

A Temporal Analysis of Cassini/VIMS Limb Brightness Profiles of Titan from 2004-2017

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We analyze the evolution of brightness profiles from Cassini Visual and Infrared Mapping Spectrometer (VIMS) spectral mapping cubes to gain insights into the vertical structure of Titan's atmosphere. Using low-phase angle observations from 2004-2017, we observe changes in the shape of limb brightening profiles with regard to wavelength and time. When densely sampled over a temporal period, these profiles can independently track the development of different atmospheric layers on a global level. We show that Titan's limb brightening depends on the wavelength and varies as a function of time. Moreover, a comparison of the brightness profiles over the northern and southern hemispheres supports existing analyses of atmospheric asymmetry behavior. By using a combination of both limb and asymmetry observations, we are able to develop a greater understanding of the underlying seasonal dynamics of Titan's haze, specifically its vertical structure.