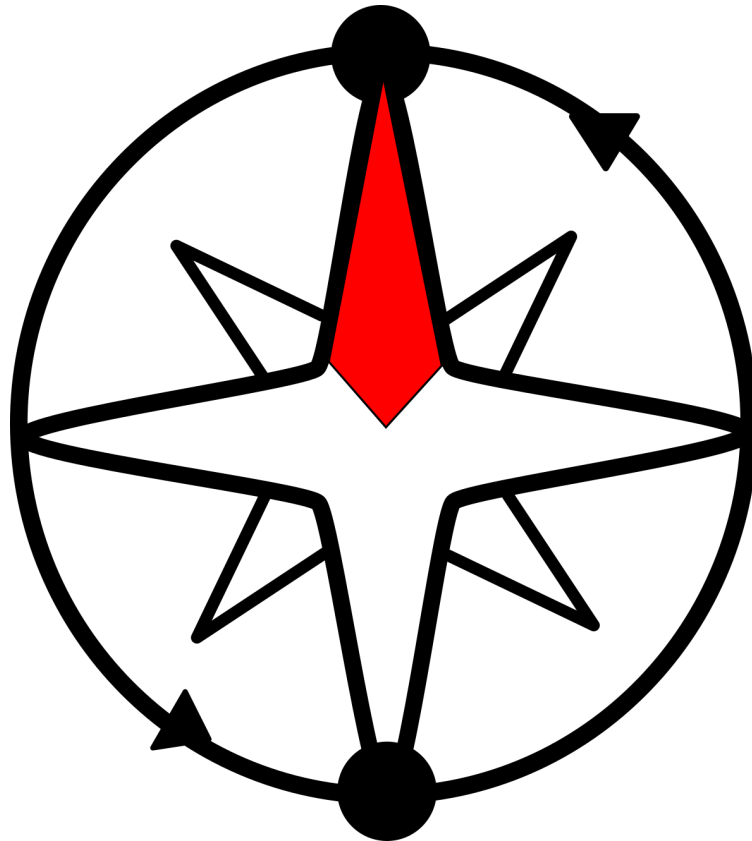




## COMPACT OBJECT MERGERS: POPULATION ASTROPHYSICS AND STATISTICS

COMPAS is a platform for the exploration of populations of compact binaries formed through isolated binary evolution (Stevenson et al., 2017; Vigna-Gómez et al., 2018; Barrett et al., 2018; Neijssel et al., 2019; Broekgaarden et al., 2019; Stevenson et al., 2019; Chattopadhyay et al., 2020). The COMPAS population synthesis code is flexible, fast and modular, allowing rapid simulation of binary star evolution. The complete COMPAS suite includes the population synthesis code together with a collection of tools for sophisticated statistical treatment of the synthesised populations.



Visit <http://compas.science> for more information.

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## Revision History

Date	Version	Description	Author
2012-2019		Original COMPAS manual	Team COMPAS
1 September 2019	0.1	Initial draft	Jeff Riley
20 September 2019	0.2	Updated compilation requirements	Jeff Riley
1 October 2019	0.3	Added (minimal) CHE documentation	Jeff Riley
21 October 2019	0.4	Added Grids documentation	
		Added Programming Style and Conventions	
		Added Appendices A-E	
		Reformatted for User Guide, Developer Guide etc.	Jeff Riley
18 December 2019	0.5	Updated Grids documentation for kick values	Jeff Riley
20 December 2019	0.6	Updated Appendices A, B, D, E	Jeff Riley
23 January 2020	0.7	Updated Grids documentation for kick values	Jeff Riley
31 March 2020	1.0	Updated for Beta release	Jeff Riley
22 April 2020	1.1	Updated to reflect pre2ndSN -> preSN variable changes	Simon Stevenson
27 April 2020	1.2	Updated to reflect SSE Grid file changes	Jeff Riley

## User Guide

Installation instructions for COMPAS and its dependencies are shown in the COMPAS Getting Started guide provided in the COMPAS suite. Please refer to that guide for dependency requirements etc.

### COMPAS Input

COMPAS provides wide-ranging functionality and affords users much flexibility in determining how the synthesis and evolution of stars (single or binary) is conducted. Users configure COMPAS functionality and provide initial conditions via the use of program options, described in Appendix A – Program Options. Initial conditions can also be provided via Grid files, described below.

A convenient method of managing the many program options provided by COMPAS (Appendix A – Program Options) is the example Python script provided in the COMPAS suite. The example script provided is `pythonSubmitDefault.py` – users should copy and modify their copy of the script to match their experimental requirements. Refer to the Getting Started guide for more details.

### Grids

Grid functionality allows users to specify a grid of initial values for both Single Star Evolution (SSE) and Binary Star Evolution (BSE).

For SSE, users can supply a text file that contains initial mass values and, optionally, metallicity and supernova kick related values, and COMPAS will evolve individual stars with those initial values (one star per record).

For BSE, users can supply a text file that contains initial mass and metallicity values for the binary constituent stars, as well as the initial separation or orbital period, and eccentricity, of the binary (one binary star per record of initial values).

### Grid File Format

Grid files are comma-separated text files, with column headers denoting the meaning of the data in the column (with an exception for SSE Grid files – see below for details).

Grid files may contain comments. Comments are denoted by the hash/pound character ('#'). The hash character and any text following it on the line in which the hash character appears is ignored. The hash character can appear anywhere on a line - if it is the first character then the entire line is a comment and ignored, or it can follow valid characters on a line, in which case the characters before the hash are processed, but the hash character and any text following it is ignored. Blank lines are ignored.

Notwithstanding the exception for SSE Grid files mentioned above, the first non-comment, non-blank line in a Grid file must be the header record. The header record is a comma-separated list of strings that denote the meaning of the data in each of the columns in the file.

Data records follow the header record. Data records, with an exception for SSE data files described below, are comma-separated lists of non-negative floating-point numbers. Any data field that contains a negative number, or characters that do not convert to floating-point numbers, is considered an error and will cause processing of the Grid file to be abandoned – an error message will be displayed.

Data records are expected to contain the same number of columns as the header record. If a data record contains more columns than the header record, data beyond the number of columns in the header record is ignored. If a data record contains fewer columns than the header record, missing data values (by position) are set equal to 0.0 – a warning message will be displayed.

## BSE Grid File Header Record

The BSE Grid file header record must be a comma-separated list of strings taken from the following list (case is not significant):

Header String	Column meaning
<b>Mass_1</b>	mass value to be assigned to the primary star, in $M_{\odot}$
<b>Mass_2</b>	mass value to be assigned to the secondary star, in $M_{\odot}$
<b>Metallicity_1</b>	metallicity value to be assigned to the primary star
<b>Metallicity_2</b>	metallicity value to be assigned to the secondary star
<b>Separation</b>	separation of the stars – the semi-major axis value to be assigned to the binary, in $AU$
<b>Eccentricity</b>	eccentricity value to be assigned to the binary
<b>Period</b>	orbital period value to be assigned to the binary, in <i>days</i>
<i>Kick_Velocity_Random_1</i>	<p>value to be used as the kick velocity magnitude random number, used to draw the kick velocity, for the primary star should it undergo a supernova event. This must be a floating-point number in the range [0.0, 1.0).</p> <p>If this column is present in the grid file, the kick velocity for the primary star, should it undergo a supernova event, will be drawn from the appropriate distribution at the time of the SN event. This column is used in preference to the <i>Kick_Velocity_1</i> column if both are present in the grid file.</p>
<i>Kick_Velocity_1</i>	value to be used as the (drawn) kick velocity for the primary star should it undergo a supernova event, in $kms^{-1}$ . If both this column and <i>Kick_Velocity_Random_1</i> are present in the grid file, <i>Kick_Velocity_Random_1</i> will be used in preference to <i>Kick_Velocity_1</i> .
<i>Kick_Theta_1</i>	value to be as the angle between the orbital plane and the 'z' axis of the supernova vector for the primary star should it undergo a supernova event, in <i>radians</i>
<i>Kick_Phi_1</i>	value to be used as the angle between 'x' and 'y', both in the orbital plane of the supernova vector, for the primary star should it undergo a supernova event, in <i>radians</i>
<i>Kick_Mean_Anomaly_1</i>	value to be used as the mean anomaly at the instant of the supernova for the primary star should it undergo a supernova event – should be uniform in [0, 2pi)
<i>Kick_Velocity_Random_2</i>	<p>value to be used as the kick velocity magnitude random number, used to draw the kick velocity, for the secondary star should it undergo a supernova event. This must be a floating-point number in the range [0.0, 1.0).</p> <p>If this column is present in the grid file, the kick velocity for the secondary star, should it undergo a supernova event, will be drawn from the appropriate distribution at the time of the SN event. This column is used in preference to the <i>Kick_Velocity_2</i> column if both are present in the grid file.</p>
<i>Kick_Velocity_2</i>	value to be used as the (drawn) kick velocity for the secondary star should it undergo a supernova event, in $kms^{-1}$ . If both this column and <i>Kick_Velocity_Random_2</i> are present in the grid file, <i>Kick_Velocity_Random_2</i> will be used in preference to <i>Kick_Velocity_2</i> .
<i>Kick_Theta_2</i>	value to be as the angle between the orbital plane and the 'z' axis of the supernova vector for the secondary star should it undergo a supernova event, in <i>radians</i>
<i>Kick_Phi_2</i>	value to be used as the angle between 'x' and 'y', both in the orbital plane of the supernova vector, for the secondary star should it undergo a supernova event, in <i>radians</i>
<i>Kick_Mean_Anomaly_2</i>	value to be used as the mean anomaly at the instant of the supernova for the secondary star should it undergo a supernova event – should be uniform in [0, 2pi)

All header strings in **bold** in the table above are required in the header record, with the exception of Separation and Period: one of Separation and Period *must* be present, but both *may* be present.

The header strings in *italics* in the table above (the kick-related header strings) are not mandatory. However, if one of the kick-related header strings is present, then *all must* be present.

The order of the columns in the BSE Grid file is not significant.

## **Data Record**

See the general description of data records above.

As for the header record, only one of Separation and Period is required to be present, but both may be present. The period may be used to calculate the separation of the binary. If the separation is present it is used as the value for the semi-major axis of the binary, regardless of whether the period is present (Separation has precedence over Period). If the period is present, but separation is not, the separation is calculated from the masses of the stars and the period given.

Also as for the header record, the kick-related values are not mandatory, but if one of the kick-related values is given, then all must be given.

## SSE Grid File

### Header Record

The SSE Grid file header record must be a comma-separated list of strings taken from the following list (case is not significant):

Header String	Column meaning
<b>Mass</b>	mass value ( ) to be assigned to the star ( $M_{\odot}$ )
<i>Metallicity</i>	metallicity value to be assigned to the star
<i>Kick_Velocity_Random</i>	<p>value to be used as the kick velocity magnitude random number, used to draw the kick velocity, for the star should it undergo a supernova event. This must be a floating-point number in the range [0.0, 1.0).</p> <p>If this column is present in the grid file, the kick velocity for the star, should it undergo a supernova event, will be drawn from the appropriate distribution at the time of the SN event. This column is used in preference to the Kick_Velocity column if both are present in the grid file.</p>
<i>Kick_Velocity</i>	value to be used as the (drawn) kick velocity for the star should it undergo a supernova event, in $\text{kms}^{-1}$ . If both this column and Kick_Velocity_Random are present in the grid file, Kick_Velocity_Random will be used in preference to Kick_Velocity.

The SSE Grid file is only required to list Mass values for each star, with Metallicity and Kick values being optional. If the Metallicity column is omitted, the metallicity value assigned to the star is the user-specified value for metallicity via the program options (OPTIONS→Metallicity()).

If the Metallicity column is omitted from the SSE Grid file, the header is optional: if there is only one column of data in the SSE Grid file it is assumed to be the Mass column, and no header is required (though may be present). If the Metallicity column header is present, the Mass column header is required.

The order of the columns in the SSE Grid file is not significant.

### Data Record

See the general description of data records above. As for the header record, only Mass is required to be present, but Metallicity may also be present. If the Metallicity is omitted, the metallicity value assigned to the star is the user-specified value for metallicity via the program options (OPTIONS→Metallicity()).



## COMPAS Output

Summary and status information during the evolution of stars is written to stdout; how much is written depends upon the value of the *quiet* program option.

Detailed information is written to log files (described below). All COMPAS output files are created inside a container directory, specified by the *output-container* program option. If BSE Detailed Output log files are created (see the *detailedOutput* program option) they will be created inside a containing directory named ‘Detailed\_Output’ within the COMPAS output container directory. Also created in the COMPAS container directory is a file named ‘Run\_Details’ in which COMPAS records some details of the run (COMPAS version, start time, program option values etc.).

COMPAS defines several standard log files that may be produced depending upon the simulation type (Single Star Evolution (SSE), or Binary Star Evolution (BSE)), and the value of various program options. The standard log files are:

- SSE Parameters log file
- BSE System Parameters log file
- BSE Detailed Output log file
- BSE Double Compact Objects log file
- BSE Common Envelopes log file
- BSE Supernovae log file
- BSE Pulsar Evolution log file

## Standard Log File Record Specifiers

Each standard log file has an associated log file record specifier that defines what data are to be written to the log files. Each record specifier is a list of known properties that are to be written as the log record for the log file associated with the record specifier. Default record specifiers for each of the standard log files are shown in Appendix E – Default Log File Record Specifications. The standard log file record specifiers can be defined by the user at run-time (see Standard Log File Record Specification below).

When specifying known properties, the property name must be prefixed with the property type. The current list of valid property types available for use is:

- STAR\_PROPERTY
- STAR\_1\_PROPERTY
- STAR\_2\_PROPERTY
- SUPERNOVA\_PROPERTY
- COMPANION\_PROPERTY
- BINARY\_PROPERTY
- PROGRAM\_OPTION

The stellar property types (all types except BINARY\_PROPERTY AND PROGRAM\_OPTION) must be paired with properties from the stellar property list, the binary property type BINARY\_PROPERTY with properties from the binary property list, and the program option type PROGRAM\_OPTION with properties from the program option property list.

## Standard Log File Record Specification

The standard log file record specifiers can be changed at run-time by supplying a definitions file via the *logfile-definitions* program option.

The syntax of the definitions file is fairly simple. The definitions file is expected to contain zero or more log file record specifications, as explained below.

For the following specification:

```
 ::=      means "expands to" or "is defined as"
{ x }     means (possible) repetition: x may appear zero or more times
[ x ]     means x is optional: x may appear, or not
<name>    is a term (expression)
"abc"     means literal string "abc"
|         means "or"
#         indicates the start of a comment
```

Logfile Definitions File specification:

```
<def_file> ::= { <rec_spec> }

<rec_spec> ::= <rec_name> <op> "{" { [ <props_list> ] } "}" <spec_delim>

<rec_name> ::= "SSE_PARMS_REC"           | # SSE only
               "BSE_SYSPARMS_REC"        | # BSE only
               "BSE_DCO_REC"              | # BSE only
               "BSE_SNE_REC"              | # BSE only
               "BSE_CEE_REC"              | # BSE only
               "BSE_PULSARS_REC"          | # BSE only
               "BSE_DETAILED_REC"         | # BSE only

<op>         ::= "=" | "+=" | "-="

<props_list> ::= <prop_spec> [ <prop_delim> <props_list> ]

<prop_spec>  ::= <prop_type> "::" <prop_name> <delim>

<spec_delim> ::= " " | EOL

<prop_delim> ::= "," | <spec_delim>

<prop_type>  ::= "STAR_PROPERTY"          | # SSE only
               "STAR_1_PROPERTY"          | # BSE only
               "STAR_2_PROPERTY"          | # BSE only
               "SUPERNOVA_PROPERTY"       | # BSE only
               "COMPANION_PROPERTY"       | # BSE only
               "BINARY_PROPERTY"          | # BSE only
               "PROGRAM_OPTION"           | # SSE or BSE

<prop_name>  ::= valid property name for specified property type
               (see definitions in constants.h)
```

The file may contain comments. Comments are denoted by the hash/pound character ('#'). The hash character and any text following it on the line in which the hash character appears is ignored by the parser. The hash character can appear anywhere on a line - if it is the first character then the entire line is a comment and ignored by the parser, or it can follow valid symbols on a line, in which case the symbols before the hash character are parsed and interpreted by the parser.

A log file specification record is initially set to its default value (see Appendix E – Default Log File Record Specifications). The definitions file informs the code as to the modifications to the default values the user wants. This means that the definitions log file is not mandatory, and if the definitions file is not present, or contains no valid record specifiers, the log file record definitions will remain at their default values.

The assignment operator given in a record specification (<op> in the file specification above) can be one of "=", "+=", and "-=". The meanings of these are:

"=" means that the record specifier should be assigned the list of properties specified in the braced-list following the "=" operator. The value of the record specifier prior to the assignment is discarded, and the new value set as described.

"+=" means that the list of properties specified in the braced-list following the "+=" operator should be appended to the existing value of the record specifier. Note that the new properties are appended to the existing list, so will appear at the end of the list (properties are printed in the order they appear in the list).

"-=" means that the list of properties specified in the braced-list following the "-=" operator should be subtracted from the existing value of the record specifier.

#### Example Log File Definitions File entries:

```
SSE_PARMS_REC = { STAR_PROPERTY::RANDOM_SEED,
                  STAR_PROPERTY::RADIUS, STAR_PROPERTY::MASS,
                  STAR_PROPERTY::LUMINOSITY }
```

```
BSE_PULSARS_REC += { STAR_1_PROPERTY::LUMINOSITY
                    STAR_2_PROPERTY::CORE_MASS
                    BINARY_PROPERTY::SEMI_MAJOR_AXIS_PRIME_RSOL
                    COMPANION_PROPERTY::RADIUS }
```

```
BSE_PULSARS_REC -= { SUPERNOVA_PROPERTY::TEMPERATURE }
```

```
BSE_PULSARS_REC += { PROGRAM_OPTION::KICK_VELOCITY_DISTRIBUTION_SIGMA_CCSN_NS,
                    BINARY_PROPERTY::ORBITAL_VELOCITY }
```

A full example Log File Record Specifications File is shown in Appendix F – Example Log File Record Specifications File.

The record specifications in the definitions file are processed individually in the sequence they appear in the file, and are cumulative: for record specifications pertaining to the same record name, the output of earlier specifications is input to later specifications.

For each record specification:

- Properties requested to be added to an existing record specification that already exist in that record specification are ignored. Properties will not appear in a record specification twice.
- Properties requested to be subtracted from an existing record specification that do not exist in that record specification are ignored.

Note that neither of those circumstances will cause a parse error for the definitions file – in both cases the user's intent is satisfied.

## Standard Log File Format

Each standard log file consists three header records followed by data records. Header records and data records are delimiter separated fields, with the delimiter being that specified by the `logfile-delimiter` program option (COMMA, TAB or SPACE), and the fields as specified by the log file record specifier.

The header records for all files are:

- Header record 1: Column Data Type Names
- Header record 2: Column Units (where applicable)
- Header record 3: Column Headings

Column Data Type Names are taken from the set {BOOL, INT, FLOAT, STRING}, where

- BOOL     the data value will be a boolean value. Boolean data values will be recorded in the log file in either numerical format (1 or 0, where 1 = TRUE and 0 = FALSE), or string format ('TRUE' or 'FALSE'), depending upon the value of the `print-bool-as-string` program option.
- INT       the data value will be an integer number
- FLOAT     the data value will be a floating-point number
- STRING    the data value will be a text string

Column Units is a string indicating the units of the corresponding data values (e.g. 'Msol\*AU^2\*yr^-1', 'Msol', 'AU', etc.). The Column Units value may be blank where units are not applicable, or may be one of:

- 'Count'    the data value is the total of a counted entity
- 'State'    the data value describes a state (e.g. 'Unbound' state is 'TRUE' or 'FALSE')
- 'Event'    the data value describes an event status (e.g. 'Simultaneous\_RLOF' is 'TRUE')

Column Headings are string labels that describe the corresponding data values. The heading strings for stellar properties of constituent stars of a binary will have appropriate identifiers appended. That is, heading strings for:

- STAR\_1\_PROPERTY::*properties* will have "\_1" appended
- STAR\_2\_PROPERTY::*properties* will have "\_2" appended
- SUPERNOVA\_PROPERTY::*properties* will have "\_SN" appended
- COMPANION\_PROPERTY::*properties* will have "\_CP" appended

## **Developer Guide**

Team COMPAS welcomes the active involvement of colleagues and others interested in the field in the ongoing development and improvement of the COMPAS software. We hope this Developer Guide helps anyone interested in contributing to the COMPAS software. We expect this guide to be a living document and improve along with the improvements made to the software.

## SINGLE STAR EVOLUTION

### Class Hierarchy

The main class for single star evolution is the **Star** class.

The Star class is a wrapper that abstracts away the details of the star and the evolution. Internally the Star class maintains a pointer to an object representing the star being evolved, with that object being an instance of one of the following classes:

**MS\_lte\_07**  
**MS\_gt\_07**  
**CH**  
**HG**  
**FGB**  
**CHeB**  
**EAGB**  
**TPAGB**  
**HeMS**  
**HeHG**  
**HeGB**  
**HeWD**  
**COWD**  
**ONeWD**  
**NS**  
**BH**  
**MR**

which track the phases from Hurley et al. 2000, with the exception of the CH class for Chemically Homogeneous stars, which is not described in Hurley et al. 2000.

Three other SSE classes are defined:

**BaseStar**  
**MainSequence**  
**GiantBranch**

These extra classes are included to allow inheritance of common functionality.

The BaseStar class is the main class for the underlying star object held by the Star class. The BaseStar class defines all member variables, and many member functions that provide common functionality. Similarly, the MainSequence and GiantBranch classes provide repositories for common functionality for main sequence and giant branch stars respectively.

The inheritance chain follows the phases described in Hurley et al. 2000 (again, with the exception of the CH, not described by Hurley et al. 2000), and is as follows:



Star

```
BaseStar → MainSequence → ( MS_lte_07 )  
                               ( MS_gt_07   ) → CH  
                               ( GiantBranch ) → HG → FGB → CheB → EAGB →  
  
→ TPAGB → HeMS → HeHG → HeGB → HeWD → COWD → OneWD → NS → BH → MR
```

CH (Chemically Homogeneous) stars inherit from MS\_gt\_07 because (in this implementation) they are just (large) main sequence stars that have a static radius.

HG (Hertzsprung Gap) stars inherit from the GiantBranch because they share the Giant Branch Parameters described in Hurley et al. 2000, section 5.2.

Each class in the inheritance chain has its own set of member functions that calculate various attributes of the star according to the phase the class represents (using the equations and parameters from Hurley et al. 2000 where applicable).

## Evolution Model

The stellar evolution model is driven by the Evolve() function in the Star class which evolves the star through its entire lifetime by doing the following:

DO:

1. calculate time step
  - calculate the giant branch parameters (as necessary)
  - calculate the timescales
  - choose time step
2. save the state of the underlying star object
3. DO:
  - a) evolve a single time step
  - b) if too much change
    - revert to the saved state
    - reduce the size of the time step

UNTIL timestep not reduced

4. resolve any mass loss
  - a) update initial mass (mass0)
  - b) update age after mass loss
  - c) apply mass transfer rejuvenation factor
5. evolve to the next stellar type if necessary

WHILE the underlying star object is not one of: { HeWD, COWD, ONeWD, NS, BH, MR }

Evolving the star through a single time step (step 3a above) is driven by the UpdateAttributesAndAgeOneTimestep() function in the BaseStar class which does the following:

1. check if the star should be a massless remnant
2. check if the star is a supernova
- if evolution on the phase should be performed
3. evolve the star on the phase – update stellar attributes
4. check if the star should evolve off the current phase to a different stellar type
- else
5. ready the star for the next time step

Evolving the star on its current phase, and off the current phase and preparing to evolve to a different stellar type, is handled by two functions in the BaseStar class: EvolveOnPhase() and ResolveEndOfPhase().

The EvolveOnPhase() function does the following:

1. Calculate Tau
2. Calculate CO Core Mass
3. Calculate Core Mass
4. Calculate He Core Mass
5. Calculate Luminosity
6. Calculate Radius
7. Calculate Envelope Mass
8. Calculate Perturbation Mu
9. Perturb Luminosity and Radius
10. Calculate Temperature
11. Resolve possible envelope loss

Each of the calculations in the EvolveOnPhase() function is performed in the context of the star evolving on its current phase. Each of the classes implements their own version of the calculations (via member functions) – some may inherit functions from the inheritance chain, while others might just return the value unchanged if the calculation is not relevant to their stellar type.

The `ResolveEndOfPhase()` function does the following:

1. Resolve possible envelope loss
2. Calculate Tau
3. Calculate CO Core Mass
4. Calculate Core Mass
5. Calculate He Core Mass
6. Calculate Luminosity
7. Calculate Radius
8. Calculate Envelope Mass
9. Calculate Perturbation Mu
10. Perturb Luminosity and Radius
11. Calculate Temperature
12. Evolve star to next phase

Each of the calculations in the `ResolveEndOfPhase()` function is performed in the context of the star evolving off its current phase to the next phase.

The remainder of the code (in general terms) supports these main driver functions.

## BINARY STAR EVOLUTION

### Class Hierarchy

The main class for single star evolution is the **BinaryStar** class. The BinaryStar class is a wrapper that abstracts away the details of the binary star and the evolution. Internally the BinaryStar class maintains a pointer to an object representing the binary star being evolved, with that object being an instance of the BaseBinaryStar class.

The BaseBinaryStar class is the main class for the underlying binary star object held by the BinaryStar class. The BaseBinaryStar class defines all member variables that pertain specifically to a binary star, and many member functions that provide binary-star specific functionality. Internally the BaseBinaryStar class maintains pointers to the two BinaryConstituentStar class objects that constitute the binary star.

The BinaryConstituentStar class inherits from the Star class, so objects instantiated from the BinaryConstituentStar class inherit the characteristics of the Star class, particularly the stellar evolution model. The BinaryConstituentStar class defines member variables and functions that pertain specifically to a constituent star of a binary system but that do not (generally) pertain to single stars that are not part of a binary system (there are some functions that are defined in the BaseStar class and its derived classes that deal with binary star attributes and behaviour – in some cases the stellar attributes that are required to make these calculations reside in the BaseStar class so it is easier and cleaner to define the functions there).

The inheritance chain is as follows:

BinaryStar → BaseBinaryStar

```
(Star → )      BinaryConstituentStar (star1)
(SStar → )      BinaryConstituentStar (star2)
```

## Evolution Model

The binary evolution model is driven by the Evolve() function in the BaseBinaryStar class which evolves the star through its entire lifetime by doing the following:

```
if (touching OR secondary too small)
    STOP = true
else
    calculate initial time step
    STOP = false

DO WHILE !STOP AND !max iterations:

    evolve a single time step
        evolve each constituent star a single time step (see SSE evolution)

    if (unbound OR touching OR Massless Remnant)
        STOP = true
    else
        evaluate the binary
            calculate lambdas if necessary
            calculate zetas if necessary

            calculate mass transfer
            calculate winds mass loss

            if common envelope
                resolve common envelope
            else if supernova
                resolve supernova
            else
                resolve mass changes

            evaluate supernovae
            resolve tides
            calculate total energy and angular momentum
            update magnetic field and spin: both constituent stars

        if (unbound OR touching OR merger)
            STOP = true
        else
            if NS+BH
                resolve coalescence

                if AIS exploratory phase
                    calculate DCO Hit

                STOP = true
            else
                if (WD+WD OR max time)
                    STOP = true
                else
                    if NOT max iterations
                        calculate new time step
```

## OBJECT IDENTIFIERS

All objects (instantiations of a class) are assigned unique object identifiers of type OBJECT\_ID (unsigned long int - see constants.h for the typedef). In the original COMPAS code, all binary star

objects were assigned unique object ids – this is just an extension of that so that all objects created in the COMPAS code are assigned unique ids. The purpose of the unique object id is to aid in object tracking and debugging.

As well as unique object ids, all objects are assigned an object type (of type OBJECT\_TYPE – see constants.h for the enum class declaring OBJECT\_TYPE), and a stellar type where applicable (of type STELLAR\_TYPE – see constants.h for the enum class declaring STELLAR\_TYPE).

Objects should expose the following functions:

OBJECT_ID	ObjectId() const	{ return m_ObjectId; }
OBJECT_TYPE	ObjectType() const	{ return m_ObjectType; }
STELLAR_TYPE	StellarType() const	{ return m_StellarType; }

If any of the functions are not applicable to the object, then they must return `"*::NONE` (all objects should implement `ObjectId()` correctly).

Any object that uses the Errors service (i.e. the `SHOW_*` macros) must expose these functions: the functions are called by the `SHOW_*` macros (the Errors service is described later in this document).

## SERVICES

A number of services have been provided to help simplify the code. These are:

- Program Options
- Random Numbers
- Logging and Debugging
- Error Handling

The code for each service is encapsulated in a singleton object (an instantiation of the relevant class). The singleton design pattern allows the definition of a class that can only be instantiated once, and that instance effectively exists as a global object available to all the code without having to be passed around as a parameter. Singletons are a little anti-OO, but provided they are used judiciously are not necessarily a bad thing, and can be very useful in certain circumstances.



## PROGRAM OPTIONS

A Program Options service is provided encapsulated in a singleton object (an instantiation of the Options class).

The Options class member variables are private, and public getter functions have been created for the program options currently used in the code.

The Options service can be accessed by referring to the Options::Instance() object. For example, to retrieve the value of the “quiet” program option, call the Quiet() getter function:

```
bool quiet = Options::Instance()→Quiet();
```

Since that could become unwieldy, there is a convenience macro to access the Options service. The macro just defines “OPTIONS” as “Options::Instance()”, so retrieving the value of the “quiet” program option can be written as:

```
bool quiet = OPTIONS→Quiet();
```

The Options service must be initialised before use. Initialise the Options service by calling the Initialise() function:

```
COMMANDLINE_STATUS programStatus = OPTIONS->Initialise(argc, argv);
```

(see constants.h for details of the COMMANDLINE\_STATUS type)

## RANDOM NUMBERS

A Random Number service is provided, with the gsl Random Number Generator encapsulated in a singleton object (an instantiation of the Rand class).

The Rand class member variables are private, and public functions have been created for random number functionality required by the code.

The Rand service can be accessed by referring to the Rand::Instance() object. For example, to generate a uniform random floating point number in the range [0, 1), call the Random() function:

```
double u = Rand::Instance()→Random();
```

Since that could become unwieldy, there is a convenience macro to access the Rand service. The macro just defines “RAND” as “Rand::Instance()”, so calling the Random() function can be written as:

```
double u = RAND→Random();
```

The Rand service must be initialised before use. Initialise the Rand service by calling the Initialise() function:

```
RAND→Initialise();
```

Dynamically allocated memory associated with the gsl random number generator should be returned to the system by calling the Free() function:

```
RAND→Free();
```

before exiting the program.

The Rand service provides the following public member functions:

### **void Initialise()**

Initialises the gsl random number generator. If the environment variable GSL\_RNG\_SEED exists, the gsl random number generator is seeded with the value of the environment variable, otherwise it is seeded with the current time.

### **void Free()**

Frees any dynamically allocated memory.

### **unsigned long int Seed(const unsigned long p\_Seed)**

Sets the seed for the gsl random number generator to p\_Seed. The return value is the seed.

**unsigned long int DefaultSeed()**

Returns the gsl default seed (gsl\_rng\_default\_seed)

**double Random(void)**

Returns a random floating point number uniformly distributed in the range [0.0, 1.0)

**double Random(const double p\_Lower, const double p\_Upper)**

Returns a random floating point number uniformly distributed in the range [p\_Lower, p\_Upper), where  $p\_Lower \leq p\_Upper$ .

(p\_Lower and p\_Upper will be swapped if  $p\_Lower > p\_Upper$  as passed)

**double RandomGaussian(const double p\_Sigma)**

Returns a Gaussian random variate, with mean 0.0 and standard deviation p\_Sigma

**int RandomInt(const int p\_Lower, const int p\_Upper)**

Returns a random integer number uniformly distributed in the range [p\_Lower, p\_Upper), where  $p\_Lower \leq p\_Upper$ .

(p\_Lower and p\_Upper will be swapped if  $p\_Lower > p\_Upper$  as passed)

**int RandomInt(const int p\_Upper)**

Returns a random integer number uniformly distributed in the range [0, p\_Upper), where  $0 \leq p\_Upper$ . Returns 0 if  $p\_Upper < 0$ .

## LOGGING & DEBUGGING

A logging and debugging service is provided encapsulated in a singleton object (an instantiation of the Log class).

The logging functionality was first implemented when the Single Star Evolution code was refactored, and the base-level of logging was sufficient for the needs of the SSE code. Refactoring the Binary Star Evolution code highlighted the need for expanded logging functionality. To provide for the logging needs of the BSE code, new functionality was added almost as a wrapper around the original, base-level logging functionality. Some of the original base-level logging functionality has almost been rendered redundant by the new functionality implemented for BSE code, but it remains (almost) in its entirety because it may still be useful in some circumstances.

When the base-level logging functionality was created, debugging functionality was also provided, as well as a set of macros to make debugging and the issuing of warning messages easier. A set of logging macros was also provided to make logging easier. The debug macros are still useful, and their use is encouraged (rather than inserting print statements using `std::cout` or `std::cerr`).

When the BSE code was refactored, some rudimentary error handling functionality was also provided in the form of the Errors service - an attempt at making error handling easier. Some of the functionality provided by the Errors service supersedes the `DBG_WARN*` macros provided as part of the Log class, but the `DBG_WARN*` macros are still useful in some circumstances (and in fact are still used in various places in the code). The `LOG*` macros are somewhat less useful, but remain in case the original base-level logging functionality (that which underlies the expanded logging functionality) is used in the future (as mentioned above, it could still be useful in some circumstances). The Errors service is described later in this document.

The expanded logging functionality introduces Standard Log Files - described later in this document.

### ***Base-Level Logging***

The Log class member variables are private, and public functions have been created for logging and debugging functionality required by the code.

The Log service can be accessed by referring to the `Log::Instance()` object. For example, to stop the logging service, call the `Stop()` function:

```
Log::Instance()→Stop();
```

Since that could become unwieldy, there is a convenience macro to access the Log service. The macro just defines “LOGGING” as “`Log::Instance()`”, so calling the `Stop()` function can be written as:

```
LOGGING→Stop();
```

The Log service must be initialised before logging and debugging functionality can be used. Initialise logging by calling the Start() function:

#### **LOGGING→Start(**

- |                  |                                                                                                   |
|------------------|---------------------------------------------------------------------------------------------------|
| outputPath,      | - location of logfiles                                                                            |
| containerName,   | - directory to be created at p_LogBasePath to hold all log files                                  |
| logfilePrefix,   | - prefix for logfile names (can be blank)                                                         |
| logLevel,        | - logging level (integer) (see below)                                                             |
| logClasses,      | - array of enabled logging classes (strings) (see below)                                          |
| debugLevel,      | - debug level (integer) (see below)                                                               |
| debugClasses,    | - array of enabled debug classes (strings) (see below)                                            |
| debugToLogfile,  | - flag (boolean) indicating whether debug statements should also be written to log file           |
| errorsToLogfile, | - flag (boolean) indicating whether error messages should also be written to log file             |
| delimiter        | - string (usually single character) to be used as the default field delimiter in log file records |

)

Start() returns nothing (void function).

The Log service should be stopped before exiting the program – this ensures all open log files are flushed to disk and closed properly. Stop logging by calling the Stop() function:

#### **LOGGING→Stop(**

- |             |                                                                      |
|-------------|----------------------------------------------------------------------|
| objectStats | - { number of objects (stars or binaries) requested, count created } |
|-------------|----------------------------------------------------------------------|

)

Stop() flushes any closes any open log files.

Stop() returns nothing (void function).

The Log service provides the following public member functions:

#### **void Start(**

- |                  |                                                                                                   |
|------------------|---------------------------------------------------------------------------------------------------|
| outputPath,      | - location of logfiles- the directory in which log files will be created                          |
| containerName,   | - directory to be created at p_LogBasePath to hold all log files                                  |
| logfilePrefix,   | - prefix for logfile names (can be blank)                                                         |
| logLevel,        | - logging level (integer) (see below)                                                             |
| logClasses,      | - array of enabled logging classes (strings) (see below)                                          |
| debugLevel,      | - debug level (integer) (see below)                                                               |
| debugClasses,    | - array of enabled debug classes (strings) (see below)                                            |
| debugToLogfile,  | - flag (boolean) indicating whether debug statements should also be written to log file           |
| errorsToLogfile, | - flag (boolean) indicating whether error messages should also be written to log file             |
| delimiter        | - string (usually single character) to be used as the default field delimiter in log file records |

)

Initialises the logging and debugging service. Logging parameters are set per the program options specified (using default values if no options are specified by the user). The log file container directory is created. If a directory with the name as given by the *containerName*

parameter already exists, a version number will be appended to the directory name. The Run\_Details file is created within the logfile container directory. Log files to which debug statements and error messages will be created and opened if required.

```
void Stop(  
    objectStats          - { number of objects (stars or binaries) requested, count created }  
)
```

Stops the logging and debugging service. All open log files are flushed to disk and closed (including and Standard Log Files open - see description of Standard Log Files later in this document). The Run\_Details file is populated and closed.

```
bool Enabled()
```

Returns a boolean indicating whether the Log service is enabled – true indicates the Log service is enable and available; false indicates the Log service is not enable and so not available.

```
int Open(
```

```
    logFileName,          - the name of the log file to be created and opened. This should be the  
                           filename only – the path, prefix and extensions are added by the  
                           logging service. If the file already exists, the logging service will  
                           append a version number to the name if necessary (see append  
                           parameter below).  
  
    append,               - flag (boolean) indicating whether the file should be opened in  
                           append mode (i.e. existing data is preserved) and new records  
                           written to the file appended, or whether a new file should be opened  
                           (with version number if necessary).  
  
    timeStamps,           - flag (boolean) indicating whether timestamps should be written with  
                           each log file record.  
  
    labels                - flag (boolean) indicating whether a label should be written with each  
                           log record. This is useful when different types of logging data is  
                           being written to the same log file file.  
  
    delimiter             - (optional) string (usually single character) to be used as the field  
                           delimiter in this log file. If delimiter is not provided the default  
                           delimiter is used (as parameter to Start()).  
)
```

Opens a log file. If the append parameter is true and a file name *logFilename* exists, the existing file will be opened and the existing contents retained, otherwise a new file will be created and opened (not a Standard Log File - see description of Standard Log Files later in this document).

The log file container directory is created at the path specified by the *outputPath* parameter passed to the Start() function. New log files are created in the logfile container directory. BSE

Detailed log files are created in the Detailed\_Output directory, which is created in the log file container directory if required.

The filename is prefixed by the *logfilePrefix* parameter passed to the Start() function.

The file extension is based on the *delimiter* parameter passed to the Start() function: if the delimiter is SPACE, the file extension is ".txt"; if the delimiter is TAB the file extension is ".tsv"; if the delimiter is COMMA the file extension is ".csv".

If a file with the name as given by the *logFilename* parameter already exists, and the *append* parameter is false, a version number will be appended to the filename before the extension (this functionality is largely redundant since the implementation of the log file container directory).

The log file identifier (integer) is returned to the caller - a value of -1 indicates the log file was not opened successfully. Multiple log files can be open simultaneously – referenced by the identifier returned.

**bool Close(**  
    logFileId,                   - the identifier of the log file to be closed (as returned by Open())  
**)**

Closes the log file specified by the *logFileId* parameter. If the log file specified by the *logFileId* parameter is open, it is flushed to disk and closed. The function returns a boolean indicating whether the file was closed successfully.

**bool Write(**  
    logFileId,                   - the identifier of the log file to be written  
  
    logClass,                   - string specifying the log class to be associated with the record to be written. Can be blank.  
  
    logLevel,                   - integer specifying the log level to be associated with the record to be written.  
  
    logString,                  - the string to be written to the log file.  
**)**

Writes an unformatted record to the specified log file. If the Log service is enabled and the specified log file is active, and the log class and log level passed are enabled (see discussion of log classes and levels), the string is written to the file.

The function returns a boolean indicating whether the record was written successfully. If an error occurred the log file will be disabled.

```

bool Put(
    logFileId,          - the identifier of the log file to be written

    logClass,           - string specifying the log class to be associated with the record to be
                        written. Can be blank.

    logLevel,           - integer specifying the log level to be associated with the record to be
                        written.

    logString,          - the string to be written to the log file.
)

```

Writes a minimally formatted record to the specified log file. If the Log service is enabled and the specified log file is active, and the log class and log level passed are enabled (see discussion of log classes and levels), the string is written to the file.

If labels are enabled for the log file, a label will be prepended to the record. The label text will be the *logClass* parameter.

If timestamps are enabled for the log file, a formatted timestamp is prepended to the record. The timestamp format is *yyyymmdd hh:mm:ss*.

The function returns a boolean indicating whether the record was written successfully. If an error occurred the log file will be disabled.

```

bool Debug(
    debugClass,         - string specifying the debug class to be associated with the record to
                        be written. Can be blank.

    debugLevel,         - integer specifying the debug level to be associated with the record to
                        be written.

    debugString,        - the string to be written to stdout (and optionally to file)
)

```

Writes *debugString* to stdout and, if logging is active and so configured (via program option debug-to-file), writes *debugString* to the debug log file.

The function returns a boolean indicating whether the record was written successfully. If an error occurred writing to the debug log file, the log file will be disabled.



**bool DebugWait(**

debugClass,	- string specifying the debug class to be associated with the record to be written. Can be blank.
debugLevel,	- integer specifying the debug level to be associated with the record to be written.
debugString,	- the string to be written to stdout (and optionally to file)

**)**

Writes *debugString* to stdout and, if logging is active and so configured (via program option debug-to-file), writes *debugString* to the debug log file, then waits for user input.

The function returns a boolean indicating whether the record was written successfully. If an error occurred writing to the debug log file, the log file will be disabled.

**bool Error(**

errorString,	- the string to be written to stdout (and optionally to file)
--------------	---------------------------------------------------------------

**)**

Writes *errorString* to stdout and, if logging is active and so configured (via program option errors-to-file), writes *errorString* to the error log file, then waits for user input.

The function returns a boolean indicating whether the record was written successfully. If an error occurred writing to the error log file, the log file will be disabled.

**void Squawk(**

squawkString,	- the string to be written to stderr
---------------	--------------------------------------

**)**

Writes *squawkString* to stderr.

**void Say(**

sayClass,	- string specifying the log class to be associated with the record to be written. Can be blank.
sayLevel,	- integer specifying the log level to be associated with the record to be written.
sayString,	- the string to be written to stdout

**)**

Writes *sayString* to stdout.

The filename to which debug records are written when Start() parameter “debugToLogfile” is true is declared in constants.h – see the LOGFILE enum class and associate descriptor map LOGFILE\_DESCRIPTOR. Currently the name is ‘Debug\_Log’.

## ***Extended Logging***

The Logging service was extended to support standard log files for Binary Star Evolution (SSE also uses the extended logging). The standard log files defined are:

- SSE Parameters log file
- BSE System Parameters log file
- BSE Detailed Output log file
- BSE Double Compact Objects log file
- BSE Common Envelopes log file
- BSE Supernovae log file
- BSE Pulsar Evolution log file

The Logging service maintains information about each of the standard log files, and will handle creating, opening, writing and closing the files. For each execution of the COMPAS program that evolves binary stars, one (and only one) of each of the log file listed above will be created, except for the Detailed Output log in which case there will be one log file created for each binary star evolved.

The Log service provides the following public member functions specifically for managing standard log files:

```
void LogSingleStarParameters(Star, Id)  
void LogBinarySystemParameters(Binary)  
void LogDetailedOutput(Binary, Id)  
void LogDoubleCompactObject(Binary)  
void LogCommonEnvelope(Binary)  
void LogSupernovaDetails(Binary)  
void LogPulsarEvolutionParameters(Binary)
```

Each of the BSE functions is passed a pointer to the binary star for which details are to be logged, and in the case of the Detailed Output log file, an integer identifier (typically the loop index of the binary star) that is appended to the log file name.

The SSE function is passed a pointer to the single star for which details are to be logged, and an integer identifier (typically the loop index of the star) that is appended to the log file name.

Each of the functions listed above will, if necessary, create and open the appropriate log file. Internally the Log service opens (creates first if necessary) once at first use, and keeps the files open for the life of the program.

The Log service provides a further two functions to manage standard log files:

**bool CloseStandardFile(LogFile)**

Flushes and closes the specified standard log file. The function returns a boolean indicating whether the log file was closed successfully.

**bool CloseAllStandardFiles()**

Flushes and closes all currently open standard log files. The function returns a boolean I indicating whether all standard log files were closed successfully.

Standard log file names are supplied via program options, with default values declared in constants.h.

## ***Logging & Debugging Macros***

### **Logging Macros**

The following macros are provide for logging:

#### **LOG(id, ...)**

Writes log record to log file specified by “id”. Use:

LOG(id, string)                      writes “string” to log file specified by “id”

LOG(id, level, string)              writes “string” to log file specified by “id” if “level” is <= “id” in Start()

LOG(id, class, level, string) writes “string” to log file specified by “id”  
if “class” is in “logClasses” in Start() and  
if “level” is <= “logLevel” in Start()

default “class” is “”; default “level” is 0

Examples:

```
LOG(SSEfileId, “This is a log record”);  
LOG(OutputFile2Id, “The value of x is “ << x << “ km”);  
LOG(MyLogfileId, 2, “Log string”);  
LOG(SSEfileId, “CHeB”, 4, “This is a CHeB only log record”);
```

**LOG\_ID(id, ...)**

Writes log record prepended with calling function name to log file. Use:

LOG_ID(id)	writes name of calling function to log file specified by “id”
LOG_ID(id, string)	writes “string” prepended with name of calling function to log file specified by “id”
LOG_ID(id, level, string)	writes “string” prepended with name of calling function to log file specified by “id” if “level” is ≤ “logLevel” in Start()
LOG_ID(id, class, level, string)	writes “string” prepended with name of calling function to log file specified by “id” if “class” is in “logClasses” in Start() and if “level” is ≤ “logLevel” in Start()

default “class” is “”; default “level” is 0

Examples:

```
LOG_ID(Outf1Id)
LOG_ID(Outf2Id, “This is a log record”);
LOG_ID(MyLogfileId, “The value of x is “ << x << “ km”);
LOG_ID(OutputFile2Id, 2, “Log string”);
LOG_ID(CHeBfileId, “CHeB”, 4, “This is a CHeB only log record”);
```

**LOG\_IF(id, cond, ...)**

Writes log record to log file if the condition given by “cond” is met. Use:

LOG_IF(id, cond, string)	writes “string” to log file specified by “id” if “cond” is true
LOG_IF(id, cond, level, string)	writes “string” to log file specified by “id” if “cond” is true and if “level” is ≤ “logLevel” in Start()
LOG_IF(id, cond, class, level, string)	writes “string” to log file specified by “id” if “cond” is true and if “class” is in “logClasses” in Start() and if “level” is ≤ “logLevel” in Start()

“cond” is any logical statement and is required; default “class” is “”; default “level” is 0

Examples:

```
LOG_IF(MyLogfileId, a > 1.0, “This is a log record”);
LOG(SSEfileId, (b == c && a > x), “The value of x is “ << x << “ km”);
LOG(CHeBfileId, flag, 2, “Log string”);
LOG(SSEfileId, (x >= y), “CHeB”, 4, “This is a CHeB only log record”);
```

**LOG\_ID\_IF(id, ...)**

Writes log record prepended with calling function name to log file if the condition given by “cond” is met. Use: see LOG\_ID(id, ...) and LOG\_IF(id, cond, ...) above.

The logging macros described above are provided in a verbose variant. The verbose macros function the same way as their non-verbose counterparts, with the added functionality that the log records written to the log file will be reflected on stdout as well. The verbose logging macros are:

**LOGV(id, ...)****LOGV\_ID(id, ...)****LOGV\_IF(id, cond, ...)****LOGV\_ID\_IF(id, cond, ...)**

A further four macros are provided that allow writing directly to stdout rather than the log file. These are:

**SAY(...)****SAY\_ID(...)****SAY\_IF(cond, ...)****SAY\_ID\_IF(cond, ...)**

The SAY macros function the same way as their LOG counterparts, but write directly to stdout instead of the log file. The SAY macros honour the logging classes and level.

## Debugging Macros

A similar set of macros is also provided for debugging purposes.

The debugging macros write directly to stdout rather than the log file, but their output can also be written to the log file if desired (see the debugToLogfile parameter of Start(), and the debug-to-file program option described above).

A major difference between the logging macros and the debugging macros is that the debugging macros can be defined away. The debugging macro definitions are enclosed in an #ifdef enclosure, and are only present in the source code if #DEBUG is defined. This means that if #DEBUG is not defined (#undef), all debugging statements using the debugging macros will be removed from the source code by the preprocessor before the source is compiled. Un-defining #DEBUG not only prevents bloat of unused code in the executable, it improves performance. Many of the functions in the code are called hundreds of thousands, if not millions, of times as the stellar evolution proceeds. Even if the debugging classes and debugging level are set so that no debug statement is displayed, just checking the debugging level every time a function is called increases the run-time of the program. The suggested use is to enable the debugging macros (#define DEBUG) while developing new code, and disable them (#undef DEBUG) to produce a production version of the executable.

The debugging macros provided are:

<b>DBG(...)</b>	analogous to the LOG(...) macro
<b>DBG_ID(...)</b>	analogous to the LOG_ID(...) macro
<b>DBG_IF(cond, ...)</b>	analogous to the LOG_IF(...) macro
<b>DBG_ID_IF(cond, ...)</b>	analogous to the LOG_ID_IF(...) macro

Two further debugging macros are provided:

<b>DBG_WAIT(...)</b>
<b>DBG_WAIT_IF(cond, ...)</b>

The DBG\_WAIT macros function in the same way as their non-wait counterparts (DBG(...) and DBG\_IF(cond, ...)) with the added functionality that they will pause execution of the program and wait for user input before proceeding.

A set of macros for printing warning message is also provided. These are the DBG\_WARN macros:

<b>DBG_WARN(...)</b>	analogous to the LOG(...) macro
<b>DBG_WARN_ID(...)</b>	analogous to the LOG_ID(...) macro
<b>DBG_WARN_IF(...)</b>	analogous to the LOG_IF(...) macro
<b>DBG_WARN_ID_IF(...)</b>	analogous to the LOG_ID_IF(...) macro

The DBG\_WARN macros write to stdout via the SAY macro, so honour the logging classes and level, and are not written to the debug or errors files.

Note that the “id” parameter of the “LOG” macros (to specify the logfileId) is not required for the DBG macros (the filename to which debug records are written is declared in constants.h – see the LOGFILE enum class and associated descriptor map LOGFILE\_DESCRIPTOR).

## ERROR HANDLING

An error handling service is provided encapsulated in a singleton object (an instantiation of the Errors class).

The Errors service provides global error handling functionality. Following is a brief description of the Errors service (full documentation coming soon...):

Errors are defined in the error catalog in constants.h (see ERROR\_CATALOG). It could be useful to move the catalog to a file so it can be changed without changing the code, or even have multiple catalogs provided for internationalisation – a task for later.

Errors defined in the error catalog have a scope and message text. The scope is used to determine when/if an error should be printed.

The current values for scope are:

NEVER	the error will not be printed
ALWAYS	the error will always be printed
FIRST	the error will be printed only on the first time it is encountered anywhere in the program
FIRST_IN_OBJECT_TYPE	the error will be printed only on the first time it is encountered anywhere in objects of the same type (e.g. Binary Star objects)
FIRST_IN_STELLAR_TYPE	the error will be printed only on the first time it is encountered anywhere in objects of the same stellar type (e.g. HeWD Star objects)
FIRST_IN_OBJECT_ID	the error will be printed only on the first time it is encountered anywhere in an object instance
FIRST_IN_FUNCTION	the error will be printed only on the first time it is encountered anywhere in the same function of an object instance (i.e. will print more than once if encountered in the same function name in different objects)

The Errors service provides methods to print both warnings and errors - essentially the same thing, but warning messages are prefixed with "WARNING:", whereas error messages are prefixed with "ERROR:".



Errors and warnings are printed by using the macros defined in ErrorsMacros.h. They are:

### **Error macros:**

#### **SHOW\_ERROR(error\_number)**

Prints "ERROR: " followed by the error message associated with "error\_number" (from the error catalog)

#### **SHOW\_ERROR(error\_number, error\_string)**

Prints "ERROR: " followed by the error message associated with "error\_number" (from the error catalog), and appends "error\_string"

#### **SHOW\_ERROR\_IF(cond, error\_number)**

If "cond" is TRUE, prints "ERROR: " followed by the error message associated with "error\_number" (from the error catalog)

#### **SHOW\_ERROR\_IF(cond, error\_number, error\_string)**

If "cond" is TRUE, prints "ERROR: " followed by the error message associated with "error\_number" (from the error catalog), and appends "error\_string"

### **Warning macros:**

#### **SHOW\_WARN(error\_number)**

Prints "WARNING: " followed by the error message associated with "error\_number" (from the error catalog)

#### **SHOW\_WARN(error\_number, error\_string)**

Prints "WARNING: " followed by the error message associated with "error\_number" (from the error catalog), and appends "error\_string"

#### **SHOW\_WARN\_IF(cond, error\_number)**

If "cond" is TRUE, prints "WARNING: " followed by the error message associated with "error\_number" (from the error catalog)

#### **SHOW\_WARN\_IF(cond, error\_number, error\_string)**

If "cond" is TRUE, prints "WARNING: " followed by the error message associated with "error\_number" (from the error catalog), and appends "error\_string"

Error and warning message always contain:

- The object id of the calling object

- The object type of the calling object

- The stellar type of the calling object (will be "NONE" if the calling object is not a star-type object)

- The function name of the calling function

Any object that uses the Errors service (i.e. the SHOW\_\* macros) must expose the following functions:

OBJECT_ID	ObjectId() const	{ return m_ObjectId; }
OBJECT_TYPE	ObjectType() const	{ return m_ObjectType; }
STELLAR_TYPE	StellarType() const	{ return m_StellarType; }

These functions are called by the SHOW\_\* macros. If any of the functions are not applicable to the object, then they must return ">::NONE (all objects should implement ObjectId() correctly).

The filename to which error records are written when Start() parameter “errorsToLogfile” is true is declared in constants.h – see the LOGFILE enum class and associate descriptor map LOGFILE\_DESCRIPTOR. Currently the name is ‘Error\_Log’.

## FLOATING-POINT COMPARISONS

Floating-point comparisons are inherently problematic. Testing floating-point numbers for equality, or even inequality, is fraught with problems due to the internal representation of floating-point numbers: floating-point numbers are stored with a fixed number of binary digits, which limits their precision and accuracy. The problems with floating-point comparisons are even more evident if one or both of the numbers being compared are the results of (perhaps several) floating-point operations (rather than comparing constants).

To avoid the problems associated with floating-point comparisons it is (almost always) better to do any such comparisons with a tolerance rather than an absolute comparison. To this end, a floating-point comparison function has been provided, and (almost all of) the floating-point comparisons in the code have been changed to use that function. The function uses both an absolute tolerance and a relative tolerance, which are both declared in constants.h. Whether the function uses a tolerance or not can be changed by #define-ing or #undef-ing the “COMPARE\_WITH\_TOLERANCE” flag in constants.h (so the change is a compile-time change, not run-time).

The compare function is defined in utils.h and is implemented as follows:

```
static int Compare(const double p_X, const double p_Y) {
#ifdef COMPARE_WITH_TOLERANCE
    return (fabs(p_X - p_Y) <= max(FLOAT_TOLERANCE_ABSOLUTE,
                                   FLOAT_TOLERANCE_RELATIVE *
                                   max(fabs(p_X),
                                       fabs(p_Y)))) ? 0 : (p_X < p_Y ? -1 : 1);
#else
    return (p_X == p_Y) ? 0 : (p_X < p_Y ? -1 : 1);
#endif
}
```

If COMPARE\_WITH\_TOLERANCE is defined, p\_X and p\_Y are compared with tolerance values, whereas if COMPARE\_WITH\_TOLERANCE is not defined the comparison is an absolute comparison.

The function returns an integer indicating the result of the comparison:

- 1 indicates that p\_X is considered to be less than p\_Y
- 0 indicates p\_X and p\_Y are considered to be equal
- +1 indicates that p\_X is considered to be greater than p\_Y

The comparison is done using both an absolute tolerance and a relative tolerance. The tolerances can be defined to be the same number, or different numbers. If the relative tolerance is defined as 0.0, the comparison is done using the absolute tolerance only, and if the absolute tolerance is defined as 0.0 the comparison is done with the relative tolerance only.

Absolute tolerances are generally more effective when the numbers being compared are small – so using an absolute tolerance of (say) 0.0000005 is generally effective when comparing single-digit numbers (or so), but is less effective when comparing numbers in the thousands or millions. For comparisons of larger numbers a relative tolerance is generally more effective (the actual tolerance is wider because the relative tolerance is multiplied by the larger absolute value of the numbers being compared).

There is a little overhead in the comparisons even when the tolerance comparison is disabled, but it shouldn't be prohibitive.

## CONSTANTS FILE – constants.h

As well as plain constant values, many distribution and prescription identifiers are declared in constants.h. These are mostly declared as enum classes, with each enum class having a corresponding

map of labels. The benefit is that the values of a particular (e.g.) prescription are limited to the values declared in the enum class, rather than any integer value, so the compiler will complain if an incorrect value is inadvertently used to reference that prescription.

For example, the Common Envelope Lambda Prescriptions are declared in constants.h thus:

```
enum class CE_LAMBDA_PRESCRIPTION: int {
    FIXED, LOVERIDGE, NANJING, KRUCKOW, DEWI
};

const std::unordered_map<CE_LAMBDA_PRESCRIPTION, std::string>
CE_LAMBDA_PRESCRIPTION_LABEL = {
    { CE_LAMBDA_PRESCRIPTION::FIXED,      "LAMBDA_FIXED" },
    { CE_LAMBDA_PRESCRIPTION::LOVERIDGE,  "LAMBDA_LOVERIDGE" },
    { CE_LAMBDA_PRESCRIPTION::NANJING,    "LAMBDA_NANJING" },
    { CE_LAMBDA_PRESCRIPTION::KRUCKOW,    "LAMBDA_KRUCKOW" },
    { CE_LAMBDA_PRESCRIPTION::DEWI,       "LAMBDA_DEWI" }
};
```

Note that the values allowed for variables of type CE\_LAMBDA\_PRESCRIPTION are limited to FIXED, LOVERIDGE, NANJING, KRUCKOW and DEWI – anything else will cause a compiler error.

The unordered map CE\_LAMBDA\_PRESCRIPTION\_LABEL is indexed by CE\_LAMBDA\_PRESCRIPTION and declares a string label for each CE\_LAMBDA\_PRESCRIPTION. The strings declared in CE\_LAMBDA\_PRESCRIPTION\_LABEL are used by the Options service to match user input to the required CE\_LAMBDA\_PRESCRIPTION. These strings can also be used if an English description of the value of a variable is required: instead of just printing an integer value that maps to a CE\_LAMBDA\_PRESCRIPTION, the string label associated with the prescription can be printed.

Stellar types are also declared in constants.h via an enum class and associate label map. This allows stellar types to be referenced using symbolic names rather than an ordinal number. The stellar types enum class is STELLAR\_TYPE, and is declared as:

```
enum class STELLAR_TYPE: int {
    MS_LTE_07,
    MS_GT_07,
    HERTZSPRUNG_GAP,
    FIRST_GIANT_BRANCH,
    CORE_HELIUM_BURNING,
    EARLY_ASYMPTOTIC_GIANT_BRANCH,
    THERMALLY_PULSING_ASYMPTOTIC_GIANT_BRANCH,
    NAKED_HELIUM_STAR_MS,
    NAKED_HELIUM_STAR_HERTZSPRUNG_GAP,
    NAKED_HELIUM_STAR_GIANT_BRANCH,
    HELIUM_WHITE_DWARF,
    CARBON_OXYGEN_WHITE_DWARF,
    OXYGEN_NEON_WHITE_DWARF,
    NEUTRON_STAR,
    BLACK_HOLE,
    MASSLESS_REMNANT,
    CHEMICALLY_HOMOGENEOUS,
    STAR,
    BINARY_STAR,
    NONE
};
```

Ordinal numbers can still be used to reference the stellar types, and because of the order of definition in the enum class the ordinal numbers match those given in Hurley et al. 2000.

The label map STELLAR\_TYPE\_LABEL can be used to print text descriptions of the stellar types, and is declared as:

```
const std::unordered_map<STELLAR_TYPE, std::string> STELLAR_TYPE_LABEL = {
    { STELLAR_TYPE::MS_LTE_07, "Main_Sequence_<= 0.7" },
    { STELLAR_TYPE::MS_GT_07, "Main_Sequence_> 0.7" },
    { STELLAR_TYPE::HERTZSPRUNG_GAP, "Hertzsprung_Gap" },
    { STELLAR_TYPE::FIRST_GIANT_BRANCH, "First_Giant_Branch" },
    { STELLAR_TYPE::CORE_HELIUM_BURNING, "Core_Helium_Burning" },
    { STELLAR_TYPE::EARLY_ASYMPTOTIC_GIANT_BRANCH, "Early_Asymptotic_Giant_Branch" },
    { STELLAR_TYPE::THERMALLY_PULSING_ASYMPTOTIC_GIANT_BRANCH, "Thermally_Pulsing_Asymptotic_Giant_Branch" },
    { STELLAR_TYPE::NAKED_HELIUM_STAR_MS, "Naked_Helium_Star_MS" },
    { STELLAR_TYPE::NAKED_HELIUM_STAR_HERTZSPRUNG_GAP, "Naked_Helium_Star_Hertzsprung_Gap" },
    { STELLAR_TYPE::NAKED_HELIUM_STAR_GIANT_BRANCH, "Naked_Helium_Star_Giant_Branch" },
    { STELLAR_TYPE::HELIUM_WHITE_DWARF, "Helium_White_Dwarf" },
    { STELLAR_TYPE::CARBON_OXYGEN_WHITE_DWARF, "Carbon-Oxygen_White_Dwarf" },
    { STELLAR_TYPE::OXYGEN_NEON_WHITE_DWARF, "Oxygen-Neon_White_Dwarf" },
    { STELLAR_TYPE::NEUTRON_STAR, "Neutron_Star" },
    { STELLAR_TYPE::BLACK_HOLE, "Black_Hole" },
    { STELLAR_TYPE::MASSLESS_REMNANT, "Massless_Remnant" },
    { STELLAR_TYPE::CHEMICALLY_HOMOGENEOUS, "Chemically_Homogeneous" },
    { STELLAR_TYPE::STAR, "Star" },
    { STELLAR_TYPE::BINARY_STAR, "Binary_Star" },
    { STELLAR_TYPE::NONE, "Not_a_Star!" }
};
```

## PROGRAMMING STYLE AND CONVENTIONS

Everyone has their own preferences and style, and the nature of a project such as COMPAS will reflect that. However, there is a need to suggest some guidelines for programming style, naming conventions etc. Following is a description of some of the elements of programming style and naming conventions used to develop COMPAS v2. These may evolve over time.

### Object-Oriented Programming

COMPAS is written in C++, an object-oriented programming (OOP) language, and OOP concepts and conventions should apply throughout the code. There are many texts and web pages devoted to understanding C++ and OOP – following is a brief description of the key OOP concepts:

#### **Abstraction**

For any entity, product, or service, the goal of abstraction is to handle the complexity of the implementation by hiding details that don't need to be known in order to use, or consume, the entity, product, or service. In the OOP paradigm, hiding details in this way enables the consumer to implement more complex logic on top of the provided abstraction without needing to understand the hidden implementation details and complexity. (There is no suggestion that consumers shouldn't understand the implementation details, but they shouldn't need to in order to consume the entity, product, or service).

Abstraction in C++ is achieved via the use of *objects* – an object is an instance of a *class*, and typically corresponds to a real-world object or entity (in COMPAS, usually a star or binary star). An object maintains the state of an object (via class member variables), and provides all necessary means of changing the state of the object (by exposing public class member functions (methods)). A class may expose public functions to allow consumers to determine the value of class member variables (“getters”), and to set the value of class member variables (“setters”).

#### **Encapsulation**

Encapsulation binds together the data and functions that manipulate the data in an attempt to keep both safe from outside interference and accidental misuse. An encapsulation paradigm that does not allow calling code to access internal object data and permits access through functions only is a strong form of abstraction. C++ allows developers to enforce access restrictions explicitly by defining class member variables and functions as *private*, *protected*, or *public*. These keywords are used throughout COMPAS to enforce encapsulation.

There are very few circumstances in which a consumer should change the value of a class member variable directly (via the use of a setter function) – almost always consumers should present new situational information to an object (via a public member function), and allow the object to respond to the new information. For example, in COMPAS, there should be almost no reason for a consumer of a star object to directly change (say) the radius of the star – the consumer should inform the star object of new circumstances or events, and allow the star object to respond to those events (perhaps changing the value of the radius of the star). Changing a single class member variable directly introduces the possibility that related class member variables (e.g. other attributes of stars) will not be changed accordingly. Moreover, developers changing the code in the future should, in almost all cases, expect that the state of an object is maintained consistently by the object, and that there should be no unexpected side-effects caused by calling non class-member functions. In short, changing the state of an object outside the object is potentially unsafe and should be avoided where possible.

#### **Inheritance**

Inheritance allows classes to be arranged in a hierarchy that represents *is-a-type-of* relationships. All *non-private* class member variables and functions of the parent (base) class are available to the child (derived) class (and, therefore, child classes of the child class). This allows easy re-use of the same

procedures and data definitions, in addition to describing real-world relationships in an intuitive way. C++ allows multiple inheritance – a class may inherit from multiple parent classes.

Derived classes can define additional class member variables (using the *private*, *protected*, and *public* access restrictions), which will be available to any descendent classes (subject to inheritance rules), but will only be available to ancestor classes via the normal access methods (getters and setters).

## **Polymorphism**

Polymorphism means *having many forms*. In OOP, polymorphism occurs when there is a hierarchy of classes and they are related by inheritance.

Following the discussion above regarding inheritance, in the OOP paradigm, and C++ specifically, derived classes can override methods defined by ancestor classes, allowing a derived class to implement functions specific to its circumstances. This means that a call to a class member function will cause a different function to be executed depending on the type of object that invokes the function. Descendent classes of a class that has overridden a base class member function inherit the overridden function (but can override it themselves).

COMPAS makes heavy use of inheritance and polymorphism, especially for the implementation of the different stellar types.

## **Programming Style**

The goal of coding to a suggested style is readability and maintainability – if many developers implement code in COMPAS with their own coding style, readability and maintainability will be more difficult than if a consistent style is used throughout the code. Strict adherence isn't really necessary, but it will make it easier on all COMPAS developers if the coding style is consistent throughout.

## **Comments**

An old, but good, rule-of-thumb is that any file that contains computer code should be about one-third code, one-third comments, and one-third white space. Adhering to this rule-of-thumb just makes the code a bit easier on the eye, and provides some description (at least of the intention) of the implementation.

## **Braces**

The placement of braces in C++ code (actually, any code that uses braces to enclose scope) is a contentious issue, with many developers having long-held, often dogmatic preferences. COMPAS (so far) uses the K&R style ("the one true brace style") - the style used in the original Unix kernel and Kernighan and Ritchie's book *The C Programming Language*.

The K&R style puts the opening brace on the same line as the control statement:

```
while (x == y) {
    something();
    somethingelse();
}
```

Note also the space between the keyword *while* and the opening parenthesis, and the closing parenthesis and the opening brace.

## ***Indentation***

There is ongoing debate in the programming community as to whether indentation should be achieved using spaces or tabs (strange, but true...). The use of spaces is more common. COMPAS (so far) has a mix of both – whatever is convenient (pragmatism is your friend...).

COMPAS uses an indentation size of 4 spaces.

## ***Function Parameters***

In most cases, function parameters should be input only – meaning that the values of function parameters should not be changed by the function. Anything that needs to be changed and returned to the caller should be returned as a functional return. There are a few exceptions to this in COMPAS – all were done for performance reasons, and are documented in the code.

To avoid unexpected side-effects, developers should expect (in most cases) that any variables they pass to a function will remain unchanged – all changes should be returned as a functional return.

## ***Performance & Optimisation***

In general COMPAS developers should code for performance – within reason. Bear in mind that many functions will be called many, many thousands of times (in some cases, millions) in one execution of the program.

- Avoid calculating values inside loops that could be calculated once outside the loop.
- Try to use constants where possible.
- Use multiplication in preference to functions such as *pow()* and *sqrt()* (note that *pow()* is very expensive computationally; *sqrt()* is expensive, but much less expensive than *pow()*).
- Don't optimise to the point that readability and maintainability is compromised. Bear in mind that most compilers are good at optimising, and are very forgiving of less-than-optimally-written code (though they are not miracle workers...).



## Naming Conventions

COMPAS (so far) uses the following naming conventions:

- all variable names should be in camelCase – don't use underscore `_` to separate words
- function names should be in camelCase, beginning with an uppercase letter. Function names should be descriptive.
- class member variable names are prefixed with “`m_`”, and the character immediately following the prefix should be uppercase (in most cases – sometimes, for well-known names or words that are always written in lowercase, lowercase might be used)
- local variable names are just camelCase, beginning with a lowercase letter (again, with the caveat that sometimes, for well-known names or words that are always written in uppercase, uppercase might be used)
- function parameter names are prefixed with “`p_`”, and the character immediately following the prefix should be uppercase (again, with the caveat that sometimes, for well-known names or words that are always written in lowercase, lowercase might be used)

## **COMPILATION & REQUIREMENTS**

Please refer to the COMPAS Getting Started guide.

## Appendix A – Program Options

### **-h [ --help ]**

Prints COMPAS help.

### **-v [ --version ]**

Print COMPAS version string.

### **--allow-rlof-at-birth**

Allow binaries that have one or both stars in RLOF at birth to evolve as overcontact systems.

Default = FALSE

### **--allow-touching-at-birth**

Allow binaries that are touching at birth to be included in the sampling.

Default = FALSE

### **--angularMomentumConservationDuringCircularisation**

Conserve angular momentum when binary is circularised when entering a Mass Transfer episode.

Default = FALSE

### **--circulariseBinaryDuringMassTransfer**

Circularise binary when it enters a Mass Transfer episode.

Default = FALSE

### **--common-envelope-allow-main-sequence-survive**

Allow main sequence donors to survive common envelope evolution.

Default = FALSE

### **--debug-to-file**

Write debug statements to file.

Default = FALSE

### **--detailedOutput**

Print detailed output to file.

Default = FALSE

### **--errors-to-file**

Write error messages to file.

Default = FALSE

### **--evolve-pulsars**

Evolve pulsar properties of Neutron Stars.

Default = FALSE

### **--evolve-unbound-systems**

Continue evolving stars even if the binary is disrupted.

Default = FALSE

### **--lambda-calculation-every-timeStep**

Calculate all values of lambda at each timestep.

Default = FALSE

### **--massTransfer**

Enable mass transfer.

Default = TRUE

**--pair-instability-supernovae**

Enable pair instability supernovae (PISN).

Default = FALSE

**--populationDataPrinting**

Print details of population.

Default = FALSE

**--print-bool-as-string**

Print boolean properties as 'TRUE' or 'FALSE'.

Default = FALSE

**--pulsational-pair-instability**

Enable mass loss due to pulsational-pair-instability (PPI).

Default = FALSE

**--quiet**

Suppress printing to stdout.

Default = FALSE

**--revised-energy-formalism-Nandez-Ivanova**

Enable revised energy formalism of Nandez & Ivanova.

Default = FALSE

**--single-star**

Evolve single star(s).

Default = FALSE

**--use-mass-loss**

Enable mass loss.

Default = FALSE

**--zeta-calculation-every-timestep**

Calculate all values of MT zetas at each timestep.

Default = FALSE

**--random-seed**

Random seed.

Default = 0

**--debug-level**

Determines which print statements are displayed for debugging.

Default = 0

**--log-level**

Determines which print statements are included in the logfile.

Default = 0

**--maximum-number-timestep-iterations**

Maximum number of timesteps to evolve binary.

Default = 99999

**-n [ --number-of-binaries ]**

Specify the number of binaries to simulate.

Default = 10

**--single-star-mass-steps**

Specify the number of mass steps for single star evolution.

Default = 100

**--common-envelope-alpha**

Common Envelope efficiency alpha.

Default = 1.0

**--common-envelope-alpha-thermal**

Thermal energy contribution to the total envelope binding energy.

Defined such that  $\lambda = \alpha_{th} * \lambda_b + (1.0 - \alpha_{th}) * \lambda_g$ .

Default = 1.0

**--common-envelope-lambda**

Common Envelope lambda.

Default = 0.1

**--common-envelope-lambda-multiplier**

Multiply lambda by some constant.

Default = 1.0

**--common-envelope-mass-accretion-constant**

Value of mass accreted by NS/BH during common envelope evolution if assuming all NS/BH accrete same amount of mass.

Used when `--common-envelope-mass-accretion-prescription = CONSTANT`, ignored otherwise.

Default = 0.0

**--common-envelope-mass-accretion-max**

Maximum amount of mass accreted by NS/BHs during common envelope evolution (Msol).

Default = 0.1

**--common-envelope-mass-accretion-min**

Minimum amount of mass accreted by NS/BHs during common envelope evolution (Msol).

Default = 0.04

**--common-envelope-recombination-energy-density**

Recombination energy density (ergs/g).

Default = 1.5E+13

**--common-envelope-slope-Kruckow**

Common Envelope slope for Kruckow lambda.

Default = -0.8

**--eccentricity-max**

Maximum eccentricity to generate.

Default = 1.0

**--eccentricity-min**

Minimum eccentricity to generate.

Default = 0.0

**--eddington-accretion-factor**

Multiplication factor for Eddington accretion for NS & BH, i.e. >1 is super-eddington and 0 is no accretion.

Default = 1.0

**--fix-dimensionless-kick-velocity**

Fix dimensionless kick velocity  $u_k$  to this value.

Default = n/a (not used if option not present)

**--initial-mass-max**

Maximum mass (in  $M_{\text{sol}}$ ) to generate using given IMF.

Default = 100.0

**--initial-mass-min**

Minimum mass (in  $M_{\text{sol}}$ ) to generate using given IMF.

Default = 8.0

**--initial-mass-power**

Single power law power to generate primary mass using given IMF.

Default = -2.3

**--kick-direction-power**

Power for power law kick direction distribution.

Default = 0.0 = isotropic, +ve = polar, -ve = in plane

**--kick-scaling-factor**

Arbitrary factor used to scale kicks.

Default = 1.0

**--kick-velocity-max**

Maximum drawn kick velocity in  $\text{km s}^{-1}$ .

Must be > 0 if using --kick-velocity-distribution=FLAT

Default = -1.0

**--kick-velocity-sigma-CCSN-BH**

Sigma for chosen kick velocity distribution for black holes.  
Default = 250.0 km s<sup>-1</sup>

**--kick-velocity-sigma-CCSN-NS**

Sigma for chosen kick velocity distribution for neutron stars.  
Default = 250.0 km s<sup>-1</sup>

**--kick-velocity-sigma-ECSN**

Sigma for chosen kick velocity distribution for ECSN.  
Default = 30.0 km s<sup>-1</sup>

**--kick-velocity-sigma-USSN**

Sigma for chosen kick velocity distribution for USSN.  
Default = 30.0 km s<sup>-1</sup>

**--luminous-blue-variable-multiplier**

Multiplicative constant for LBV mass loss.  
Default = 1.5, use 10 for Mennekens & Vanbeveren 2014

**--mass-ratio-max**

Maximum mass ratio m<sub>2</sub>/m<sub>1</sub> to generate.  
Default = 1.0

**--mass-ratio-min**

Minimum mass ratio m<sub>2</sub>/m<sub>1</sub> to generate.  
Default = 0.0

**--mass-transfer-fa**

Mass Transfer fraction accreted, when the FIXED mass transfer prescription is used.  
Default = 1.0 (fully conservative)

**--mass-transfer-jloss**

Specific angular momentum with which the non-accreted system leaves the system.  
Used when --mass-transfer-angular-momentum-loss-prescription = ARBITRARY, ignored otherwise.  
Default = 1.0

**--mass-transfer-thermal-limit-C**

Mass Transfer Thermal rate factor for the accretor.  
Default = 10.0

**--maximum-evolution-time**

Maximum time to evolve binaries (Myr).  
Default = 13700.0

**--maximum-mass-donor-Nandez-Ivanova**

Maximum donor mass allowed for the revised common envelope formalism of Nandez & Ivanova (M<sub>sol</sub>).  
Default = 2.0

**--maximum-neutron-star-mass**

Maximum mass of a neutron star (Msol).

Default = 3.0

**--MCBUR1**

minimum core mass at base of the AGB to avoid fully degenerate CO core formation (Msol); e.g., 1.6 in Hurley prescription, 1.83 in Fryer and Belczynski models to

Default = 1.6

**-z [ --metallicity ]**

Metallicity (Zsol)

Default = 0.02

**--minimum-secondary-mass**

Minimum mass of secondary to generate (Msol).

Default = 0.0

**--neutrino-mass-loss-bh-formation-value**

Amount of mass lost in neutrinos during BH formation (either as fraction or in solar masses, depending on neutrino-mass-loss-bh-formation setting).

Default = 0.1

**--orbital-period-max**

Maximum period to generate (days)

Default = 1000.0

**--orbital-period-min**

Minimum period to generate (days).

Default = 1.1

**--PISN-lower-limit**

Minimum core mass for PISN (Msol).

Default = 60.0

**--PISN-upper-limit**

Maximum core mass for PISN (Msol).

Default = 135.0

**--PPI-lower-limit**

Minimum core mass for PPI (Msol).

Default = 35.0

**--PPI-upper-limit**

Maximum core mass for PPI (Msol).

Default = 60.0

**--pulsar-birth-magnetic-field-distribution-max**

Maximum (log10) pulsar birth magnetic field.

Default = 13.0

**--pulsar-birth-magnetic-field-distribution-min**

Minimum (log10) pulsar birth magnetic field.

Default = 11.0



**--pulsar-birth-spin-period-distribution-max**

Maximum pulsar birth spin period (ms).

Default = 100.0

**--pulsar-birth-spin-period-distribution-min**

Minimum pulsar birth spin period (ms).

Default = 0.0

**--pulsar-magnetic-field-decay-massscale**

Mass scale on which magnetic field decays during accretion (Msol).

Default = 0.025

**--pulsar-magnetic-field-decay-timescale**

Timescale on which magnetic field decays (Myr).

Default = 1000.0

**--pulsar-minimum-magnetic-field**

log10 of the minimum pulsar magnetic field (Gauss).

Default = 8.0

**--semi-major-axis-max**

Maximum semi-major axis to generate (AU).

Default = 1000.0

**--semi-major-axis-min**

Minimum semi-major axis to generate (AU).

Default = 0.1

**--single-star-mass-max**

Maximum mass for single star evolution (Msol).

Default = 100.0

**--single-star-mass-min**

Minimum mass for single star evolution (Msol).

Default = 5.0

**--wolf-rayet-multiplier**

Multiplicative constant for WR winds.

Default = 1.0

**--zeta-adiabatic-arbitrary**

Value of logarithmic derivative of radius with respect to mass, zeta adiabatic.

Default = 1.0E+04

**--zeta-radiative-giant-star**

Value of logarithmic derivative of radius with respect to mass, zeta for radiative-envelope giant-like stars, including Hertzsprung gap stars.

Default = 6.5

**--zeta-main-sequence**

Value of logarithmic derivative of radius with respect to mass, zeta on the main sequence.

Default = 2.0

**--black-hole-kicks**

Black hole kicks relative to NS kicks (options: FULL, REDUCED, ZERO, FALLBACK).

Default = FALLBACK

**--case-bb-stability-prescription**

Prescription for the stability of case BB/BC mass transfer (options: ALWAYS\_STABLE, ALWAYS\_STABLE\_ONTO\_NSBH, TREAT\_AS\_OTHER\_MT, NEVER\_STABLE).

Default = ALWAYS\_STABLE

**--chemically-homogeneous-evolution**

Chemically Homogeneous Evolution mode (options: NONE, OPTIMISTIC, PESSIMISTIC).

Default = NONE

**--common-envelope-lambda-prescription**

CE lambda prescription.

(options:

LAMBDA\_FIXED, LAMBDA\_LOVERIDGE, LAMBDA\_NANJING, LAMBDA\_KRUCKOW, LAMBDA\_DEWI)

Default = LAMBDA\_NANJING

**--common-envelope-mass-accretion-prescription**

Assumption about whether NS/BHs can accrete mass during common envelope evolution.

(options: ZERO, CONSTANT, UNIFORM, MACLEOD)

Default = ZERO

**--envelope-state-prescription**

Prescription for determining whether the envelope of the star is convective or radiative.

(options: LEGACY, HURLEY, FIXED\_TEMPERATURE)

Default = LEGACY

**--stellar-zeta-prescription**

Prescription for stellar zeta.

(options: STARTRACK, SOBERMAN, HURLEY, ARBITRARY)

Default = SOBERMAN

**-e [ --eccentricity-distribution ]**

Initial eccentricity distribution, e.

(options:

ZERO, FIXED, FLAT, THERMALISED, GELLER+2013, THERMAL, DUQUENNOY MAYOR1991, SANA2012, IMPORTANCE).

Default = ZERO

**--fryer-supernova-engine**

Supernova engine type if using Fryer et al. 2012 fallback prescription.

(options: DELAYED, RAPID).

Default = DELAYED

**--grid**

Grid filename.

Default = ''

**-i [ --initial-mass-function ]**

Initial mass function, (options: SALPETER, POWERLAW, UNIFORM, KROUPA)

Default = KROUPA

**--kick-direction**

Natal kick direction distribution.

(options: ISOTROPIC, INPLANE, PERPENDICULAR, POWERLAW, WEDGE, POLES)

Default = ISOTROPIC

**--kick-velocity-distribution**

Natal kick velocity distribution.

(options:

ZERO, FIXED, FLAT, MAXWELLIAN, MUELLER2016, MUELLER2016MAXWELLIAN, BRAYEL-  
DRIDGE, MULLERMANDEL)

Default = MAXWELLIAN

**--logfile-BSE-common-envelopes**

Filename for BSE Common Envelopes logfile.

Default = 'BSE\_Common\_Envelopes'

**--logfile-BSE-detailed-output**

Filename for BSE Detailed Output logfile.

Default = 'BSE\_Detailed\_Output'

**--logfile-BSE-double-compact-objects**

Filename for BSE Double Compact Objects logfile.

Default = 'BSE\_Double\_Compact\_Objects'

**--logfile-BSE-pulsar-evolution**

Filename for BSE Pulsar Evolution logfile.

Default = 'BSE\_Pulsar\_Evolution'

**--logfile-BSE-supernovae**

Filename for BSE Supernovae logfile.

Default = 'BSE\_Supernovae'

**--logfile-BSE-system-parameters**

Filename for BSE System Parameters logfile.

Default = 'BSE\_System\_Parameters'

**--logfile-definitions**

Filename for logfile record definitions.

Default = ''

**--logfile-delimiter**

Field delimiter for logfile records.

Default = TAB

**--logfile-name-prefix**

Prefix for logfile names.

Default = ''

**--logfile-SSE-parameters**

Filename for SSE Parameters logfile.

Default = 'SSE\_Parameters'

**--mass-loss-prescription**

Mass loss prescription, (options: NONE, HURLEY, VINK).

Default = VINK

**-q [ --mass-ratio-distribution ]**

Initial mass ratio distribution for  $q=m_2/m_1$ .

(options: FLAT, DUQUENNOY MAYOR1991, SANA2012).

Default = FLAT

**--mass-transfer-accretion-efficiency-prescription**

Mass Transfer Accretion Efficiency prescription, (options: THERMAL, FIXED).

Default = THERMAL

**--mass-transfer-angular-momentum-loss-prescription**

Mass Transfer Angular Momentum Loss prescription.

(options: JEANS, ISOTROPIC, CIRCUMBINARY, ARBITRARY)

Default = ISOTROPIC

**--mass-transfer-rejuvenation-prescription**

Mass Transfer Rejuvenation prescription, (options: NONE, STARTRACK).

Default = NONE

**--mass-transfer-thermal-limit-accretor**

Mass Transfer Thermal Accretion limit multiplier, (options: CFACTOR, ROCHELOBE).

Default = CFACTOR

**--neutrino-mass-loss-bh-formation**

Assumption about neutrino mass loss during BH formation.

(options: FIXED\_FRACTION, FIXED\_MASS).

Default = FIXED\_FRACTION

**--neutron-star-equation-of-state**

Neutron star equation of state, (options: SSE, ARP3).

Default = SSE

**-c [ --output-container ]**

Container (directory) name for output files.

Default = 'COMPAS\_Output'

**-o [ --outputPath ]**

Path to which output is saved (i.e. directory in which the output container is created).

Default = CWD ('.')

**--pulsar-birth-magnetic-field-distribution**

Pulsar Birth Magnetic Field distribution.

(options: ZERO, FIXED, FLATINLOG, UNIFORM, LOGNORMAL).

Default = ZERO

**--pulsar-birth-spin-period-distribution**

Pulsar Birth Spin Period distribution, (options: ZERO, FIXED, UNIFORM, NORMAL).

Default = ZERO

**--pulsational-pair-instability-prescription**

Pulsational Pair Instability prescription, (options: COMPAS, STARTRACK, MARCHANT).

Default = COMPAS

**--remnant-mass-prescription**

Choose remnant mass prescription.

(options: HURLEY2000, BELCZYNSKI2002, FRYER2012, MULLER2016, MULLERMANDEL)  
Default = FRYER2012

**--rotational-velocity-distribution**

Initial rotational velocity distribution, (options: ZERO, HURLEY, VLTFLAMES).  
Default = ZERO

**-a [ --semi-major-axis-distribution ]**

Initial semi-major axis distribution, a.  
(options: FLATINLOG, CUSTOM, DUQUENNOYMAYOR1991, SANA2012)  
Default = FLATINLOG

**--debug-classes**

Debug classes enabled.  
Default = '' (None)

**--log-classes**

Logging classes enabled.  
Default = '' (None)

## Appendix B – Log File Record Specification: Stellar Properties

As described in **Standard Log File Record Specifiers**, when specifying known properties in a log file record specification record, the property name must be prefixed with the property type.

The current list of valid stellar property types available for use is:

- STAR\_PROPERTY for SSE
- STAR\_1\_PROPERTY for the primary star of a binary for BSE
- STAR\_2\_PROPERTY for the secondary star of a binary for BSE
- SUPERNOVA\_PROPERTY for the exploding star in a supernova for BSE
- COMPANION\_PROPERTY for the companion of the exploding star in a supernova for BSE

For example, to specify the property TEMPERATURE for an individual star being evolved for SSE, use:

STAR\_PROPERTY::TEMPERATURE

To specify the property TEMPERATURE for the primary star in a binary star being evolved for BSE, use:

STAR\_1\_PROPERTY::TEMPERATURE

To specify the property TEMPERATURE for the supernova star in a binary star being evolved for BSE, use:

SUPERNOVA\_PROPERTY::TEMPERATURE

Following is the list of stellar properties available for inclusion in log file record specifiers.

### Stellar Properties

#### AGE

Header string: Age

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Age

Effective age (changes with mass loss/gain) (Myr)

#### ANGULAR\_MOMENTUM

Header string: Ang\_Momentum

Data type: DOUBLE

COMPAS variable: BaseStar::m\_AngularMomentum

Angular momentum ( $M_{\text{sol}} \text{ AU}^2 \text{ yr}^{-1}$ )

#### BINDING\_ENERGY\_AT\_COMMON\_ENVELOPE

Header string: Binding\_Energy@CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.bindingEnergy

Absolute value of the envelope binding energy (ergs) at the onset of unstable RLOF, used for calculating post-CE separation.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

#### BINDING\_ENERGY\_FIXED

Header string: BE\_Fixed  
Data type: DOUBLE  
COMPAS variable: BaseStar::m\_BindingEnergies.fixed  
Absolute value of the envelope binding energy (ergs) calculated using a fixed lambda parameter.  
Calculated using  $\lambda = m\_Lambdas.fixed$

#### **BINDING\_ENERGY\_KRUCKOW**

Header string: BE\_Kruckow  
Data type: DOUBLE  
COMPAS variable: BaseStar::m\_BindingEnergies.kruckow  
Absolute value of the envelope binding energy (ergs) calculated using the fit by Vigna-Gomez et al. (2018) to Kruckow et al. (2016).  
Calculated using  $\alpha = \text{OPTIONS} \rightarrow \text{CommonEnvelopeSlopeKruckow}()$

#### **BINDING\_ENERGY\_LOVERIDGE**

Header string: BE\_Loveridge  
Data type: DOUBLE  
COMPAS variable: BaseStar::m\_BindingEnergies.loveridge  
Absolute value of the envelope binding energy (ergs) calculated as per Loveridge et al. (2001).  
Calculated using  $\lambda = m\_Lambdas.loveridge$

#### **BINDING\_ENERGY\_LOVERIDGE\_WINDS**

Header string: BE\_Loveridge\_Winds  
Data type: DOUBLE  
COMPAS variable: BaseStar::m\_BindingEnergies.loveridgeWinds  
Absolute value of the envelope binding energy (ergs) calculated as per Webbink 1984 & Loveridge et al. 2011 including winds.  
Calculated using  $\lambda = m\_Lambdas.loveridgeWinds$

#### **BINDING\_ENERGY\_NANJING**

Header string: BE\_Nanjing  
Data type: DOUBLE  
COMPAS variable: BaseStar::m\_BindingEnergies.nanjing  
Absolute value of the envelope binding energy (ergs) calculated as per Xu & Li, 2010.  
Calculated using  $\lambda = m\_Lambdas.nanjing$

#### **BINDING\_ENERGY\_POST\_COMMON\_ENVELOPE**

Header string: Binding\_Energy>CE  
Data type: DOUBLE  
COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.bindingEnergy  
Absolute value of the binding energy (ergs) immediately after CE.  
Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

#### **BINDING\_ENERGY\_PRE\_COMMON\_ENVELOPE**

Header string: Binding\_Energy<CE  
Data type: DOUBLE  
COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.bindingEnergy  
Absolute value of the envelope binding energy (ergs) at the onset of unstable RLOF leading to the CE.  
Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

#### **CHEMICALLY\_HOMOGENEOUS\_MAIN\_SEQUENCE**

Header string: CH\_on\_MS  
Data type: BOOL

COMPAS variable: BaseStar::m\_CHE

Flag to indicate whether the star evolved as a CH star for its entire MS lifetime.

TRUE indicates star evolved as CH star for entire MS lifetime

FALSE indicates star spun down and switched from CH to a normal MS

### **CO\_CORE\_MASS**

Header string: Mass\_CO\_Core

Data type: DOUBLE

COMPAS variable: BaseStar::m\_COCoreMass

Carbon-Oxygen core mass (Msol).

### **CO\_CORE\_MASS\_AT\_COMMON\_ENVELOPE**

Header string: Mass\_CO\_Core@CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.COCoreMass

Carbon-Oxygen core mass (Msol) at the onset of unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **CO\_CORE\_MASS\_AT\_COMPACT\_OBJECT\_FORMATION**

Header string: : Mass\_CO\_Core@CO

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.COCoreMassAtCOFormation

Carbon-Oxygen core mass (Msol) immediately prior to a supernova.

### **CORE\_MASS**

Header string: Mass\_Core

Data type: DOUBLE

COMPAS variable: BaseStar::m\_CoreMass

Core mass (Msol)

### **CORE\_MASS\_AT\_COMMON\_ENVELOPE**

Header string: Mass\_Core@CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.CoreMass

Core mass (Msol) at the onset of unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **CORE\_MASS\_AT\_COMPACT\_OBJECT\_FORMATION**

Header string: Mass\_Core@CO

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.CoreMassAtCOFormation

Core mass (Msol) immediately prior to a supernova .



**DRAWN\_KICK\_VELOCITY**

Header string: Drawn\_Kick\_Velocity

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.drawnKickVelocity

Magnitude of natal kick velocity ( $\text{kms}^{-1}$ ) without accounting for fallback ( $\text{km s}^{-1}$ ).

Supplied by user in grid file or drawn from distribution (default).

This value is used to calculate the actual kick velocity

**DT**

Header string: dT

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Dt

Current timestep (Myr).

**DYNAMICAL\_TIMESCALE**

Header string: Tau\_Dynamical

Data type: DOUBLE

COMPAS variable: BaseStar::m\_DynamicalTimescale

Dynamical time (Myr) .

**DYNAMICAL\_TIMESCALE\_POST\_COMMON\_ENVELOPE**

Header string: Tau\_Dynamical>CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.dynamicalTimescale

Dynamical time (Myr) immediately following common envelope event.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

**DYNAMICAL\_TIMESCALE\_PRE\_COMMON\_ENVELOPE**

Header string: Tau\_Dynamical<CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.dynamicalTimescale

Dynamical timescale (Myr) immediately prior to common envelope event.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

**ECCENTRIC\_ANOMALY**

Header string: Eccentric\_Anomaly

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.eccentricAnomaly

Eccentric anomaly calculated using Kepler's equation.

**ENV\_MASS**

Header string: Mass\_Env

Data type: DOUBLE

COMPAS variable: BaseStar::m\_EnvMass

Envelope mass ( $M_{\text{sol}}$ ) calculated using Hurley et al. 2000.

**ERROR**

Header string: Error

Data type: INT

COMPAS variable: <derived from BaseStar::m\_Error>

Error number (if error condition exists, else 0).

**EXPERIENCED\_CCSN**

Header string: Experienced\_CCSN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the star exploded as a core-collapse supernova at any time prior to the current timestep.

### **EXPERIENCED\_ECSN**

Header string: Experienced\_ECSN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past >

Flag to indicate whether the star exploded as an electron-capture supernova at any time prior to the current timestep.

### **EXPERIENCED\_PISN**

Header string: Experienced\_PISN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the star exploded as a pair-instability supernova at any time prior to the current timestep.

### **EXPERIENCED\_PPISN**

Header string: Experienced\_PPISN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the star exploded as a pulsational pair-instability supernova at any time prior to the current timestep.

### **EXPERIENCED\_RLOF**

Header string: Experienced\_RLOF

Data type: BOOL

COMPAS variable: <derived from BinaryConstituentStar::m\_RLOFDetails.experiencedRLOF>

Flag to indicate whether the star has overflowed its Roche Lobe at any time prior to the current timestep.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **EXPERIENCED\_SN\_TYPE**

Header string: Experienced\_SN\_Type

Data type: INT

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

The type of supernova event experienced by the star prior to the current timestep.

Printed as one of {NONE = 0, CCSN = 1, ECSN = 2, PISN = 4, PPISN = 8, USSN = 16}

(see section **Supernova events/states** below for explanation)

### **EXPERIENCED\_USSN**

Header string: Experienced\_USSN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the star exploded as an ultra-stripped supernova at any time prior to the current timestep.

**FALLBACK\_FRACTION**

Header string: Fallback\_Fraction

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.fallbackFraction

Fallback fraction during a supernova.

**HE\_CORE\_MASS**

Header string: Mass\_He\_Core

Data type: DOUBLE

COMPAS variable: BaseStar::m\_HeCoreMass

Helium core mass (Msol).

**HE\_CORE\_MASS\_AT\_COMMON\_ENVELOPE**

Header string: Mass\_He\_Core@CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.HeCoreMass

Helium core mass (Msol) at the onset of unstable RLOF leading to the CE..

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

**HE\_CORE\_MASS\_AT\_COMPACT\_OBJECT\_FORMATION**

Header string: Mass\_He\_Core@CO

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.HeCoreMassAtCOFormation

Helium core mass (Msol) at immediately prior to a of supernova.

**HYDROGEN\_POOR**

Header string: Hydrogen\_Poor

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.hydrogenContent>

Flag to indicate if star is hydrogen poor.

**HYDROGEN\_RICH**

Header string: Hydrogen\_Rich

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.hydrogenContent>

Flag to indicate if star is hydrogen rich.

**ID**

Header string: ID

Data type: UNSIGNED LONG INT

COMPAS variable: BaseStar::m\_ObjectId

Unique object identifier for c++ object – used in debugging to identify objects.

**INITIAL\_STELLAR\_TYPE**

Header string: Stellar\_Type@ZAMS

Data type: INT

COMPAS variable: BaseStar::m\_InitialStellarType

Stellar type at zero age main-sequence (per Hurley et al. 2000).

*Note that this property has the same header string as INITIAL\_STELLAR\_TYPE\_NAME below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

**INITIAL\_STELLAR\_TYPE\_NAME**

Header string: Stellar\_Type@ZAMS

Data type: STRING

COMPAS variable: <derived from BaseStar::m\_InitialStellarType>

Stellar type at zero age main-sequence name (per Hurley at al. 2000).

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as INITIAL\_STELLAR\_TYPE above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **IS\_CCSN**

Header string: CCSN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.current>

Flag to indicate whether the star is currently a core-collapse supernova.

### **IS\_ECSN**

Header string: ECSN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.current>

Flag to indicate whether the star is currently an electron-capture supernova.

### **IS\_PISN**

Header string: PISN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.current>

Flag to indicate whether the star is currently a pair-instability supernova.

### **IS\_PPISN**

Header string: PPISN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.current>

Flag to indicate whether the star is currently a pulsational pair-instability supernova.

### **IS\_RLOF**

Header string: RLOF

Data type: BOOL

COMPAS variable: <derived from BinaryConstituentStar::m\_RLOFDetails.isRLOF>

Flag to indicate whether the star is currently undergoing Roche Lobe overflow.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **IS\_USSN**

Header string: USSN

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.current>

Flag to indicate whether the star is currently an ultra-stripped supernova.

### **KICK\_VELOCITY**

Header string: Applied\_Kick\_Velocity

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.kickVelocity

Magnitude of natal kick velocity ( $\text{kms}^{-1}$ ) received during a supernova.

Calculate using the drawn kick velocity (see BaseStar::CalculateSNKickVelocity() for details)

### **LAMBDA\_AT\_COMMON\_ENVELOPE**

Header string: Lambda@CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.lambda

Common-envelope lambda parameter calculated at the unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **LAMBDA\_DEWI**

Header string: Dewi

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.dewi

Common envelope lambda parameter calculated as per Dewi & Tauris (2000) using the fit from Appendix A of Claeys et al. (2014).

### **LAMBDA\_FIXED**

Header string: Lambda\_Fixed

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.fixed

Universal common envelope lambda parameter specified by the user (program option --common-envelope-lambda).

### **LAMBDA\_KRUCKOW**

Header string: Kruckow

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.kruckow

Common envelope lambda parameter calculated as per Kruckow et al. (2016) with the alpha exponent set by `OPTIONS->CommonEnvelopeSlopeKruckow()`.

From Kruckow et al. 2016 (arXiv:1610.04417), fig 1.

Spectrum fit to the region bounded by the upper and lower limits as shown in Kruckow+ 2016

### **LAMBDA\_KRUCKOW\_BOTTOM**

Header string: Kruckow\_Bottom

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.kruckowBottom

Common envelope lambda parameter calculated as per Kruckow et al. (2016) with the alpha exponent set to -1.0.

From Kruckow et al. 2016 (arXiv:1610.04417), fig 1.

Spectrum fit to the region bounded by the upper and lower limits as shown in Kruckow+ 2016

### **LAMBDA\_KRUCKOW\_MIDDLE**

Header string: Kruckow\_Middle

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.kruckowMiddle

Common envelope lambda parameter calculated as per Kruckow et al. (2016) with the alpha exponent set to -4/5.

From Kruckow et al. 2016 (arXiv:1610.04417), fig 1.

Spectrum fit to the region bounded by the upper and lower limits as shown in Kruckow+ 2016

### **LAMBDA\_KRUCKOW\_TOP**

Header string: Kruckow\_Top

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.kruckowTop

Common envelope lambda parameter as per Kruckow et al. (2016) with the alpha exponent set to -2/3.

From Kruckow et al. 2016 (arXiv:1610.04417), fig 1.

Spectrum fit to the region bounded by the upper and lower limits as shown in Kruckow+ 2016

### **LAMBDA\_LOVERIDGE**

Header string: Loveridge

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.loveridge

Common envelope lambda parameter calculated per Webbink 1984 & Loveridge et al. 2011.

### **LAMBDA\_LOVERIDGE\_WINDS**

Header string: Loveridge\_Winds

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.loveridgeWinds

Common envelope lambda parameter calculated per Webbink 1984 & Loveridge et al. 2011 including winds.

Note: currently this evaluates the same as BINDING\_ENERGY\_LOVERIDGE

### **LAMBDA\_NANJING**

Header string: Lambda\_Nanjing

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Lambdas.nanjing

Common envelope lambda parameter calculated as per Xu & Li (2010).

From X.-J. Xu and X.-D. Li arXiv:1004.4957 (v1, 28Apr2010) as implemented in STARTRACK

### **LBV\_PHASE\_FLAG**

Header string: LBV\_Phase\_Flag

Data type: BOOL

COMPAS variable: BaseStar::m\_LBVphaseFlag

Flag to indicate if the star ever entered the luminous blue variable phase.

### **LUMINOSITY**

Header string: Luminosity

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Luminosity

Luminosity (Lsol).

### **LUMINOSITY\_POST\_COMMON\_ENVELOPE**

Header string: Luminosity>CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.luminosity

Luminosity ( $L_{\text{sol}}$ ) immediately following common envelope event.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **LUMINOSITY\_PRE\_COMMON\_ENVELOPE**

Header string: Luminosity<CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.luminosity

Luminosity ( $L_{\text{sol}}$ ) at the onset of unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **MASS**

Header string: Mass

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Mass

Mass ( $M_{\text{sol}}$ ).

### **MASS\_0**

Header string: Mass\_0

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Mass0

Effective initial mass ( $M_{\text{sol}}$ ).

### **MASS\_LOSS\_DIFF**

Header string: dmWinds

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_MassLossDiff

The amount of mass lost due to winds ( $M_{\text{sol}}$ )

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **MASS\_TRANSFER\_CASE\_INITIAL**

Header string: MT\_Case

Data type: INT

COMPAS variable: BinaryConstituentStar::m\_MassTransferCase

Indicator of mass transfer type when first RLOF occurs, if any (one of {NONE, A, B, C, D})

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **MASS\_TRANSFER\_DIFF**

Header string: dmMT

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_MassTransferDiff

Amount of mass ( $M_{\text{sol}}$ ) accreted or donated during a mass transfer episode.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **MDOT**

Header string: Mdot

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Metallicity

Mass loss rate ( $M_{\text{sol}} \text{ yr}^{-1}$ ).

### **MEAN\_ANOMALY**

Header string: SN\_Kick\_Mean\_Anomaly

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.meanAnomaly

Mean anomaly of supernova kick, supplied by user in grid file, default = random number between 0 and  $2\pi$ .

See [https://en.wikipedia.org/wiki/Mean\\_anomaly](https://en.wikipedia.org/wiki/Mean_anomaly) for explanation

## **METALLICITY**

Header string: Metallicity@ZAMS

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Metallicity

Metallicity.

## **MZAMS**

Header string: Mass@ZAMS

Data type: DOUBLE

COMPAS variable: BaseStar::m\_MZAMS

ZAMS Mass ( $M_{\text{sol}}$ ).

## **NUCLEAR\_TIMESCALE**

Header string: Tau\_Nuclear

Data type: DOUBLE

COMPAS variable: BaseStar::m\_NuclearTimescale

Nuclear timescale (Myr).

## **NUCLEAR\_TIMESCALE\_POST\_COMMON\_ENVELOPE**

Header string: Tau\_Nuclear>CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.nuclearTimescale

Nuclear timescale (Myr) immediately following common envelope event.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

## **NUCLEAR\_TIMESCALE\_PRE\_COMMON\_ENVELOPE**

Header string: Tau\_Nuclear<CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.nuclearTimescale

Nuclear timescale (Myr) at the onset of unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

## **OMEGA**

Header string: Omega

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Omega

Angular frequency ( $\text{yr}^{-1}$ ).

## **OMEGA\_BREAK**

Header string: Omega\_Break

Data type: DOUBLE

COMPAS variable: BaseStar::m\_OmegaBreak

Break-up angular frequency ( $\text{yr}^{-1}$ ).

## **OMEGA\_ZAMS**

Header string: Omega@ZAMS

Data type: DOUBLE

COMPAS variable: BaseStar::m\_OmegaZAMS



Angular frequency at ZAMS ( $\text{yr}^{-1}$ ).

#### **ORBITAL\_ENERGY\_POST\_SUPERNOVA**

Header string: Orbital\_Energy>SN

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_PostSNeOrbitalEnergy

Absolute value of orbital energy immediately following supernova event ( $\text{Msol AU}^2 \text{yr}^{-2}$ ).

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

#### **ORBITAL\_ENERGY\_PRE\_SUPERNOVA**

Header string: Orbital\_Energy<SN

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_PreSNeOrbitalEnergy

Orbital energy immediately prior to supernova event ( $\text{Msol AU}^2 \text{yr}^{-2}$ ).

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

#### **PULSAR\_MAGNETIC\_FIELD**

Header string: Pulsar\_Mag\_Field

Data type: DOUBLE

COMPAS variable: BaseStar::m\_PulsarDetails.magneticField

Pulsar magnetic field strength (G).

#### **PULSAR\_SPIN\_DOWN\_RATE**

Header string: Pulsar\_Spin\_Down

Data type: DOUBLE

COMPAS variable: BaseStar::m\_PulsarDetails.spinDownRate

Pulsar spin-down rate.

#### **PULSAR\_SPIN\_FREQUENCY**

Header string: Pulsar\_Spin\_Freq

Data type: DOUBLE

COMPAS variable: BaseStar::m\_PulsarDetails.spinFrequency

Pulsar spin angular frequency ( $\text{rads s}^{-1}$ ).

#### **PULSAR\_SPIN\_PERIOD**

Header string: Pulsar\_Spin\_Period

Data type: DOUBLE

COMPAS variable: BaseStar::m\_PulsarDetails.spinPeriod

Pulsar spin period (ms).

#### **RADIAL\_EXPANSION\_TIMESCALE**

Header string: Tau\_Radial

Data type: DOUBLE

COMPAS variable: BaseStar::m\_RadialExpansionTimescale

e-folding time (Myr) of stellar radius.

### **RADIAL\_EXPANSION\_TIMESCALE\_POST\_COMMON\_ENVELOPE**

Header string: Tau\_Radial>CE

Data type: DOUBLE

COMPAS variable: : BinaryConstituentStar::m\_CEDetails.postCEE.radialExpansionTimescale  
e-folding time (Myr) of stellar radius immediately following common envelope event.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **RADIAL\_EXPANSION\_TIMESCALE\_PRE\_COMMON\_ENVELOPE**

Header string: Tau\_Radial<CE

Data type: DOUBLE

COMPAS variable: : BinaryConstituentStar::m\_CEDetails.preCEE.radialExpansionTimescale  
e-folding time (Myr) of stellar radius at the onset of unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **RADIUS**

Header string: Radius

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Radius

Radius (Rsol).

### **RANDOM\_SEED**

Header string: SEED

Data type: UNSIGNED LONG

COMPAS variable: BaseStar::m\_RandomSeed

Seed for random number generator for this star.

### **RECYCLED\_NEUTRON\_STAR**

Header string: Recycled\_NS

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the object was a recycled neutron star at any time prior to the current timestep (was a neutron star accreting mass).

### **RLOF\_ONTO\_NS**

Header string: RLOF->NS

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the star transferred mass to a neutron star at any time prior to the current timestep.

### **RUNAWAY**

Header string: Runaway

Data type: BOOL

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.past>

Flag to indicate whether the star was unbound by a supernova event at any time prior to the current timestep.

(Unbound after supernova event and not a WD, NS, BH or MR)

## **RZAMS**

Header string: R@ZAMS

Data type: DOUBLE

COMPAS variable: BaseStar::m\_RZAMS

ZAMS Radius (Rsol)

## **SN\_TYPE**

Header string: SN\_Type

Data type: INT

COMPAS variable: <derived from BaseStar::m\_SupernovaDetails.events.current>

The type of supernova event currently being experienced by the star.

Printed as one of {NONE = 0, CCSN = 1, ECSN = 2, PISN = 4, PPISN = 8, USSN = 16}

(see section **Supernova events/states** below for explanation)

## **STELLAR\_TYPE**

Header string: Stellar\_Type

Data type: INT

COMPAS variable: BaseStar::m\_StellarType

Stellar type (per Hurley at al. 2000)

*Note that this property has the same header string as STELLAR\_TYPE\_NAME below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

## **STELLAR\_TYPE\_NAME**

Header string: Stellar\_Type

Data type: STRING

COMPAS variable: <derived from BaseStar::m\_StellarType>

Stellar type name (per Hurley at al. 2000).

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as STELLAR\_TYPE above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

## **STELLAR\_TYPE\_PREV**

Header string: Stellar\_Type\_Prev

Data type: INT

COMPAS variable: BaseStar::m\_StellarTypePrev

Stellar type (per Hurley at al. 2000) at previous timestep.

*Note that this property has the same header string as STELLAR\_TYPE\_PREV\_NAME below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

## **STELLAR\_TYPE\_PREV\_NAME**

Header string: Stellar\_Type\_Prev

Data type: STRING

COMPAS variable: <derived from BaseStar::m\_StellarTypePrev>

Stellar type name (per Hurley at al. 2000) at previous timestep.

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as STELLAR\_TYPE\_PREV above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

## **SUPERNOVA\_KICK\_VELOCITY\_MAGNITUDE\_RANDOM\_NUMBER**

Header string: SN\_Kick\_Magnitude\_Random\_Number

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.kickVelocityRandom

Random number for drawing the supernova kick velocity magnitude (if required).

Either supplied by user in grid file, default = drawn from uniform random distribution 0..1

### **SUPERNOVA\_PHI**

Header string: SN\_Kick\_Phi

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.phi

Angle between 'x' and 'y', both in the orbital plane of supernovae vector (rad).

Either supplied by user in grid file, default = drawn from uniform random distribution 0..2pi

### **SUPERNOVA\_THETA**

Header string: SN\_Kick\_Theta

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.theta

Angle between the orbital plane and the 'z' axis of supernovae vector (rad).

Either supplied by user in grid file, default = drawn from specified distribution (option – kick\_direction)

### **TEMPERATURE**

Header string: Teff

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Temperature

Effective temperature (Tsol).

### **TEMPERATURE\_POST\_COMMON\_ENVELOPE**

Header string: Teff>CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.temperature

Effective temperature (Tsol) immediately following common envelope event.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **TEMPERATURE\_PRE\_COMMON\_ENVELOPE**

Header string: Teff<CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.temperature

Effective temperature (Tsol) at the unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **THERMAL\_TIMESCALE**

Header string: Tau\_Thermal

Data type: DOUBLE

COMPAS variable: BaseStar::m\_ThermalTimescale

Thermal timescale (Myr).

### **THERMAL\_TIMESCALE\_POST\_COMMON\_ENVELOPE**

Header string: Tau\_Thermal>CE

Data type: DOUBLE

COMPAS variable: : BinaryConstituentStar::m\_CEDetails.postCEE.thermalTimescale

Thermal timescale immediately following common envelope event (Myr).

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **THERMAL\_TIMESCALE\_PRE\_COMMON\_ENVELOPE**

Header string: Tau\_Thermal<CE

Data type: DOUBLE

COMPAS variable: : BinaryConstituentStar::m\_CEDetails.preCEE.thermalTimescale

Thermal timescale (Myr) at the onset of the unstable RLOF leading to the CE.

Applies only to constituent stars of a binary system (i.e. does not apply to SSE star)

### **TIME**

Header string: Time

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Time

Time since ZAMS (Myr).

### **TIMESCALE\_MS**

Header string: tMS

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Timescales[tMS]

Main Sequence timescale (Myr).

### **TOTAL\_MASS\_AT\_COMPACT\_OBJECT\_FORMATION**

Header string: Mass\_Total@CO

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.totalMassAtCOFormation

Total mass of the star at the beginning of a supernova event (Msol).

### **TRUE\_ANOMALY**

Header string: True\_Anomaly(psi)

Data type: DOUBLE

COMPAS variable: BaseStar::m\_SupernovaDetails.trueAnomaly

True anomaly (rad) calculated using Kepler's equation .

See [https://en.wikipedia.org/wiki/True\\_anomaly](https://en.wikipedia.org/wiki/True_anomaly) for explanation

### **ZETA\_HURLEY**

Header string: Zeta\_Hurley

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Zetas.hurley

Adiabatic exponent calculated per Hurley et al. (2002) using core mass.

### **ZETA\_HURLEY\_HE**

Header string: Zeta\_Hurley\_He

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Zetas.hurleyHe

Adiabatic exponent calculated per (Hurley et al. 2002) using He core mass.

**ZETA\_SOBERMAN**

Header string: Zeta\_Soberman

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Zetas.soberman

Adiabatic exponent calculated per Soberman, Phinney, van den Heuvel (1997) using core mass.

**ZETA\_SOBERMAN\_HE**

Header string: Zeta\_Soberman\_He

Data type: DOUBLE

COMPAS variable: BaseStar::m\_Zetas.sobermanHe

Adiabatic exponent calculated per Soberman, Phinney, van den Heuvel (1997) using He-core mass.

## Supernova events/states

Supernova events/states, both current (“is”) and past (“experienced”), are stored within COMPAS as bit maps. That means different values can be ORed or ANDed into the bit map, so that various events or states can be set concurrently.

The values shown below for the SN\_EVENT type are powers of 2 so that they can be used in a bit map and manipulated with bit-wise logical operators. Any of the individual supernova event/state types that make up the SN\_EVENT type can be set independently of any other event/state.

```
enum class SN_EVENT: int {
    NONE          = 0,
    CCSN          = 1,
    ECSN          = 2,
    PISN          = 4,
    PPISN         = 8,
    USSN          = 16,
    RUNAWAY       = 32,
    RECYCLED_NS   = 64,
    RLOF_ONTO_NS  = 128
};

const COMPASUnorderedMap<SN_EVENT, std::string> SN_EVENT_LABEL = {
    { SN_EVENT::NONE,          "No Supernova" },
    { SN_EVENT::CCSN,          "Core Collapse Supernova" },
    { SN_EVENT::ECSN,          "Electron Capture Supernova" },
    { SN_EVENT::PISN,          "Pair Instability Supernova" },
    { SN_EVENT::PPISN,         "Pulsational Pair Instability Supernova" },
    { SN_EVENT::USSN,          "Ultra Stripped Supernova" },
    { SN_EVENT::RUNAWAY,       "Runaway Companion" },
    { SN_EVENT::RECYCLED_NS,   "Recycled Neutron Star" },
    { SN_EVENT::RLOF_ONTO_NS,  "Donated Mass to Neutron Star through RLOF" }
};
```

A convenience function (shown below) is provided in utils.cpp to interpret the bit map. Given an SN\_EVENT bitmap (current or past), it returns (in priority order):

SN_EVENT::CCSN	iff CCSN bit is set and USSN bit is not set
SN_EVENT::ECSN	iff ECSN bit is set
SN_EVENT::PISN	iff PISN bit is set
SN_EVENT::PPISN	iff PPISN bit is set
SN_EVENT::USSN	iff USSN bit is set
SN_EVENT::NONE	otherwise

```

/*
 * Returns a single SN type based on the SN_EVENT parameter passed
 *
 * Returns (in priority order):
 *
 *     SN_EVENT::CCSN   iff CCSN   bit is set and USSN bit is not set
 *     SN_EVENT::ECSN   iff ECSN   bit is set
 *     SN_EVENT::PISN   iff PISN   bit is set
 *     SN_EVENT::PPISN  iff PPISN  bit is set
 *     SN_EVENT::USSN   iff USSN   bit is set
 *     SN_EVENT::NONE   otherwise
 *
 *
 * @param   [IN]    p_SNEvent    SN_EVENT mask to check for SN event type
 * @return                                SN_EVENT
 */
SN_EVENT SNEventType(const SN_EVENT p_SNEvent) {

    if ((p_SNEvent & (SN_EVENT::CCSN | SN_EVENT::USSN)) == SN_EVENT::CCSN )
        return SN_EVENT::CCSN;

    if ((p_SNEvent & SN_EVENT::ECSN ) == SN_EVENT::ECSN ) return SN_EVENT::ECSN;
    if ((p_SNEvent & SN_EVENT::PISN ) == SN_EVENT::PISN ) return SN_EVENT::PISN;
    if ((p_SNEvent & SN_EVENT::PPISN) == SN_EVENT::PPISN ) return SN_EVENT::PPISN;
    if ((p_SNEvent & SN_EVENT::USSN ) == SN_EVENT::USSN ) return SN_EVENT::USSN;

    return SN_EVENT::NONE;
}

```



## Appendix C – Log File Record Specification: Binary Properties

As described in **Standard Log File Record Specifiers**, when specifying known properties in a log file record specification record, the property name must be prefixed with the property type.

Currently there is a single binary property type available for use: BINARY\_PROPERTY.

For example, to specify the property SEMI\_MAJOR\_AXIS\_PRE\_COMMON\_ENVELOPE for a binary star being evolved in BSE, use:

BINARY\_PROPERTY::SEMI\_MAJOR\_AXIS\_PRE\_COMMON\_ENVELOPE

Following is the list of binary properties available for inclusion in log file record specifiers.

### Binary Properties

#### CIRCULARIZATION\_TIMESCALE

Header string: Tau\_Circ

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CircularizationTimescale

Tidal circularisation timescale (Myr).

#### COMMON\_ENVELOPE\_ALPHA

Header string: CE\_Alpha

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.alpha

Common envelope alpha (efficiency) parameter

User-supplied via command line option --common-envelope-alpha.

#### COMMON\_ENVELOPE\_AT\_LEAST\_ONCE

Header string: CEE

Data type: BOOL

COMPAS variable: <derived from BaseBinaryStar::m\_CEDetails.CEEcount>

Flag to indicate if there has been at least one common envelope event.

#### COMMON\_ENVELOPE\_EVENT\_COUNT

Header string: CE\_Event\_Count

Data type: UNSIGNED INT

COMPAS variable: BaseBinaryStar::m\_CEDetails.CEEcount

The number of common envelope events.

#### DIMENSIONLESS\_KICK\_VELOCITY

Header string: Kick\_Velocity(uK)

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_uK

Dimensionless kick velocity supplied by user (see option --fix-dimensionless-kick-velocity).

#### DOUBLE\_CORE\_COMMON\_ENVELOPE

Header string: Double\_Core\_CE

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_CEDetails.doubleCoreCE

Flag to indicate double-core common envelope.

**DT**

Header string: dT  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_Dt  
Current timestep (Myr).

#### **ECCENTRICITY**

Header string: Eccentricity  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_Eccentricity  
Orbital eccentricity.

#### **ECCENTRICITY\_AT\_DCO\_FORMATION**

Header string: Eccentricity@DCO  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_EccentricityAtDCOFormation  
Orbital eccentricity at DCO formation.

#### **ECCENTRICITY\_INITIAL**

Header string: Eccentricity@ZAMS  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_EccentricityInitial  
Orbital eccentricity at ZAMS.  
Supplied by user via grid file or sampled from distribution (see `–eccentricity-distribution` option)  
(default)

#### **ECCENTRICITY\_POST\_COMMON\_ENVELOPE**

Header string: Eccentricity>CE  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_CEDetails.postCEE.eccentricity  
Eccentricity immediately following common envelope event.

#### **ECCENTRICITY\_PRE\_SUPERNOVA**

Header string: Eccentricity<SN  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_EccentricityPreSN  
Current timestep (Myr)  
Eccentricity of the binary immediately prior to supernova event

#### **ECCENTRICITY\_PRE\_COMMON\_ENVELOPE**

Header string: Eccentricity  
Data type: DOUBLE<CE  
COMPAS variable: BaseBinaryStar::m\_CEDetails.preCEE.eccentricity  
Eccentricity at the onset of RLOF leading to the CE.

#### **ECCENTRICITY\_PRIME**

Header string: Eccentricity  
Data type: DOUBLE  
COMPAS variable: BaseBinaryStar::m\_EccentricityPrime  
Post-SN eccentricity.

**ERROR**

Header string: Error

Data type: INT

COMPAS variable: <derived from BaseBinaryStar::m\_Error>

Error number (if error condition exists, else 0).

**ID**

Header string: ID

Data type: UNSIGNED LONG INT

COMPAS variable: BaseBinaryStar::m\_ObjectId

Unique object identifier for c++ object – used in debugging to identify objects.

**IMMEDIATE\_RLOF\_POST\_COMMON\_ENVELOPE**

Header string: Immediate\_RLOF>CE

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_RLOFDetails.immediateRLOFPostCEE

Flag to indicate if either star overflows its Roche lobe immediately following common envelope event.

**LUMINOUS\_BLUE\_VARIABLE\_FACTOR**

Header string: LBV\_Multiplier

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_LBVfactor

Luminous blue variable wind mass loss multiplier

User-supplied via option --luminous-blue-variable-multiplier.

**MASS\_1\_FINAL**

Header string: Core\_Mass\_1

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_Mass1Final

Mass of the primary star (Msol) after losing its envelope (assumes complete loss of envelope).

**MASS\_1\_POST\_COMMON\_ENVELOPE**

Header string: Mass\_1>CE

Data type: DOUBE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.mass

Mass of the primary star (Msol) immediately following common envelope event.

**MASS\_1\_PRE\_COMMON\_ENVELOPE**

Header string: Mass\_1<CE

Data type: DOUBE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.mass

Mass of the primary star (Msol) immediately prior to common envelope event.

**MASS\_2\_FINAL**

Header string: Core\_Mass\_2

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_Mass2Final

Mass of the secondary star (Msol) after losing its envelope (assumes complete loss of envelope).

### **MASS\_2\_POST\_COMMON\_ENVELOPE**

Header string: Mass\_2>CE

Data type: DOUBE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.mass

Mass of the secondary star (Msol) immediately following common envelope event.

### **MASS\_2\_PRE\_COMMON\_ENVELOPE**

Header string: Mass\_2<CE

Data type: DOUBE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.mass

Mass of the secondary star (Msol) immediately prior to common envelope event.

### **MASS\_ENV\_1**

Header string: Mass\_Env\_1

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_MassEnv1

Envelope mass of the primary star (Msol).

### **MASS\_ENV\_2**

Header string: Mass\_Env\_2

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_MassEnv2

Envelope mass of the secondary star (Msol).

### **MASSES\_EQUILIBRATED**

Header string: Equilibrated

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_MassesEquilibrated

Flag to indicate whether chemically homogeneous stars had masses equilibrated and orbit circularised due to Roche lobe overflow during evolution.

### **MASSES\_EQUILIBRATED\_AT\_BIRTH**

Header string: Equilibrated\_At\_Birth

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_MassesEquilibratedAtBirth

Flag to indicate whether stars had masses equilibrated and orbit circularised at birth due to Roche lobe overflow.

### **MASS\_TRANSFER\_TRACKER\_HISTORY**

Header string: MT\_History

Data type: INT

COMPAS variable: <derived from BaseBinaryStar::m\_MassTransferTrackerHistory>

Indicator of mass transfer history for the binary.

Will be printed as one of:

NO\_MASS\_TRANSFER = 0

STABLE\_FROM\_1\_TO\_2 = 1

STABLE\_FROM\_2\_TO\_1 = 2

CE\_FROM\_1\_TO\_2 = 3

CE\_FROM\_2\_TO\_1 = 4

CE\_DOUBLE\_CORE = 5

CE\_BOTH\_MS = 6

CE\_MS\_WITH\_CO = 7

### **MERGES\_IN\_HUBBLE\_TIME**

Header string: Merges\_Hubble\_Time

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_MergesInHubbleTime

Flag to indicate if the binary compact remnants merge within a Hubble time.

### **OPTIMISTIC\_COMMON\_ENVELOPE**

Header string: Optimistic\_CE

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_CEDetails.optimisticCE

Flag that returns TRUE if we have a Hertzsprung-gap star, and we allow it to survive the CE.

### **ORBITAL\_VELOCITY**

Header string: Orbital\_Velocity

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_OrbitalVelocity

Orbital velocity ( $\text{km s}^{-1}$ ).

### **ORBITAL\_VELOCITY\_PRE\_SUPERNOVA**

Header string: Orbital\_Velocity<SN

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_OrbitalVelocityPreSN

Orbital velocity ( $\text{km s}^{-1}$ ) immediately prior to supernova event.

### **RADIUS\_1\_POST\_COMMON\_ENVELOPE**

Header string: Radius\_1>CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.radius

Radius of the primary star ( $R_{\text{sol}}$ ) immediately following common envelope event.

### **RADIUS\_1\_PRE\_COMMON\_ENVELOPE**

Header string: Radius\_1<CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.radius

Radius of the primary star ( $R_{\text{sol}}$ ) at the onset of RLOF leading to the common-envelope episode.

### **RADIUS\_2\_POST\_COMMON\_ENVELOPE**

Header string: Radius\_2>CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.radius

Radius of the secondary star ( $R_{\text{sol}}$ ) immediately following common envelope event.

### **RADIUS\_2\_PRE\_COMMON\_ENVELOPE**

Header string: Radius\_2<CE

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_CEDetails.preCEE.radius

Radius of the secondary star ( $R_{\text{sol}}$ ) at the onset of RLOF leading to the common-envelope episode.

**RANDOM\_SEED**

Header string: SEED

Data type: UNSIGNED LONG

COMPAS variable: BaseBinaryStar::m\_RandomSeed

Seed for random number generator for this binary star.

(supplied by user via option—random-seed, default generated from system time)

**ROCHE\_LOBE\_RADIUS\_1**

Header string: RocheLobe\_1/a

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_RocheLobeRadius

Roche radius of the primary star ( $R_{\text{sol}}$ ).

**ROCHE\_LOBE\_RADIUS\_2**

Header string: RocheLobe\_2/a

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_RocheLobeRadius

Roche radius of the secondary star ( $R_{\text{sol}}$ ).

**ROCHE\_LOBE\_RADIUS\_1\_POST\_COMMON\_ENVELOPE**

Header string: RocheLobe\_1>CE

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.postCEE.rocheLobe1to2

Roche radius of the primary star ( $R_{\text{sol}}$ ) immediately following common envelope event.

**ROCHE\_LOBE\_RADIUS\_2\_POST\_COMMON\_ENVELOPE**

Header string: RocheLobe\_2>CE

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.postCEE.rocheLobe2to1

Roche radius of the secondary star ( $R_{\text{sol}}$ ) immediately following common envelope event.

**ROCHE\_LOBE\_RADIUS\_1\_PRE\_COMMON\_ENVELOPE**

Header string: RocheLobe\_1<CE

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.preCEE.rocheLobe1to2

Roche radius of the primary star ( $R_{\text{sol}}$ ) at the onset of RLOF leading to the common-envelope episode.

**ROCHE\_LOBE\_RADIUS\_2\_PRE\_COMMON\_ENVELOPE**

Header string: RocheLobe\_2<CE

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.preCEE.rocheLobe2to1

Roche radius of the secondary star ( $R_{\text{sol}}$ ) at the onset of RLOF leading to the common-envelope episode.

**ROCHE\_LOBE\_TRACKER\_1**

Header string: Radius\_1/RL

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_RocheLobeTracker

Ratio of the primary star's stellar radius to Roche radius ( $R/RL$ ), evaluated at periapsis

### **ROCHE\_LOBE\_TRACKER\_2**

Header string: Radius\_2/RL

Data type: DOUBLE

COMPAS variable: BinaryConstituentStar::m\_RocheLobeTracker

Ratio of the secondary star's stellar radius to Roche radius (R/RL), evaluated at periapsis.

### **SECONDARY\_TOO\_SMALL\_FOR\_DCO**

Header string: Secondary<<DCO

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_SecondaryTooSmallForDCO

Flag to indicate that the secondary star was born too small for the binary to evolve into a DCO.

### **SEMI\_MAJOR\_AXIS\_AT\_DCO\_FORMATION**

Header string: Separation@DCO

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_SemiMajorAxisAtDCOFormation

Semi-major axis(AU) at DCO formation.

### **SEMI\_MAJOR\_AXIS\_INITIAL**

Header string: Separation@ZAMS

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_SemiMajorAxisInitial

Semi-major axis (AU) at ZAMS.

### **SEMI\_MAJOR\_AXIS\_POST\_COMMON\_ENVELOPE**

Header string: Separation>CE

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.postCEE.semiMajorAxis

Semi-major axis (AU) immediately following common envelope event.

### **SEMI\_MAJOR\_AXIS\_PRE\_SUPERNOVA**

Header string: Separation<SN

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_SemiMajorAxisPreSN

Semi-major axis (AU) immediately prior to supernova event.

*Note that this property has the same header string as*

*SEMI\_MAJOR\_AXIS\_PRE\_SUPERNOVA\_RSOL below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **SEMI\_MAJOR\_AXIS\_PRE\_SUPERNOVA\_RSOL**

Header string: Separation<SN

Data type: DOUBLE

COMPAS variable: <derived from BaseBinaryStar::m\_SemiMajorAxisPreSN>

Semi-major axis (Rsol) immediately prior to supernova event.

*Note that this property has the same header string as SEMI\_MAJOR\_AXIS\_PRE\_SUPERNOVA above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **SEMI\_MAJOR\_AXIS\_PRE\_COMMON\_ENVELOPE**

Header string: Separation<CE

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_CEDetails.preCEE.semiMajorAxis

Semi-major axis(AU) at the onset of RLOF leading to the common-envelope episode.

### **SEMI\_MAJOR\_AXIS\_PRIME**

Header string: Separation

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_SemiMajorAxisPrime

Semi-major axis of the binary (AU).

*Note that this property has the same header string as SEMI\_MAJOR\_AXIS\_PRIME\_RSOL below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **SEMI\_MAJOR\_AXIS\_PRIME\_RSOL**

Header string: Separation

Data type: DOUBLE

COMPAS variable: <derived from BaseBinaryStar::m\_SemiMajorAxisPrime>

Semi-major axis of the binary (RSOL).

*Note that this property has the same header string as SEMI\_MAJOR\_AXIS\_PRIME above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **SIMULTANEOUS\_RLOF**

Header string: Simultaneous\_RLOF

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_RLOFDetails.simultaneousRLOF

Flag to indicate that both stars are undergoing RLOF.

### **STABLE\_RLOF\_POST\_COMMON\_ENVELOPE**

Header string: Stable\_RLOF>CE

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_RLOFDetails.stableRLOFPostCEE

Flag to indicate stable mass transfer after common envelope event.

### **STELLAR\_MERGER**

Header string: Merger

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_StellarMerger

Flag to indicate the stars merged (stars were touching) during evolution.

### **STELLAR\_MERGER\_AT\_BIRTH**

Header string: Merger\_At\_Birth

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_StellarMergerAtBirth

Flag to indicate the stars merged (stars were touching) at birth.



### **STELLAR\_TYPE\_1\_POST\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_1>CE

Data type: INT

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type (per Hurley et al. 2000) of the primary star immediately following common envelope event.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_NAME\_1\_POST\_COMMON\_ENVELOPE below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **STELLAR\_TYPE\_1\_PRE\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_1<CE

Data type: INT

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type (per Hurley et al. 2000) of the primary star at the onset of RLOF leading to the common-envelope episode.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_NAME\_1\_PRE\_COMMON\_ENVELOPE below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **STELLAR\_TYPE\_2\_POST\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_2>CE

Data type: INT

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type of the secondary star immediately following common envelope event.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_NAME\_2\_POST\_COMMON\_ENVELOPE below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **STELLAR\_TYPE\_2\_PRE\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_2<CE

Data type: INT

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type (per Hurley et al. 2000) of the secondary star at the onset of RLOF leading to the common-envelope episode.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_NAME\_2\_PRE\_COMMON\_ENVELOPE below. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

#### **STELLAR\_TYPE\_NAME\_1\_POST\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_1>CE

Data type: STRING

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type name (per Hurley et al. 2000) of the primary star immediately following common envelope event.

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_1\_POST\_COMMON\_ENVELOPE above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

#### **STELLAR\_TYPE\_NAME\_1\_PRE\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_1<CE

Data type: STRING

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type name (per Hurley et al. 2000) of the primary star at the onset of RLOF leading to the common-envelope episode.

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_1\_PRE\_COMMON\_ENVELOPE above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

#### **STELLAR\_TYPE\_NAME\_2\_POST\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_2>CE

Data type: STRING

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type name (per Hurley et al. 2000) of the secondary star immediately following common envelope event.

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_2\_POST\_COMMON\_ENVELOPE above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

#### **STELLAR\_TYPE\_NAME\_2\_PRE\_COMMON\_ENVELOPE**

Header string: Stellar\_Type\_2<CE

Data type: STRING

COMPAS variable: BinaryConstituentStar::m\_CEDetails.postCEE.stellarType

Stellar type name (per Hurley et al. 2000) of the secondary star at the onset of RLOF leading to the common-envelope episode.

e.g. "First\_Giant\_Branch", "Core\_Helium\_Burning", "Early\_Asymptotic\_Giant\_Branch", etc.

*Note that this property has the same header string as*

*STELLAR\_TYPE\_2\_PRE\_COMMON\_ENVELOPE above. It is expected that one or the other is printed in any file, but not both. If both are printed then the file will contain two columns with the same header string.*

### **SUPERNOVA\_STATE**

Header string: Supernova\_State

Data type: INT

COMPAS variable: <derived from BaseBinaryStar::m\_SupernovaState>

Indicates which star(s) went supernova.

Will be printed as one of:

0 = no supernova

1 = Star 1 is the supernova

2 = Star 2 is the supernova

3 = Both stars are supernovae

### **SYNCHRONIZATION\_TIMESCALE**

Header string: Tau\_Sync

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_SynchronizationTimescale

Tidal synchronisation timescale (Myr).

### **SYSTEMIC\_VELOCITY**

Header string: Systemic\_Velocity

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_SystemicVelocity

Post-supernova systemic (center-of-mass) velocity ( $\text{km s}^{-1}$ ).

### **TIME**

Header string: Time

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_Time

Time (Myrs) since ZAMS.

### **TIME\_TO\_COALESCENCE**

Header string: Coalescence\_Time

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_TimeToCoalescence

Time between formation of double compact object and gravitational-wave driven merger (Myr).

### **TOTAL\_ANGULAR\_MOMENTUM\_PRIME**

Header string: Ang\_Momentum\_Total

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_TotalAngularMomentumPrime

Total angular momentum calculated using regular conservation of energy ( $\text{Msol AU}^2 \text{yr}^{-1}$ ).

### **TOTAL\_ENERGY\_PRIME**

Header string: Energy\_Total

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_TotalAngularMomentumPrime

Total energy calculated using regular conservation of energy ( $\text{Msol AU}^2 \text{yr}^{-1}$ ).

**UNBOUND**

Header string: Unbound

Data type: BOOL

COMPAS variable: BaseBinaryStar::m\_Unbound

Flag to indicate the binary is unbound (or has become unbound after a supernova event).

**WOLF\_RAYET\_FACTOR**

Header string: WR\_Multiplier

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_WolfRayetFactor

Multiplicative constant for Wolf-Rayet mass loss rate

User-supplied via option --wolf-rayet-multiplier.

**ZETA\_LOBE**

Header string: Zeta\_Lobe

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_ZetaLobe

The logarithmic derivative of Roche lobe radius with respect to donor mass for  $q = M_d / M_a$  at the onset of the RLOF.

**ZETA\_STAR**

Header string: Zeta\_Star

Data type: DOUBLE

COMPAS variable: BaseBinaryStar::m\_ZetaStar

Mass-radius exponent of the star at the onset of the RLOF. Calculated differently based on the value of  $-zeta$ -prescription.

## Appendix D – Log File Record Specification: Program Options

As described in **Standard Log File Record Specifiers**, when specifying known properties in a log file record specification record, the property name must be prefixed with the property type.

Currently there is a single program option property type available for use: PROGRAM\_OPTION.

For example, to specify the program option property RANDOM\_SEED, use:

PROGRAM\_OPTION::RANDOM\_SEED

Following is the list of program option properties available for inclusion in log file record specifiers.

### Program Option Properties

#### **KICK\_VELOCITY\_DISTRIBUTION\_SIGMA\_CCSN\_BH**

Header string: Sigma\_Kick\_CCSN\_BH

Data type: DOUBLE

COMPAS variable: Options::kickVelocityDistributionSigmaCCSN\_BH

Value of program option --kick-velocity-sigma-CCSN-BH

#### **KICK\_VELOCITY\_DISTRIBUTION\_SIGMA\_CCSN\_NS**

Header string: Sigma\_Kick\_CCSN\_NS

Data type: DOUBLE

COMPAS variable: Options::kickVelocityDistributionSigmaCCSN\_NS

Value of program option --kick-velocity-sigma-CCSN-NS

#### **KICK\_VELOCITY\_DISTRIBUTION\_SIGMA\_FOR\_ECSN**

Header string: Sigma\_Kick\_ECSN

Data type: DOUBLE

COMPAS variable: Options::kickVelocityDistributionSigmaForECSN

Value of program option --kick-velocity-sigma-ECSN

#### **KICK\_VELOCITY\_DISTRIBUTION\_SIGMA\_FOR\_USSN**

Header string: Sigma\_Kick\_USSN

Data type: DOUBLE

COMPAS variable: Options::kickVelocityDistributionSigmaForUSSN

Value of program option --kick-velocity-sigma-USSN

#### **RANDOM\_SEED**

Header string: SEED\_(ProgramOption)

Data type: UNSIGNED LONG INT

COMPAS variable: Options::randomSeed

Value of program option --random-seed

## Appendix E – Default Log File Record Specifications

Following are the default log file record specifications for each of the standard log files. These specifications can be overridden by the use of a log file specifications file via the logfile-definitions program option.

### SSE Parameters

```
const ANY_PROPERTY_VECTOR SSE_PARAMETERS_REC = {  
    STAR_PROPERTY::AGE,  
    STAR_PROPERTY::DT,  
    STAR_PROPERTY::TIME,  
    STAR_PROPERTY::STELLAR_TYPE,  
    STAR_PROPERTY::METALLICITY,  
    STAR_PROPERTY::MASS_0,  
    STAR_PROPERTY::MASS,  
    STAR_PROPERTY::RADIUS,  
    STAR_PROPERTY::RZAMS,  
    STAR_PROPERTY::LUMINOSITY,  
    STAR_PROPERTY::TEMPERATURE,  
    STAR_PROPERTY::CORE_MASS,  
    STAR_PROPERTY::CO_CORE_MASS,  
    STAR_PROPERTY::HE_CORE_MASS,  
    STAR_PROPERTY::MDOT,  
    STAR_PROPERTY::TIMESCALE_MS  
};
```

## BSE System Parameters

```
const ANY_PROPERTY_VECTOR BSE_SYSTEM_PARAMETERS_REC = {  
    BINARY_PROPERTY::ID,  
    BINARY_PROPERTY::RANDOM_SEED,  
    STAR_1_PROPERTY::MZAMS,  
    STAR_2_PROPERTY::MZAMS,  
    BINARY_PROPERTY::SEMI_MAJOR_AXIS_INITIAL,  
    BINARY_PROPERTY::ECCENTRICITY_INITIAL,  
    STAR_1_PROPERTY::SUPERNOVA_KICK_VELOCITY_MAGNITUDE_RANDOM_NUMBER,  
    STAR_1_PROPERTY::SUPERNOVA_THETA,  
    STAR_1_PROPERTY::SUPERNOVA_PHI,  
    STAR_1_PROPERTY::MEAN_ANOMALY,  
    STAR_2_PROPERTY::SUPERNOVA_KICK_VELOCITY_MAGNITUDE_RANDOM_NUMBER,  
    STAR_2_PROPERTY::SUPERNOVA_THETA,  
    STAR_2_PROPERTY::SUPERNOVA_PHI,  
    STAR_2_PROPERTY::MEAN_ANOMALY,  
    STAR_1_PROPERTY::OMEGA_ZAMS,  
    STAR_2_PROPERTY::OMEGA_ZAMS,  
    PROGRAM_OPTION::KICK_VELOCITY_DISTRIBUTION_SIGMA_CCSN_NS,  
    PROGRAM_OPTION::KICK_VELOCITY_DISTRIBUTION_SIGMA_CCSN_BH,  
    PROGRAM_OPTION::KICK_VELOCITY_DISTRIBUTION_SIGMA_FOR_ECSN,  
    PROGRAM_OPTION::KICK_VELOCITY_DISTRIBUTION_SIGMA_FOR_USSN,  
    BINARY_PROPERTY::LUMINOUS_BLUE_VARIABLE_FACTOR,  
    BINARY_PROPERTY::WOLF_RAYET_FACTOR,  
    BINARY_PROPERTY::COMMON_ENVELOPE_ALPHA,  
    STAR_1_PROPERTY::METALLICITY,  
    STAR_2_PROPERTY::METALLICITY,  
    BINARY_PROPERTY::UNBOUND,  
    BINARY_PROPERTY::STELLAR_MERGER,  
    BINARY_PROPERTY::STELLAR_MERGER_AT_BIRTH,  
    STAR_1_PROPERTY::INITIAL_STELLAR_TYPE,  
    STAR_1_PROPERTY::STELLAR_TYPE,  
    STAR_2_PROPERTY::INITIAL_STELLAR_TYPE,  
    STAR_2_PROPERTY::STELLAR_TYPE,  
    BINARY_PROPERTY::ERROR  
};
```

## BSE Detailed Output

```
const ANY_PROPERTY_VECTOR BSE_DETAILED_OUTPUT_REC = {  
  BINARY_PROPERTY::ID,  
  BINARY_PROPERTY::RANDOM_SEED,  
  BINARY_PROPERTY::DT,  
  BINARY_PROPERTY::TIME,  
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_PRIME_RSOL,  
  BINARY_PROPERTY::ECCENTRICITY_PRIME,  
  STAR_1_PROPERTY::MZAMS,  
  STAR_2_PROPERTY::MZAMS,  
  STAR_1_PROPERTY::MASS_0,  
  STAR_2_PROPERTY::MASS_0,  
  STAR_1_PROPERTY::MASS,  
  STAR_2_PROPERTY::MASS,  
  STAR_1_PROPERTY::ENV_MASS,  
  STAR_2_PROPERTY::ENV_MASS,  
  STAR_1_PROPERTY::CORE_MASS,  
  STAR_2_PROPERTY::CORE_MASS,  
  STAR_1_PROPERTY::HE_CORE_MASS,  
  STAR_2_PROPERTY::HE_CORE_MASS,  
  STAR_1_PROPERTY::CO_CORE_MASS,  
  STAR_2_PROPERTY::CO_CORE_MASS,  
  STAR_1_PROPERTY::RADIUS,  
  STAR_2_PROPERTY::RADIUS,  
  BINARY_PROPERTY::ROCHE_LOBE_RADIUS_1,  
  BINARY_PROPERTY::ROCHE_LOBE_RADIUS_2,  
  BINARY_PROPERTY::ROCHE_LOBE_TRACKER_1,  
  BINARY_PROPERTY::ROCHE_LOBE_TRACKER_2,  
  STAR_1_PROPERTY::OMEGA,  
  STAR_2_PROPERTY::OMEGA,  
  STAR_1_PROPERTY::OMEGA_BREAK,  
  STAR_2_PROPERTY::OMEGA_BREAK,  
  STAR_1_PROPERTY::INITIAL_STELLAR_TYPE,  
  STAR_2_PROPERTY::INITIAL_STELLAR_TYPE,  
  STAR_1_PROPERTY::STELLAR_TYPE,  
  STAR_2_PROPERTY::STELLAR_TYPE,  
  STAR_1_PROPERTY::AGE,  
  STAR_2_PROPERTY::AGE,  
  STAR_1_PROPERTY::LUMINOSITY,  
  STAR_2_PROPERTY::LUMINOSITY,  
  STAR_1_PROPERTY::TEMPERATURE,  
  STAR_2_PROPERTY::TEMPERATURE,  
  STAR_1_PROPERTY::ANGULAR_MOMENTUM,  
  STAR_2_PROPERTY::ANGULAR_MOMENTUM,  
  STAR_1_PROPERTY::DYNAMICAL_TIMESCALE,  
  STAR_2_PROPERTY::DYNAMICAL_TIMESCALE,  
  STAR_1_PROPERTY::THERMAL_TIMESCALE,  
  STAR_2_PROPERTY::THERMAL_TIMESCALE,  
  STAR_1_PROPERTY::NUCLEAR_TIMESCALE,  
  STAR_2_PROPERTY::NUCLEAR_TIMESCALE,  
  STAR_1_PROPERTY::ZETA_SOBERMAN,  
  STAR_2_PROPERTY::ZETA_SOBERMAN,  
  STAR_1_PROPERTY::ZETA_SOBERMAN_HE,  
  STAR_2_PROPERTY::ZETA_SOBERMAN_HE,  
  STAR_1_PROPERTY::ZETA_HURLEY,  
  STAR_2_PROPERTY::ZETA_HURLEY,  
  STAR_1_PROPERTY::ZETA_HURLEY_HE,  
  STAR_2_PROPERTY::ZETA_HURLEY_HE,  
  STAR_1_PROPERTY::MASS_LOSS_DIFF,  
  STAR_2_PROPERTY::MASS_LOSS_DIFF,  
  STAR_1_PROPERTY::MASS_TRANSFER_DIFF,  
  STAR_2_PROPERTY::MASS_TRANSFER_DIFF,  
  BINARY_PROPERTY::TOTAL_ANGULAR_MOMENTUM_PRIME,  
}
```



```

    BINARY_PROPERTY::TOTAL_ENERGY_PRIME,
    STAR_1_PROPERTY::LAMBDA_NANJING,
    STAR_2_PROPERTY::LAMBDA_NANJING,
    STAR_1_PROPERTY::LAMBDA_LOVERIDGE,
    STAR_2_PROPERTY::LAMBDA_LOVERIDGE,
    STAR_1_PROPERTY::LAMBDA_KRUCKOW_TOP,
    STAR_2_PROPERTY::LAMBDA_KRUCKOW_TOP,
    STAR_1_PROPERTY::LAMBDA_KRUCKOW_MIDDLE,
    STAR_2_PROPERTY::LAMBDA_KRUCKOW_MIDDLE,
    STAR_1_PROPERTY::LAMBDA_KRUCKOW_BOTTOM,
    STAR_2_PROPERTY::LAMBDA_KRUCKOW_BOTTOM,
    STAR_1_PROPERTY::METALLICITY,
    STAR_2_PROPERTY::METALLICITY,
    BINARY_PROPERTY::MASS_TRANSFER_TRACKER_HISTORY,
    STAR_1_PROPERTY::PULSAR_MAGNETIC_FIELD,
    STAR_2_PROPERTY::PULSAR_MAGNETIC_FIELD,
    STAR_1_PROPERTY::PULSAR_SPIN_FREQUENCY,
    STAR_2_PROPERTY::PULSAR_SPIN_FREQUENCY,
    STAR_1_PROPERTY::PULSAR_SPIN_DOWN_RATE,
    STAR_2_PROPERTY::PULSAR_SPIN_DOWN_RATE,
    STAR_1_PROPERTY::RADIAL_EXPANSION_TIMESCALE,
    STAR_2_PROPERTY::RADIAL_EXPANSION_TIMESCALE
};

```

## BSE Double Compact Objects

```
const ANY_PROPERTY_VECTOR BSE_DOUBLE_COMPACT_OBJECTS_REC = {  
  BINARY_PROPERTY::ID,  
  BINARY_PROPERTY::RANDOM_SEED,  
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_AT_DCO_FORMATION,  
  BINARY_PROPERTY::ECCENTRICITY_AT_DCO_FORMATION,  
  STAR_1_PROPERTY::MASS,  
  STAR_1_PROPERTY::STELLAR_TYPE,  
  STAR_2_PROPERTY::MASS,  
  STAR_2_PROPERTY::STELLAR_TYPE,  
  BINARY_PROPERTY::TIME_TO_COALESCENCE,  
  BINARY_PROPERTY::TIME,  
  STAR_1_PROPERTY::MASS_TRANSFER_CASE_INITIAL,  
  STAR_2_PROPERTY::MASS_TRANSFER_CASE_INITIAL,  
  BINARY_PROPERTY::MERGES_IN_HUBBLE_TIME,  
  STAR_1_PROPERTY::RECYCLED_NEUTRON_STAR,  
  STAR_2_PROPERTY::RECYCLED_NEUTRON_STAR,  
};
```

## BSE Common Envelopes

```
const ANY_PROPERTY_VECTOR BSE_COMMON_ENVELOPES_REC = {
  BINARY_PROPERTY::ID,
  BINARY_PROPERTY::RANDOM_SEED,
  BINARY_PROPERTY::TIME,
  STAR_1_PROPERTY::LAMBDA_AT_COMMON_ENVELOPE,
  STAR_2_PROPERTY::LAMBDA_AT_COMMON_ENVELOPE,
  STAR_1_PROPERTY::BINDING_ENERGY_PRE_COMMON_ENVELOPE,
  STAR_2_PROPERTY::BINDING_ENERGY_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::ECCENTRICITY_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::ECCENTRICITY_POST_COMMON_ENVELOPE,
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_POST_COMMON_ENVELOPE,
  BINARY_PROPERTY::ROCHE_LOBE_RADIUS_1_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::ROCHE_LOBE_RADIUS_1_POST_COMMON_ENVELOPE,
  BINARY_PROPERTY::ROCHE_LOBE_RADIUS_2_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::ROCHE_LOBE_RADIUS_2_POST_COMMON_ENVELOPE,
  BINARY_PROPERTY::MASS_1_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::MASS_ENV_1,
  BINARY_PROPERTY::MASS_1_FINAL,
  BINARY_PROPERTY::RADIUS_1_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::RADIUS_1_POST_COMMON_ENVELOPE,
  BINARY_PROPERTY::STELLAR_TYPE_1_PRE_COMMON_ENVELOPE,
  STAR_1_PROPERTY::STELLAR_TYPE,
  STAR_1_PROPERTY::LAMBDA_FIXED,
  STAR_1_PROPERTY::LAMBDA_NANJING,
  STAR_1_PROPERTY::LAMBDA_LOVERIDGE,
  STAR_1_PROPERTY::LAMBDA_LOVERIDGE_WINDS,
  STAR_1_PROPERTY::LAMBDA_KRUCKOW,
  STAR_1_PROPERTY::BINDING_ENERGY_FIXED,
  STAR_1_PROPERTY::BINDING_ENERGY_NANJING,
  STAR_1_PROPERTY::BINDING_ENERGY_LOVERIDGE,
  STAR_1_PROPERTY::BINDING_ENERGY_LOVERIDGE_WINDS,
  STAR_1_PROPERTY::BINDING_ENERGY_KRUCKOW,
  BINARY_PROPERTY::MASS_2_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::MASS_ENV_2,
  BINARY_PROPERTY::MASS_2_FINAL,
  BINARY_PROPERTY::RADIUS_2_PRE_COMMON_ENVELOPE,
  BINARY_PROPERTY::RADIUS_2_POST_COMMON_ENVELOPE,
  BINARY_PROPERTY::STELLAR_TYPE_2_PRE_COMMON_ENVELOPE,
  STAR_2_PROPERTY::STELLAR_TYPE,
  STAR_2_PROPERTY::LAMBDA_FIXED,
  STAR_2_PROPERTY::LAMBDA_NANJING,
  STAR_2_PROPERTY::LAMBDA_LOVERIDGE,
  STAR_2_PROPERTY::LAMBDA_LOVERIDGE_WINDS,
  STAR_2_PROPERTY::LAMBDA_KRUCKOW,
  STAR_2_PROPERTY::BINDING_ENERGY_FIXED,
  STAR_2_PROPERTY::BINDING_ENERGY_NANJING,
  STAR_2_PROPERTY::BINDING_ENERGY_LOVERIDGE,
  STAR_2_PROPERTY::BINDING_ENERGY_LOVERIDGE_WINDS,
  STAR_2_PROPERTY::BINDING_ENERGY_KRUCKOW,
  BINARY_PROPERTY::MASS_TRANSFER_TRACKER_HISTORY,
  BINARY_PROPERTY::STELLAR_MERGER,
  BINARY_PROPERTY::OPTIMISTIC_COMMON_ENVELOPE,
  BINARY_PROPERTY::COMMON_ENVELOPE_EVENT_COUNT,
  BINARY_PROPERTY::DOUBLE_CORE_COMMON_ENVELOPE,
  STAR_1_PROPERTY::IS_RLOF,
  STAR_1_PROPERTY::LUMINOSITY_PRE_COMMON_ENVELOPE,
  STAR_1_PROPERTY::TEMPERATURE_PRE_COMMON_ENVELOPE,
  STAR_1_PROPERTY::DYNAMICAL_TIMESCALE_PRE_COMMON_ENVELOPE,
  STAR_1_PROPERTY::THERMAL_TIMESCALE_PRE_COMMON_ENVELOPE,
  STAR_1_PROPERTY::NUCLEAR_TIMESCALE_PRE_COMMON_ENVELOPE,
  STAR_2_PROPERTY::IS_RLOF,
```

```

STAR_2_PROPERTY::LUMINOSITY_PRE_COMMON_ENVELOPE,
STAR_2_PROPERTY::TEMPERATURE_PRE_COMMON_ENVELOPE,
STAR_2_PROPERTY::DYNAMICAL_TIMESCALE_PRE_COMMON_ENVELOPE,
STAR_2_PROPERTY::THERMAL_TIMESCALE_PRE_COMMON_ENVELOPE,
STAR_2_PROPERTY::NUCLEAR_TIMESCALE_PRE_COMMON_ENVELOPE,
BINARY_PROPERTY::ZETA_STAR_COMPARE,
BINARY_PROPERTY::ZETA_RLOF_ANALYTIC,
BINARY_PROPERTY::SYNCHRONIZATION_TIMESCALE,
BINARY_PROPERTY::CIRCULARIZATION_TIMESCALE,
STAR_1_PROPERTY::RADIAL_EXPANSION_TIMESCALE_PRE_COMMON_ENVELOPE,
STAR_2_PROPERTY::RADIAL_EXPANSION_TIMESCALE_PRE_COMMON_ENVELOPE,
BINARY_PROPERTY::IMMEDIATE_RLOF_POST_COMMON_ENVELOPE,
BINARY_PROPERTY::SIMULTANEOUS_RLOF
};

```

## BSE Supernovae

```
const ANY_PROPERTY_VECTOR BSE_SUPERNOVAE_REC = {
  BINARY_PROPERTY::ID,
  BINARY_PROPERTY::RANDOM_SEED,
  SUPERNOVA_PROPERTY::DRAWN_KICK_VELOCITY,
  SUPERNOVA_PROPERTY::KICK_VELOCITY,
  SUPERNOVA_PROPERTY::FALLBACK_FRACTION,
  BINARY_PROPERTY::ORBITAL_VELOCITY_PRE_SUPERNOVA,
  BINARY_PROPERTY::DIMENSIONLESS_KICK_VELOCITY,
  SUPERNOVA_PROPERTY::TRUE_ANOMALY,
  SUPERNOVA_PROPERTY::SUPERNOVA_THETA,
  SUPERNOVA_PROPERTY::SUPERNOVA_PHI,
  SUPERNOVA_PROPERTY::SN_TYPE,
  BINARY_PROPERTY::UNBOUND,
  SUPERNOVA_PROPERTY::TOTAL_MASS_AT_COMPACT_OBJECT_FORMATION,
  COMPANION_PROPERTY::MASS,
  SUPERNOVA_PROPERTY::CO_CORE_MASS_AT_COMPACT_OBJECT_FORMATION,
  SUPERNOVA_PROPERTY::MASS,
  SUPERNOVA_PROPERTY::EXPERIENCED_RLOF,
  SUPERNOVA_PROPERTY::STELLAR_TYPE,
  BINARY_PROPERTY::SUPERNOVA_STATE,
  SUPERNOVA_PROPERTY::STELLAR_TYPE_PREV,
  COMPANION_PROPERTY::STELLAR_TYPE_PREV,
  SUPERNOVA_PROPERTY::CORE_MASS_AT_COMPACT_OBJECT_FORMATION,
  SUPERNOVA_PROPERTY::HE_CORE_MASS_AT_COMPACT_OBJECT_FORMATION,
  BINARY_PROPERTY::TIME,
  BINARY_PROPERTY::ECCENTRICITY_PRE_SUPERNOVA,
  BINARY_PROPERTY::ECCENTRICITY,
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_PRE_SUPERNOVA_RSOL,
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_PRIME_RSOL,
  BINARY_PROPERTY::SYSTEMIC_VELOCITY,
  SUPERNOVA_PROPERTY::HYDROGEN_RICH,
  SUPERNOVA_PROPERTY::HYDROGEN_POOR,
  COMPANION_PROPERTY::RUNAWAY
};
```

## BSE Pulsar Evolution

```
const ANY_PROPERTY_VECTOR BSE_PULSAR_EVOLUTION_REC = {
  BINARY_PROPERTY::ID,
  BINARY_PROPERTY::RANDOM_SEED,
  STAR_1_PROPERTY::MASS,
  STAR_2_PROPERTY::MASS,
  STAR_1_PROPERTY::STELLAR_TYPE,
  STAR_2_PROPERTY::STELLAR_TYPE,
  BINARY_PROPERTY::SEMI_MAJOR_AXIS_PRIME_RSOL,
  BINARY_PROPERTY::MASS_TRANSFER_TRACKER_HISTORY,
  STAR_1_PROPERTY::PULSAR_MAGNETIC_FIELD,
  STAR_2_PROPERTY::PULSAR_MAGNETIC_FIELD,
  STAR_1_PROPERTY::PULSAR_SPIN_FREQUENCY,
  STAR_2_PROPERTY::PULSAR_SPIN_FREQUENCY,
  STAR_1_PROPERTY::PULSAR_SPIN_DOWN_RATE,
  STAR_2_PROPERTY::PULSAR_SPIN_DOWN_RATE,
  BINARY_PROPERTY::TIME,
  BINARY_PROPERTY::DT
};
```

## Appendix F – Example Log File Record Specifications File

Following is an example log file record specifications file. COMPAS can be configured to use this file via the logfile-definitions program option.

This file (COMPAS\_Output\_Definitions.txt) is also delivered as part of the COMPAS github repository.

```
# sample standard log file specifications file

# the '#' character and anything following it on a single line is considered a comment
# (so, lines starting with '#' are comment lines)

# case is not significant
# specifications can span several lines
# specifications for the same log file are cumulative
# if a log file is not specified in this file, the default specification is used

# SSE Parameters

# start with the default SSE Parameters specification and add ENV_MASS

sse_parms_rec += { STAR_PROPERTY::ENV_MASS }

# take the updated SSE Parameters specification and add ANGULAR_MOMENTUM

sse_parms_rec += { STAR_PROPERTY::ANGULAR_MOMENTUM }

# take the updated SSE Parameters specification and subtract MASS_0 and MDOT

sse_parms_rec -= { STAR_PROPERTY::MASS_0, STAR_PROPERTY::MDOT }

# BSE System Parameters

bse_sysparms_rec = {
    BINARY_PROPERTY::ID,          # set the BSE System Parameters specification to:
    BINARY_PROPERTY::RANDOM_SEED,  # ID of the binary
    STAR_1_PROPERTY::MZAMS,       # RANDOM_SEED for the binary
    STAR_2_PROPERTY::MZAMS       # MZAMS for Star1
}
    STAR_1_PROPERTY::MZAMS       # MZAMS for Star2

# ADD to the BSE System Parameters specification:
# SEMI_MAJOR_AXIS_INITIAL for the binary
# ECCENTRICITY_INITIAL for the binary
# SUPERNOVA_THETA for Star1 and SUPERNOVA_PHI for Star1

bse_sysparms_rec += {
    BINARY_PROPERTY::SEMI_MAJOR_AXIS_INITIAL,
    BINARY_PROPERTY::ECCENTRICITY_INITIAL,
    STAR_1_PROPERTY::SUPERNOVA_THETA, STAR_1_PROPERTY::SUPERNOVA_PHI
}

bse_sysparms_rec += {
    SUPERNOVA_PROPERTY::IS_ECSN,    # ADD to the BSE System Parameters specification:
    SUPERNOVA_PROPERTY::IS_SN,      # IS_ECSN for the supernova star
    SUPERNOVA_PROPERTY::IS_USSN,    # IS_SN for the supernova star
    SUPERNOVA_PROPERTY::IS_USSN,    # IS_USSN for the supernova star
    SUPERNOVA_PROPERTY::EXPERIENCED_PISN, # EXPERIENCED_PISN for the supernova star
    SUPERNOVA_PROPERTY::EXPERIENCED_PPISN, # EXPERIENCED_PPISN for the supernova star
    BINARY_PROPERTY::SURVIVED_SUPERNOVA_EVENT, # SURVIVED_SUPERNOVA_EVENT for the binary
    SUPERNOVA_PROPERTY::MZAMS,      # MZAMS for the supernova star
    COMPANION_PROPERTY::MZAMS       # MZAMS for the companion star
}

# SUBTRACT from the BSE System Parameters specification:
# RANDOM_SEED for the binary
# ID for the binary
bse_sysparms_rec -= {
    BINARY_PROPERTY::RANDOM_SEED,    # SUBTRACT from the BSE System Parameters specification:
    BINARY_PROPERTY::ID             # RANDOM_SEED for the binary
}
    BINARY_PROPERTY::ID             # ID for the binary

# BSE Double Compas Objects
```

```

# set the BSE Double Compact Objects specifivation to MZAMS for Star1, and MZAMS for Star2
BSE_DCO_Rec = {STAR_1_PROPERTY::MZAMS, STAR_2_PROPERTY::MZAMS}

# set the BSE Double Compact Objects specification to empty - nothing will be printed
# (file will not be created)
BSE_DCO_Rec = {}

# BSE Supernovae
BSE_SNE_Rec = {}          # set spec empty - nothing will be printed (file will not be created)

# BSE Common Envelopes
BSE_CEE_Rec = {   }      # set spec empty - nothing will be printed (file will not be created)

# BSE Pulsars
# line ignored (comment). BSE Pulsars specification will be default
# BSE_Pulsars_Rec= { STAR_1_PROPERTY::MASS, STAR_2_PROPERTY::MASS }

# BSE Detailed Output
BSE_Detailed_Rec={}      # set spec empty - nothing will be printed (file will not be created)

```