STELLAR CLASSIFICATION BASED ON SPECTRAL CHARACTERISTICS

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Summary: In this study, Stellar Classification Dataset - SDSS17 was explored. This is a multiclass classification problem with three classes: galaxies, stars, and quasars. The logistic regression model with multinom function from nnet package was utilised to solve this problem. The obtained accuracy for distinguishing stars from other classes is 99.4%. The obtained φ coefficient for distinguishing quasars from other classes is 89.2%. The obtained accuracy for distinguishing galaxies from other classes is 96.2%.

Introduction

Stellar Classification Dataset - SDSS17 [1] provides data for classification of stars, galaxies, and quasars based on their spectral characteristics. This dataset contains 18 attributes, one of which is class. Other 17 attributes include various spectral characteristics, object identifiers, and measurement-related parameters. Overall, the dataset contains 100000 observations of space taken by the Sloan Digital Sky Survey (SDSS).

The distribution of stars, quasars, and galaxies is shown in Fig. 1.

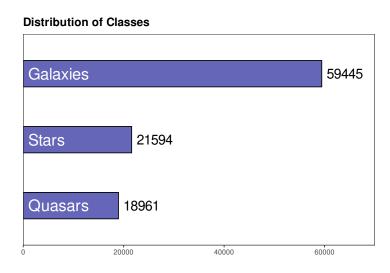


Figure 1: Distribution of stars, galaxies, and guasars in Stellar Classification Dataset - SDSS17

As one can observe from Fig. 1, the problem is quite imbalanced because the number of quasars is low with respect to the number of galaxies. The same concerns the number of stars.

Feature Selection

Full description of the attributes can be found in [1]. We need to select meaningful attributes for use in the model. To this, exploratory analysis is performed.

First, obj-ID is the object identifier, the unique value that identifies the object in the image catalog used by the CAS. Therefore, it should not be used in further analysis because it is not a spectral characteristic.

Second, alpha and delta specify the right ascension and declination angles, respectively. These angles specify the position of the observed object in the sky, making their use irrelevant for the model.

Third, u, r, g, i, redshift, and z are spectral characteristics. They should be included into the model.

Fourth, <code>run-ID</code>, <code>rerun-ID</code>, <code>cam-col</code>, <code>fiber-ID</code>, <code>plate</code>, <code>MJD</code>, and <code>field-ID</code> are features related to the measurement setup, which makes them irrelevant for stellar classification based on spectroscopic data.

Fifth, spec-obj-ID is related to spectroscopic observations. Hence, we will use this feature in the model.

Let us now calculate the correlation between the selected features to understand their relationship between each other.

```
##
                                             z spec obj ID redshift
## u
               1.000 0.999 0.999 0.046 0.998
                                                     0.030
                                                              0.014
## r
               0.999 1.000 1.000 0.056 0.999
                                                     0.039
                                                              0.023
## g
                                                              0.023
               0.999 1.000 1.000 0.056 0.999
                                                     0.039
                                                     0.662
## i
               0.046 0.056 0.056 1.000 0.056
                                                              0.492
               0.998 0.999 0.999 0.056 1.000
                                                     0.038
## Z
                                                              0.030
## spec obj ID 0.030 0.039 0.039 0.662 0.038
                                                     1.000
                                                              0.389
               0.014 0.023 0.023 0.492 0.030
                                                     0.389
## redshift
                                                              1.000
```

We observe that [r], [g], [u], and [z] are strongly correlated. However, further analysis of the model performance has shown that inclusion of all these features into the model improves the results.

Data Preprocessing

The selected features have been scaled so that they all have a mean of 0 and a standard deviation of 1. The dataset is split to 80% train and 20% test data randomly.

Problem Visualisation

We selected two meaningful features spec-obj-ID and redshift and visualised the distribution of classes. The results are shown in Fig. 2.

Clusters of galaxies, quasars, and stars in the redshift-spec obj ID coordinate system

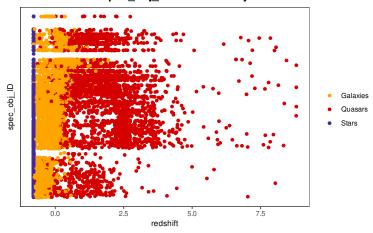


Figure 2: Visualised distribution of classes in the coordinate system of spec-obj-ID and redshift

Figure 2 shows three distinct clusters that should be easily identified using logistic regression. We see that the red shift value that corresponds to quasars is the highest because quasars are the most distant objects humans can observe. Stars have the lowest red shift value. The red shift value of galaxies is in between.

Model and Results

We use the logistic regression model with multinom function from nnet package. To evaluate the model performance, we use one-versus rest approach and investigate the following three cases: 1) galaxies-versus-rest classification, 2) quasars-versus-rest classification, and 3) stars-versus-rest classification. The performance of the model is assessed on the test data.

For the classification of stars versus other objects, the obtained train and test accuracies are 99.4% and 99.5%, respectively. Figure 3 shows the corresponding confusion matrices. The accuracy metric is a good choice in this case because the number of misclassification cases is low. Hence, stars are easy to distinguish from galaxies and quasars.

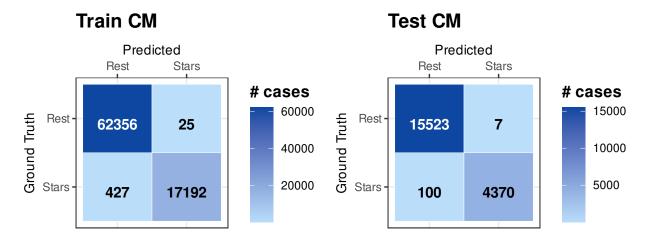


Figure 3: Confusion matrices for train and test data for classification of stars versus other classes.

For the classification of quasars versus other objects, the obtained train and test ϕ coefficients are 89.2% and 89.2%, respectively. Figure 4 shows the corresponding confusion matrices. The ϕ coefficient metric is a good choice in this case because of the considerable imbalance of quasars with respect to other classes and difficulty in distinguishing quasars from galaxies based on the studied spectral characteristics.

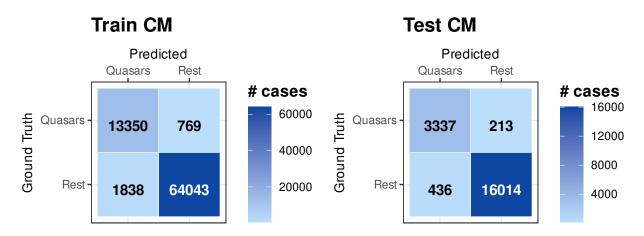


Figure 4: Confusion matrices for train and test data for classification of quasars versus other classes.

Considering the considerable imbalance, the model performance is decent in distinguishing quasars from other classes.

Finally, we assess the model performance in distinguishing galaxies from other classes.

For the classification of galaxies versus other objects, the obtained train and test accuracies are 96.2% and 96.2%, respectively. Figure 5 shows the corresponding confusion matrices. The accuracy metric is a good choice in this case because the number of false positives and false negatives is similar and the number of galaxies is quite balanced with respect to the number of other classes.

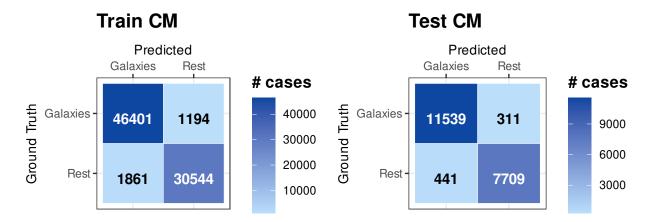


Figure 5: Confusion matrices for train and test data for classification of galaxies versus other classes.

Conclusion

Summary: In this study, Stellar Classification Dataset - SDSS17 was explored. This is a multiclass classification problem with three classes: galaxies, stars, and quasars. The logistic regression model with multinom function from nnet package was utilised to solve this problem. The obtained accuracy for distinguishing stars from other classes is 99.4%. The obtained φ coefficient for distinguishing quasars from other classes is 89.2%. The obtained accuracy for distinguishing galaxies from other classes is 96.2%.

Strengths and limitations: We can observe that the model performance is excellent for distinguishing stars from other classes. The model performance is slightly worse but still good for distinguishing galaxies from other classes. In contrast, the model performance is bad for distinguishing quasars from other classes.

Future scope: We will attempt to increase the performance of quasars' classification without compromising the performance of stars' and galaxies' classification by improving the model. Furthermore, the performance of the model selected herein will be compared with those of other models, and the best model will be used. In addition, the features used in the model should be investigated further; for example, basis expansion can be used, and the correlated features should be explored further. Finally, the model evaluation criteria will be adjusted; for example, ROC and AUC characteristics will be explored.

References

[1] https://www.kaggle.com/datasets/fedesoriano/stellar-classification-dataset-sdss17 (The data released by the SDSS is under public domain.)

Appendix

Full code

```
library(nnet)
library(ggplot2)
# READING THE DATA
data <- read.csv("data/star_classification.csv", stringsAsFactors = TRUE)</pre>
# FEATURE SELECTION
# Now. let us see the correlation between features
features <- data.frame(u = data$u, r = data$q, q = data$q, i = data$i,
                z = data$z, spec obj ID = data$spec obj ID,
                redshift = data$redshift)
target <- data$class</pre>
# Compute the correlation matrix
cor mat <- cor(features)</pre>
print(cor mat)
# We observe that r, g, u, and z are strongly correlated.
# DATA PREPROCESSING
features <- scale(features)</pre>
data prep <- as.data.frame(cbind(target, features))</pre>
n <- dim(data prep)[1]</pre>
# Data division
set.seed(12345)
tr ind \leftarrow sample(1:n, floor(0.8*n))
tr <- data prep[tr ind, ]</pre>
te <- data prep[-tr ind, ]
# logistic regression with multinom
# DATA VISUALISATION
# Let us visualise the distribution of data points as spec obj ID versus red shift
# colored by class
plot <- ggplot(data = te) + theme bw() +
 geom\ point(mapping = aes(x = redshift, y = spec obj ID,
                   color = as.factor(target))) +
 scale color manual(values = c("orange", "#d50000", "#4527a0"),
               name = NULL,
               labels = c("Galaxies", "Quasars", "Stars")) +
 labs(title = paste0("Clusters of galaxies, guasars, and stars\nin the",
                "redshift-spec obj ID coordinate system")) +
 theme(panel.grid.major = element blank(),
     panel.grid.minor = element blank(),
     axis.text.y = element blank(),
     axis.ticks.y = element blank(),
     title = element text(size = 12, face = 'bold'),
```

```
axis.title = element text(size = 10, face = "plain"))
print(plot)
# LOGISTIC REGRESSION
logreg <- multinom(formula = target ~ ., data = tr)</pre>
pred_tr <- predict(logreg, tr, type = "class")</pre>
pred te <- predict(logreg, te, type = "class")</pre>
# PERFORMANCE EVALUATION
CM tr <- table(pred tr, tr$target, deparse.level = 0)
CM te <- table(pred te, te$target, deparse.level = 0)</pre>
rownames(CM_te) <- c("GAL_pr", "QUA_pr", "STA_pr")
colnames(CM_te) <- c("GAL_tr", "QUA_tr", "STA_tr")</pre>
cat("The confusion matrix for train data is\n"); print(CM tr)
cat("The confusion matrix for test data is\n"); print(CM te)
# 1 is GALAXY, 2 is QUASAR, 3 is STAR
# Let us introduce this renewed stars-versus-rest CM
CM_tr_stars_vs_rest \leftarrow matrix(c(CM tr[1, 1] + CM tr[1, 2] + CM tr[2, 1] + CM tr[2, 1])
                                CM tr[2, 2],
                              CM tr[2, 3] + CM tr[1, 3],
                              CM tr[3, 1] + CM tr[3, 2],
                              CM tr[3, 3]), byrow = TRUE, ncol = 2)
colnames(CM tr stars vs rest) <- c("rest tr", "STA tr")</pre>
rownames(CM tr stars vs rest) <- c("rest pr", "STA pr")
print("For train data")
print(CM tr stars vs rest)
GT <- factor(c("Rest", "Stars", "Rest", "Stars"))
predicted <- factor(c("Rest", "Rest", "Stars", "Stars"))</pre>
values <- c(CM tr stars vs rest)</pre>
df <- data.frame(GT, predicted, values)</pre>
# Visualise this CM
plot <- qqplot(data = df, mapping = aes(x = predicted, y = GT)) +
 geom tile(aes(fill = values), color = "white") +
 geom text(data = subset(df, values > 30000),
           aes(label = values),
           vjust = 1, color = "white", fontface = "bold", family = "arial") +
 geom text(data = subset(df, values < 30000),</pre>
           aes(label = values),
           vjust = 1, color = "black", fontface = "bold", family = "arial") +
 scale fill gradient(low = "#bbdefb", high = "#0d47a1") +
 theme bw() +
 scale x discrete(position = "top") +
 scale y discrete(limits=rev) +
```

```
labs(x = "Predicted", y = "Ground Truth", fill = "# cases",
                 title = "Train CM for stars-versus-rest case") +
     theme(title = element text(size = 12, face = 'bold'),
                   axis.title = element text(size = 10, face = "plain"))
print(plot)
CM te stars vs rest \leftarrow matrix(c(CM te[1, 1] + CM te[1, 2] + CM te[2, 1] +
                                                                                 CM te[2, 2],
                                                                            CM \ te[2, 3] + CM \ te[1, 3],
                                                                            CM te[3, 1] + CM te[3, 2],
                                                                            CM_{te[3, 3]}, byrow = TRUE, ncol = 2)
colnames(CM_te_stars_vs_rest) <- c("rest_tr", "STA_tr")</pre>
rownames(CM te stars vs rest) <- c("rest pr", "STA pr")
print("For test data")
print(CM te stars vs rest)
GT <- factor(c("Rest", "Stars", "Rest", "Stars"))</pre>
predicted <- factor(c("Rest", "Rest", "Stars", "Stars"))</pre>
values <- c(CM te stars vs rest)</pre>
df <- data.frame(GT, predicted, values)</pre>
# Visualise this CM
plot <- ggplot(\frac{data}{data} = \frac{df}{data} = \frac{df}{data}
     geom tile(aes(fill = values), color = "white") +
     geom text(data = subset(df, values > 10000),
                             aes(label = values),
                             vjust = 1, color = "white", fontface = "bold", family = "arial") +
     geom text(data = subset(df, values < 10000),</pre>
                            aes(label = values),
                            vjust = 1, color = "black", fontface = "bold", family = "arial") +
     scale fill gradient(low = "#bbdefb", high = "#0d47a1") +
     theme bw() +
     scale x discrete(position = "top") +
     scale y discrete(limits=rev) +
     labs(x = "Predicted", y = "Ground Truth", fill = "# cases",
                 title = "Test CM for stars-versus-rest case") +
     theme(title = element text(size = 12, face = 'bold'),
                   axis.title = element text(size = 10, face = "plain"))
print(plot)
cat("Train accuracy for distinguishing between stars and other classes\n")
MCR tr <- (CM tr stars vs rest[1, 1] + CM tr stars vs rest[2, 2]) /
     sum(CM tr stars vs rest)
cat(round(\overline{MCR}_{tr*100}, 1), "%", sep="")
cat("Test accuracy for distinguishing between stars and other classes\n")
MCR te \leftarrow (CM te stars vs rest[1, 1] + CM te stars vs rest[2, 2]) /
     sum(CM te stars vs rest)
cat(round(MCR_te*100, 1), "%", sep="")
```

```
# Now, let us consider distinguishing quasars from other data. The confusion
# matrix is as follows:
CM tr quasars vs rest \leftarrow matrix(c(CM tr[1, 1] + CM_tr[1, 3] + CM_tr[3, 1] +
                                                                             CM tr[3, 3],
                                                                    CM tr[1, 2] + CM tr[3, 2],
                                                                    CM_{tr[2, 1]} + CM_{tr[2, 3]}
                                                                    CM tr[2, 2]), byrow = TRUE, ncol = 2)
colnames(CM_tr_quasars vs rest) <- c("rest tr", "QUA tr")</pre>
rownames(CM_tr_quasars_vs_rest) <- c("rest_pr", "QUA_pr")
print("For train data")
print(CM tr quasars vs rest)
GT <- factor(c("Rest", "Quasars", "Rest", "Quasars"))</pre>
predicted <- factor(c("Rest", "Rest", "Quasars", "Quasars"))</pre>
values <- c(CM tr quasars vs rest)</pre>
df <- data.frame(GT, predicted, values)</pre>
# Visualise this CM
plot <- ggplot(data = df, mapping = aes(x = predicted, y = GT)) +
    geom tile(aes(fill = values), color = "white") +
    geom text(data = subset(df, values > 40000),
                         aes(label = values),
                         vjust = 1, color = "white", fontface = "bold", family = "arial") +
    geom text(data = subset(df, values < 40000),</pre>
                         aes(label = values),
vjust = 1, color = "black", fontface = "bold", family = "arial") +
    scale_fill_gradient(low = "#bbdefb", high = "#0d47a1") +
    theme bw() +
    scale x discrete(position = "top") +
    scale y discrete(limits=rev) +
    labs(x = "Predicted", y = "Ground Truth", fill = "# cases",
               title = "Train CM for quasars-versus-rest case") +
    theme(title = element text(size = 12, face = 'bold'),
                 axis.title = element text(size = 10, face = "plain"))
print(plot)
CM_{te}_{quasars} = -matrix(c(CM_{te}[1, 1] + CM_{te}[1, 3] + CM_{te}[3, 1] 
                                                                             CM te[3, 3],
                                                                    CM te[1, 2] + CM te[3, 2],
                                                                    CM te[2, 1] + CM te[2, 3],
                                                                    CM_{te[2, 2]}, byrow = TRUE, ncol = 2)
colnames(CM te quasars vs rest) <- c("rest tr", "QUA tr")</pre>
rownames(CM te quasars vs rest) <- c("rest pr", "QUA pr")
print("For test data")
print(CM te quasars vs rest)
GT <- factor(c("Rest", "Quasars", "Rest", "Quasars"))</pre>
predicted <- factor(c("Rest", "Rest", "Quasars", "Quasars"))</pre>
```

```
values <- c(CM te quasars vs rest)</pre>
df <- data.frame(GT, predicted, values)</pre>
# Visualise this CM
plot <- ggplot(data = df, mapping = aes(x = predicted, y = GT)) +
  geom_tile(aes(fill = values), color = "white") +
  geom text(data = subset(df, values > 6000),
            aes(label = values),
            vjust = 1, color = "white", fontface = "bold", family = "arial") +
  geom_text(data = subset(df, values < 6000),</pre>
            aes(label = values),
vjust = 1, color = "black", fontface = "bold", family = "arial") +
  scale_fill_gradient(low = "#bbdefb", high = "#0d47a1") +
  theme bw() +
  scale x discrete(position = "top") +
  scale y discrete(limits=rev) +
  labs(x = "Predicted", y = "Ground Truth", fill = "# cases",
       title = "Test CM for quasars-versus-rest case") +
  theme(title = element text(size = 12, face = 'bold'),
        axis.title = element text(size = 10, face = "plain"))
print(plot)
# Here, we observe the large number of misclassification cases. To assess the
# model performance, considering the data imbalance (low number of guasars wrt
# other classes, as shown in Figure X), we use the Phi coefficient, which is a
# recommended choice for binary classification in the case of imbalanced data
phi coef <- function(M){</pre>
  numerator \leftarrow (M[1, 1]*M[2, 2] - M[1, 2]*M[2, 1])
 MCC \leftarrow numerator/sqrt(M[1, 1] + M[1, 2])
 MCC \leftarrow MCC/sqrt(M[2, 1] + M[2, 2])
 MCC \leftarrow MCC/sqrt(M[1, 1] + M[2, 1])
 MCC \leftarrow MCC/sqrt(M[1, 2] + M[2, 2])
  return(MCC)
}
cat("Train Phi coeff for distinguishing between quasars and other classes\n",
    round(phi coef(CM tr quasars vs rest)*100, 1), "%", sep = "")
cat("Test Phi coeff for distinguishing between quasars and other classes\n",
    round(phi coef(CM te quasars vs rest)*100, 1), "%", sep = "")
# Considering the considerable imbalance, the model performance is decent in
# distinguishing quasars from other classes
# Finally, we want to assess the model performance in distinguishing galaxies
# from other classes.
CM tr galaxies vs rest \leftarrow matrix(c(CM tr[2, 2] + CM tr[2, 3] + CM tr[3, 2] +
                                       CM_tr[3, 3],
                                    CM tr[1, 2] + CM tr[1, 3],
                                    CM tr[2, 1] + CM tr[3, 1],
```

```
CM tr[1, 1]), byrow = TRUE, ncol = 2)
colnames(CM tr galaxies vs rest) <- c("rest tr", "GAL tr")</pre>
rownames(CM tr galaxies vs rest) <- c("rest pr", "GAL pr")</pre>
print("For train data")
print(CM tr galaxies vs rest)
GT <- factor(c("Rest", "Galaxies", "Rest", "Galaxies"))</pre>
predicted <- factor(c("Rest", "Rest", "Galaxies", "Galaxies"))</pre>
values <- c(CM tr galaxies vs rest)</pre>
df <- data.frame(GT, predicted, values)</pre>
# Visualise this CM
plot <- ggplot(\frac{data}{data} = \frac{df}{data} = \frac{df}{data}
     geom tile(aes(fill = values), color = "white") +
     geom text(data = subset(df, values > 30000),
                             aes(label = values),
                              vjust = 1, color = "white", fontface = "bold", family = "arial") +
     geom text(data = subset(df, values < 30000),</pre>
                             aes(label = values),
vjust = 1, color = "black", fontface = "bold", family = "arial") +
     scale fill gradient(low = "#bbdefb", high = "#0d47a1") +
     theme bw() +
     scale x discrete(position = "top") +
     scale_y discrete(limits=rev) +
     labs(x = "Predicted", y = "Ground Truth", fill = "# cases",
                 title = "Train CM for galaxies-versus-rest case") +
     theme(title = element text(size = 12, face = 'bold'),
                   axis.title = element text(size = 10, face = "plain"))
print(plot)
CM_{te}_{galaxies}_{vs}_{rest} \leftarrow matrix(c(CM_{te}_{2}, 2] + CM te_{2}, 3] + CM te_{3} + CM te_{3} + CM te_{3}
                                                                                           CM te[3, 3],
                                                                                   CM \ te[1, 2] + CM \ te[1, 3],
                                                                                   CM te[2, 1] + CM te[3, 1],
                                                                                   CM te[1, 1]), byrow = TRUE, ncol = 2)
colnames(CM te galaxies vs rest) <- c("rest tr", "GAL tr")</pre>
rownames(CM te galaxies vs rest) <- c("rest pr", "GAL pr")</pre>
print("For test data")
print(CM te galaxies vs rest)
GT <- factor(c("Rest", "Galaxies", "Rest", "Galaxies"))
predicted <- factor(c("Rest", "Rest", "Galaxies", "Galaxies"))</pre>
values <- c(CM te galaxies_vs_rest)</pre>
df <- data.frame(GT, predicted, values)</pre>
# Visualise this CM
plot <- ggplot(data = df, mapping = aes(x = predicted, y = GT)) +
     geom tile(aes(fill = values), color = "white") +
     geom text(data = subset(df, values > 6000),
                             aes(label = values),
```

```
vjust = 1, color = "white", fontface = "bold", family = "arial") +
 geom text(data = subset(df, values < 6000),</pre>
           aes(label = values),
           vjust = 1, color = "black", fontface = "bold", family = "arial") +
 scale fill gradient(low = "#bbdefb", high = "#0d47a1") +
 theme bw() +
 scale x discrete(position = "top") +
 scale y discrete(limits=rev) +
 labs(x = "Predicted", y = "Ground Truth", fill = "# cases",
      title = "Test CM for galaxies-versus-rest case") +
 theme(title = element text(size = 12, face = 'bold'),
       axis.title = element text(size = 10, face = "plain"))
print(plot)
# For assessing the model performance for distinguishing galaxies from other classes,
# we can use the accuracy metric because the imbalance is not a problem here.
cat("Train accuracy for distinguishing between galaxies and other classes\n")
MCR tr <- (CM tr galaxies vs rest[1, 1] + CM tr galaxies vs rest[2, 2]) /
  sum(CM tr galaxies vs rest)
cat(round(MCR tr*100, 1), "%", sep="")
cat("Test accuracy for distinguishing between galaxies and other classes\n")
MCR te <- (CM te galaxies vs rest[1, 1] + CM te galaxies vs rest[2, 2]) /
 sum(CM te galaxies vs rest)
cat(round(MCR te*100, 1), "%", sep="")
# CLASS DISTRIBUTION
# Visualise the distribution of classes
classes = c("GALAXY", "STAR", "QSO")
counts = numeric(3)
for (i in 1:3){
 counts[i] = length(which(data$class == classes[i]))
}
df classes <- data.frame(classes = classes, counts = counts)</pre>
plot <- ggplot(data = df classes, mapping = aes(x = counts,</pre>
                                             v = reorder(classes, counts))) +
 theme bw() +
 geom bar(stat = "identity", fill = "darkblue", alpha = .6, width = .4,
          color = "black") +
 labs(title = "Distribution of Classes", x = "", y = "") +
 theme(panel.grid.major = element blank(),
       panel.grid.minor = element blank(),
       axis.text.y = element blank(),
       axis.ticks.y = element blank(),
       title = element text(size = 12, face = 'bold')) +
 scale x continuous(expand = c(0, 10), limits = c(0, 70000)) +
 geom text(
   aes(1000, y = classes, label = c("Galaxies", "Stars", "Quasars")),
   hjust = 0,
```

```
nudge_x = 0,
  color = "white",
  family = "Calibri",
  size = 7
) +
  geom_text(
   aes(counts + 1000, y = classes, label = counts),
  hjust = 0,
  nudge_x = 0.3,
  colour = "black",
  family = "Calibri",
  size = 6
)
print(plot)
```