



## NCES unit-1

Non-Conventional Energy Sources (Jawaharlal Nehru Technological University,  
Hyderabad)



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\* NON-CONVENTIONAL SOURCES OF ENERGY \*

\* CREDITS - 3 \*

\* SUBJECT CODE :- ME8000E \*

\* UNIT - I \*

\* NOTES \*

\* PRINCIPLES OF SOLAR RADIATION \*

# \* NON-CONVENTIONAL SOURCES OF ENERGY \* ①

## \* Energy :-

- \* It is the primary and most universal measure of all kinds of work by human beings and Nature.
- \* Everything what happens in the world is the expression of flow of energy in one of its forms.
- \* The energy sources available can be divided into Three Types :-

### ① Primary energy sources :-

It can be defined as sources which provide a net supply of energy.  
ex:- coal, oil, uranium etc.

### ② Secondary fuels :-

This sources produce no net energy. though it may be necessary for the economy.  
ex:- Intensive agricultural.

### ③ Supplementary sources :-

This sources are those whose net energy yield is zero.  
ex:- Thermal insulation

- \* DEFINITION OF NON-CONVENTIONAL SOURCE OF ENERGY:-
- \* Non-conventional sources of energy is also known as Renewable sources of energy. all energy sources that are continually renewed by natural processes
- \* It includes such as sunlight, wind, the-movement of water and geothermal heat.

### \* DEFINITION OF CONVENTIONAL SOURCE OF ENERGY:-

- \* Source of energy that has been in use for centuries are called conventional source of energy.
- \* It includes wood, coal, petroleum and flowing water.

### \* UNIT-1 \*

#### \* PRINCIPLES OF SOLAR RADIATION \*

- \* Solar Radiation:  
convert sunlight into electricity either through photo voltaic cell (or) through mirror that concentrates solar radiation.

- \* Role and potential of new and renewable source:-  
Attention of scientists and engineers all over the world has been drawn to develop alternative energy technologies, since the oil crisis in 1973.

- \* With technological advancements, people have now become aware of the demerits of burning fossil-fuels and thus it is observed that by end of 21st century fossil fuels get exhausted.

\* As population increases, demand for energy- ② increases.

\* Hence, non conventional sources of energy must be given much significance, which can fulfill the future-power demand.

\* Potential of New and Renewable sources:-

\* India ranks 3<sup>rd</sup> place in production of electricity through Renewable Sources.

\* High potential for generation of renewable energy from - wind, solar, Biomass etc.

\* In November 2021, India has Renewable energy capacity of 150GW consisting of

\* Solar - 48.55 Gigawatts (GW)

\* Wind - 40.03 GW

\* Small Hydro - 4.83 GW

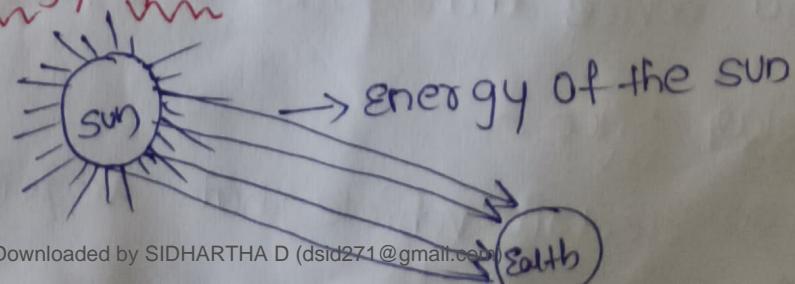
\* Biomass - 10.62 GW

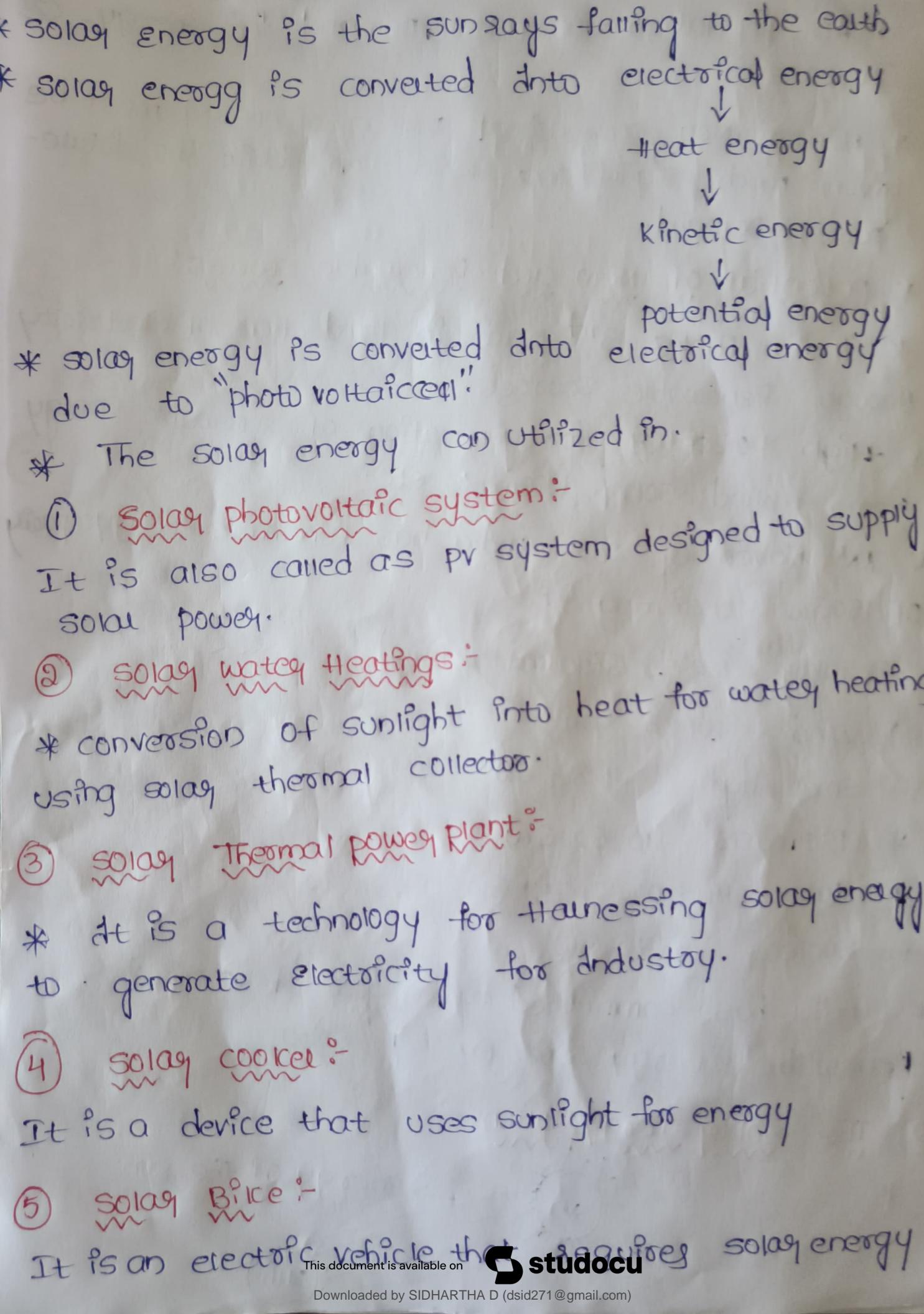
\* Large Hydro - 46.51 GW

\* Nuclear - 6.78 GW

\* India has committed for a goal of 450(GW) Renewable energy capacity by 2030

② The solar energy option:-





to charge the battery.

### ⑥ Solar pump:-

It is a pumping running on electricity generated by photo voltaic panels.

### ⑦ Solar power tower:-

It is a type of solar furnace using a tower to receive the focussed light.

### ⑧ Environmental Impact of solar power:-

- \* The sun provides a tremendous resource for generating clean and sustainable electricity without toxic pollution.
- \* The potential environmental impacts associated with solar power - land use and habitat loss, water use and the use of hazardous materials in manufacturing can vary greatly depending on the technology, which includes two broad categories.
- \* photovoltaic cells (or) concentrating solar thermal plants.
- \* These PV and CSP projects also play a significant role in the level of environmental impact.
- \* Land use:
  - \*) It requires the large area of land this include

Inland grabbing in high populated areas.

- 2) Estimates for utility-scale PV systems range from 3.5 to 10 acres per megawatt.
- 3) While estimates for CSP facilities are between 4 and 16.5 acres per megawatt.
- 4) There is less opportunity for solar projects to share land with agricultural uses.
- 5) Smaller scale solar PV arrays, which can be built on homes (or) commercial buildings, also have minimal land use impact.

## ② Water use:

- \* Solar PV cells do not use water for generating electricity. However as in all manufacturing processes, some water is used to manufacture solar PV components.
- \* Concentrating solar thermal plants, like all thermal electric plants, require water for cooling.
- \* Water use depends on the plant design, plant location, and the type of cooling system.
- \* CSP plants that use wet-recirculating technology with cooling towers withdraw b/w 600 and 650 gallons of water per megawatt-hour of electricity produced.

- \* CSP plants with once-through cooling technology have higher levels of water withdrawal, but lower total water consumption.
- \* Dry-cooling Technology can reduce water use at CSP plants by approximately 90%.
- \* Many of the regions in the United States that have the highest potential for solar energy also tend to be those with the driest climates, so careful consideration of these water trade-offs is essential.

### ③ Hazardous materials:

- \* The PV cell manufacturing process includes a number of hazardous materials most of which are used to clean and purify the semiconductor surface.
- \* These chemicals, similar to those used in the general semiconductor industry include hydrochloric acid, sulfuric acid, nitric acid, hydrogen fluoride, 1,1,1-trichloroethane and acetone.
- \* The amount and type of chemicals used depends on the type of cell.
- \* Workers also face risks associated with inhaling silicon dust.
- \* Thus, PV manufacturers must follow U.S. law's to ensure that workers are not harmed by

exposure to these chemicals. and that manufacturing waste products are disposed of properly.

\* Thin-film PV cells contain a number of more toxic materials than those used in traditional silicon photovoltaic cells, including gallium arsenide, copper indium-gallium-[~~selenide~~, and ~~arsenide~~] and cadmium-telluride.

\* If not handled and disposed properly, these materials could pose serious environmental or public health threats.

#### ④ Life-cycle Global warming emissions:

\* While there are no global warming emissions associated with generating electricity from solar-energy, there are emissions associated with other stages of the solar life-cycle. Including manufacturing, materials transportation, installation, maintenance and decommissioning and dismantlement.

\* Most estimates of life-cycle emissions for photovoltaic system are between 0.07 and 0.18 pounds of carbon dioxide equivalent per kilo-watt-hour.

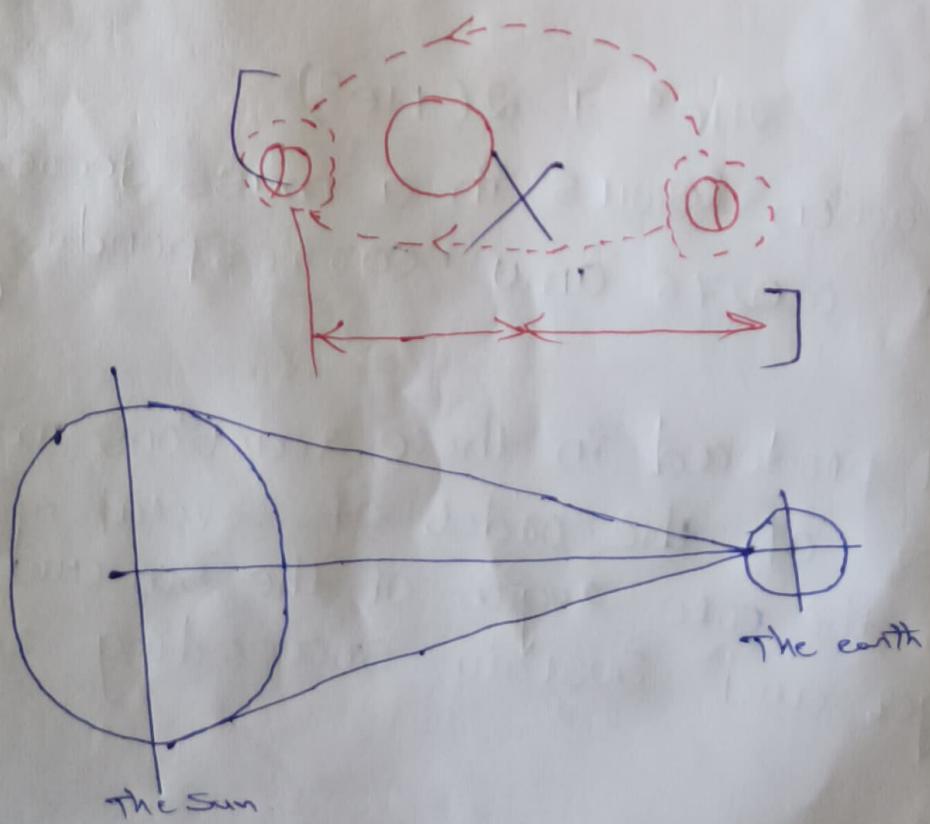
\* Most estimates for concentrating solar power-range from 0.08 to 0.2 pounds of carbon dioxide equivalent per kilowatt-hour.

\* In both cases, this is far less than the

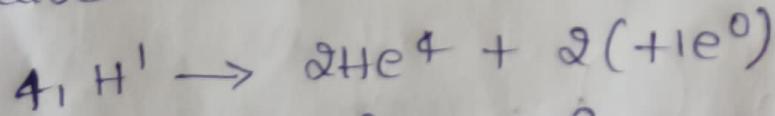
life cycle emission rates for Natural gas (0.6-2 lbs of CO<sub>2</sub> e/kwh) and coal (1.4-3.6 lbs of CO<sub>2</sub> e/kwh) ⑤

### \* Physics of the Sun:-

- \* Sun is a medium sized yellow star that may be considered as a sphere of intensely hot gaseous matter with an average diameter of  $1.39 \times 10^9$  m at an average distance of  $1.495 \times 10^{11}$  m (or 1 Astronomical Unit) from the earth.
- \* The sun coalesced from a cloud of gas and dust formed 4.5 billion years ago.
- \* Earth goes around the sun in an elliptic orbit.



- \* The earth is closest to the sun at  $1.47 \times 10^{11} \text{ m}$  the perihelion around January 2 each year.
- \* It is farthest at  $1.526 \times 10^{11} \text{ m}$  around July 2 each year Sun.
- \* Most important fusion reaction is hydrogen combining to form helium.
- \* Energy is produced in the sun by continuous-fusion in which four nuclei of hydrogen fuse in a series of reactions involving other-particles that continually appear and disappear in the course of the reactions.
- \* such as  $\text{He}_3$ , nitrogen, carbon and other nuclei but culminating in one nucleus of helium and two photons. Resulting in a mass decrease.



- \* This reaction results in a mass decrease of about 0.0276 amu corresponding to 25.7 Mev.
- \* The heat produced in these reactions maintains temperatures of the order of several million degrees in the core region of the sun and several to trigger and sustain succeeding reactions.

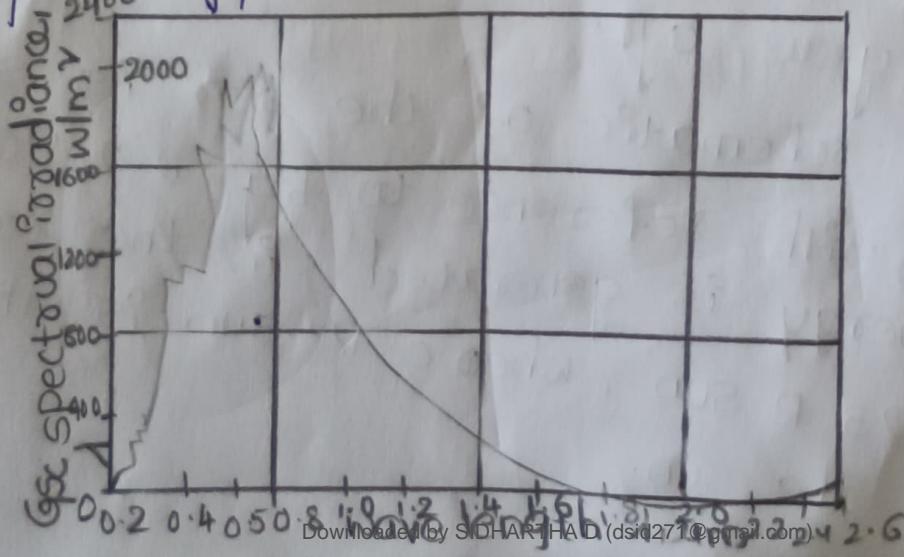
\* The Temperature at  $0.7 R_{\text{SUN}}$  is the order of  $130 \times 10^3$  K. from the region  $r > R_{\text{SUN}}$ , convection process begins and is referred to as the convective zone.

\* The upper layer of the convective zone is the photosphere.

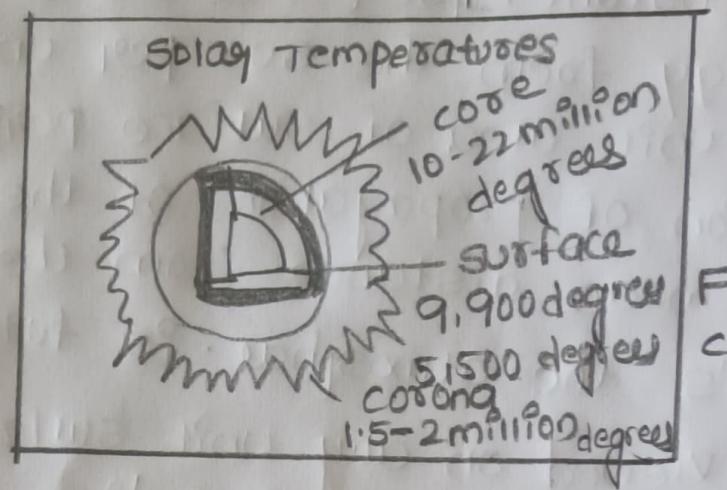
\* which is the source of most solar radiation. other layers following the leveling layer are, chromosphere, and lastly, the corona, of very low density and at a temperature of 106K.

\* Solar Radiation is the composite result of several layers which absorb and emit at various wave lengths.

\* for thermal purposes it is adequate to consider the sun to be a black body at an effective temperature of 5762K. This information is adequate for many solar energy calculations.



- \* The sun is made up of about  $2 \times 10^{30}$  kg of gas. It is composed of about 75% hydrogen and 25% helium.
- \* About 0.1% is metals made from hydrogen via nuclear fusion.
- \* It has been estimated that the sun has used up about half of its initial hydrogen available 4.5 billion years ago.
- \* A simple schematic of the sun is shown in fig 2.2 with reference to fig 2.2, 90% of the energy is generated in the region  $0 < r < 0.23 R_{\text{sun}}$ , where  $R_{\text{sun}}$  is the radius of the sun.



The details of the sun's temperature are depicted in fig 2.4. the sun's core can reach  $5.5 \times 10^6^\circ\text{C}$  to  $12.5 \times 10^6^\circ\text{C}$ . the estimates vary, the higher one going upto  $40 \times 10^6^\circ\text{C}$ . the surface temperature of the sun goes up to  $1.5 \times 10^6^\circ\text{C}$  outer atmosphere of the sun is  $2 \times 10^6^\circ\text{C}$  degrees.

\* The effective temperature of the sun is determined by measuring how much energy it emits.

- \* The solar constant.
- \* The solar constant,  $I_{SC}$  [it is the flux density]  
It is the energy from the sun per unit time,  
received on a unit area of surface perpendicular  
to the direction of propagation of the radiation  
at the earth's mean distance from the sun.  
at the earth's the atmosphere is fully Transpalet
- \* If the earth's the atmosphere is fully Transpalet
- \* It may be viewed on any unit surface normal  
to sun rays on a sphere of radius equal to  
the sun-earth mean distance.
- \* The recently reported value of the solar constant  
is  $1367 \text{ W/m}^2$
- \* The physical idea can be obtained by considering  
the sun to be a sphere of diameter  $1.39 \times 10^9 \text{ m}$ .  
emitting as a black body at an effective temperature  
of  $5762 \text{ K}$ .
- \* The amount of radiation emitted by the sun  
as  $T^4$  would be  $\sigma$  is the Stefan-Boltzmann  
constant ( $= 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$ )
- \* As  $\sigma$  is the surface area of the sun. the  
same amount of radiation passes through any  
sphere surrounding the sun as center.
- \* Thus, the surface of a sphere of radius  
equal to the sun-earth mean distance will  
experience Intensity ( $\text{W/m}^2$ ) Inversely proportional

to the square of the radius of the sphere<sup>(3)</sup> enclosing the sun. This proportioning leads approximately to  $1349 \text{ W/m}^2$ .

\* The variation in the total radiation emitted by sun is less than the energy in large spectrum of the entire solar spectrum and when the transmittance of atmosphere is a major uncertainty, the emitted energy by the sun may be considered as constant.

### \* Extra-terrestrial solar radiation :-

\* Solar radiation on a surface normal to sun rays kept at a distance of sun-earth.

\*  $G_{on}$ , will essentially be the solar constant, modified to take into account the varying distance b/w the sun and the earth.  $G_{on}$  is given by  $(G_{on}) = I_{sc} \left( 1 + 0.033 \cos \left[ \frac{360n}{365} \right] \right)$

In eq(3.1) n, is the day of the year, i.e. the sequential number of the day counting January 1 as 1.

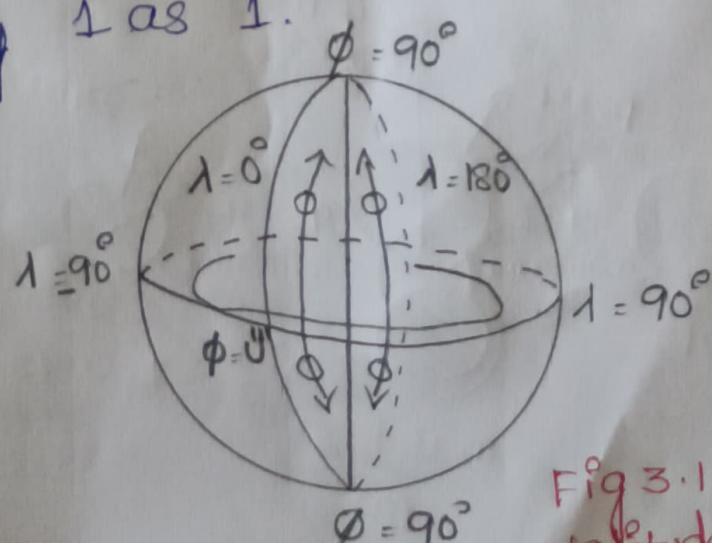
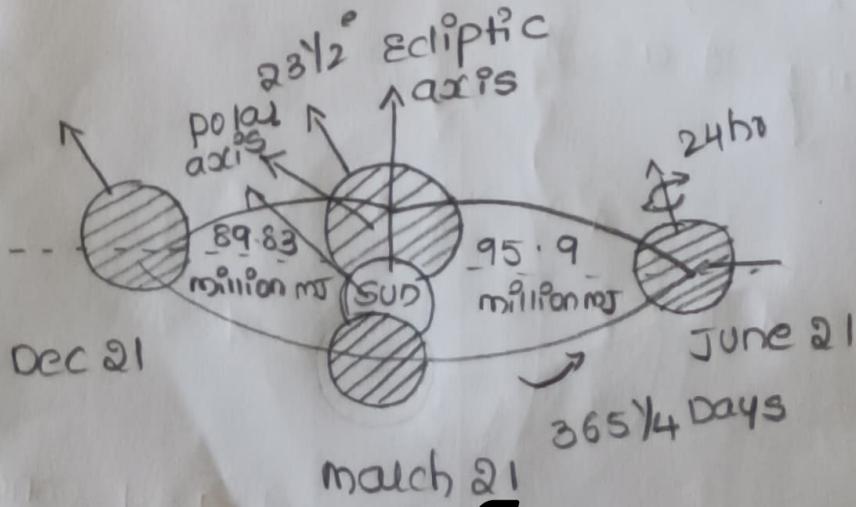


Fig 3.1 concept of latitude, and longitude

- \* Latitude and longitude are the coordinates of a point on earth's surface.
- \* A location on the earth surface may be specified by latitude, longitude and elevation.
- \* Longitudes, i.e., is described along with "east" (or) "west", meaning that the location is situated to east (or) to west of the Greenwich meridian.
- \* Thus the latitude values from  $-90^{\circ}$  to  $+90^{\circ}$  and the longitude values from  $0^{\circ}$  to  $180^{\circ}\text{E}$  (or)  $0^{\circ}$  to  $180^{\circ}\text{W}$ .
- \* The latitude of a point on the earth's surface is the angle b/w the equatorial plane and a line that passes through that point and is normal to the surface.
- \* The North pole is  $90^{\circ}\text{N}$ , the South pole is  $90^{\circ}\text{S}$ .
- \* TERRESTRIAL Solar Radiation



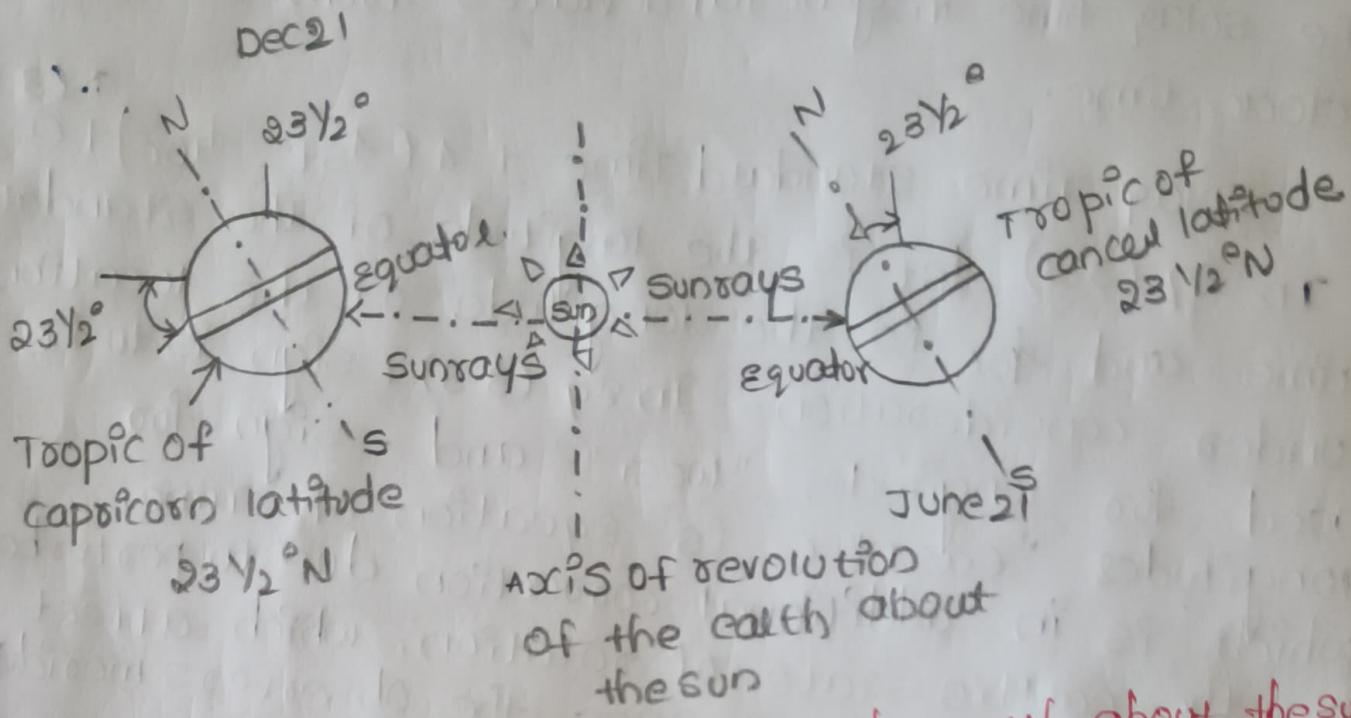


Fig 3.2 Motion of the earth about the sun

\* Declination, is the angular position of the Sun at solar noon with respect to the place of the equator. Declination varies from  $-23.45^{\circ}$  to  $23.45^{\circ}$  in degrees, for the day  $D$  of the year, it can be calculated from:

$$\delta = 23.45 \sin \left( 360 \frac{284+D}{365} \right) \quad (3.2)$$

positions of the earth relative to Sun at different times of the year are shown in fig 3.2:

\* Solar Time is the time based on the apparent motion of the Sun across the sky, with solar noon being the time the Sun crosses the meridian of the observer.

\* Solar Time is used in all sun angle relationships  
The solar time in general deviates from the local-  
clock time.

Solar Time = standard time  $\pm 4(1st - loc) + \epsilon$  Eq (3.3)

\* In Eq (3.3) 1st is the longitude of the standard time and loc is the longitude of the location under consideration. In Eq (3.3) + sign is to be used for longitudes "West" and - sign for longitudes "East"  $\epsilon$  is equation of time, which is in additional correction due to perturbations in the earth's rate of rotation, which affect the time the sun crosses the observer's meridian. "E" in minutes can be calculated from

$$\epsilon = 9.87 \sin 2B - 7.53 \cos B - 1.5 \sin B \quad (3.4)$$

\* In Eq (3.4)

$$B = \frac{360(n - 81)}{364}, \text{ where } n \text{ is the day of the year, } 1 \leq n \leq 365$$

\* Hour angle,  $h$ , is the angular displacement of the sun east (or) west of the local meridian due to rotation.

## \* SOLAR RADIATION ON TITLED SURFACE :-

- \* The solar radiation received by a tilted surface comprises of Direct, diffuse and ground reflected components.
- \* A horizontal surface receives the direct and diffuse components of radiation.
- \* While the titled surface receives in addition, a ground reflected component of the surface 'sees' the ground.
- \* The procedure describes to evaluate a factor referred to as the Tilt factor for direct radiation which when multiplies the direct radiation available on a horizontal surface yields the direct radiation on the titled surface.
- \* The ground reflected radiation would be the product of Solar Radiation on the horizontal surface yields the direct radiation on the titled surface.
- \* It is conceivable that "G" comprises of, & designated by the declination.
- \* Where  $q_b$  and  $q_d$  are the direct and diffuse components of Intensity. the total radiation comprising of the three components at the instant on the titled surface  $q_r$  can be formally expressed as,

$$G_t = q_b R_b + F_s q_d + F_g P_g q \quad [8.2]$$

In Eq(8.2) "G<sub>T</sub>" is the radiation on the tilted surface, "P<sub>G</sub>" is the ground reflectivity and "R<sub>b</sub>" is the instantaneous tilt factor for direct radiation, defined as the ratio of direct radiation on the surface under consideration and the direct radiation on a horizontal surface. Thus

$$R_b = \frac{G_{bT}}{G_b}$$

\* F<sub>S</sub> and F<sub>G</sub> are the appropriate factors for sky diffuse and ground reflected components of radiation, assuming an isotropic distribution of diffuse radiation, the factor F<sub>S</sub> can be obtained as,

$$F_S = \frac{(1 + \cos \beta)}{2} \rightarrow (8.4)$$

\* The view factor F<sub>G</sub> for the ground reflected component of the radiation is given by,

$$F_G = \frac{(1 - \cos \beta)}{2} \rightarrow (8.5)$$

\* EXPRESSION FOR R<sub>b</sub>:

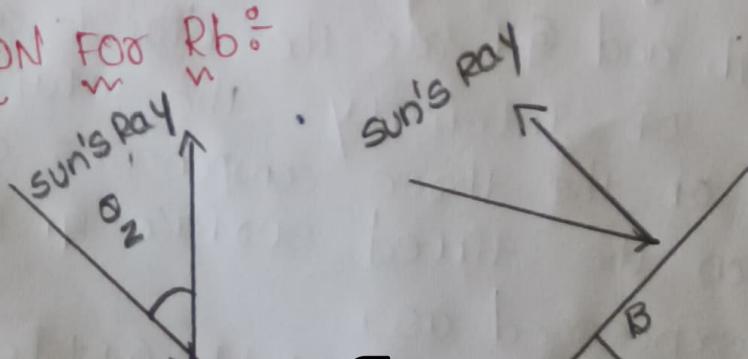


Fig 8.1-

Direct radiation on horizontal & tilted surface

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- \* consider a horizontal and a tilted surface ⑪  
as shown in fig 8.1.
- \* Let  $G_{b,D}$  be the Intensity of direct radiation "z"  
with the zenith.  $\theta$  on a plane perpendicular to the  
Ray.
- \* The Ray as shown in Fig 8.1 (a) makes an angle  
with the outer normal to the tilted surface as  
"θ" similarly, the Ray makes an angle with the  
horizontal from the geometry as shown B.  
shown in fig 8.1 (b). The tilted surface makes  
an angle  $\phi$  in fig 8.1, it follows.

$$G_b = G_{b,D}$$

Thus  $R_b$  can be expressed as,

$$R_b = \frac{G_{b,D}}{G_b} = \frac{\cos \theta}{\cos \theta_2} \rightarrow (8.6)$$

The expressions for  $\cos \theta$  are given by eqs  
(3.6) (or) (3.7)  $\cos \theta$  for south facing surface  
is given by eq (3.12)  $\cos \theta_2$  is expressed  
by eq (3.11). Using eqs (3.12) and (3.11) in  
eq (8.6),  $R_b$  for a south facing surface is  
given by

$$R_b = \frac{\cos(\phi - B) \cos \theta \cos \omega + \sin(\phi - B) \sin \theta}{\cos \theta \cos \phi \cos \omega + \sin \phi \sin \theta}$$

expression given by eq (8.7), strictly speaking, is valid for an instant by the value of the hour angle  $\omega$  employed. However, eq (8.7) can be used with no significant loss of accuracy for a time period of one hour ( $\omega_1$  to  $\omega_2$ ) using  $\omega = \frac{(\omega_1 + \omega_2)}{2}$ .

### \* INSTRUMENTS FOR SOLAR RADIATION

- \* There are two types of infrared detectors for solar radiation measurement.
- \* Thermal detectors and photon detectors.
- \* Thermal detectors and photon detectors.
- \* Heating effect of incident radiation causing a change in some physical property of the detector is the principle underlying the detectors.
- \* Thermal
- \* The time constant of the detectors should be small for responding to quick changes in the incident radiation.
- \* Time constant of several seconds can be accepted in solar radiation measurements.
- \* Photon detectors converts some of the incident radiation directly into electricity, which is proportional to the incident radiation.
- \* The detecting capability in general, of photon

detectors is one (or) two orders of magnitude 12 greater.

\* However, the penalty associated with photon detectors is that their spectral response is non-uniform.

\* Alternatively, instruments to measure solar radiation broadly fall into three categories.

- They are 1) measure global radiation  
2) diffuse radiation  
3) Direct radiation

\* It is easy to envisage that it is best to measure normal incident direct radiation and in order to do this,

\* The instrument sensor needs to face the sun always.

\* This calls for a tracking instrument with consequent inconvenience and additional equipment to follow the sun.

\* The instruments used to measure global radiation are referred to as pyranometers.

\* The instruments that measure direct radiation are called pyrheliometers.

## \* Pyranometers:

- \* Eppley pyranometer designed by Kimball and Hobbs of US Weather Bureau has become the most widely used working pyranometer's.
- \* The Eppley pyranometer's detector (or) the working surface consists of two concentric silvered rings.
- \* The inner ring is coated with Parsons optical black lacquer and the outer one is coated with white magnesium oxide. The temperature difference between the two is an indication of the incident solar radiation and is measured by a thermopile.
- \* The sensor is placed inside a hemispherically sealed spherical lamp bulb filled with dry air.
- \* The detector is best when used to measure horizontal radiation.
- \* A reduction of 5% output is caused when the instrument is mounted vertically due to convective currents in the glass enclosure.
- \* The new Eppley pyranometer equipped with a thermister compensated electrical circuit reduced the temperature dependence and improved the cosine response.

## \* SUNSHINE RECORDER :-

The sunshine recorders essentially static do not require power supply and alleviate the difficulties faced by the recorders where uninterrupted power supply is not available.

\* They are essentially spherical lenses, which blacken a sensitive strip.

\* The length of the charred portion of the strip is related to the solar radiation with the help of measurements made by alternate.

Instruments.

\* The empirical constants derived in the relations are expected to be valid for the type of climate for which they are derived.

\* They are simple to use.

\* Indeed, significant numbers of studies have been made in deriving reliable constants and examining their applicability for different climates.

\* Campbell-Bell-Stokes Recorder focuses solar-radiation to burn a trace in a chart.

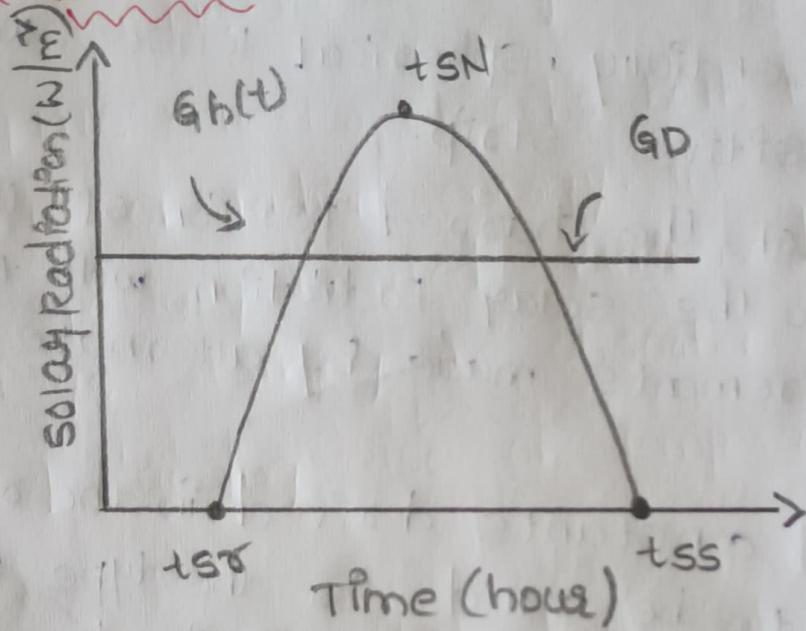
\* Jordan Recorder focuses sunlight on to photographic paper.

\* Malvin Recorder makes use of a thermoelectric switch to actuate a chronograph to trace the.

## Sunshine hours.

- \* The daily global solar radiation on a horizontal plane is related to the number of hours of bright sunshine.

## \* SOLAR RADIATION DATA :-



- \* The solar radiation data is widely used as the source of renewable energy.
- \* It is highly required to formulate the estimation model and forecasting the meteorological data.
- \* However, solar data play a key role in various domains.
- \* The application of AI includes modeling hourly, monthly, and daily solar radiation data.

## \* SOLAR ENERGY COLLECTION:-



\* Solar collectors are used to collect the solar energy and convert the incident radiations into thermal energy by absorbing them.

\* This heat is extracted by flowing fluid in the tube of the collector for further utilization in different applications.

## \* TYPES OF SOLAR COLLECTORS:-

\* 1) Non-concentrating collectors

2) Concentrating (focusing) collectors.

### 1) Non-concentrating collectors:-

\* In these type, the collector area is the same as the absorber area.

\* It is generally used for low & medium Temperature Requirements.

### 2) Concentrating (focusing) collectors:-

\* It have a larger Interceptor than absorber.

\* It is generally used for high Temperature Requirements.

### \* FLAT PLATE COLLECTORS:-

\* Flat plate collector is most important part of any solar Thermal energy system.

\* It is simplest in design and both direct and

~~Diffuse components of Radiant energy.~~



## The absorber plate

diffuse Radiations are absorbed by collector and converted into useful heat. These collectors are suitable for heating to temperature below 100°C.

- \* Flat plate collector absorbs both beam and diffuse components of radiant energy. The absorber plate is a specially treated blackened metal surface.
- \* Sun rays striking the absorber plate are absorbed causing rise of temperature of transport fluid.
- \* Thermal insulation behind the absorber plate and transparent cover sheets prevent loss of heat to surroundings.
- \* The constructional details of flat plate collector is given below.

### i) Insulated Box:-

The Rectangular box is made of thin Galvanised sheet and is insulated from sides and bottom using glass (or) mineral wool of thickness 5 to 8 cm to reduce losses from conduction to back and side wall.

- \* The box is tilted at due south and a tilt

angle depends on the latitude of location. (15)  
The face area of the collector box is kept between 1 to 2 m<sup>2</sup>.

### a) Transparent cover:

This allows solar energy to pass through and reduces the convective heat losses from the absorber plate through air space.

\* The transparent tempered glass cover is placed on top of rectangular box to trap the solar energy and sealed by rubber gaskets to prevent the leakage of hot air.

\* It is made of plastic/glass but glass is most favorable because of its transmittance and low degradation.

\* The plastics are available at low cost, light in weight and can be used to make tubes, plates and covers, but are suitable for low temperature application 70-120°C with single cover plate (or) upto 150°C using double cover plate.

\* The thickness of glass cover 3 to 4 mm is commonly used and 1 to 2 covers with spacing 1.5 to 3 cm are generally used b/w plates.

\* The temperature of glass cover is lower than the absorber plate, and is a good absorber

