Class Diagram for Density Splits Code

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```
Class Universe
private:
 int expansion_or_collapse;
 int number_of_time_steps;
 double t initial;
 double eta_initial;
 double H_initial;
 double a_initial;
 double a_final;
 vector<double> t;
 vector<double> eta;
 vector<double> a;
 vector<double> H;
 vector<double> H_prime;
 cosmological_model cosmology;
public:
Universe(cosmological_model cosmo, double a_min, double a_max, int expand_or_collapse);
~Universe();
double return_a_initial();
double return_a_final();
double return_eta_initial();
double return eta final();
cosmological_model return_cosmology();
void print background cosmology(string filename);
double f k(double w);
double t_at_eta(double e);
double a_at_eta(double e);
double H_at_eta(double e);
double H_prime_at_eta(double e);
double eta_at_a(double a);
vector<vector<double> > return_background_expansion();
vector<vector<double> > return_background_expansion(int conformal_time_steps);
double rho_m_of_a(double scale); // All in units of TODAYS critical density
double rho_r_of_a(double scale);
double rho_L_of_a(double scale);
double w_L_of_a(double scale);
static void expansion_in_flat_matter_dominated_universe(double a, double *t_phys, double *eta, double *H_conformal, double *H_conformal_prime);
static void expansion_in_flat_radiation_dominated_universe(double a, double *t_phys, double *eta, double *H_conformal, double *H_conformal_prime);
static void expansion_in_flat_Lambda_dominated_universe(double a, double *t_phys, double *eta, double *H_conformal, double *H_conformal_prime);
private:
 void set initial conditions();
 void set_background_cosmology();
 void set_number_of_time_steps(int n_entries);
 double hubble_from_Friedmann(double a_start);
```

int collapse; double Omega_m; double Omega_r; double Omega_L; double Omega_b; double Omega_k; double theta_27; double w0;

double w1;

double n_s;

double h_100;

double sigma_8;