

## CWs

(\*)

### SMBH intro

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Supermassive black holes are key components in the assembly and evolution of cosmic structure

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Structure formation is thought to have occurred via a hierarchical clustering process

→ Small galaxies merge together to form larger ones ad infinitum.

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Most galaxies are thought to harbour a massive black hole at their centre.

→ The evolution of black-hole + host is symbiotic.

→ There are well-publicised relationships between the mass of the black-hole and galaxy properties

e.g.  $M_{BH} - \sigma$ ,  $M_{BH} - M_{BULGE}$   
→  $M_{BH}$  (of interest to us)  $\sim 10^7 - 10^{10} M_{\odot}$

\*

So, when galaxies merge together, so do the resident black-holes.

\*

The individual black-holes sink together by dynamical friction (from stellar distribution) within the common galactic merger remnant.

\*

Dynamical friction not efficient enough to bring black-holes close enough together to sink into a BHs and merge within a Hubble time → "final parsec problem" [Chiara will discuss]

\*

Other mechanisms in the galactic nuclear environment can couple to the binary, evolve it quickly.

\* The binary is hardened if the orbital velocity of the binary is  $\sim$  the same (or greater) than the velocity dispersion of ambient star distribution.

\* At this point the evolution of the binary is driven by the emission of GWs.

→ We have a distribution of masses with a non-zero quadrupole moment.

→ The masses orbiting each gives an accelerating quadrupole moment.

→ Assuming circular binaries,  $f_{\text{GW}} = 2 f_{\text{orb}}$

→ With mpc orbital separations,  $f_{\text{GW}} \sim \text{nHz}$ .

## (\*) Astrophysical distribution

\* The distribution of sources in mass + redshift which are resolvable is different from the population dist. of the underlying population → selection effects

\* Masses of resolvable sources are weighted to higher mass  $\sim 10^9 M_{\odot}$

\* Redshift peaked @  $z \sim 1$  ... favours close binaries, but if  $\frac{dN}{dz} \sim \text{const}$  then the distribution favours higher redshift.

\* The orbital frequency distribution favours  $f_{\text{orb}} \geq 10^{-8} \text{ Hz}$ , where the time spent by each binary in each frequency bin is smaller, so stochasticity of BG breaks down.

## ⑧ Gravitational waveform

\* For our target population, systems are pretty far from merger,  $\rightarrow f_{\text{iso}} \sim 4.4 \times 10^{-6} (M/10^9 M_\odot)^{-1} \text{ Hz}$

\* We are in the early adiabatic inspiral phase of the binary merging process.

$\rightarrow$  System emits GWs, loses energy from the orbit,  
 $\rightarrow$  Evolves smoothly through geodesics.

\* Newtonian leading order is a sufficient description of the waveform.

$$\rightarrow v/c \sim 1.73 \times 10^{-2} \left( \frac{M}{10^9 M_\odot} \right)^{2/3} \left( \frac{f}{50 \text{ nHz}} \right)^{1/3}$$

\* Also, frequency evolution during observation time ( $\sim 10-20$  years) is v. small.

$$\rightarrow \Delta f \sim \dot{f} T \sim 0.05 M^{5/3} f^{11/3} T_{10} \text{ nHz}$$

[Mingarelli et al. 2012]

\* Spin enters phase evolution at  $1.5 \text{ PN } (v^3) + \text{orbital plane precession}$

⑧  $\Rightarrow$  So for circular binaries, our waveform is a simple sinusoid described by leading order in  $(v/c)$ .

$$\underline{h}(t, \hat{\alpha}) = h_+(t, \hat{\alpha}) \underline{e}_+^T(\hat{\alpha}) + h_\times(t, \hat{\alpha}) \underline{e}_\times^T$$

$$\underline{z}(t, \hat{\alpha}) = \frac{\delta v}{2} = \underline{d} : \underline{h}$$

$$\therefore z(t, \hat{\Omega}) = h_+ \underbrace{(d : \underline{e}^+)}_{F^+(\hat{\Omega})} + h_x \underbrace{(d : \underline{e}^x)}_{F^x(\hat{\Omega})}$$

⇒ The residuals ~~are~~ (or delay-signal, or whatever) are the cumulative redshift of the pulse arrival rate over observation time.

$$\Rightarrow S_0, \quad s(t) = \int_{t_0}^t z(t') dt'$$

$$s(t) = F^+ s_+(t) + F^x s_x(t)$$

$$* \quad s_+(t) \propto \frac{M^{\frac{8}{3}}}{D_L \omega^{\frac{1}{3}}} \left[ (1 + \cos^2 i) \cos 2\varphi \sin 2\omega t + 2 \cos i \sin 2\varphi \cos 2\omega t \right]$$

$$s_x(t) \propto \frac{M^{\frac{8}{3}}}{D_L \omega^{\frac{1}{3}}} \left[ (1 + \cos^2 i) \sin 2\varphi \sin 2\omega t - 2 \cos i \cos 2\varphi \cos 2\omega t \right]$$

\* Pulsar term

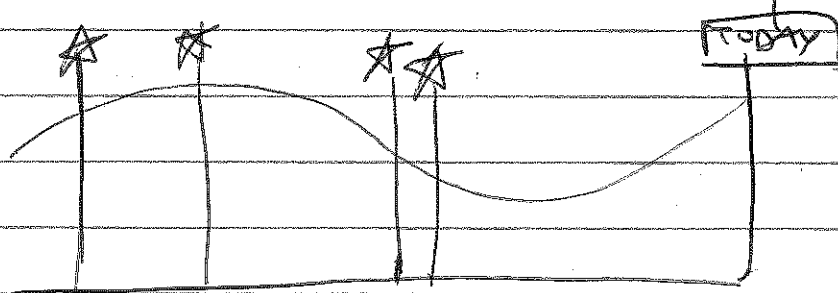
⇒ I said that the redshift was just the projection of the metric tensor along the detector's sensitivity.  
 ⇒ But really  $z(t) = \underline{d} : \Delta \underline{h}$

- \* The signal is really the integrated influence of the GW over the entire photon path.
- \* The result is an influence from GW at the Earth and at each pulsar.

$$S(t) = F^+ \Delta S_+ + F^x \Delta S_x$$

$$\rightarrow \Delta S_{t,x} = S_{t,x}^e - S_{t,x}^p$$

signature of the binary in the past.



- $\rightarrow$  ~~separate~~ separate  $(M, D_L)$  + improve source localization.
- $\rightarrow$  Justin will talk about how the pulsar term is a blessing and a curse.

## \* Extension to eccentricity

- The environmental mechanisms to drive the binary components close together (and thus mitigate the "final parsec problem") can also leave the binary with significant eccentricity.
- Eccentricity leads to the GW ~~being~~ <sup>characteristic</sup> having a distribution of frequencies which are integer multiples of  $f_{\text{orb}}$ .
- Taylor et al. [2015] gives closed form signal model assuming (i) no evolution over observation time (ii) no evolution of the direction of periastron passage.
- At high eccentricity, instead of the dominant GW frequency being " $2 \times f_{\text{orb}}$ ", it is  $f_{\text{orb}}$  itself.
- Eccentricity is a double-edged sword in terms of detection prospects
  - pushes systems below band into band.
  - but pushes systems already in sensitive region out of band.



## EM counterparts

### \* Burke-Spencer (2013)

- show plot.
- Radio AGN
- periodic jet flux.
- X-shaped sources.
- circumbinary disk activity.