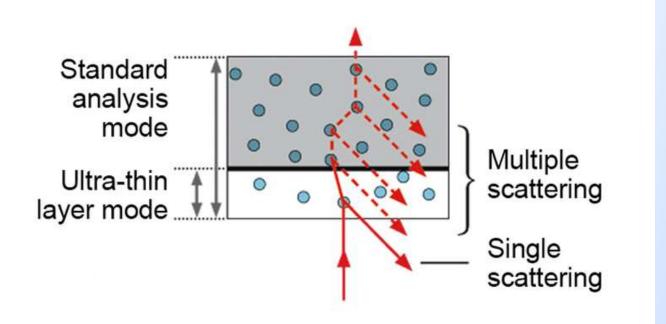
Green Laser (original)



### Types of Scattering

Single scattering: photons scattered only once prevail in optically thin media, since photons have a high probability of exiting the medium (e.g., a thin cloud) before being scattered again. Also favored in strongly absorbing media

Multiple scattering: prevails in optically thick, strongly scattering and non-absorbing media. Photons may be scattered hundreds of times before emerging





## Practical example: Variation in sky brightness

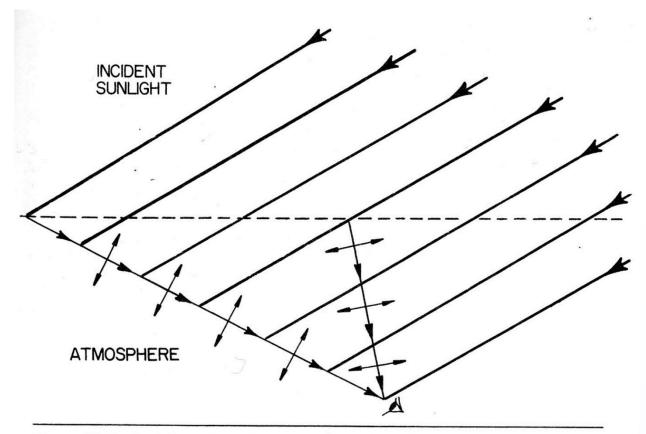


Figure 20.3 Path lengths in the atmosphere. An observer receives light scattered by all the molecules and particles along the line of sight. Paths near the horizon are longer than those near the zenith, hence the horizon sky is brighter. From *The Physics Teacher*, C. F. Bohren and A. B. Fraser, May 1985.

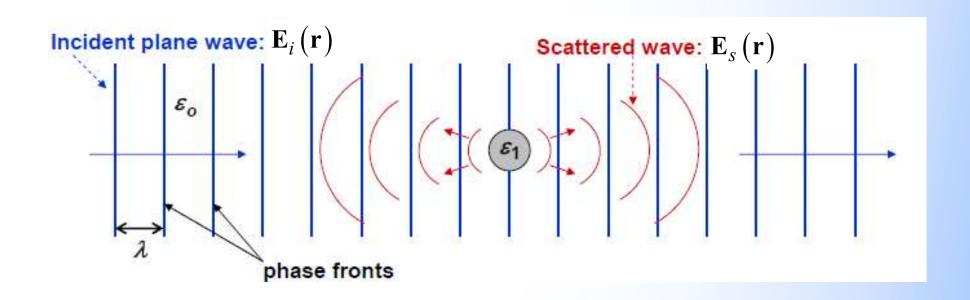
NOTE: The horizon sky is usually brighter than the zenith sky.

This is a result of single scattering (zenith) vs. multiple scattering (horizon)



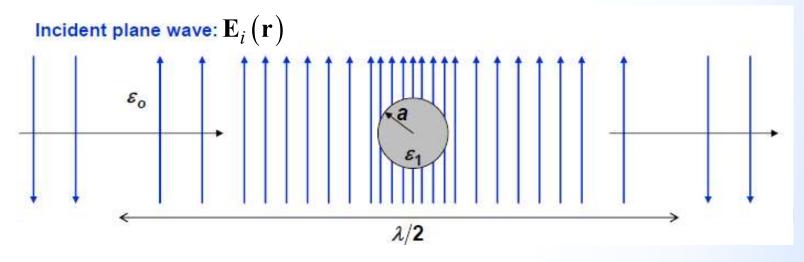
### Parameters governing scattering

- (1) The wavelength ( $\lambda$ ) of the incident radiation
- (2) The size of the scattering particle
- (3) The particle optical properties relative to the surrounding medium





### Scattering from spherical particles



Different scattering conditions can be identified depending on their geometrical size in relation with the incident wavelength. Let's define the adimensional parameter:

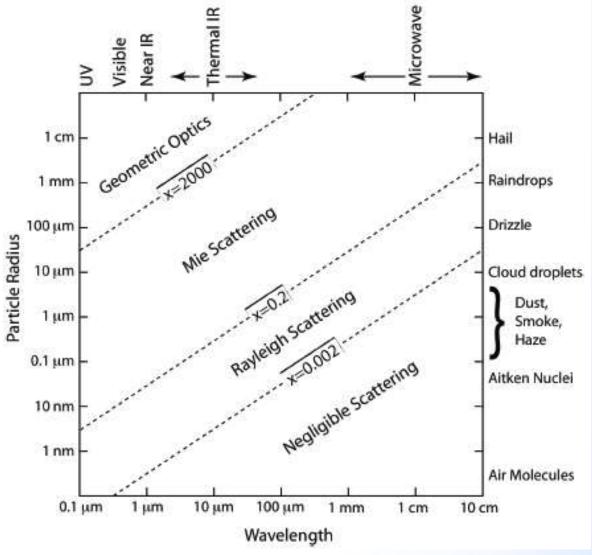
$$x = \frac{2\pi a}{\lambda} = ka$$

If ka << 1 the scattering is called Rayleigh Scattering
If ka ~ 1 the scattering is called Mie Scattering
If ka >> 1 the scattering is called Geometrical Scattering

When ka << 1, the particle sees a uniform E-field that is slowly oscillating in time



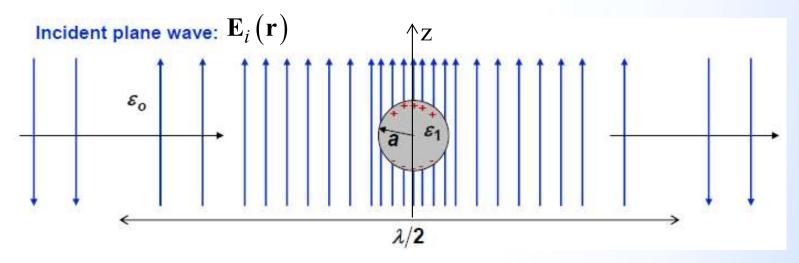
### Light scattering regimes



NOTE: This plot considers only single scattering by spheres. Multiple scattering and scattering by non-spherical objects can get really complex!



### Rayleigh Scattering



One way to understand scattering is as follows:

- i) The incident E-field induces a time-varying dipole moment in the sphere
- ii) The time-varying dipole radiates like a Hertzian dipole and this is the scattered radiation

Let's now suppose the z-directed E-field phasor for the incident plane wave at the location of the particle is:

$$\mathbf{E}(\mathbf{r}=0) = \hat{z}E_i$$

The z-directed dipole moment p induced in a sphere in the presence of E-field E is:

$$p = 4\pi\varepsilon_0 a^3 \left( \frac{\varepsilon_1 - \varepsilon_0}{\varepsilon_1 + 2\varepsilon_0} \right) E_i$$



### Rayleigh Scattering

Total scattered power  $P_s$  from a dielectric sphere is:

$$P_{s} = \int_{0}^{2\pi} \int_{0}^{\pi} \frac{\left|\mathbf{E}_{s}(\mathbf{r})\right|^{2}}{2\eta_{0}} r^{2} \sin(\theta) d\theta d\phi = \frac{4\pi}{3\eta_{0}} k^{4} a^{6} \left(\frac{\varepsilon_{1} - \varepsilon_{0}}{\varepsilon_{1} + 2\varepsilon_{0}}\right)^{2} \left|E_{i}\right|^{2}$$

The incident power per unit area is the Poynting vector of the incident wave:  $\frac{\left|E_i\right|^2}{2\eta_0}$ 

The scattering cross-section  $\sigma_s$  of a scatterer is defined as the area of a plane oriented perpendicular to the direction of incident wave that would intercept the same total incident power as the power  $P_s$  that the scatterer radiates:

$$\sigma_{s} = \frac{P_{s}}{\left|\mathbf{E}_{i}(\mathbf{r})\right|^{2}/2\eta_{0}}$$

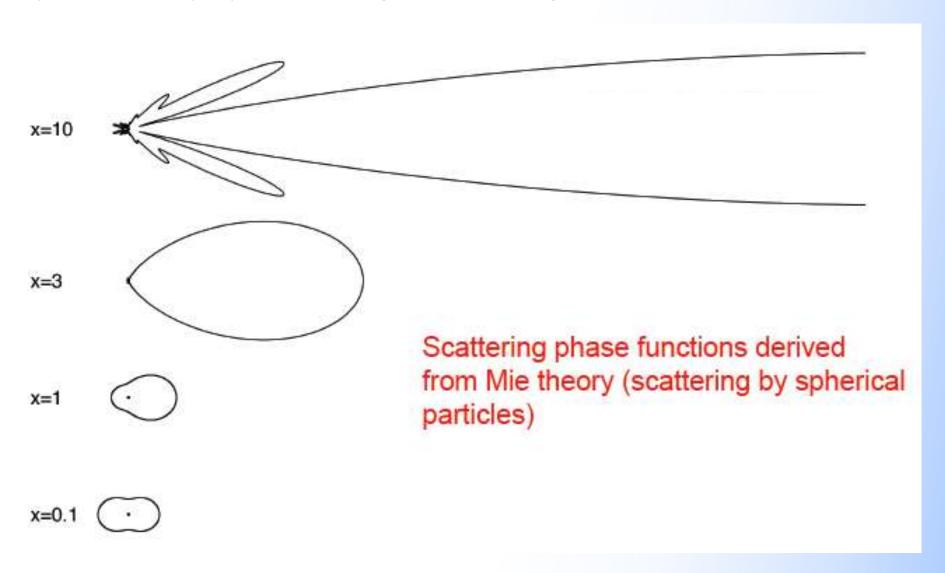
 $\sigma_s$  is also the ratio of the total scattered power to the power per unit area of the incident wave at the location of the scatterer

For the dielectric sphere:  $\sigma_{s} = \frac{8\pi}{3} k^{4} a^{6} \left( \frac{\varepsilon_{1} - \varepsilon_{0}}{\varepsilon_{1} + 2\varepsilon_{0}} \right)^{2}$ 



### Scattering phase functions

The scattering phase function, or phase function, gives the angular distribution of light intensity scattered by a particle at a given wavelength.



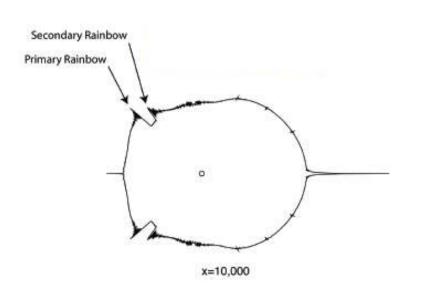


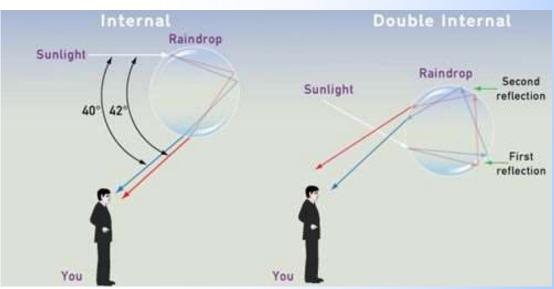
### Practical example: Rainbows



Rainbow: for large particles (x = 10,000), the forward and backward peaks in the scattering phase function become very narrow (almost non-existent). Light paths are best predicted using geometric optics and ray tracing.

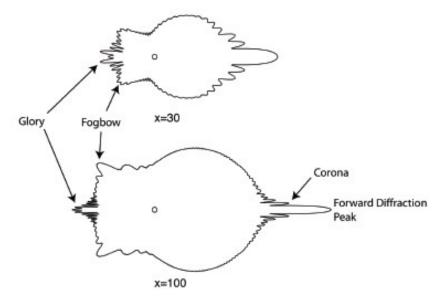
Primary rainbow: single internal reflection Secondary rainbow: double internal reflection







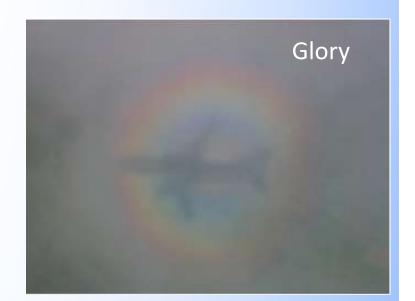
Fogbow: spikes in scattering phase function present but not sharp as for rainbows. Hence the separation of colors (due to varying refractive index) is not as vivid as a normal rainbow. A whitish ring centered on one's shadow (i.e. opposite the sun) is seen. Arises when water droplets have a size characteristic of fog and clouds rather than rain.



Glory: opposite end of the phase function from the corona. Seen as a 'halo' around one's shadow when looking at a fog bank with the sun at your back. Also seen from aircraft. Glories have vivid colors if the range of drop sizes in the fog is relatively narrow, otherwise they are whitish.

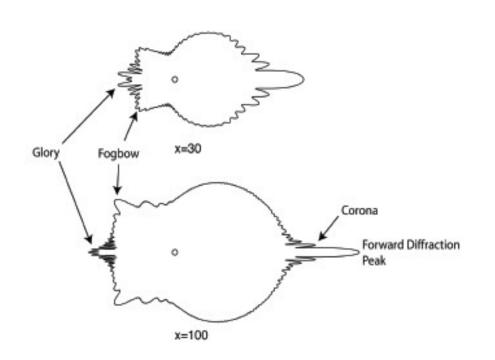
## Practical example: Glory and Fogbow







### Practical example: Corona





Corona: for intermediate values of the size parameter (x), the forward scattering peak is accompanied by weaker *sidelobes*. If you were to view the sun through a thin cloud composed of identical spherical droplets (with x = 100 or less), you would see closely spaced rings around the light source. The angular position of the rings depends on wavelength, so the rings would be colored. This is a *corona*. Because few real clouds have a sufficiently narrow distribution of drop sizes, coronas are usually more diffuse and less brightly colored.