

# COMS 30115

## Classic Radiosity

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March 18th, 2018

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# Introduction

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# Last Time

- Hemisphere
- BRDFs
- Rendering Equation



- The BRDF tells us how to represent reflected radiance in terms of incoming radiance

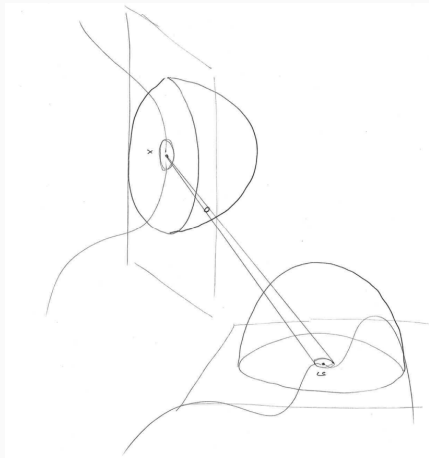
$$L_r(x \rightarrow \Theta) = \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi$$

- Putting these together gives the rendering equation,

$$\begin{aligned} L(x \rightarrow \Theta) &= L_e(x \rightarrow \Theta) \\ &+ \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi \end{aligned}$$

- We want to solve for the radiance  $L(\forall x \rightarrow \forall \theta)$  for the whole scene, why is this complicated?

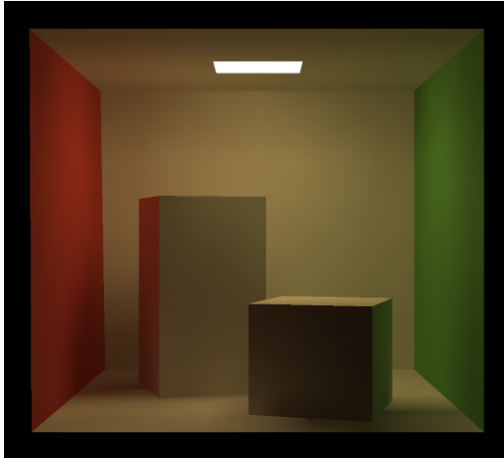
# Rendering Equation



$$L(x \rightarrow \Theta) = L_e(x \rightarrow \Theta) + \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi$$

- Area formulation of the rendering equation
- Classic Radiosity

# Cornell Box





- Comparing two Global Illumination Models
- Radiosity
- Equation Compendium

# Transport Problem

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# Transport Problem

- Our eyes sensitive to radiance
- Want to compute radiance from all “surfaces” & solid angles
- Flux is emitted from light sources
- The rendering equation tells us how this flux is distributed in the scene

$$L(x \rightarrow \Theta) = L_e(x \rightarrow \Theta) + \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi$$

$$\begin{aligned}L(x \rightarrow \Theta) &= L_e(x \rightarrow \Theta) \\&\quad + \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi \\&= L_e(x \rightarrow \Theta) + \mathcal{T}(L(x \rightarrow \Theta))\end{aligned}$$

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$$\mathcal{T}(L(x \rightarrow \Theta)) = \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi$$

$$\mathbf{x}^T \mathbf{a} = \sum_{i=1}^D x_i \cdot a_i, \quad \langle f, g \rangle = \int f(x)g(x)dx$$

- Rendering Equation formulates recursive transport
- Transport is a linear operator in radiance

$$\langle F^{\rightarrow}, G^{\leftarrow} \rangle = \int_A \int_{\Omega} F(x \rightarrow \Theta) G(x \leftarrow \Theta) \cos(\mathbf{n}_x, \Theta) d\omega_{\Theta} dA$$

## Inner product formulation

- Transport problem can be written as the inner-product of two functions



$$L(x \rightarrow \Theta) = L = L_e + \langle T, L \rangle$$

$$\begin{aligned} L(x \rightarrow \Theta) &= L = L_e + \langle T, L \rangle \\ &= L_e + \langle T, L_e \rangle + \langle T, TL \rangle \end{aligned}$$

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- Adjoint operator:

$$\forall F, G \in V : \langle \mathcal{O}_1 F, G \rangle = \langle F, \mathcal{O}_2 G \rangle$$

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- *Think raytracer!*
- Computing light transport back to light source or from light source are adjoint operators
- Remember dual formulation of raytracing

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5. Different algorithms approach this in different way

# Transport Problem

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3. What is the set of points & solid angles?
  - what is the perceptual importance of each part of the geometry?
4. How do we perform series expansion?
  - How does normal raytracing do this?
  - Perceptual importance?
5. Different algorithms approach this in different way
6. We know that energy is balanced in the scene

# Radiosity

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- Initially developed for solving heat transfer
- Formulation for Computer Graphics<sup>1</sup>
- Game engine Geomerics<sup>2</sup>

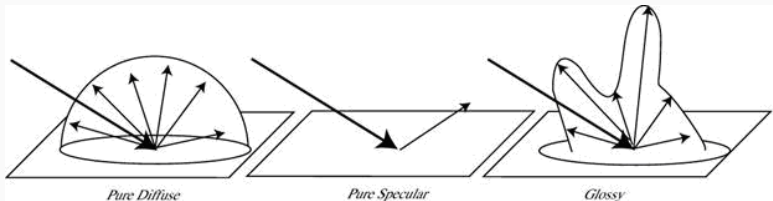
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<sup>1</sup>Goral, C. M., Torrance, K. E., Greenberg, D. P., & Battaile, B. (1984).

<sup>2</sup><http://www.geomerics.com/>



# Radiosity



- Diffuse reflections means that reflection is the **same** in all directions
- This means radiance is the same in each direction
- This simplifies the problem rendering problem hugely
- Classic Radiosity assumes all surfaces are perfectly diffuse

**Radiosity** the **exitant** radiant power on a surface per unit surface area

- $B(x) - [Watt/m^2]$

**Radiance** the radiant power per unit projected area per unit solid angle

- $L(x) - [Watt/steradian \cdot m^2]$

If we are assuming each surface to be diffuse we do not need the angle

$$B_i = \frac{1}{A_i} \int_{S_i} \int_{\Omega_x} L(x \rightarrow \Theta) \cos(\mathbf{n}_x, \Theta) d\omega_{\Theta} dA_x$$

$$L(x \rightarrow \Theta) = L_e(x \rightarrow \Theta) + \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi$$

- Angles does not matter if we have perfectly diffuse surfaces
- Write the rendering equation in terms of surfaces instead

$$\begin{aligned}L(x \rightarrow \Theta) &= L_e(x \rightarrow \Theta) \\&+ \int_{\Omega_x} f_r(x, \Psi \rightarrow \Theta) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi \\&= L_e(x) + \int_{\Omega_x} f_r(x) L(x \leftarrow \Psi) \cos(\mathbf{n}_x, \Psi) d\omega_\Psi\end{aligned}$$

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- If we have only perfectly diffuse surfaces we don't need full BRDF

$\rho(x)$  – reflectivity at point  $x$

- The light that arrives at point  $x$  from surface  $y$  only depends on visibility

$K(x, y)$  – how much of light leaving point  $y$  arrives at  $x$

$$\frac{1}{A_i} \int_{S_i} B(x) dA_x = \frac{1}{A_i} \int_{S_i} B_e(x) dA_x + \frac{1}{A_i} \int_{S_i} \rho(x) \int_S K(x, y) B(y) dA_y dA_x$$

**Assumption** the radiosity is **constant** across a patch

$$\forall x \in S_i : B(x) = B_i$$

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$$B_i = B_{ei} + \rho_i \sum_j \frac{1}{A_i} \int_{S_i} \int_{S_j} K(x, y) B(y) dA_y dA_x$$

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- Discretise the scene into patches
- $K(x, y)$  constant over patch
- $F_{ij}$  is depends only on the geometry
- referred to as a **form factor**

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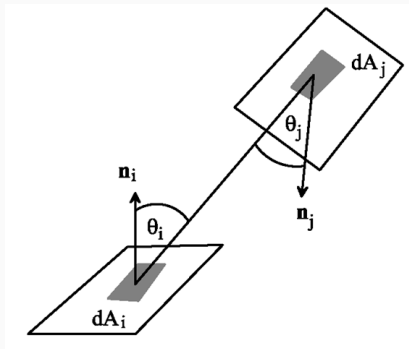
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- Discretise the scene into patches
- $K(x, y)$  constant over patch
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# Form Factor



- "How large portion of the *view* from *i* is blocked by *j*"

$$B_i = B_{ei} + \rho_i \sum_j F_{ij} B_j$$

$$B_{ei} = B_i - \rho_i \sum_j F_{ij} B_j$$

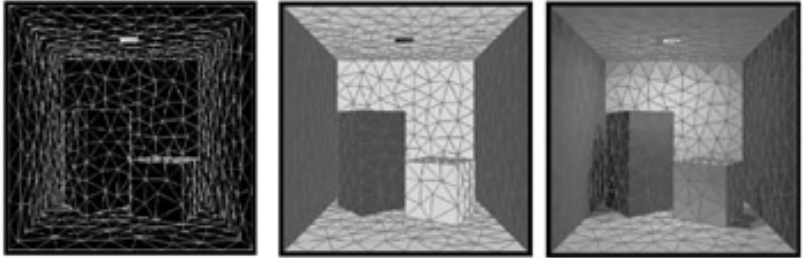
- The radiosity of patch  $B_i$  is a linear combination of radiosities of all other patches  $B_j$
- Importantly the coefficients of this system depends **only** on geometry which is known and self emitted radiosity which is **known**

$$B_{ei} = B_i - \rho_i \sum_j F_{ij} B_j$$

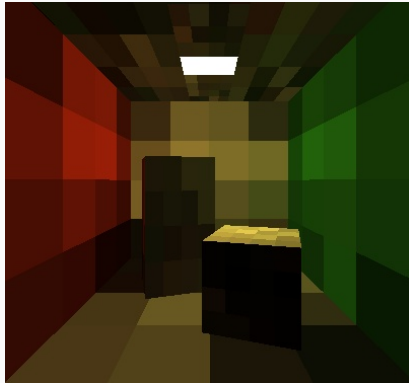
$$\begin{bmatrix} B_{e1} \\ B_{e2} \\ \vdots \\ B_{en} \end{bmatrix} = \begin{bmatrix} 1 - \rho_1 F_{11} & -\rho_1 F_{12} & \dots & \rho_1 F_{1n} \\ -\rho_2 F_{21} & 1 - \rho_2 F_{22} & \dots & \rho_2 F_{2n} \\ \vdots & \vdots & \vdots & \vdots \\ -\rho_n F_{n1} & -\rho_n F_{n2} & \dots & 1 - \rho_n F_{nn} \end{bmatrix} \begin{bmatrix} B_1 \\ B_2 \\ \vdots \\ B_n \end{bmatrix}$$

- We know  $B_{ei}$  and we can compute  $F_{ij}$
- Solve for  $B_i$

# Radiosity



- Compute form factors
- Solve system of equations
- Render image (Raycaster)



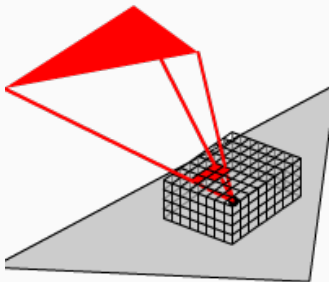
**Discretisation** Computational and storage cost

**Form Factor Computations 1** Cubic storage in number of patches

**Form Factor Computations 2** Complicated integral

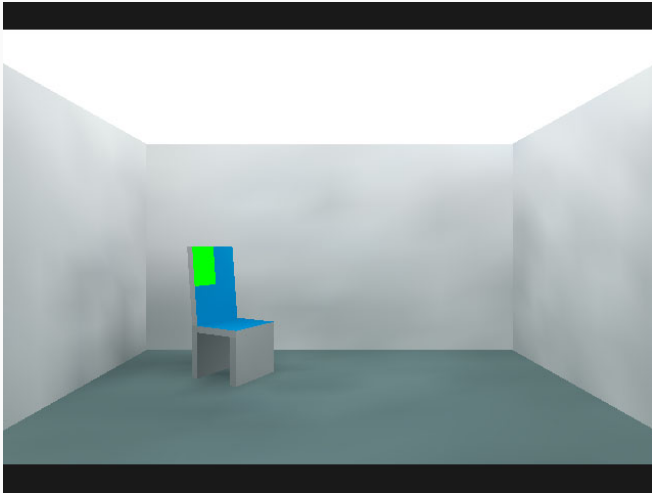
**Numerical Solution** Structure in coefficient to exploit

# Hemicube approximation



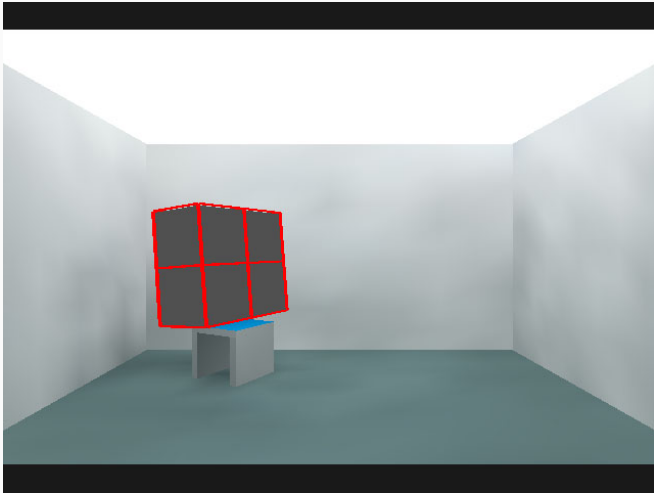
- Form factor computations are expensive
- Approximate with projections onto sphere

# Hemicube approximation<sup>3</sup>



<sup>3</sup><http://www.deluxerender.com/2014/11/hemicube-form-factors/>

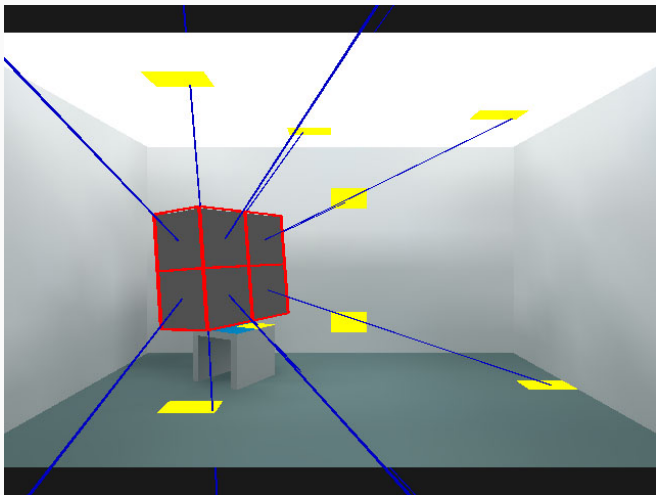
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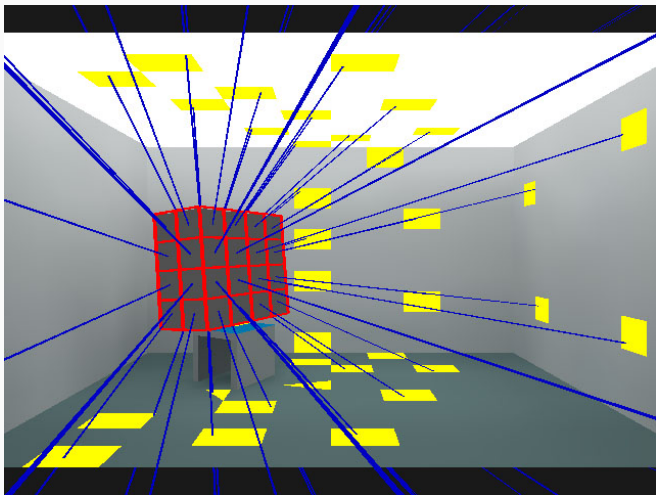


## Hemicube approximation<sup>3</sup>



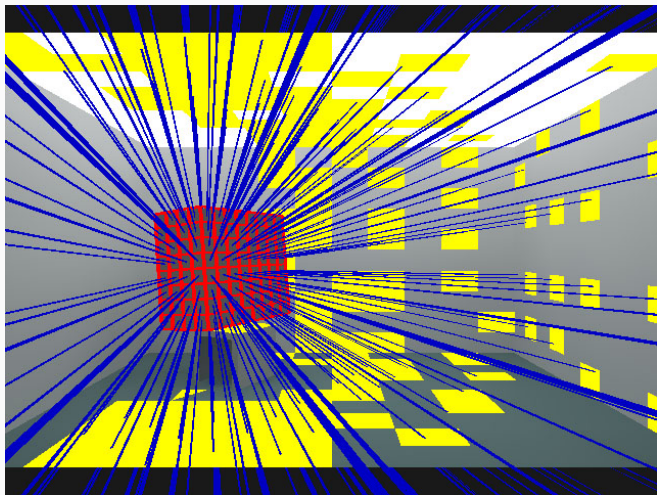
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## Hemicube approximation<sup>3</sup>



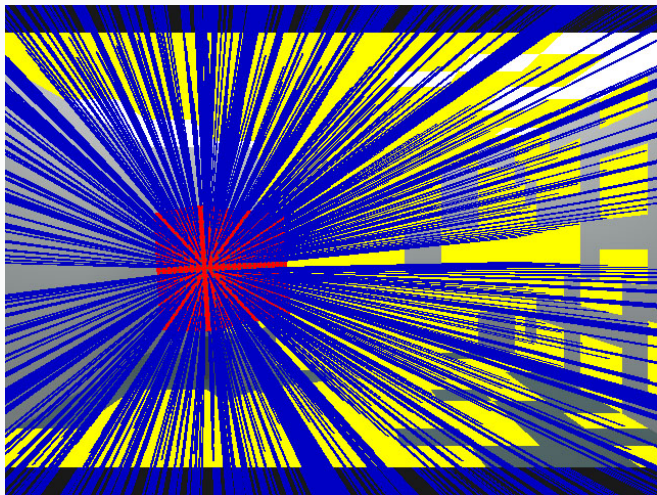
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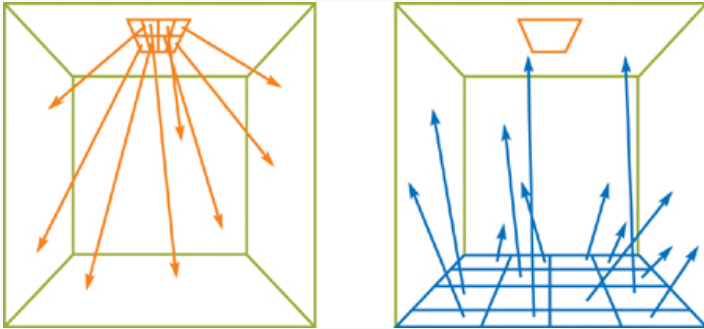
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## Hemicube approximation<sup>3</sup>



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# Progressive Radiosity<sup>4</sup>



- Solve equation iteratively
- Each patch stores, **accumulated** and **residual** energy
- Iterate until residual energy low

<sup>4</sup>[http:](http://http.developer.nvidia.com/GPUGems2/gpugems2_chapter39.html)

[//http.developer.nvidia.com/GPUGems2/gpugems2\\_chapter39.html](http://http.developer.nvidia.com/GPUGems2/gpugems2_chapter39.html)

## Summary

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- Intuition into rendering equation
- Cast raytracing as a solution to the rendering equation
- Classic Radiosity
  - very clear what assumptions are
  - images look very "radiosity" ;-)

**Lecture** Monday 23rd of March

- Path Tracing (Stochastic Methods)

**Lecture** Thursday 16th of April

- Photon Mapping (Caustics)

**Lecture** Monday 27th of April

- Last lecture
- Unit Summary



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