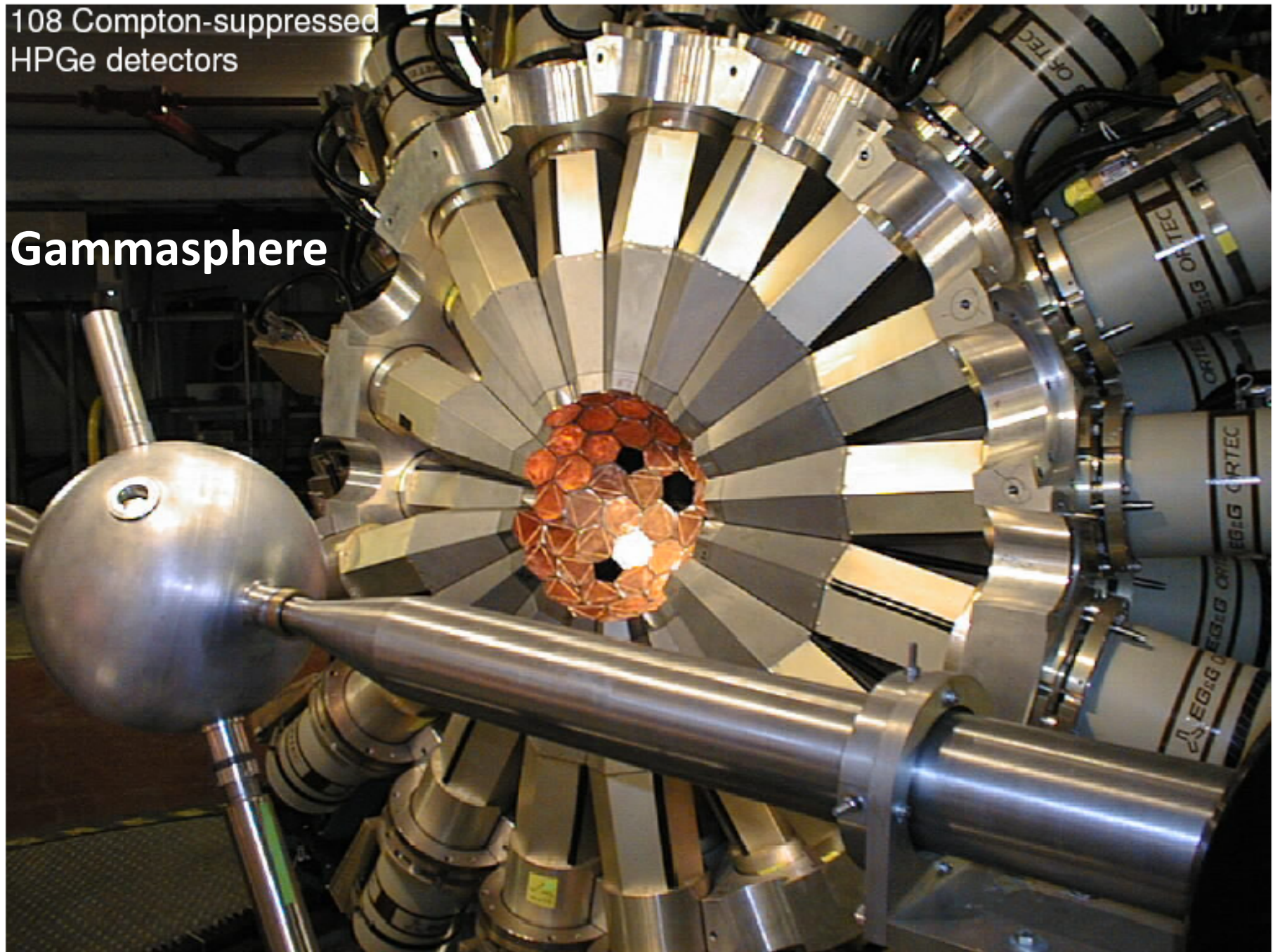
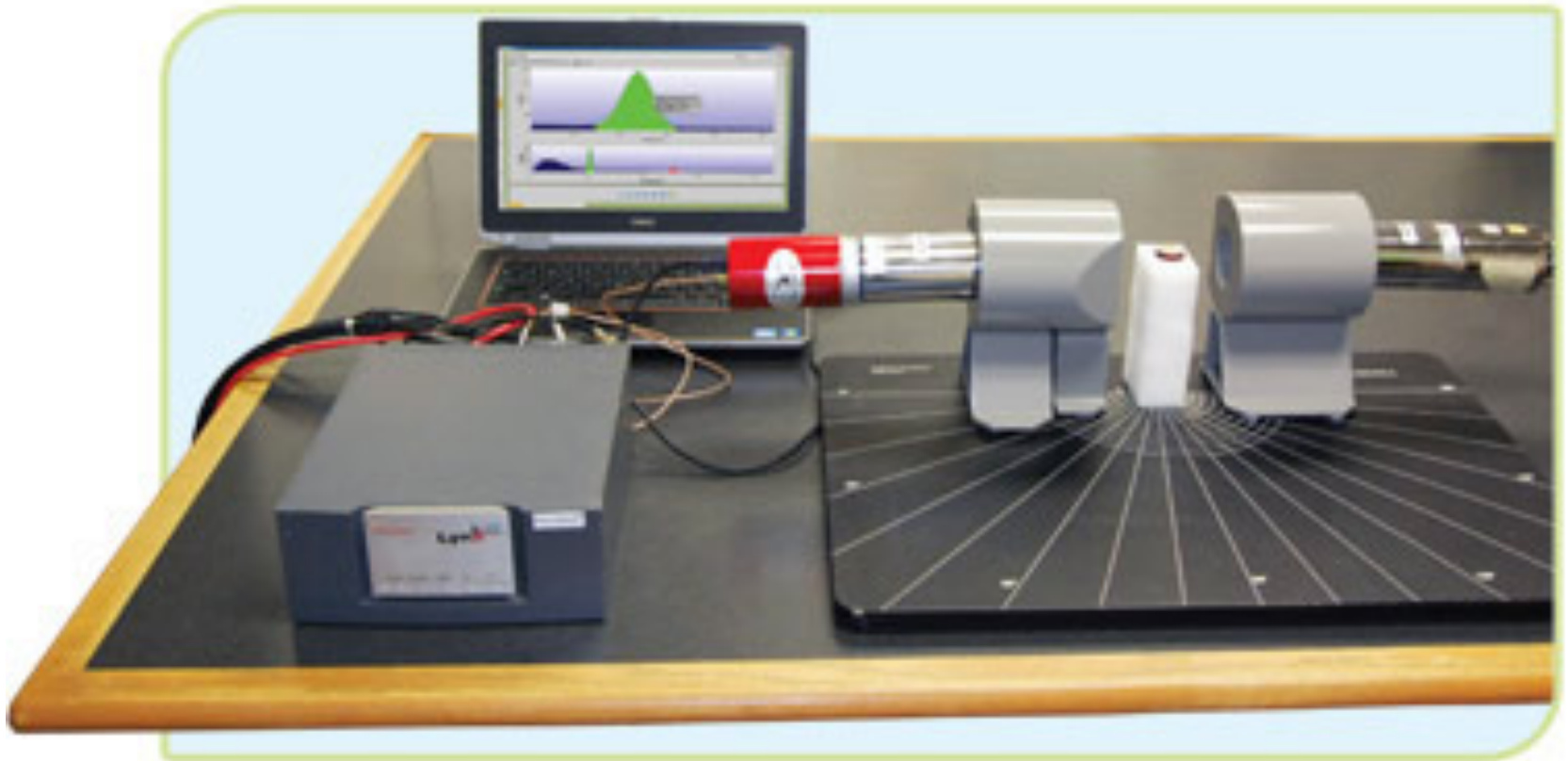


108 Compton-suppressed
HPGe detectors

Gammasphere



2) Gamma Ray Spectroscopy Coincidence Techniques



Principle

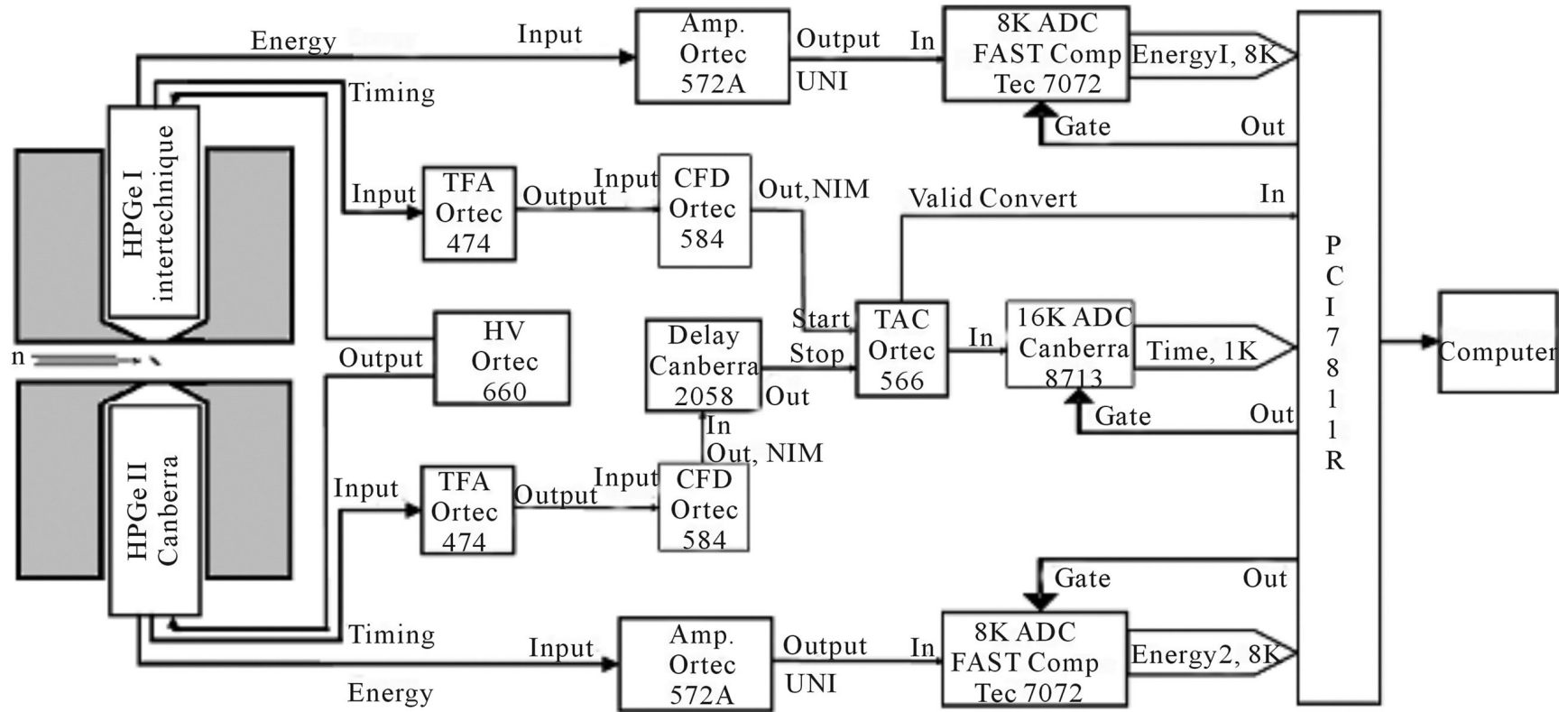
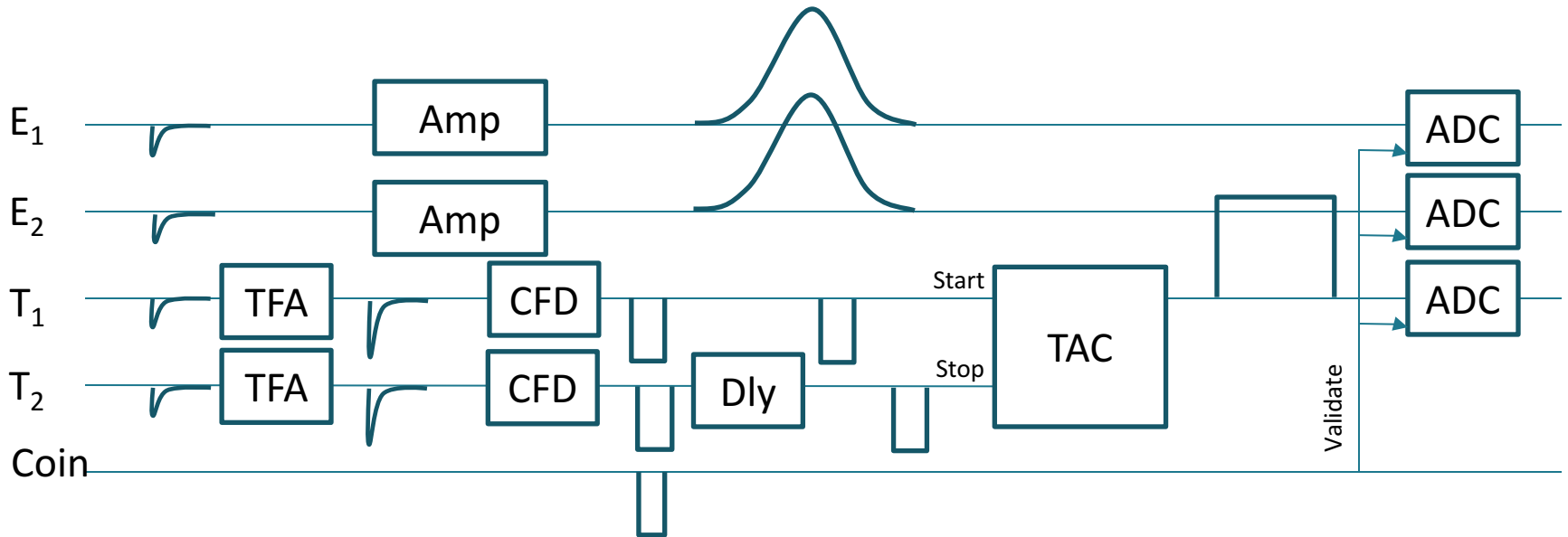
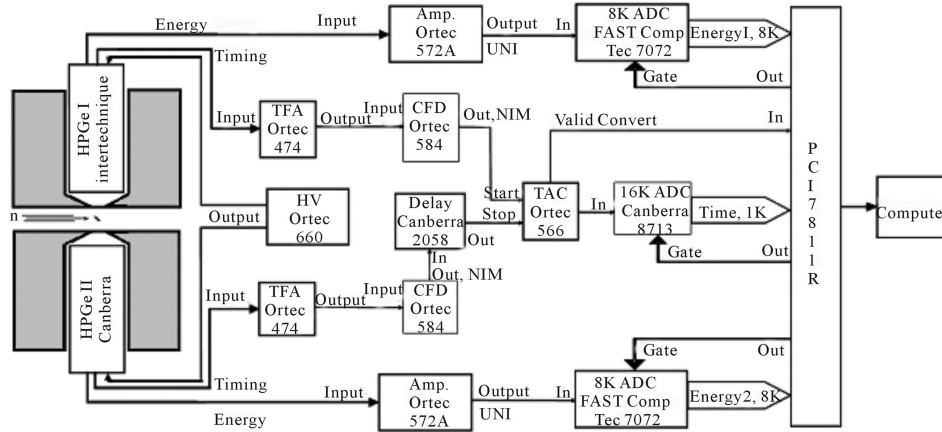


Image taken from WJNST 42520 (14), doi: [10.4236/wjnst.2014.41007](https://doi.org/10.4236/wjnst.2014.41007)

Principle

Image taken from WJNST 42520 (14), doi: [10.4236/wjnst.2014.41007](https://doi.org/10.4236/wjnst.2014.41007)



Here we know we
have a coincidence

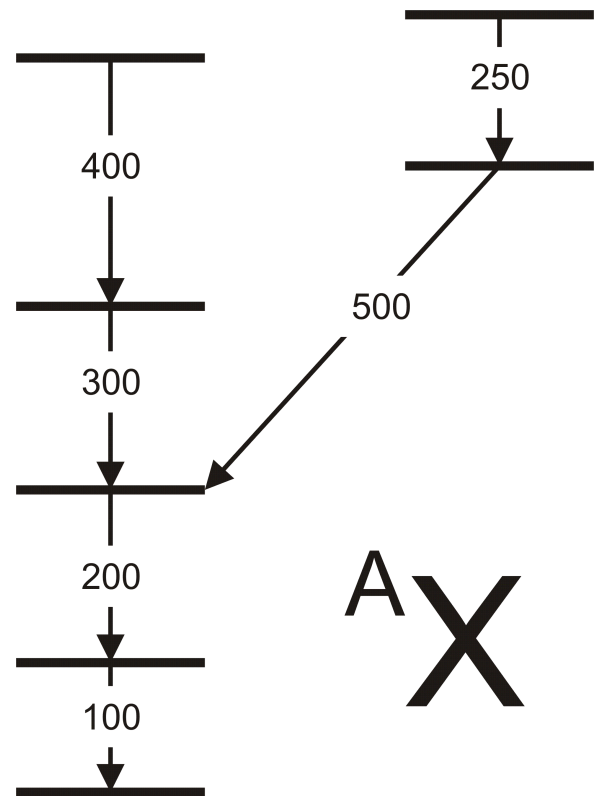
Coincidence timing

- Depends on:
 - the time resolution of the detector
 - the physics (lifetimes of levels)
 - electronic response
- Typical values for Ge detectors
 - Coincidence window of 100 ns

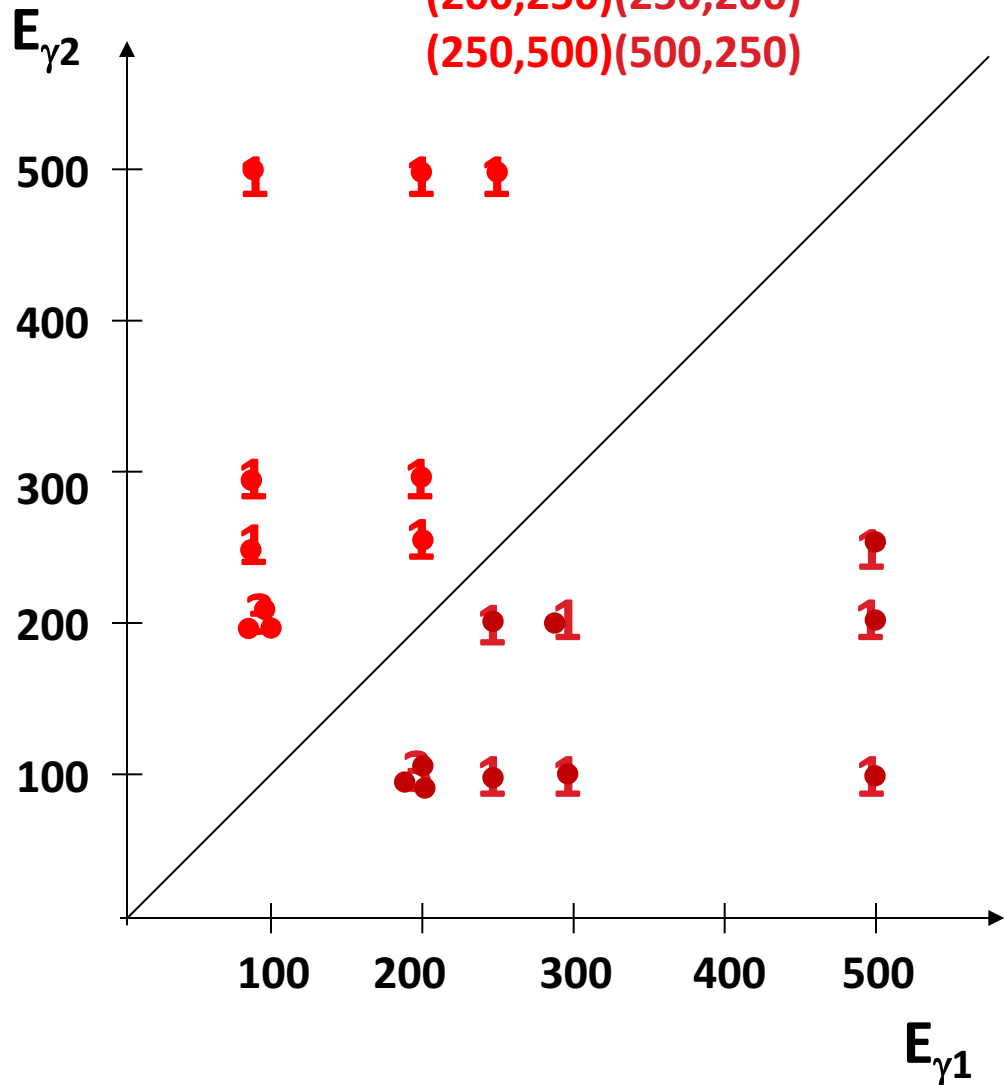
Generating spectra from high-fold coincidence data.

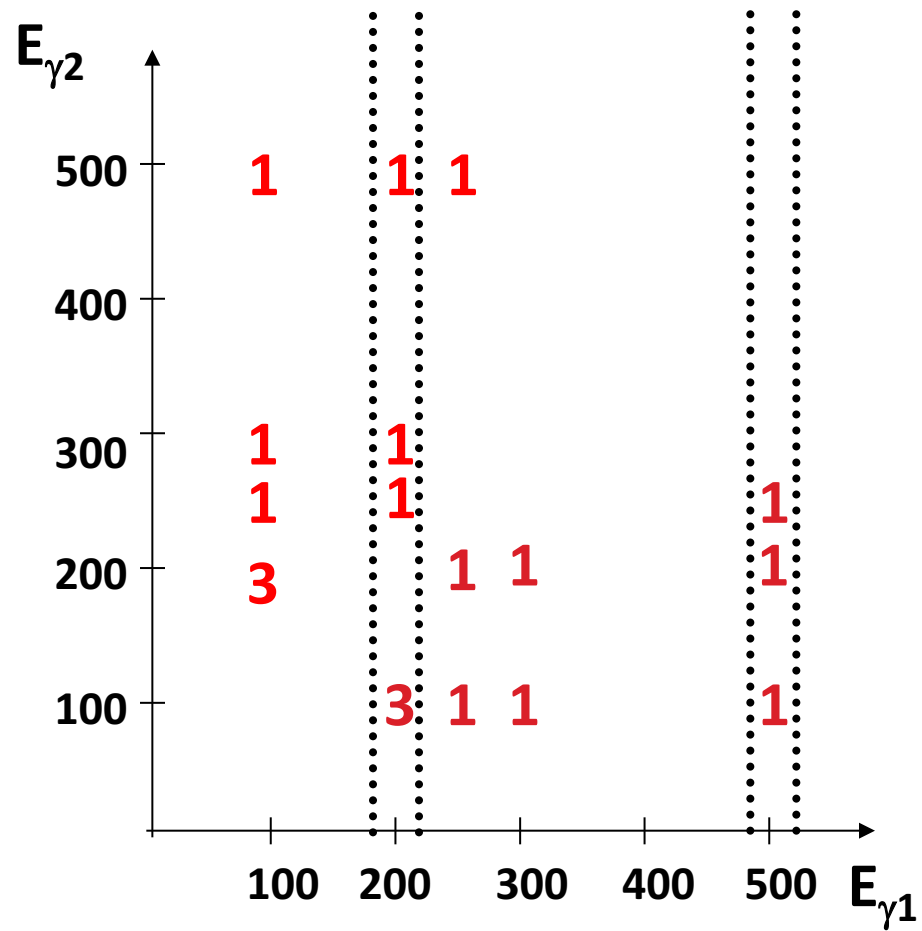
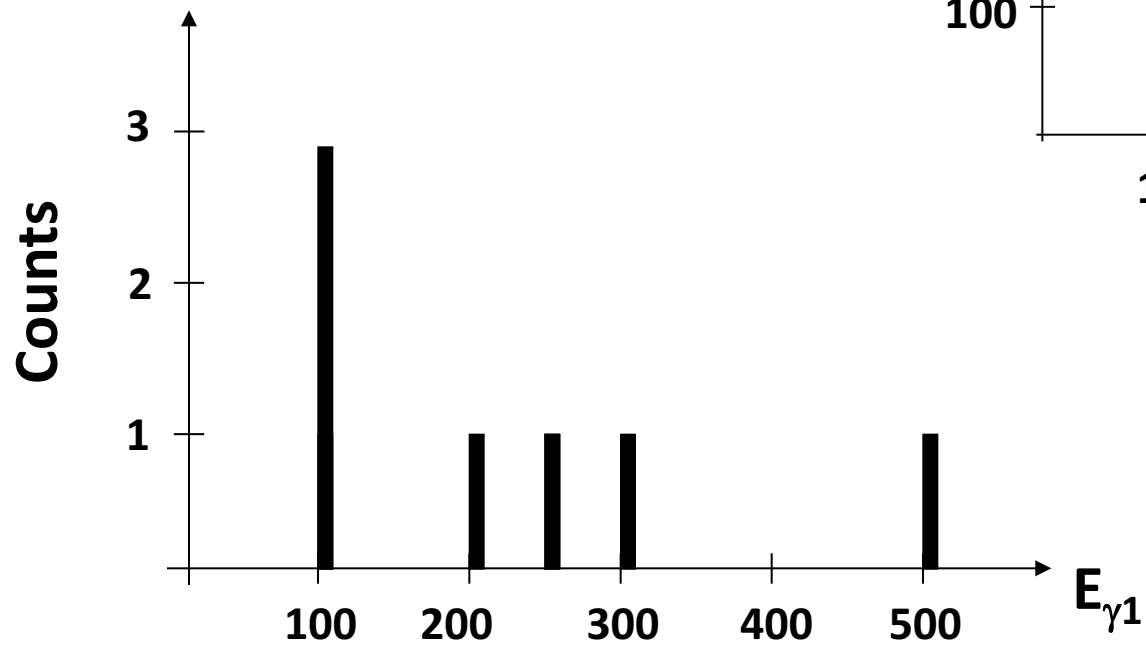
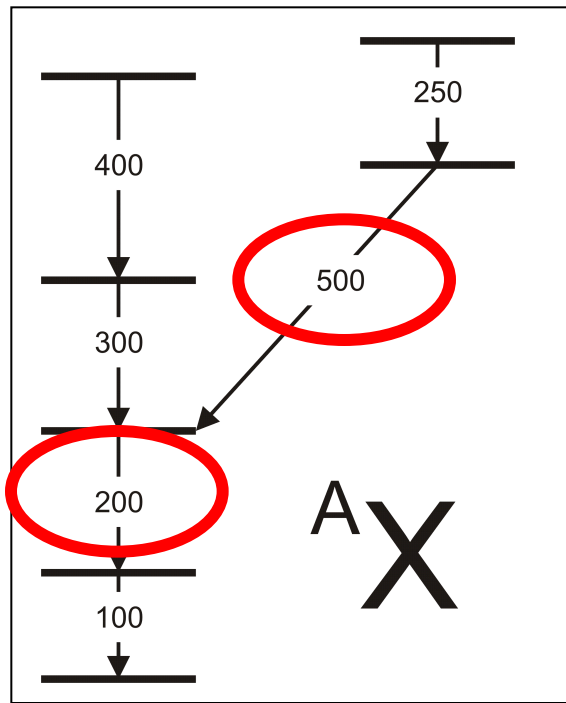
Raw Event 1: 100, 200.
 Raw Event 2: 100, 200, 300.
 Raw Event 3: 100, 200, 500, 250.

Event 1 (100, 200)(200,100)
 Event 2 (100,200)(200,100)
 (100,300)(300,100)
 (200,300)(300,200)

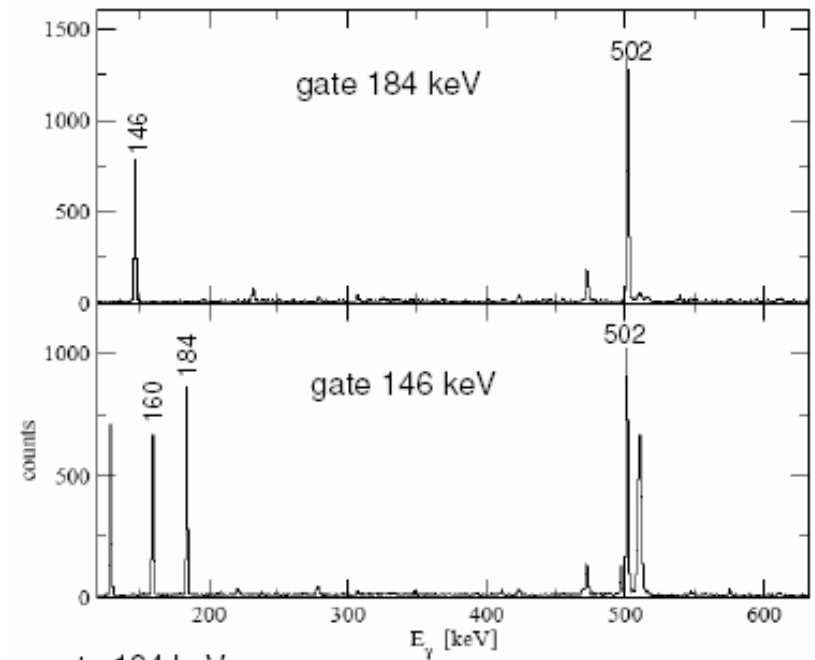
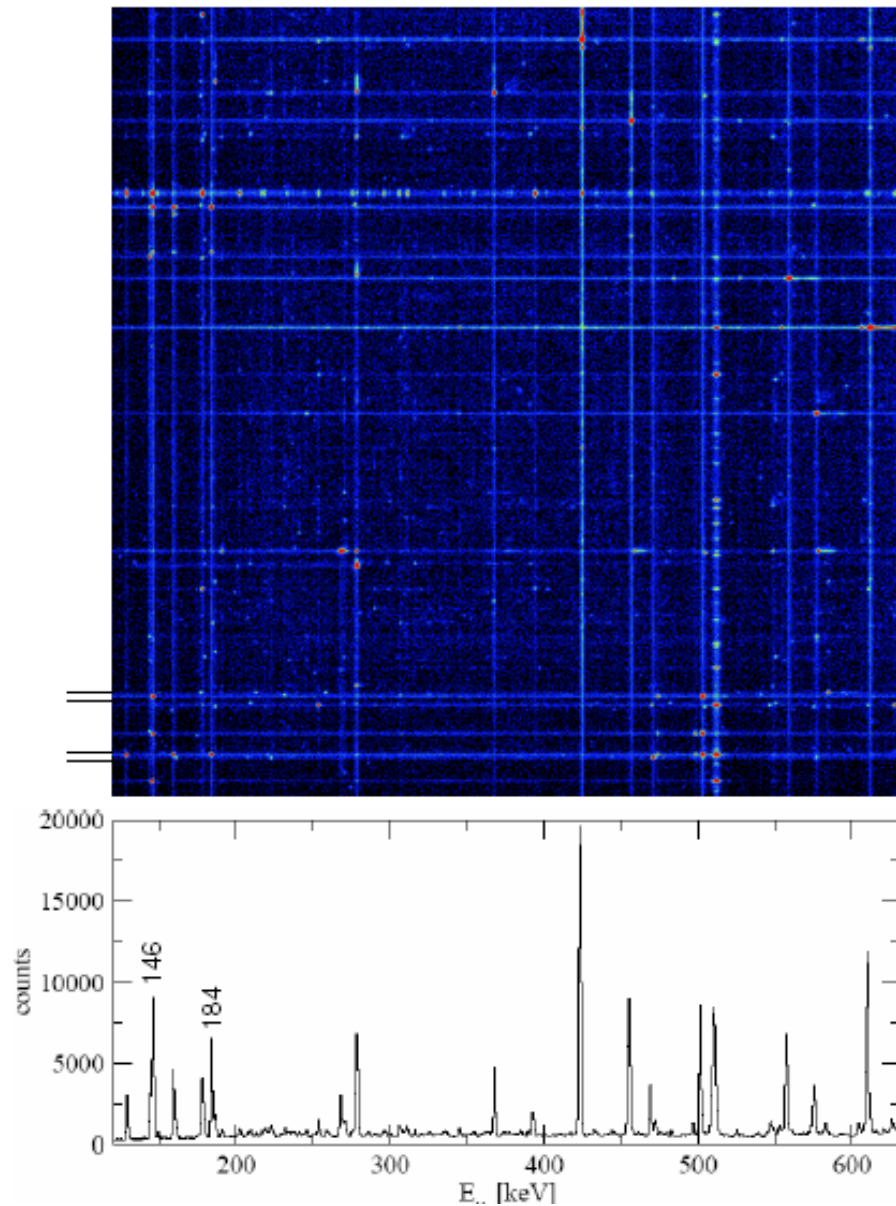


Event 3 (100,200)(200,100)
 (100,500)(500,100)
 (100,250)(250,100)
 (200,500)(500,200)
 (200,250)(250,200)
 (250,500)(500,250)



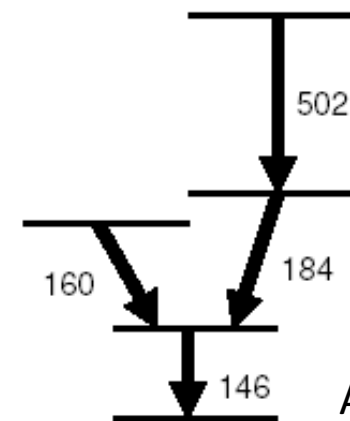


Coincidence technique



gate 184 keV

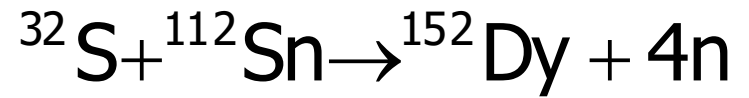
gate 146 keV



A. Goergen

Big example

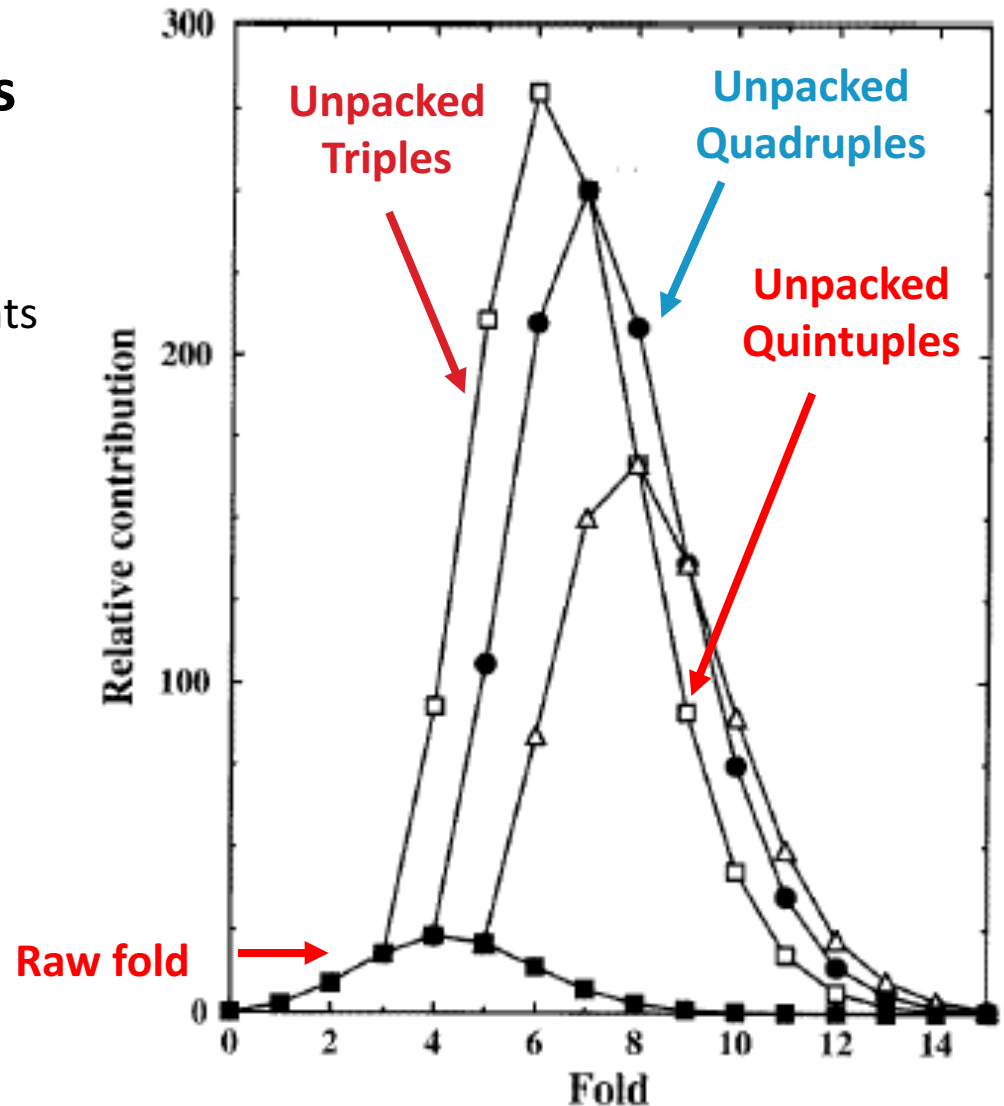
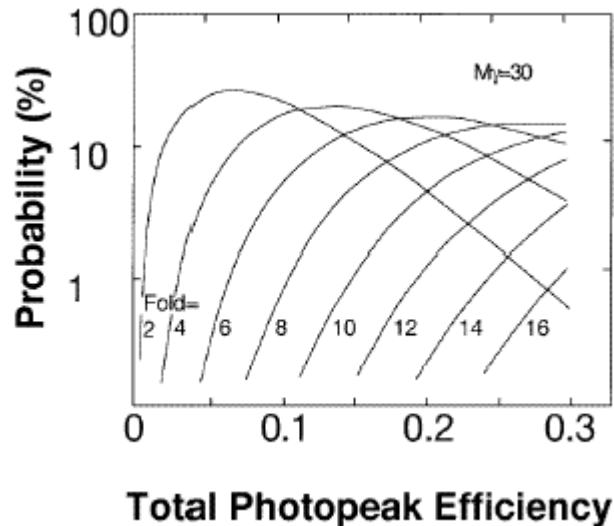
- Homework Challenge!



EUROGAM II

High-fold Data & Multidimensional Analysis

Probability of detecting high-fold events increases with increased photopeak detection efficiency.

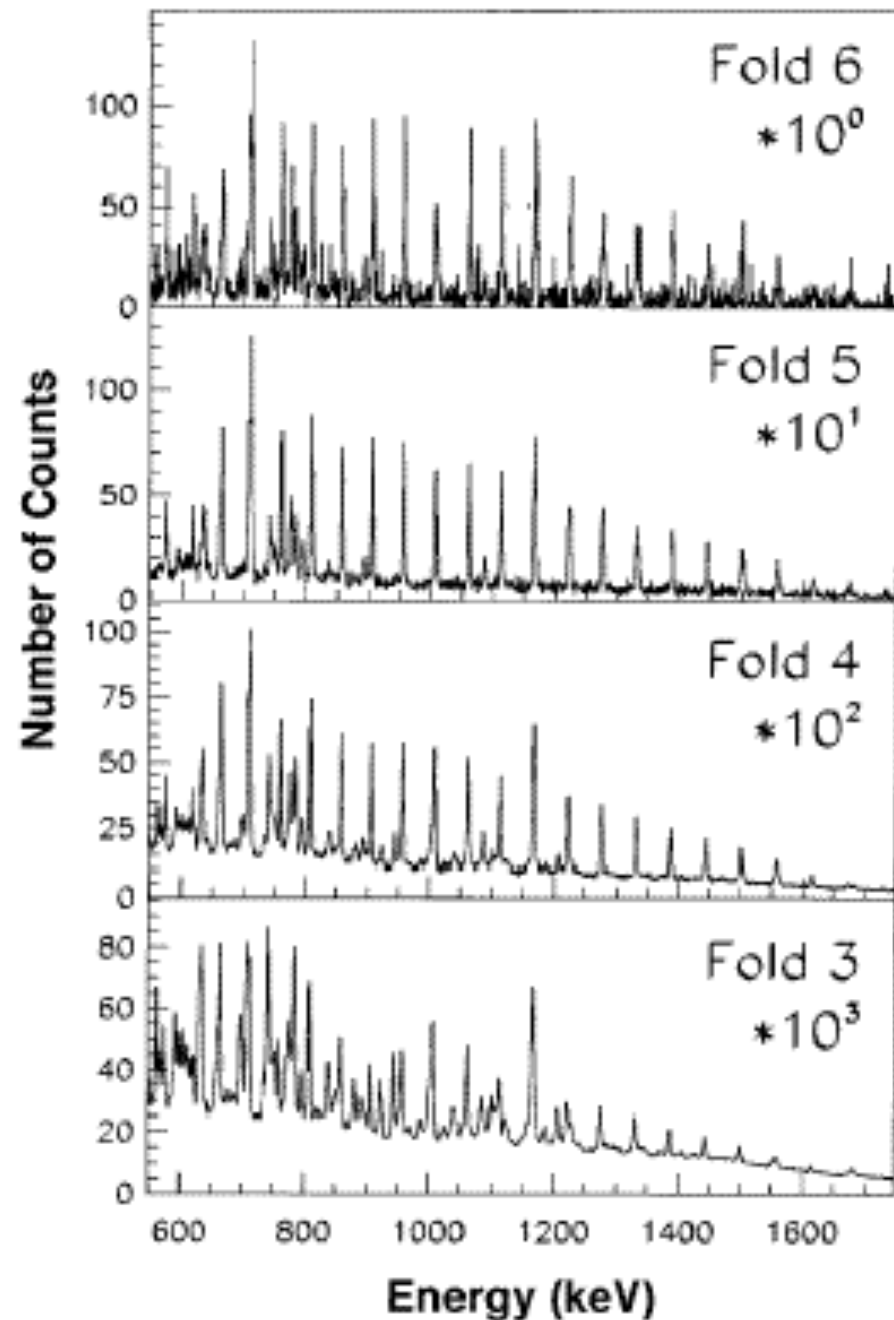


A superdeformed band in ^{149}Gd .

Spectra showing the effect
of 'gating' on coincident transitions.

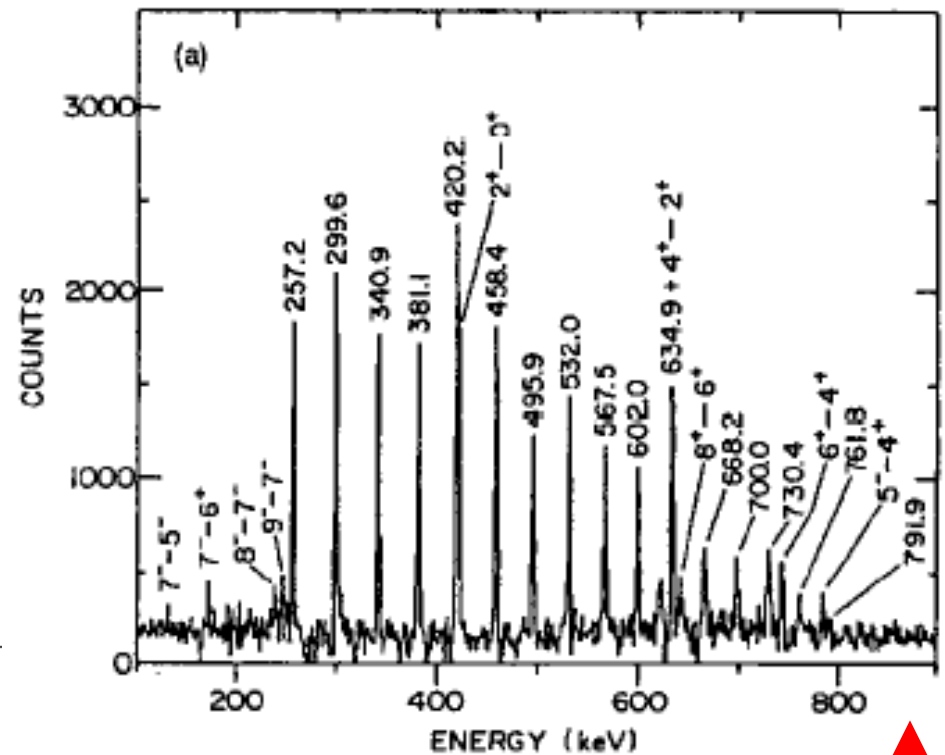
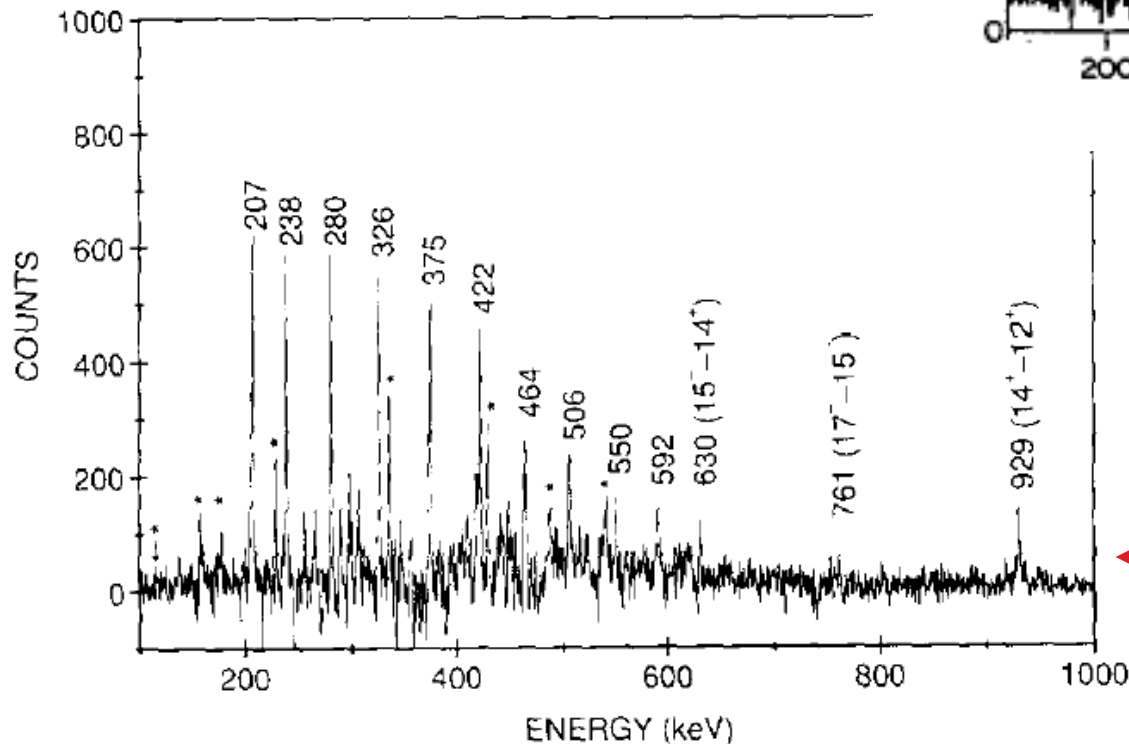
For a review of spectrometer arrays
and their properties see

C.W. Beausang & J.Simpson,
J. Phys. G22 (1996) 527.



Can you tell the difference between these bands?

Differences revealed
by
angular distributions

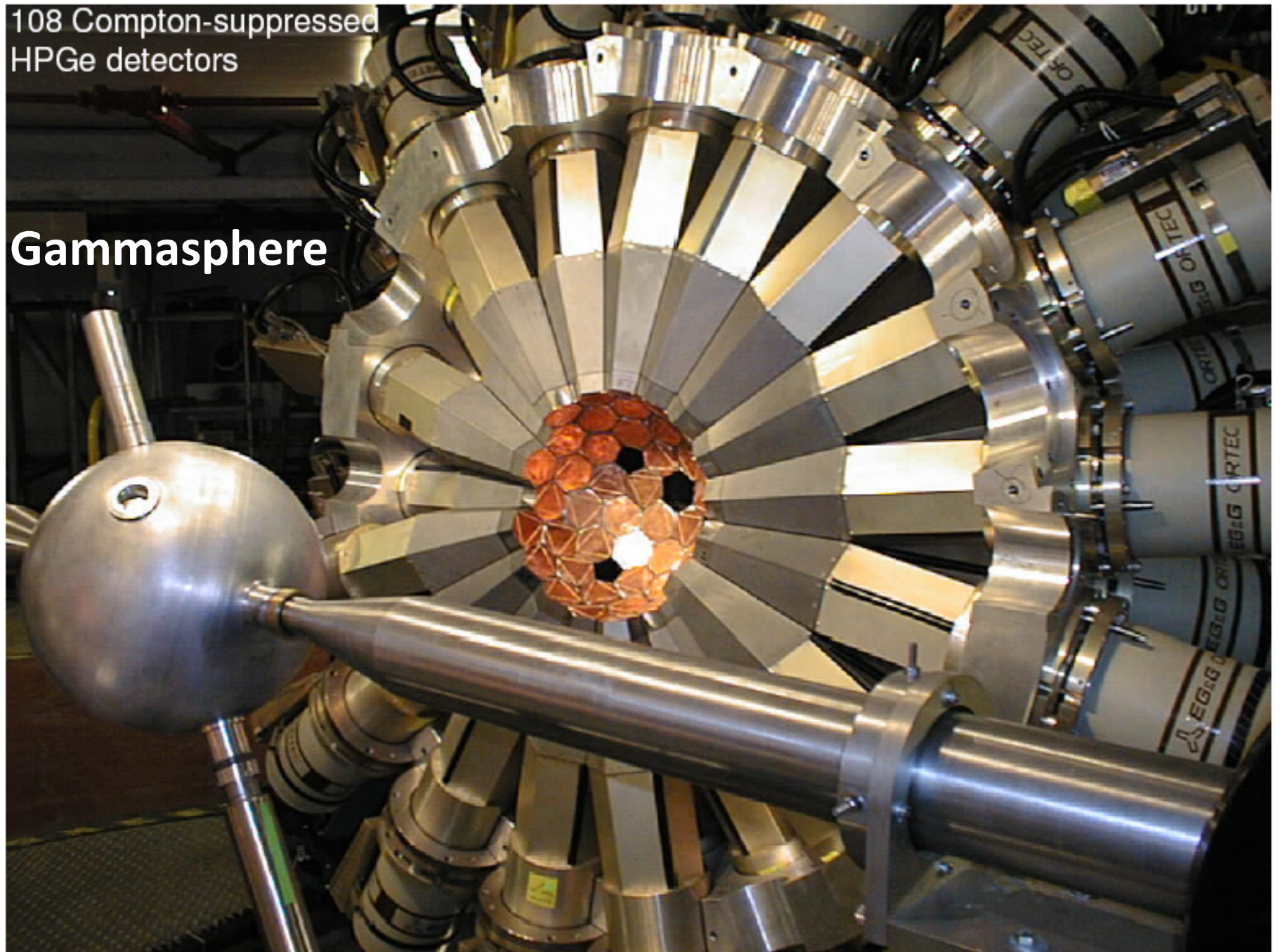


Superdeformed
Band in ^{192}Hg
(Cascade of E2 transitions)

M1 cascade in ^{198}Pb

108 Compton-suppressed
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Aspects of spectrometer array design

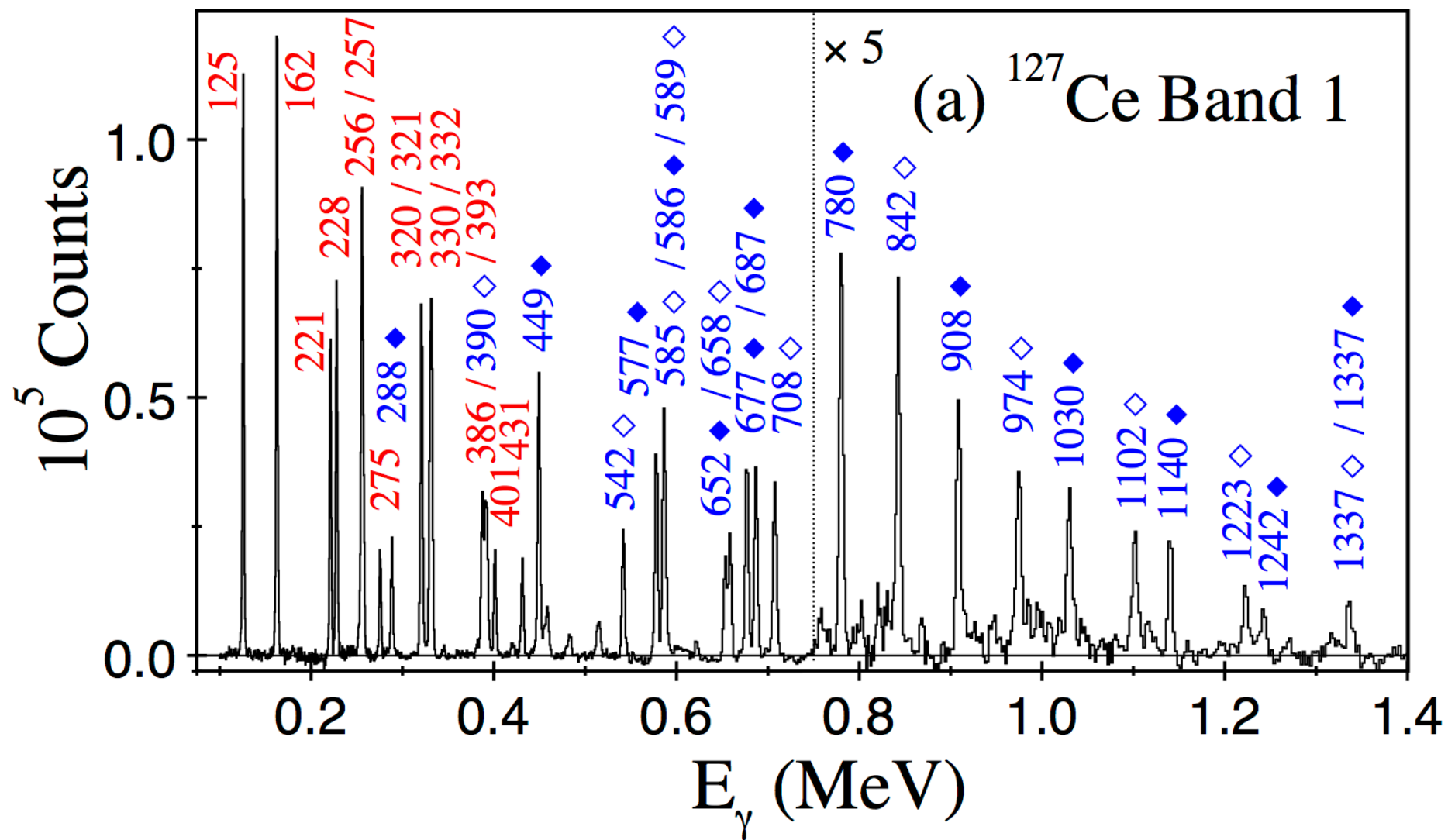
There are many gamma-ray spectrometers in use around the world.

Requirements for a good spectrometer.

- (a) High photopeak detection efficiency.
- (b) Excellent resolution.
- (c) Good peak-to-total signal.
- (d) High granularity.

Spectrometers are designed for specific reaction mechanisms / applications.

e.g. High gamma-ray multiplicities and large Doppler shifts require high granularity.



Sensitivity

Average separation of consecutive γ rays in a cascade.

Resolving power
(as large as possible)

$$R = \left(\frac{SE_{\gamma}}{\Delta E_{\gamma}^{\text{final}}} \right) PT$$

Peak-to-Total ratio

FWHM of detector
resolution for
 γ rays in spectrum

$$\left(\frac{N_P}{N_B} \right)_F$$

$$= \alpha_0 (0.76R)^F$$

Fold

Limit of observation (array)

Final peak-to-total ratio with F-1 gating condition.

Factors effecting resolution

The major factors affecting the final energy resolution (FWHM) of the array at a particular energy are as follows.

$$\Delta E_{\gamma}^{\text{final}} = \left(\Delta E_{\text{Int}}^2 + \Delta E_{\text{Open}}^2 + \Delta E_{\text{Ang}}^2 + \Delta E_{\text{Vel}}^2 \right)^{\frac{1}{2}}$$

ΔE_{Int} - The intrinsic resolution of the detector system.

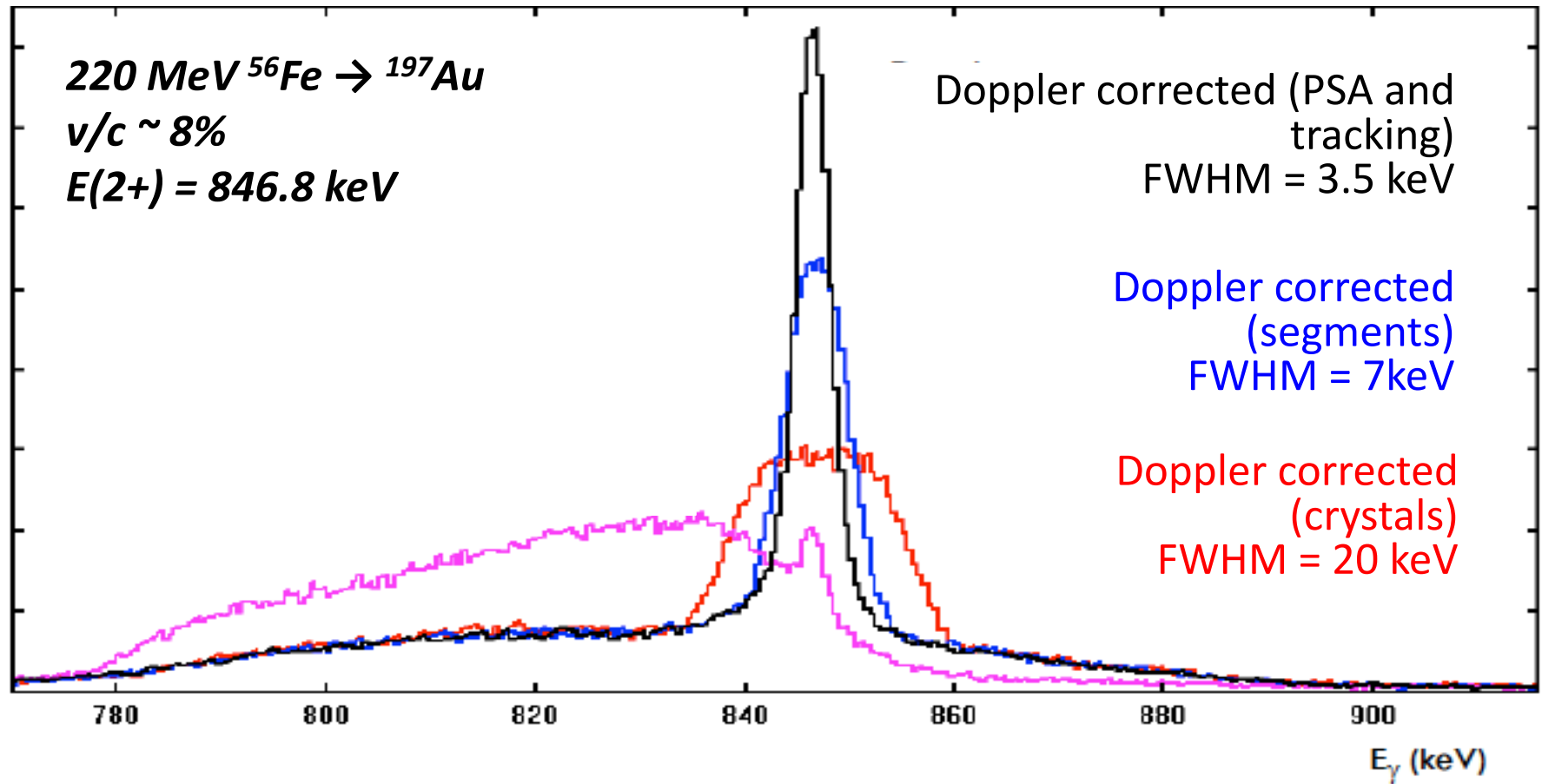
This includes contributions from the detector itself and the electronic components used to process the signal.

ΔE_{Open} - The Doppler broadening arising from the opening angle of the detectors.

ΔE_{Ang} - The Doppler broadening arising from the angular spread of recoils in the target.

ΔE_{Vel} - The Doppler broadening arising from the velocity (energy) variation of the recoils across the target.

The main contribution to poor resolution arise from **Doppler effects**. Require detector granularity!



Tutorial question: Derive Doppler broadening formula for Ge detector at 90° to beam direction.

Resolving Power through the ages

