

# Tutorial 2 - Gaussian Beams

A. O. B

Reminder:

$$E(x, y, z) = \frac{E_0}{\sqrt{1 + z^2/d^2}} \cdot e^{-\frac{(x^2+y^2)}{w^2(z)}} \cdot e^{-j\frac{\pi(x^2+y^2)}{\lambda R(z)}} \cdot e^{-jkz} \cdot e^{-j\text{Gouy phase shift.}}$$

Gaussian amplitude
spherical phase term
propagation term
Gouy phase shift.

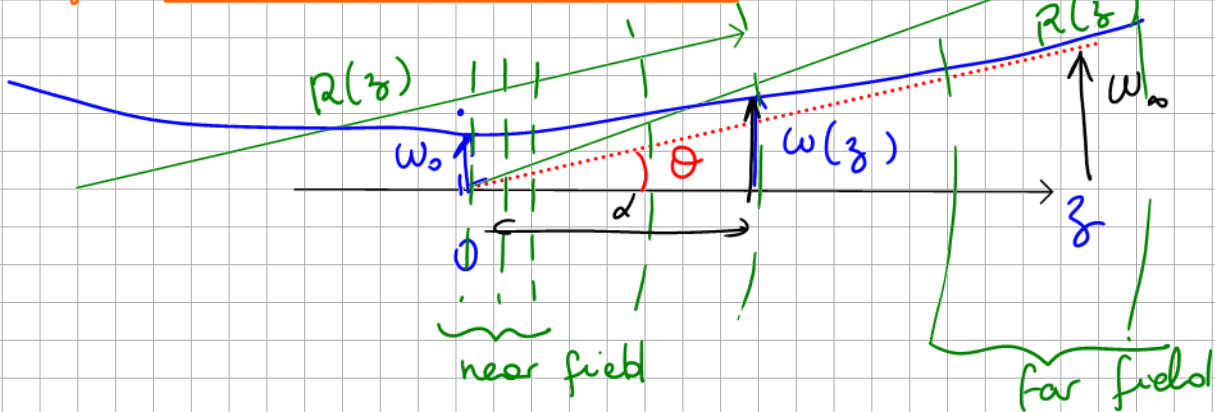
with

$$d = \frac{\pi w_0^2}{\lambda}$$

$$w^2(z) = w_0^2 \left(1 + \frac{z^2}{d^2}\right)$$

$$R(z) = z + \frac{d^2}{z}$$

$$z \left(1 + \frac{d^2}{z^2}\right)$$



Ex 1

1)  $w^2(z=d) = w_0^2 (1+1) \Rightarrow \boxed{w(d) = \sqrt{2} w_0}$

2)

$w_0$	0.01 mm	1 mm	10 mm	1 m
$d$	0.3 mm	3 m	300 m	3 000 km.

Single mode fiber

He-Ne Laser

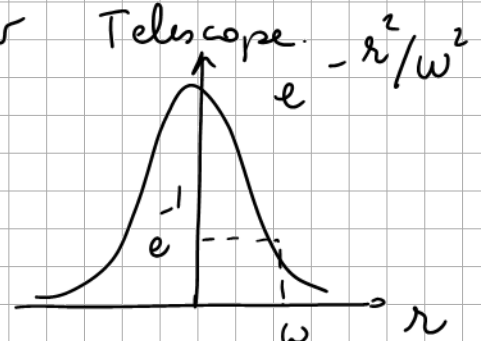
Expander

Telescope

$$d = \frac{\pi w_0^2}{\lambda}$$

$z \ll d \rightarrow$  Near field

$z \gg d \rightarrow$  Far field



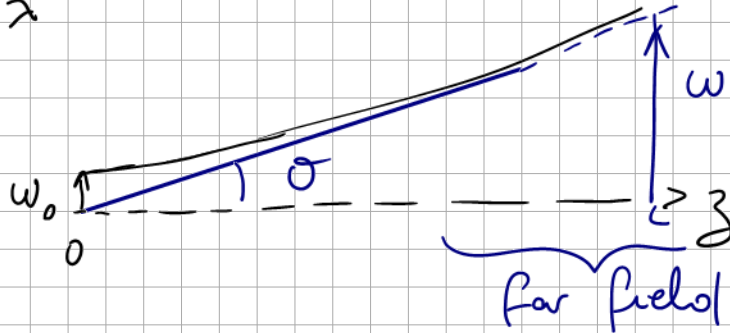


Ex 2:

1) Far field  $\Rightarrow w(L) = \theta \cdot L$  with  $\theta = \frac{\lambda}{\pi w_0}$   
 $\approx 25 \text{ km}$   
 $\phi(L) \approx 50 \text{ km}$

$$\alpha = \frac{\pi w_0^2}{\lambda}$$

$$L \gg \alpha$$



$$\frac{w}{L} = \theta$$

$$\theta = \frac{\lambda}{\pi w_0}$$

$$2) P_0 = \kappa \iint_{-\infty}^{+\infty} |E|^2 ds$$

$$P_0 = \frac{\mathcal{E}}{\Delta t} = \frac{1}{2 \cdot 10^{-9}} = 0.5 \text{ GW}$$

$$P_m = \kappa \iint_M |E|^2 ds$$

$$- \left( \frac{x^2 + y^2}{w_0^2} \right)$$

$$E(z=0, x, y) = E_0 e^{-\frac{x^2 + y^2}{w_0^2}}$$

$$P_0 = \kappa \iint_{-\infty}^{+\infty} e^{-2 \left( \frac{x^2 + y^2}{w_0^2} \right)} dx dy$$

$$= \kappa \int_0^{2\pi} \int_0^{+\infty} e^{-2 \frac{\rho^2}{w_0^2}} \rho d\rho d\varphi$$

$$P_0 = \kappa \cdot 2\pi \left[ -\frac{w_0^2}{4} e^{-2 \rho^2 / w_0^2} \right]_0^{+\infty}$$

$$\boxed{P_0 = \frac{\pi}{2} \kappa w_0^2}$$

Whole power on Moon:  $\mathcal{E} \cdot P_0 = P_{\text{Moon}}$

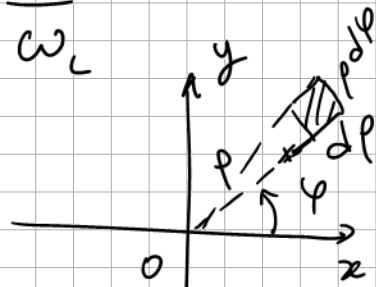
Power on the mirror (without losses due to the atmosphere)  $-(x^2 + y^2)/w_L^2$

$$|E_L(z=L, x, y)| = \frac{E_0}{\sqrt{1 + L^2/\alpha^2}} e^{-\frac{(x^2 + y^2)}{w_L^2}} = \frac{w_0}{w_L} e^{-\frac{(x^2 + y^2)}{w_L^2}}$$

$$w_L^2 = w_0^2 \left(1 + \frac{L^2}{\alpha^2}\right) \Rightarrow \frac{1}{\sqrt{1 + L^2/\alpha^2}} = \frac{w_0}{w_L}$$

$$P_{\text{mirror}} = k \epsilon \cdot \iint_{\text{Mirror}} |E_{\text{Moon}}|^2 dS$$

$$= k \epsilon \frac{w_0^2}{w_L^2} \int_0^{2\pi} \int_0^{\pi} e^{-2\rho^2/w_L^2} \rho d\rho d\varphi$$



$r$  = radius of the mirror

$$P_{\text{mirror}} = 2\pi k \epsilon \frac{w_0^2}{w_L^2} \left[ -\frac{w_L^2}{4} e^{-2\rho^2/w_L^2} \right]_0^r$$

$$P_{\text{mirror}} = \frac{\pi}{2} k \epsilon w_0^2 \left( 1 - e^{-2r^2/w_L^2} \right)$$

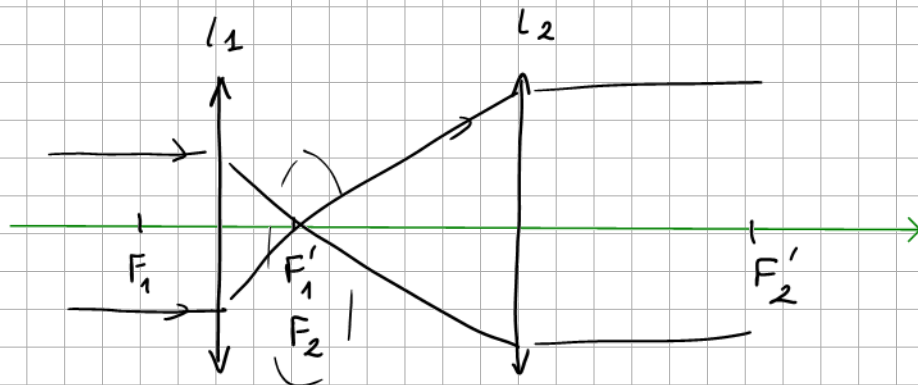
$$P_{\text{mirror}} = \epsilon P_0 \left( 1 - e^{-2r^2/w_L^2} \right)$$

$$r \ll w_L \quad e^{-2r^2/w_L^2} \approx 1 - \frac{2r^2}{w_L^2} \quad \text{Taylor expansion}$$

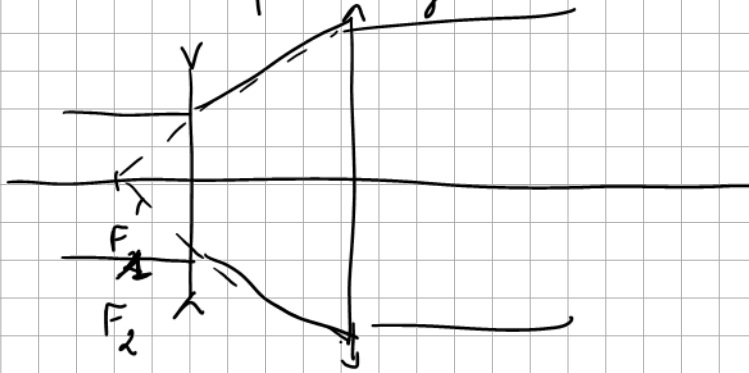
$$P_{\text{mirror}} = \epsilon P_0 \cdot \frac{2r^2}{w_L^2} \Rightarrow 2\epsilon P_0 \left( \frac{r}{w_L} \right)^2 = P_{\text{mirror}}$$

$$\rightarrow \boxed{P_{\text{mirror}} = 0.18 W}$$

3) Telescope to increase  $w_0$  (i.e. increase  $d$  and decrease  $\theta$ )



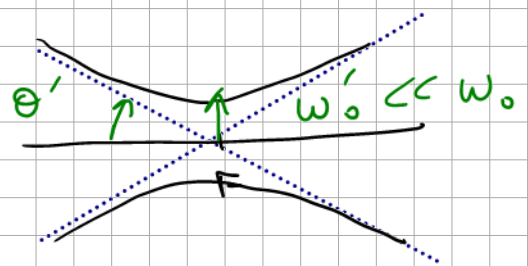
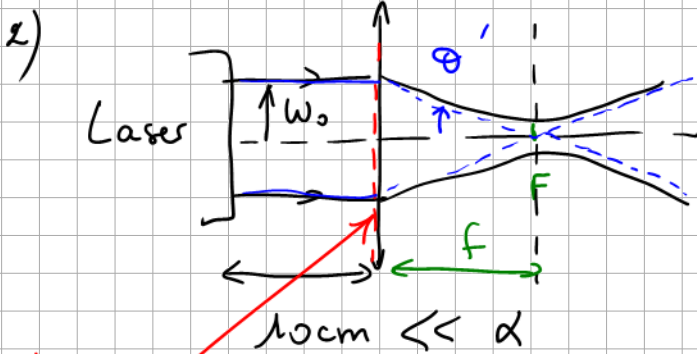
Afocal system



Ex:  $\lambda_0 = 1.55 \mu\text{m}$

1)  $\alpha = \frac{\pi w_0^2}{\lambda} = 2 \text{ m}$

$w_0 = 1 \text{ mm}$



plane wavefront  $\longrightarrow$  focusing of the laser light in the focal plane

a)  $w_0' = 4 \mu\text{m} \Rightarrow \theta' = \frac{\lambda}{\pi w_0'} = 0.123 \text{ rad}$

b)  $\theta' < \text{N.A.} \rightarrow \text{OK}$

c)  $\theta' = \frac{w_0}{f} \Rightarrow f = \frac{w_0}{\theta'} = \frac{\pi w_0 w_0'}{\lambda} = \boxed{18 \text{ mm}}$

### Ex 4

1.  $d = \frac{\pi \omega_0^2}{\lambda} = 38.7 \mu\text{m}$

2.  $z \gg d$

$$E(x, y, z) = j \frac{E_0 d}{z} e^{-\frac{(x^2 + y^2)}{\sigma^2} \frac{z^2}{d^2}} e^{-j \pi \frac{(x^2 + y^2)}{\lambda} \frac{z}{d}} e^{-j k z}$$

3.  $E_L = j \frac{E_0 d}{d} e^{-\frac{(x^2 + y^2)}{\sigma^2} d^2} e^{-j \pi \frac{(x^2 + y^2)}{\lambda} \left( \frac{1}{d} - \frac{1}{f} \right)} e^{-j k z}$

4.  $d = f$

$$E_L = j \frac{E_0 d}{f} e^{-\frac{(x^2 + y^2)}{\sigma^2} f^2} e^{-j k f}$$

5.  $\omega_L = \sigma f = 30 \text{ mm}$

$$d_L = \frac{\pi \omega_L^2}{\lambda} = 2,2 \text{ km} \rightarrow \text{Collimated beam}$$

