## Problem Set 4

# 1 Radiative Forcing with CO<sub>2</sub>

Carbon dioxide is one of the major gases driving the anthropogenic greenhouse effect (the dominant greenhouse gas is  $H_2O$ , but it is not anthropogenic). One of the most impactful transition in CO2 for our climate is the 15  $\mu$ m line, which corresponds to the bending oscillation of the linear CO2 molecule. The main line (there are rotational/vibrational branches that we'll discuss in a later class) has an Einstein A of about  $A \sim 10 \text{ s}^{-1}$ .

#### 1.1

Assume the gas temperature of the CO2 is about the same as the Earth's blackbody (255 K), and plot the line profile of this transition with both Doppler and intrinsic components. Use a logrithmic y axis.

#### 1.2

Model the Earth's upgoing emission as a 255 K blackbody. Plot this spectrum, and overlay on it the spectrum including absorption from the 15  $\mu$ m line of CO2. At line center, the optical depth of this transition in our atmosphere is about  $\tau \sim 3$ , so set the number density of CO2 accordingly.

### 1.3

Calculate the flux of emission absorbed by CO2 over the frequency interval where this transition is optically thick ( $\tau > 1$ ). Compare this to the amount of energy absorbed by CO2 over the entire blackbody spectrum. What fraction of the total energy absorbed is in the optically thick regime?

#### 1.4

Assume that, for all the flux absorbed by CO2, half of it is radiated upward (i.e. it escapes the atmosphere) and half of it is radiated downward. The flux radiated downward is effectively subracted from the total flux radiated off by the Earth's blackbody. To radiate off the same about of flux as it did without CO2, by how many degrees must the blackbody of the Earth increase? This is called radiative forcing, and you have now roughly calculated the contribution of the 15  $\mu$ m line of CO2 to the greenhouse effect!