

Classical Optics - Light matter interactions.

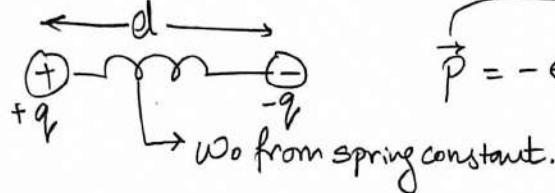
Reading: Siegman Lasers
Ch: 2

Horentz: Model atomic response as harmonic oscillator.

Classical model works well with fitting parameters which come from quantum theory: resonance freq ω_0 , oscillator strength f , and damping γ .

> Transitions b/w atomic or molecular energy levels corresponds to an oscillation in space of charge.

Horentz model:



$\vec{p} = -e \vec{r}$ → dipole moment

→ "Strength of a dipole"

→ How much torque it would experience in an electric field

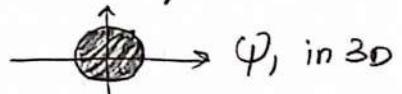
→ How much force it would experience in a gradient field.

Qualitative QM approach to atomic dipoles.

> $E_1 \rightarrow$ ground state with symmetric wavefn. Ψ_1 like S orbital.

> $E_2 \rightarrow$ excited state with antisymmetric wavefn. Ψ_2 like p orbital.

q_n 1D: $\Psi_1(-x) = \Psi_1(x)$; $\Psi_2(-x) = -\Psi_2(x)$.



q_n 3D: $\Psi_1(-\vec{r}) = \Psi_1(\vec{r})$; $\Psi_2(-\vec{r}) = -\Psi_2(\vec{r})$



> Probability density fn. is given by $\Psi^* \Psi$

> Electronic charge density $\rho(\vec{r}) = -e \Psi^*(\vec{r}) \Psi(\vec{r})$ by definition.
(since $|\Psi|^2$ is probability per unit volume)

⇒ Average position of the electron is $\langle \vec{r} \rangle = \int \Psi^*(\vec{r}) \Psi(\vec{r}) d^3 r$

⇒ Dipole moment of the system $\vec{P} = -e \langle \vec{r} \rangle$

$$\Rightarrow \vec{P} = \int \vec{r} (-e \psi^* \psi) d^3 \vec{r}$$

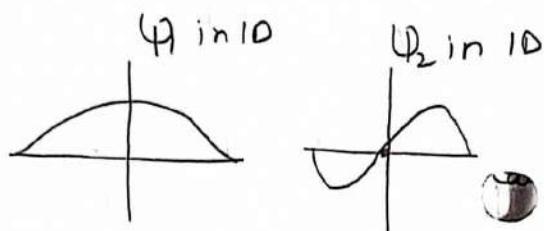
$$\boxed{\vec{P} = \int \vec{r} \rho(\vec{r}) d^3 \vec{r}}$$

$$\text{where } \rho(\vec{r}) = -e \psi^* \psi$$

- > If system is in Ψ_1 or Ψ_2 , $\rho(\vec{r})$ is symmetric and \vec{P} is antisymmetric so $\vec{P} = 0$ which just says the centers of the two +ve charges cancel.
- > If the atom is in a superposition state, assume the amplitudes are equal

$$\Rightarrow \Psi_{\text{tot}} = \Psi_1 + \Psi_2$$

Recall, $\Psi_1(\vec{r}, t) = \Psi_1(\vec{r}) e^{-iE_1 t/\hbar}$
 $\Psi_2(\vec{r}, t) = \Psi_2(\vec{r}) e^{-iE_2 t/\hbar}$



$|\Psi_1|^2$ & $|\Psi_2|^2$ are symmetric
 but $\Psi_1 \Psi_2^*$ is antisymmetric
 $\Rightarrow \vec{r} \Psi_1 \Psi_2^*$ is symmetric.

$$\Rightarrow \vec{P} = e \int_0^{\vec{r}} |\Psi_1(\vec{r})|^2 d^3 \vec{r} + e \int_0^{\vec{r}} |\Psi_2(\vec{r})|^2 d^3 \vec{r} + e \int \vec{r} \Psi_2^* \Psi_1 e^{-i(E_1-E_2)t/\hbar} d^3 \vec{r}$$

+ c.c. of ↑

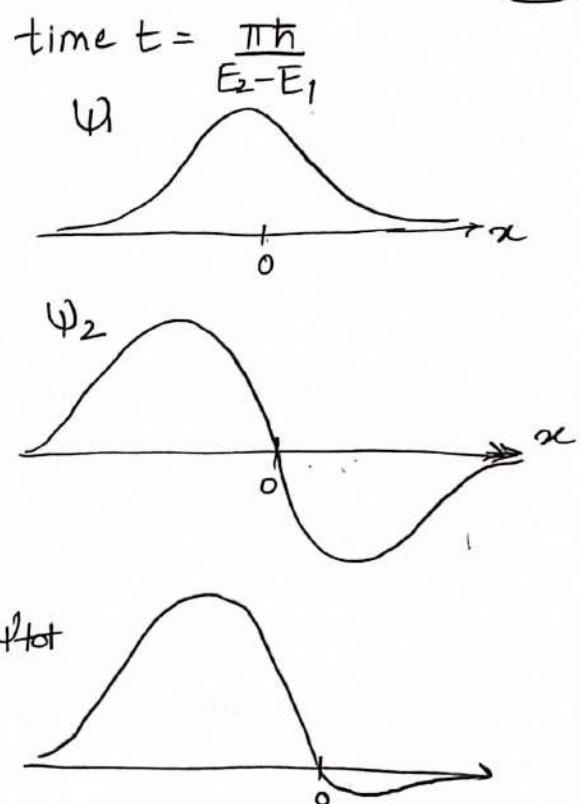
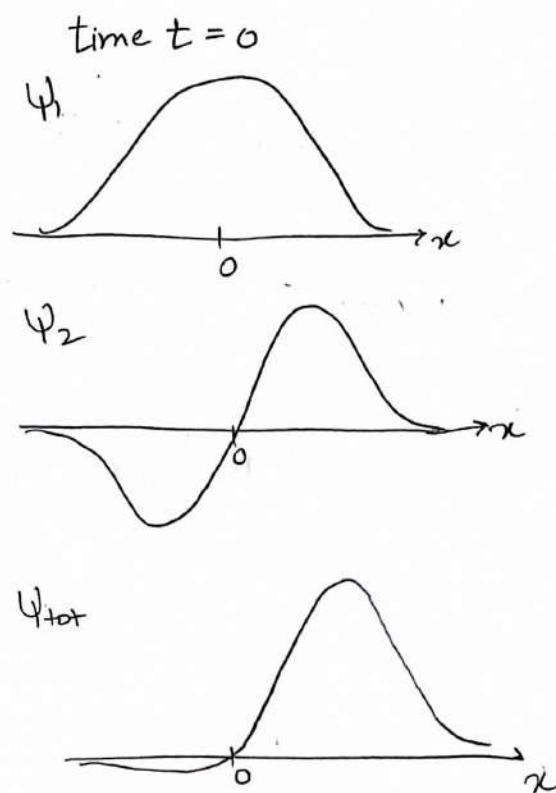
$\Rightarrow \vec{P}$ oscillates with a frequency given by $E_2 - E_1$! \rightarrow Hertzian dipole

> Assume Energy origin to be the ground state $\Rightarrow E_1 = 0$

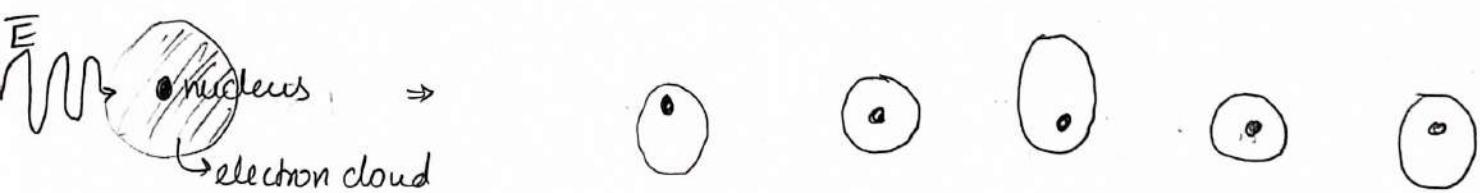
$$\Rightarrow \Psi_1(\vec{r}, t) = \Psi_1(\vec{r}) = \text{constant in time.}$$

> Let us look at a 1D slice for pictorial insight.

(2-3)



- Recall charge density $\rho(\vec{r}) = -e |\Psi_{\text{tot}}(x, t)|^2 \rightarrow$ oscillates periodically.
- How do we put the atom in a superposition state? By illuminating it with light..



This displacement of the electron from S-orbital can mathematically be written as "mixing" in some amount of p-orbital. More generally we can perform a mode expansion of the displaced electron cloud

$$\begin{aligned} \text{as, } \Psi_{\text{dis}} &= \sum_{n,l,m} a_{nlm} \Psi_{nlm} \\ &= a_{1s} \Psi_{1s} + a_{2s} \Psi_{2s} + a_{2p} \Psi_{2p} + \dots \end{aligned}$$

$$\text{where } a_{nlm} = \int \Psi_{nlm}^* \Psi_{\text{dis}} d^3\vec{r}$$

} where n, l, m are the quantum numbers of the valence electron.

Eg:- For hydrogen atom,

- > Ψ_{1s} is the ground state,
- > Ψ_{2p}, Ψ_{2s} are the first excited states.
- > By symmetry a_{1s} will be zero, ^{why?} and the largest coefficients will be $a_{1s} = \int \Psi_{1s}^* \Psi_{\text{disp}} d^3\vec{r}$ and $a_{2p} = \int \Psi_{2p}^* \Psi_{\text{disp}} d^3\vec{r}$
- > If $\hbar\omega = E_2 - E_1$, for the optical field, these will be the only 2 states.

Review of Electric Dipole Radiation.

Given sources $\bar{J}(\vec{r}, t)$ and $P(\vec{r}, t)$, \vec{A} satisfies-

$$\nabla^2 \vec{A} - \mu_0 \epsilon_0 \frac{\partial^2 \vec{A}}{\partial t^2} = -\mu_0 \bar{J}$$

General soln:

$$\vec{A}(\vec{r}, t) = \frac{\mu_0}{4\pi} \int_{\vec{r}'} d^3\vec{r}' \int dt' \frac{\bar{J}(\vec{r}', t')}{|\vec{r} - \vec{r}'|} \delta(t' + \frac{|\vec{r} - \vec{r}'|}{c} - t)$$

When $t' = t - \frac{|\vec{r} - \vec{r}'|}{c}$ \Rightarrow when $t = \frac{|\vec{r} - \vec{r}'|}{c}$ we have $t' = 0$

\Rightarrow t' clock starts when then & follows what happened at $t=0$ from then on.

\Rightarrow Basically you remove $\frac{|\vec{r} - \vec{r}'|}{c}$ from t \times that time is t' when the δ function fires & extracts the value of current accordingly. It extracts a "retarded" value of the current density. Each pt. in space of the current distribution provides a contribution at \vec{r} which are added up!

If the current is harmonic,

$$\bar{J}(r, t) = \bar{J}(\vec{r}) e^{i\omega t}$$

$$\bar{A}(\vec{r}) = \frac{\mu_0}{4\pi} \int_{V'} \bar{J}(\vec{r}') \frac{e^{-ik|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} d^3\vec{r}'$$

In optics $|\vec{r}'| \ll |\vec{r}| \Rightarrow |\vec{r}-\vec{r}'| = r - \underbrace{\vec{r}' \cdot \hat{r}}_{\approx 0} = r$

$$\Rightarrow \bar{A}(\vec{r}) = \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \int_{V'} \bar{J}(\vec{r}') d^3\vec{r}'$$

Note $\bar{J} = 0$ on the surface since V' bounds the charges completely.

$$\int_{V'} \bar{J}(\vec{r}') d^3\vec{r}' = - \int_{V'} \vec{r}' (\nabla \cdot \bar{J}) d^3\vec{r}' \rightarrow \text{integration by parts in 3D.}$$

Digression

This expression & the necessary condition $\bar{J} = 0$ are not a trivial formulation. Proof:

Consider,

$$\nabla \cdot (x \bar{J}) = \nabla_x \cdot \bar{J} + x \nabla \cdot \bar{J} = J_x + x \nabla \cdot \bar{J}$$

$$\text{Similarly } \nabla \cdot (y \bar{J}) = J_y + y \nabla \cdot \bar{J}; \quad \nabla \cdot (z \bar{J}) = J_z + z \nabla \cdot \bar{J}$$

$$\Rightarrow \sum_i \hat{x}_i (\nabla \cdot (x_i \bar{J})) = \bar{J} + \vec{r} (\nabla \cdot \bar{J})$$

$$\Rightarrow \int_{R^3} \sum_i \hat{x}_i (\nabla \cdot (x_i \bar{J})) d^3R = \int_{R^3} \bar{J} d^3R + \int_{R^3} \vec{r} (\nabla \cdot \bar{J}) d^3R$$

$$\Rightarrow \sum_i \hat{x}_i \oint (\hat{n} \cdot (x_i \bar{J})) d^3R = \text{RHS.} \quad \text{But } \hat{n} \cdot \bar{J} = 0 \text{ on the surface} \\ \Rightarrow \oint \hat{x}_i \phi d^3R = \text{RHS.} \quad \Rightarrow \phi_{-n} \Rightarrow \text{DHC} - n \quad \blacksquare$$

This is analogous to the integration by parts in 1D.

$$\int_a^b d(uv) = \int_a^b u dv + \int_a^b v du$$

$$\Rightarrow [uv]_a^b = \int_a^b u dv + \int_a^b v du$$

Therefore when uv is 0 at the "boundaries" $uv|_a^b = 0$

$$\Rightarrow \int_a^b u dv = - \int_a^b v du$$

- > In 3D instead of $d(uv) = u dv + v du$ we used $\nabla \cdot (\vec{x} \vec{J}) = \nabla_x \cdot \vec{J} + x \nabla \cdot \vec{J}$
- & used the div. thm to form a surface integral that vanished.
- > However it is not always a trivial matter to get the right form of the 3D diff eqn.

Coming back.

$$\int_{V'} \vec{J}(\vec{r}') d^3 \vec{r}' = - \int_{V'} \vec{r}' (\nabla \cdot \vec{J}) d^3 \vec{r}' = i\omega \int_{V'} \vec{r}' \rho(\vec{r}') d^3 \vec{r}'$$

But we know that $\vec{P} = \int_{V'} \vec{r}' \rho(\vec{r}') d^3 \vec{r}'$ from earlier derivation.

$$\Rightarrow \boxed{\vec{A} = i\omega \frac{\mu_0}{4\pi} \frac{e^{-ikr}}{r} \vec{P}}$$

where \vec{P} is the dipole moment of the atomic system with the charges in their "cloud" distribution

Assume $\vec{P} = P_0 \hat{z} \cos \omega t$ we can find the fields.

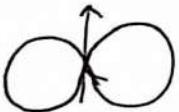
(2.7)

$$\vec{E} = (E_r, E_\theta, E_\phi) = \left(0, -\frac{\omega^2 P_0}{4\pi\epsilon_0 c^2} \sin\theta \frac{\cos\omega(t - \frac{r}{c})}{r}, 0 \right)$$

$$\vec{H} = (H_r, H_\theta, H_\phi) = \left(0, 0, -\frac{\omega^2 P_0}{4\pi c} \sin\theta \frac{\cos\omega(t - \frac{r}{c})}{r} \right)$$

$$\vec{S} = (S_r, S_\theta, S_\phi) = \left(\frac{\omega^4 P_0^2}{16\pi^2 \epsilon_0 c^3} \sin^2\theta \frac{\cos^2\omega(t - \frac{r}{c})}{r^2}, 0, 0 \right)$$

Note: Radiated power \propto square of dipole moment magnitude
 " " " $\propto \omega^4$



Power pattern is that of a Hertzian dipole

Average power radiated

$$\langle w \rangle = \frac{P_0^2 \omega^4}{12\pi\epsilon_0 c^3}$$

Radiative decay

Energy of an oscillator,

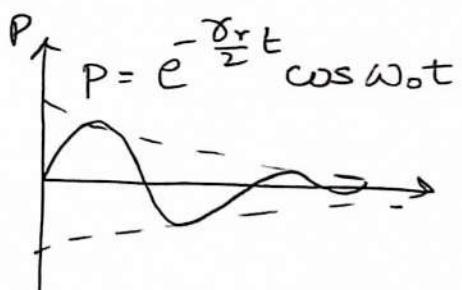
$$U(t) = \frac{1}{2} m \omega_0^2 z(t)^2 + \frac{1}{2} m \left(\frac{dz}{dt} \right)^2$$

$$\text{Where } z(t) = -\frac{1}{e} P = -\frac{1}{e} e^{-\frac{\gamma_r t}{2}} \cos\omega_0 t$$

$$\text{We can show } U(t) = -\frac{1}{\gamma_r} \frac{dU(t)}{dt}$$

Equating $\frac{dU(t)}{dt}$ to $\langle w \rangle$ we get

$$\gamma_r = \frac{1}{T_r} = \frac{1}{4\pi\epsilon_0} \left(\frac{2\pi^2 e^2}{m} \right) \cdot \frac{1}{3} \cdot \left(\frac{\omega_0^2}{\pi^2 c^3} \right)$$



replaced in QM by atomic dipole matrix element

spatial avg.

DOS in 3D of free space.

> τ_r is the classic radiative lifetime.

& energy decays exponentially as $U(t) = U(0) e^{-t/\tau_r}$.

In general $\tau_r = \frac{1}{4\pi\epsilon_0} P_o(\omega)$

\hookrightarrow DOS seen by the dipole (including an angular average)

- => Decay rate can be engineered by modifying the density of EM modes with which the dipole interacts.
- > The modification of radiative decay via control of the DOS is known as the Purcell effect in Quantum theory of radiation.
- > Modern photonics research focuses a lot on controlling the τ_r through DOS using waveguides, crystals, metamaterials etc.
- > $\tau_r \approx 1\text{ns}$ but for lasers we can get $\approx \mu\text{s}/\text{ms}$ since they can operate in "forbidden regions".

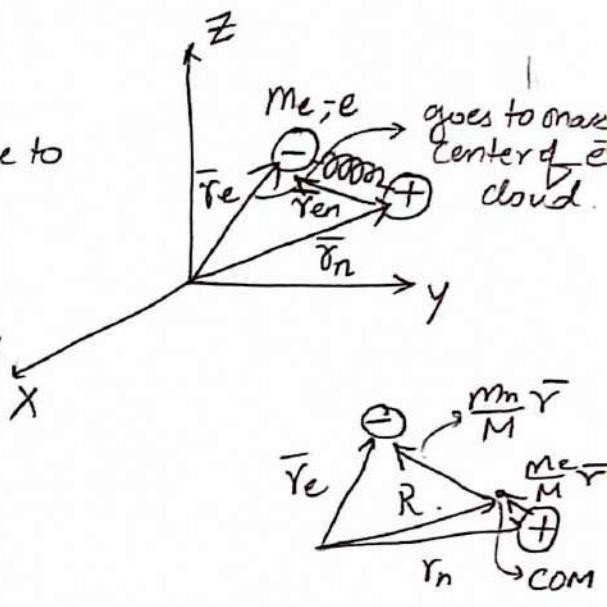
Driven Lorentz Oscillator

2.9

$$\vec{r}_{en} = \vec{r}_e - \vec{r}_n$$

$\vec{F}_{en} = -\vec{F}_{ne}$ is the force on the electron due to the ion/nucleus.

This is not a coulomb force but rather a harmonic restoring force when slightly displaced. This can be justified for small displacements in QM.



$$\vec{F} = q(\vec{E} + \vec{v} \times \vec{B}) \approx q\vec{E}$$

$$\Rightarrow m_e \frac{d^2 \vec{r}_e}{dt^2} = -e \vec{E}(\vec{r}_e, t) + \vec{F}_{en}(\vec{r}_{en}) \quad \begin{matrix} \text{Newton's Law} \\ F=ma. \end{matrix}$$

$$m_n \frac{d^2 \vec{r}_n}{dt^2} = e \vec{E}(\vec{r}_n, t) + \vec{F}_{ne}(\vec{r}_{en})$$

Introducing a center of mass coordinate $\vec{R} = \frac{m_e \vec{r}_e + m_n \vec{r}_n}{m_e + m_n}$

we have $\vec{r}_e = \vec{R} + \frac{m_n}{M} \vec{r}$ & $\vec{r}_n = \vec{R} - \frac{m_e}{M} \vec{r}$ where $\vec{r} = \vec{r}_{en}$ & $M = m_e + m_n$

Subs. into Newton's Eqs.

$$m_e \frac{d^2 \vec{R}}{dt^2} + m \frac{d^2 \vec{r}}{dt^2} = -e \vec{E}(\vec{R} + \frac{m_n}{M} \vec{r}, t) + \vec{F}_{en}(\vec{r})$$

$$m_n \frac{d^2 \vec{R}}{dt^2} - m \frac{d^2 \vec{r}}{dt^2} = e \vec{E}(\vec{R} - \frac{m_e}{M} \vec{r}, t) - \vec{F}_{ne}(\vec{r})$$

where, $m = \frac{m_e m_n}{M}$ (close to m_e). We want to decouple \vec{r} & \vec{R}

Center of mass motion

Adding the 2 eqns \Rightarrow

$$M \frac{d^2 \bar{R}}{dt^2} = e \left[\bar{E}(\bar{R} - \frac{m_e}{M} \bar{r}, t) - \bar{E}(\bar{R} + \frac{m_p}{M} \bar{r}, t) \right]$$

Taylor expansion, (It is valid since $\frac{\bar{r} \cdot \nabla E}{E}$ is small due to dipole approx)

$$E_j(x + \Delta x) \approx E_j(x) + \frac{\partial E_j}{\partial x} \Delta x.$$

$$\text{In 3D, } \frac{\partial E_j}{\partial x} = \hat{x} \cdot \nabla E_j$$

$$\Rightarrow E_j(\bar{R} + \Delta \bar{R}) \approx E_j(\bar{R}) + \left(\sum_{j=1}^3 \Delta R_j \hat{x}_j \right) \cdot \nabla E_j \quad j=x, y, z$$

In vector form

$$\bar{E}(\bar{R} + \Delta \bar{R}) \approx \bar{E}(\bar{R}) + \Delta \bar{R} \cdot \nabla \bar{E}$$

$$\Rightarrow \bar{E}(R \pm \frac{m_l}{M} \bar{r}, t) \approx \bar{E}(\bar{R}, t) \pm \frac{m_l}{M} \bar{r} \cdot \nabla \bar{E}(\bar{R}, t) ; \text{ where } l=n, e$$

$$\Rightarrow M \frac{d^2 \bar{R}}{dt^2} = e \left[-\frac{m_e}{M} \bar{r} \cdot \nabla \bar{E}(\bar{R}, t) - \frac{m_p}{M} \bar{r} \cdot \nabla \bar{E}(\bar{R}, t) \right]$$

$$\Rightarrow \boxed{M \frac{d^2 \bar{R}}{dt^2} = -e \bar{r} \cdot \nabla \bar{E}(\bar{R}, t)}$$

The COM is accelerated by a gradient in the incident field.

> The dipole moment of a Lorentz oscillator depends on \bar{E} by a factor α called polarizability.

$$\Rightarrow \vec{P} = -e \bar{r} = \alpha \bar{E} \rightarrow \text{linear assumption.}$$

$$\Rightarrow M \frac{d^2 \vec{R}}{dt^2} = \alpha \vec{E} \cdot \nabla \vec{E}$$

Note, $\nabla(\vec{E} \cdot \vec{E}) = \vec{E} \cdot \nabla \vec{E} + (\nabla \vec{E}) \cdot \vec{E} = 2 \vec{E} \cdot \nabla \vec{E}$

$$\Rightarrow \boxed{M \frac{d^2 \vec{R}}{dt^2} = \frac{\alpha}{2} \nabla(\vec{E} \cdot \vec{E})}$$

\vec{E} is a real field & \therefore
we have \vec{E}, \vec{E} & not $\vec{E} \cdot \vec{E} = |\vec{E}|^2$

\therefore the force oscillates at $\cos^2 \omega t$ but this is too high so let's find time average over one cycle.

$$\vec{E} = \hat{E} E_0(\vec{R}) \cos \omega t$$

$$\Rightarrow \langle \vec{F} \rangle = \frac{\alpha}{4} \nabla E_0^2(\vec{R})$$

$\downarrow E_0^2$
 Force along gradient of Intensity
 \Rightarrow pushes atom/molecule towards higher Intensity.

$\rightarrow \alpha$ for a small dielectric sphere is $\alpha = n_m^2 \left(\frac{n_r^2 - 1}{n_r^2 + 2} \right)$ where $n_r = n/n_m$ where n is ior of particle & n_m is i.o.r of medium.

Lorentz oscillator - internal coordinate

Subtracting (ii) from (i),

$$2m \frac{d^2 \vec{r}}{dt^2} = -e \left[2\vec{E}(\vec{R}, t) + \underbrace{\left(\frac{m_e - m_n}{M} \right) \vec{r} \cdot \nabla \vec{E}(\vec{R}, t)}_{=0} \right] + 2\vec{F}_{en}(\vec{r}) + \underbrace{(m_e - m_n) \frac{d^2 \vec{R}}{dt^2}}_{=0}$$

dipole approx

$$\Rightarrow \frac{d^2 \vec{r}}{dt^2} = -\frac{e}{m} \vec{E}(\vec{R}, t) + \frac{1}{m} \vec{F}_{en}(\vec{r})$$

Lorentz argued that $\vec{F}_{en}(\vec{r}) = -\mu_{ab} \vec{r}$ \rightarrow harmonic restoring force.

$$\Rightarrow \frac{d^2\bar{r}}{dt^2} + \omega_0^2 \bar{r} = -\frac{e}{m} \bar{E}(R, t) \rightarrow \text{undamped harmonic oscillator}$$

> Even a single atom decays due to radiation.

$$\Rightarrow \boxed{\frac{d^2\bar{r}}{dt^2} + \gamma_r \frac{dr}{dt} + \omega_0^2 \bar{r} = -\frac{e}{m} \bar{E}(R, t)}$$

Let $\bar{r}(t=0) = \bar{r}_0$ is the initial amplitude.

Trial solution $\bar{r}(t) = \bar{r}_0 e^{(-\Gamma+i\omega)t}$. Solve for Γ & ω .

$$\Rightarrow (-\Gamma + i\omega)^2 + \gamma_r (-\Gamma + i\omega) + \omega_0^2 = 0$$

$$\Rightarrow (\Gamma^2 - \omega^2 + \gamma_r \Gamma + \omega_0^2) + i(-2\omega\Gamma + \omega\gamma_r) = 0$$

Imaginary part $\Rightarrow \omega(-2\Gamma + \gamma_r) = 0$

$$\Rightarrow \boxed{\Gamma = +\frac{\gamma_r}{2}} \rightarrow \text{dipole decay rate.}$$

Real part $\Rightarrow \boxed{\omega = \sqrt{\omega_0^2 - \left(\frac{\gamma_r}{2}\right)^2}}$ \rightarrow shifted or "renormalized" (red shifted) from ω_0 , (almost negligible).

$$\Rightarrow \bar{r}(t) = \bar{r}_0 e^{i\omega t - \frac{\gamma_r t}{2}} \text{ where } \omega \approx \omega_0$$

$$U(t) = \frac{1}{2} m \omega_0^2 |\bar{r}(t)|^2 + \frac{1}{2} m |\frac{d\bar{r}}{dt}|^2$$

$$\boxed{U(t) = U(t_0) e^{-\gamma_r t}} \text{ Energy decays at } \gamma_r. \rightarrow \text{twice as fast as polarization.}$$

Polarizability of a single isolated oscillator

Since $\bar{P} = -\epsilon \bar{r}$

We can rewrite the Lorentz equation as,

$$\frac{d^2 \bar{P}}{dt^2} + \gamma_r \frac{d \bar{P}}{dt} + \omega_0^2 \bar{P} = \frac{e^2}{m} \bar{E}(\bar{R}, t).$$

Let $\bar{R} = 0$ and \bar{E} be the incident polarization, $\bar{E}(t) = E_0 e^{i\omega t} \hat{e}$

> we expect a solution of the form $\bar{P}(t) = \hat{e} P_0 e^{i\omega t}$
 where \bar{P}_0 is the complex dipole amplitude allowing for phase shifts
 wrt \bar{E} .

$$\Rightarrow -\omega^2 P_0 + i\omega \gamma_r P_0 + \omega_0^2 P_0 = \frac{e^2}{m} E_0$$

$$\Rightarrow P_0 = \left[\frac{e^2}{m} \cdot \frac{1}{\omega_0^2 - \omega^2 + i\omega \gamma_r} \right] E_0 = \alpha E_0$$

where

$$\boxed{\alpha = \frac{e^2}{m} \cdot \frac{1}{\omega_0^2 - \omega^2 + i\omega \gamma_r}}$$

Polarizability!
 (note linear response)

> At resonance, the dipole moment lags the driving field by 90° .

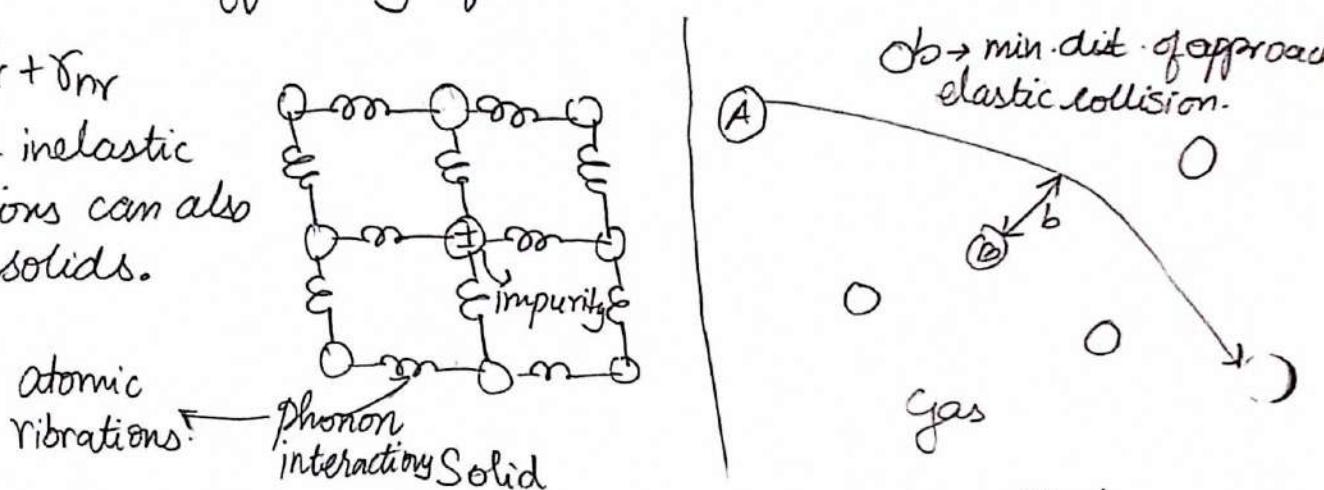
> Far from resonance we can ignore $i\omega \gamma_r \Rightarrow$ no damping & the \bar{P} is
 in phase with

Dipole-environment interactions and dephasing

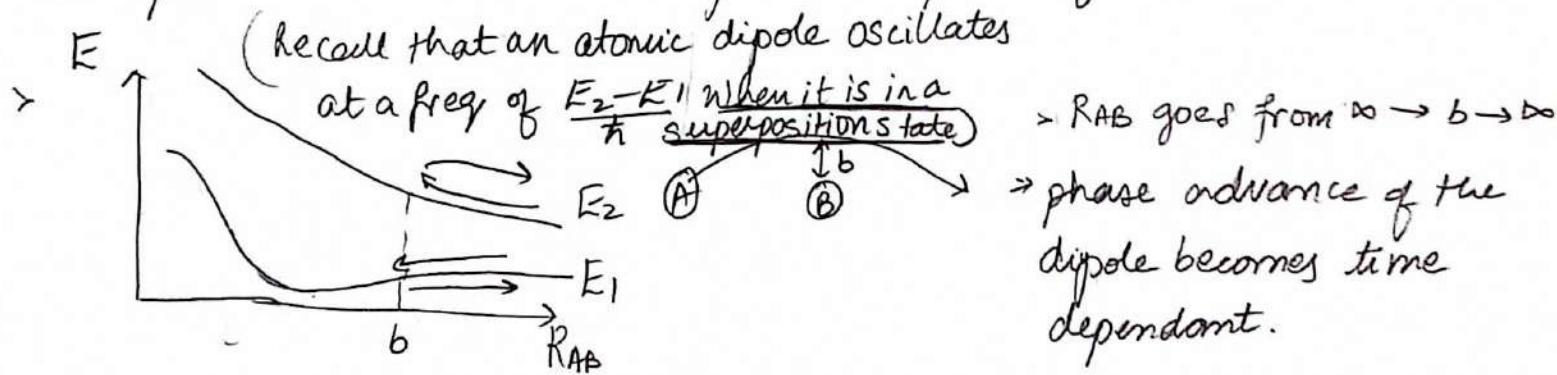
- > Environmental interactions dominate the radiative response of the atoms.
- > Collisions happen very frequently. Elastic collisions merely change the trajectory but do not take energy away from the dipole oscillation. Inelastic collisions provide a nonradiative channel for energy decay of the oscillator, in which case

$$\gamma = \gamma_r + \gamma_{nr}$$

- > Elastic & inelastic interactions can also occur in solids.



- > The phonons modulate the amplitude & phase of the dipole.



$$\Rightarrow e^{-i(E_2 - E_1)t/\hbar} \rightarrow e^{-i\Delta E(t)t/\hbar} = e^{-i\phi(t)}$$

where $\Delta E(t) = E_2(t) - E_1(t)$

This is a phenomenological argument to extend the Lorentz model.

Therefore the total phase change from the collision $\phi_{\text{collision}} = \int_{t_{\text{coll}}}^{\infty} \frac{\Delta E(t)}{\hbar} dt$ $T_{\text{coll}} \approx \frac{\text{atoms}}{V_m} \approx 100 \text{ fs.}$

(2.15)

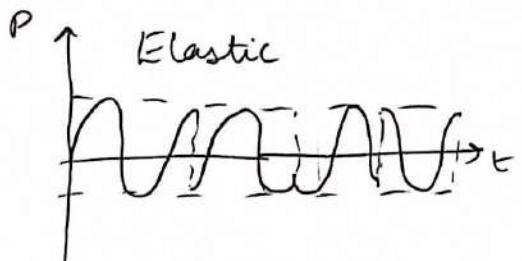
- > The argument is that in a medium each atom undergoes these phase shifts completely at random and the ϕ_{coll} is thus a random variable equally distributed from $[0, 2\pi]$.
- > The effect of the collision is thus to randomize the phase of dipole oscillations. The collision duration is much longer than the time period of oscillation. So the interaction is more of a random (AM, PM) modulation.
- > In the Lorentz model we can add this effect as a random force apart from the harmonic restoring force.

$$\Rightarrow \boxed{\frac{d^2 \bar{p}}{dt^2} + \gamma \frac{dp}{dt} + \omega_0^2 \bar{p} = \frac{e^2}{m} \bar{E} + \bar{F}_r(t)}$$

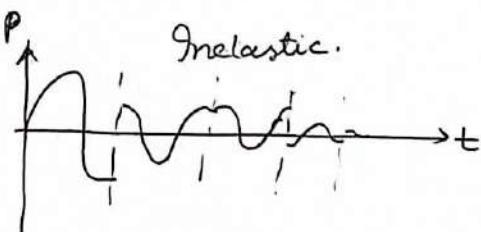
→ decay rate due to γ_r , γ_{nr} & collision interactions.

- > Only the statistical properties of $\bar{F}_r(t)$ are known.

-
- The equation describes the \bar{p} for a single atom/oscillator that interacts with a surrounding "bath".
described by $\bar{F}_r(t)$
 - A wrong picture where interactions are instant looks as follows.



no decay of $\langle\langle p^2(t) \rangle\rangle$



decay of $\langle\langle p(t) \rangle\rangle$

Macroscopic Polarization.

- Maxwell's Equations for macroscopic fields are the spatial averages over atomic scale fluctuations.
- Polarization enters through $\vec{D} = \epsilon_0 \vec{E} + \vec{P}$

where $\vec{P} = \frac{1}{V} \sum_{j=1}^N \vec{p}_j$

- total no. of "oscillators"
- Lorentz dipole moment of each atom.
- volume

Let $N = \frac{N}{V}$ where N is the number density \Rightarrow no. of oscillators per unit vol.

$$\vec{P} = N \left[\frac{1}{N} \sum_{j=1}^N \vec{p}_j \right]$$

$\Rightarrow \boxed{\vec{P} = N \cdot \langle\langle \vec{p} \rangle\rangle}$ → ensemble average dipole moment of all the oscillators.

(in)

In the language of stat mech, we can consider each atom to be different "realisations" of identical oscillators to justify the ensemble average interpretation.

→ We can use the eqn. $\vec{P}(t) = \vec{P}_0 e^{i\omega t}$ and extend it when $F_r(t)$ is present as $\vec{P}(t) = \vec{P}_0(t) e^{i\omega t}$ since $F_r(t)$ adds a "modulation".

→ Real experiments have long averaging times so we really only observe over one averaging time period,

$$\langle \vec{P}_0(t) \rangle = \frac{1}{T} \int_t^{t+T} \vec{P}_0(t') dt'. \text{ For an ergodic process } \langle \vec{P}_0(t) \rangle = \underline{\underline{\langle \vec{P}_0(t) \rangle}}$$

Dephasing: Decay of the Macroscopic Polarization.

- > For a system with no collisions if an impulse incident field is applied, the macroscopic polarization is maximum since all atoms/dipoles line up $\vec{P} = N p_0 \hat{e}$.
 - > Introducing collisions causes an immediate dissipation due to the random phase & amplitude modulations. The macroscopic polarization decays (even without energy loss in the elastic case).
 - > Big assumption made of this being a rate process → heuristic argument
 Assume collision rate $\gamma_c = \frac{1}{T_2'}$, T_2' is the avg time b/w randomizing collisions.
 - > By the definition of a rate process the fractional change in the ensemble average $\langle\langle p \rangle\rangle$ in a time Δt will be,
- $$\frac{\Delta \langle\langle p(t) \rangle\rangle}{\langle\langle p(t) \rangle\rangle} = - \frac{\Delta t}{T_2'} \xrightarrow{\text{decays!}}$$
- this ensure exponential decay.
- $$\Rightarrow \frac{d}{dt} \langle\langle p(t) \rangle\rangle = - \frac{1}{T_2'}, \quad \langle\langle p(t) \rangle\rangle$$
- $$\Rightarrow \langle\langle p(t) \rangle\rangle = p_0 e^{-t/T_2'}$$
- $\Rightarrow \vec{P}(t) = \hat{e} N p_0 e^{-t/T_2'}$
- This decay of macroscopic polarization due to interactions of oscillators at random is called dephasing.
- Phase coherence is lost. T_2' is dephasing time $\times \frac{1}{T_2'}$, is the dephasing rate.

→ So far we have not considered decay in macroscopic polarization due to radiative & nonradiative (inelastic) loss. Adding them in we get,

$$\boxed{\bar{P}(t) = \hat{e} N P_0 e^{-t/T_2}}$$

Where, $\frac{1}{T_2} = \frac{\gamma_r + \gamma_{nr}}{2} + \frac{1}{T_2'}$ is the total decay rate.

use, $\frac{1}{T_1} = \gamma_r + \gamma_{nr} \rightarrow T_1$ is the energy decay time. since only these processes cause energy decay.

$$\Rightarrow \frac{1}{T_2} = \frac{1}{2T_1} + \frac{1}{T_2'}$$

> Generally $T_2' \ll 2T_1 \Rightarrow$ dephasing dominates macroscopic pol-decay.

Equation of motion for average polarization.

$$\gamma = \gamma_r + \gamma_{nr} + \frac{2}{T_2'} = \frac{2}{T_2} \quad \& \text{ note } \langle\langle \bar{F}_r(t) \rangle\rangle = 0 \rightarrow \text{must!}$$

In most case a long time average $\langle\langle \bar{F}_r(t) \rangle\rangle = 0 \rightarrow$ ergodic

$$\Rightarrow \frac{d^2 \langle\langle \bar{p}(t) \rangle\rangle}{dt^2} + \gamma \frac{d}{dt} \langle\langle \bar{p}(t) \rangle\rangle + \omega_0^2 \langle\langle p(t) \rangle\rangle = \frac{e^2}{m} \vec{E}(\vec{R}, t).$$

$$\text{use } \langle\langle \bar{p}(t) \rangle\rangle \cdot N = \bar{P}$$

Solve this for index of refraction.
Lorentz eqn - for Macroscopic polarization.

$$\boxed{\frac{d^2 \bar{P}}{dt^2} + \gamma \frac{d \bar{P}}{dt} + \omega_0^2 \bar{P} = \frac{Ne^2}{m} \vec{E}(\vec{R}, t)}$$

Optical Susceptibility

Recall,

$$\frac{d^2 \bar{P}}{dt^2} + \gamma \frac{d \bar{P}}{dt} + \omega_0^2 \bar{P} = \frac{Ne^2}{m} \bar{E}(t).$$

Consider the steady state response to a time harmonic incident field,

$$\bar{E}(t) = \hat{e} E_0 e^{i\omega t}$$

We expect a soln. of the form.

$$\bar{P}(t) = \hat{e} P_0 e^{i\omega t} \quad \text{where } P_0 \text{ is complex.}$$

Substituting -

$$-\omega^2 P_0 + i\omega \gamma P_0 + \omega_0^2 P_0 = \frac{Ne^2}{m} E_0$$

So

$$P_0 = \frac{Ne^2}{m} \left[\frac{1}{\omega_0^2 - \omega^2 + i\omega \gamma} \right] E_0$$

$$\text{Write } \bar{P} = \epsilon_0 \chi \bar{E}$$

$$\Rightarrow \boxed{\chi(\omega) = \frac{Ne^2}{\epsilon_0 m} \cdot \frac{1}{\omega_0^2 - \omega^2 + i\omega \gamma}} \rightarrow \text{electric susceptibility.}$$

Here the field $\bar{E}(t)$ is not the incident field, it is the resultant field which must be derived self consistently with $\bar{P}(t)$ so we need an equation for \bar{E} & \bar{P} that describes propagation of \bar{E} .

Wave equation with sources.

Fact: Any vector field \bar{A} can be split into a longitudinal field \bar{A}^L and a transverse field \bar{A}^T such that

$$\nabla \cdot \bar{A}^T = 0 \quad \nabla \times \bar{A}^L = 0.$$

$$\Rightarrow \bar{E} = \bar{E}^L + \bar{E}^T$$

$$\text{Since } \nabla \times \bar{E}^L = 0 \quad \Rightarrow \bar{E}^L = -\nabla \phi$$

$$\Rightarrow \nabla \cdot \bar{E} = \nabla \cdot \bar{E}^L = -\nabla^2 \phi$$

Gauss' Law $\Rightarrow \nabla^2 \phi = -\frac{\rho}{\epsilon_0}$ \rightarrow longitudinal field arise from response to charges in electrostatics problems.

> Waves that propagate into the far field and that can be described by the wave eqn. are transverse.

$$\begin{aligned} \nabla \times (\nabla \times \bar{E}^T) &= -\mu_0 \frac{\partial}{\partial t} (\nabla \times \bar{H}^T) \\ &= \nabla(\nabla \cdot \bar{E}^T) - \nabla^2 \bar{E}^T = -\nabla^2 \bar{E}^T \end{aligned}$$

$$\nabla \times \bar{H}^T = \bar{J}^T + \frac{\partial \bar{D}^T}{\partial t} = \bar{J}^T + \frac{\partial}{\partial t} (\epsilon_0 \bar{E}^T + \bar{P})$$

Therefore the sources of radiated fields are $\frac{\partial \bar{J}^T}{\partial t}$ and $\frac{\partial^2 \bar{P}}{\partial t^2}$

\bar{J} is usually used for free charges & \bar{P} for bound charges but it can be hard to distinguish them in optics.

In fact, $\frac{\partial \bar{P}}{\partial t} = \bar{P} \nabla = \bar{J}$ from earlier $\Rightarrow \bar{J} = i \omega \bar{P}$ for time harmonic charges.

(90° phase shift
Physically obvious)

(2.21)

> We will lump \bar{J} & \bar{P} into \bar{P} . (Don't double count)

$$\Rightarrow \nabla^2 \bar{E}^T = \mu_0 \frac{\partial^2}{\partial t^2} (\epsilon_0 \bar{E}^T + \bar{P})$$

Dropping the transverse superscript since we only care about travelling waves,

$$\boxed{\nabla^2 \bar{E} - \mu_0 \epsilon_0 \frac{\partial^2 \bar{E}}{\partial t^2} = \mu_0 \frac{\partial^2 \bar{P}}{\partial t^2}}$$

This equation must be solved simultaneously with the Lorentz equation for $\bar{E}(t) \propto \bar{P}(t)$.

• Time Harmonic Fields - Helmholtz Equation.

Let $\bar{E} = \bar{E}(\vec{r}) e^{i\omega t} \Rightarrow \bar{P} = \epsilon_0 \chi \bar{E}$ where $\chi = \chi(\omega) = \frac{N e^2}{\epsilon_0 m} \left[\frac{1}{\omega_b^2 - \omega^2 + i\gamma\omega} \right]$

Inserting into above equation.

$$\nabla^2 \bar{E}(\vec{r}) + \frac{\omega^2}{c^2} (1+\chi) \bar{E}(\vec{r}) = 0$$

Plane wave solutions

$$\Rightarrow \bar{E}(\vec{r}) = \hat{E} E_0 e^{-i \vec{k} \cdot \vec{r}} \quad \Rightarrow \boxed{k^2 = \frac{\omega^2}{c^2} [1 + \chi(\omega)]} \rightarrow \text{Dispersion relation.}$$

→ Note that k is complex due to $\chi(\omega)$.

→ Define, $\boxed{\tilde{n}^2(\omega) = 1 + \chi(\omega)}$

$$\Rightarrow \boxed{k^2 = \frac{\tilde{n}^2(\omega)^2}{c^2}}$$

→ We can write $\tilde{n} = n(1 - i\chi)$

where n is the familiar refractive index

and $-n\chi$ is the imaginary part that accounts for loss.

$$e^{-ikz} = e^{-in\omega z/c} e^{-n\chi \omega z/c}$$

$$= e^{-in\omega z/c} e^{-z/d}$$

where $d = \frac{c}{n\chi\omega}$ is the skin depth. That is the length over which amplitude decays by $1/e$.

Index of refraction in transparent media.

↪ no loss $\Rightarrow \gamma_r \rightarrow 0$
 $\omega \ll \omega_0$

$$\omega_0^2 - \omega^2 \gg \omega \gamma$$

then,

$$\chi(\omega) \approx \frac{Ne^2}{Gm} \cdot \frac{1}{\omega_0^2 - \omega^2} \quad \text{which is real}$$

$$n^2(\omega) = 1 + \chi(\omega) = 1 + \frac{Ne^2}{Gm} \cdot \frac{1}{\omega_0^2 - \omega^2} \quad \Rightarrow \text{Dispersion since } n \text{ is a fn. of } \omega.$$

i) Visible frequencies are s.t. $\omega_0 > \omega \Rightarrow n > 1$.

ii) $n \propto \omega^1$ or $\propto \lambda^{-1} \Rightarrow$ "Normal dispersion". \Rightarrow "red is faster than blue"

In real materials several resonances may exist so an empirical model which assumes ≈ 3 resonances is used which fits the data well.

$$n^2 = 1 + \sum_j \frac{A_j}{\omega_{0j}^2 - \omega^2}$$

$A_j \propto N_j$ \Rightarrow number density of the j^{th} resonator.

Alternative form:

$$n^2 = 1 + \sum_j \frac{B_j \lambda^2}{\lambda^2 - \lambda_j^2}$$

$$\lambda_j = \frac{2\pi c}{\omega_{0j}}$$

B_j & λ_j tables exist

\rightarrow Note that $n < 0$ or $n < 1$ are possible. $n=0$ aka ENR epsilon near zero materials.

Resonant absorption

What happens as $\omega \rightarrow \omega_0$?

$$X(\omega) = \frac{Ne^2}{6\sigma m} \frac{1}{\omega_0^2 - \omega^2 + i\omega\gamma}$$

$$\omega \approx \omega_0 \Rightarrow \omega_0^2 - \omega^2 = (\omega_0 + \omega)(\omega_0 - \omega) \approx 2\omega_0(\omega_0 - \omega)$$

$$\Rightarrow X(\omega) \approx \frac{Ne^2}{2\sigma m \omega_0} \cdot \frac{1}{\omega_0 - \omega + i\frac{\gamma}{2}} \quad \begin{array}{l} \text{Complex} \\ \text{Horentzian} \\ \text{only valid for } \omega \approx \omega_0 \end{array}$$

$$X(\omega) = X'(\omega) + iX''(\omega) = \frac{Ne^2}{6\sigma m \omega_0 \gamma} \left[\frac{(\omega_0 - \omega) \gamma/2}{(\omega_0 - \omega)^2 + (\gamma/2)^2} - i \frac{(\gamma/2)^2}{(\omega_0 - \omega)^2 + (\gamma/2)^2} \right]$$

Let $X_0'' = \frac{Ne^2}{6m\omega_0\gamma} \rightarrow$ strength of the light matter interactions.

Using γ_r from earlier,

$$X_0'' = \frac{3}{4\pi^2} (N\lambda^3) \frac{\gamma_r}{\gamma}$$

$$\text{Recall } \gamma = \gamma_\gamma + \gamma_{nr} + \frac{2}{T_2'}$$

$$\Rightarrow \frac{\gamma_r}{\gamma} = \frac{\gamma_\gamma}{\gamma_\gamma + \gamma_{nr} + \frac{2}{T_2'}} \leq 1$$

The equality occurs if polarization decays purely through radiation.

↓
photons phonons ↪ dephasing.

> In reality dephasing dominates $\Rightarrow \frac{\gamma_r}{\gamma} \ll 1$ & the strength of the interactions is greatly reduced.

> In heroic experiments you can have $\frac{\gamma_r}{\gamma} \rightarrow 1$ & $X_0'' = \text{constant} \times N\lambda^3$

and the max. strength of light matter interactions is therefore given by the number of oscillators in a volume of cubic wavelength. Later we will see that this strength can be expressed as a "cross section" of an atom (similar to RCS) that the light interacts with.

Real atomic dipole transitions.

$$\text{Recall, } \gamma_r = \frac{e^2 \omega_0^2}{6\pi G m c^3}$$

Real atoms have γ_r that vary over many orders of magnitude, and this is accounted for phenomenologically as

$$\gamma_r = \frac{e^2 f \omega_0^2}{6\pi G m c^3} \quad \text{where } f \text{ is the "oscillator strength" of the transition.}$$

→ In the classical model f is a fitting parameter. In QM it can be derived from first principles.

Optical Susceptibility

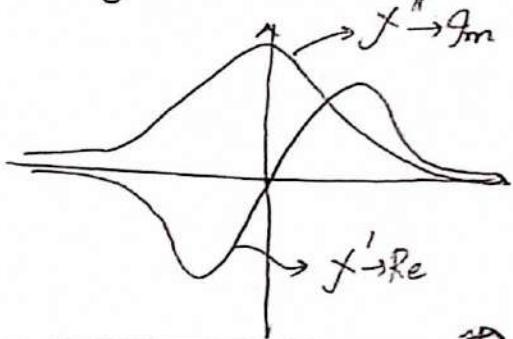
$$\chi(\omega) = -\chi_0'' \left[\frac{(\omega - \omega_0) \gamma_{1/2}}{(\omega - \omega_0)^2 + (\frac{\gamma}{2})^2} + i \frac{(\frac{\gamma}{2})^2}{(\omega - \omega_0)^2 + (\frac{\gamma}{2})^2} \right]$$

$\omega - \omega_0$ is called "detuning". Also the dephasing rate γ will turn out to be the width of the absorption spectra.

⇒ Let us define $\Delta = \frac{\omega - \omega_0}{\gamma_{1/2}}$ as the "normalised detuning" ie. the freq relative to the resonance freq, normalised to the spectral width.

$$X(\Delta) = -X_0'' \left[\frac{\Delta}{1+\Delta^2} + i \frac{1}{1+\Delta^2} \right]$$

- > The real part of X is responsible for dispersion and the function $\frac{\Delta}{1+\Delta^2}$ is called a "dispersive line shape".
- > The imaginary part of X is responsible for optical absorption and the function $\frac{1}{1+\Delta^2}$ is called the "Lorentzian line shape".



Absorption and Dispersion.

$\chi(\omega)$ is complex $\Rightarrow n(\omega)$ is complex.

$$\Rightarrow \tilde{n}(\omega) = \sqrt{1 + \chi(\omega)} = n(\omega)[1 - i\kappa(\omega)]$$

Consider,

$$\tilde{E}(\vec{r}, t) = \hat{E} \underbrace{\mathcal{E}(z)}_{\text{pol}} e^{-iBz} e^{i\omega t}$$

envelope carrier.

$\mathcal{E}(z)$ is complex and hence accounts for both amplitude change & phase change w.r.t carrier e^{-iBz}

Helmholtz \Rightarrow

$$\frac{d^2 \mathcal{E}(z)}{dz^2} - 2i\beta \frac{d \mathcal{E}(z)}{dz} - \beta^2 \mathcal{E}(z) + \frac{\omega^2}{c^2} (1 + \chi) \mathcal{E}(z) = 0$$

Assume envelope varies slowly along Z w.r.t wavelength.

$$\left| \frac{d^2 \mathcal{E}}{dz^2} \right| \ll 2\beta \left| \frac{d\mathcal{E}}{dz} \right| \quad \text{where } \beta = \frac{\omega}{c} = \frac{2\pi}{\lambda}$$

$$\Rightarrow -2i\beta \frac{d\mathcal{E}}{dz} + \beta^2 \chi \mathcal{E} = 0$$

$$\Rightarrow \boxed{\frac{d\mathcal{E}}{dz} = -i \frac{\beta}{2} \chi \mathcal{E}}$$

Here by assigning $\beta = \frac{\omega}{c}$ the damping appears in the envelope & not as the imaginary part of k .

$$\text{Let } X = X' + iX''$$

$$\Rightarrow \mathcal{E}(z) = E_0 e^{(\alpha_m z - i\Delta\beta_m z)}$$

decay of field amp.

$$\text{where } \alpha_m = \frac{\beta}{2} X''(\omega)$$

$\alpha_m \rightarrow$ absorption coefficient

$$\Delta \beta_m = \frac{\beta}{2} X'(\omega)$$

$\Delta\beta_m \rightarrow$ phase shift factor.

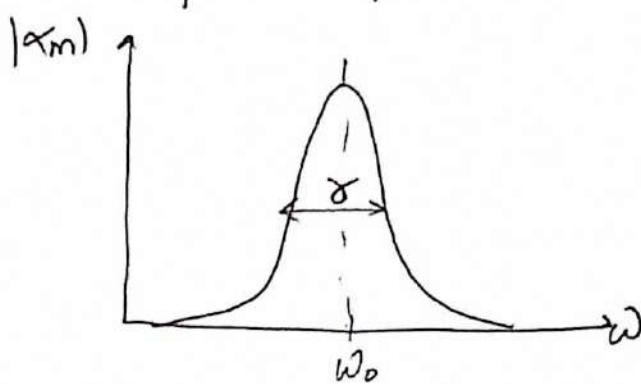
phase shift $\omega -$
vacuum phase shift

Beer's Law of absorption.

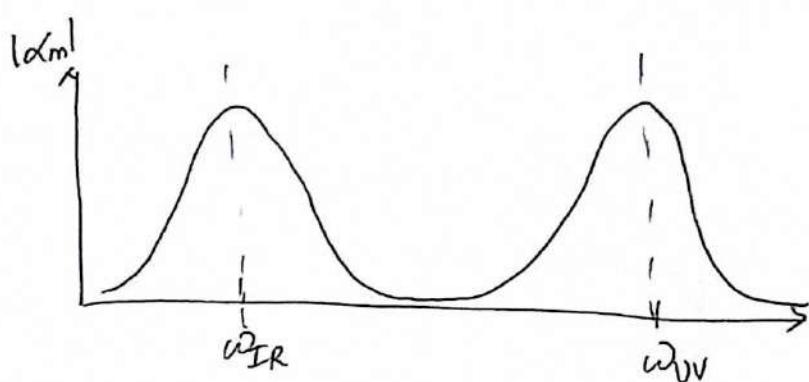
$$\chi'' \sim -\chi_0'' < 0 \quad \text{and } \alpha_m < 0 \Rightarrow |\mathcal{E}(z)| = |E_0| e^{-\alpha_m z}$$

$$\Rightarrow |\mathcal{E}(z)| = |E_0| e^{-|\alpha_m| z} \Rightarrow \text{exp. decay with propagation.}$$

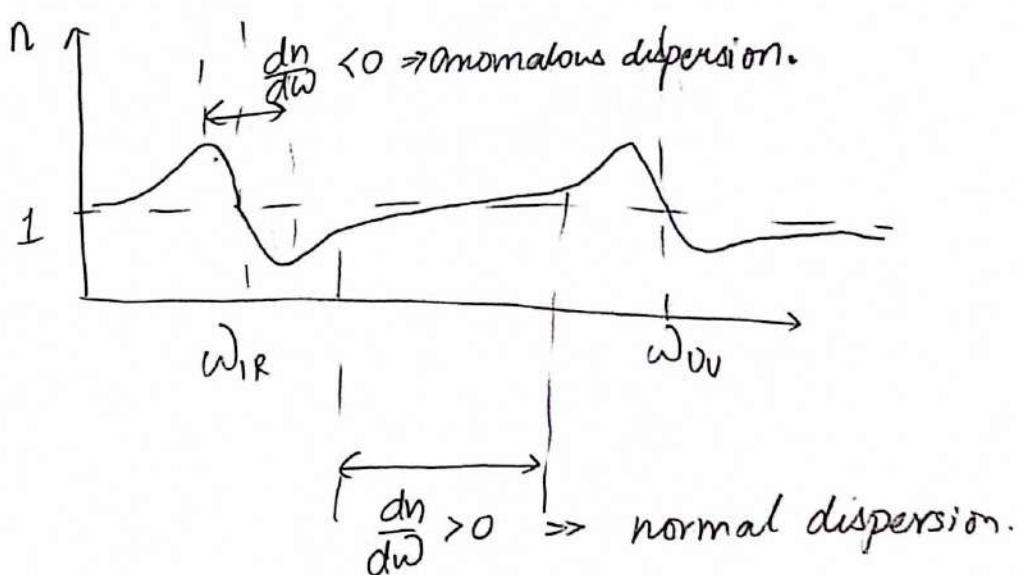
> The absorption spectra is a 'Lorentzian' with FWHM of γ .



Ex.



Different materials have diff. & multiple resonances from diff. modes of vibration of the electrons, diff atomic transitions & so on.

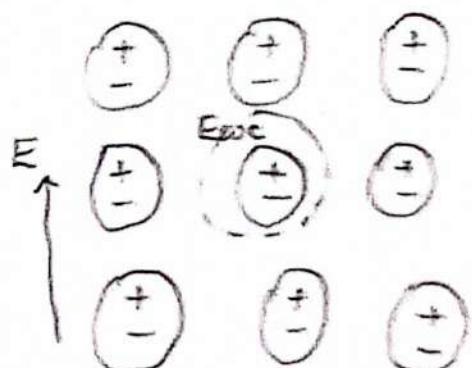


Dense Dielectrics

- > In solids if atoms can interact via dipole-dipole interactions (most solids) we refer to them as dense dielectrics.
- > The dipole induced on a single atom is the sum of the dipole induced by the external field plus the polarization induced by the other dipoles surrounding it.

$$\Rightarrow \vec{P} = \alpha \vec{E}_{loc}$$

→ external field + field due
to neighbouring dipoles



- > This is modeled as a small spherical cavity surrounded by a uniform dielectric polarized by an external electric field.

$$\Rightarrow \vec{E}_{loc} = \vec{E} + \frac{1}{3\epsilon_0} \vec{P}$$

→ macroscopic polarization of medium.

Recall $\vec{P} = N \ll \vec{p} \gg$

$$\Rightarrow \frac{d^2 \vec{P}}{dt^2} + \gamma \frac{d\vec{P}}{dt} + \omega_0^2 \vec{P} = \frac{Ne^2}{m} \vec{E}_{loc} = \frac{Ne^2}{m} \left[\vec{E} + \frac{1}{3\epsilon_0} \vec{P} \right]$$

or

$$\frac{d^2 \vec{P}}{dt^2} + \gamma \frac{d\vec{P}}{dt} + \left(\omega_0^2 - \frac{Ne^2}{3\epsilon_0 m} \right) \vec{P} = \frac{Ne^2}{m} \vec{E}$$

So,

$$\frac{d^2 \vec{P}}{dt^2} + \gamma \frac{d \vec{P}}{dt} + \omega_0'^2 \vec{P} = \frac{Ne^2}{m} \vec{E}$$

Here we have defined a new resonance frequency ω_0' by

$$\omega_0' = \sqrt{\omega_0^2 - \frac{Ne^2}{36m}}$$

dipole-dipole interactions cause a red-shift. AKA Lorentz-Lorentz red-shift.

In classical theory ω_0 is anyway a measured/fitted value so there isn't a distinction between ω_0' & ω_0 .

Metals and Plasmons.

- Since it is difficult to distinguish bound & free charges, the formalism of $\frac{d^2\bar{P}}{dt^2}$ and $\frac{d\bar{J}}{dt}$ are equivalent.

$$\chi(\omega) = \frac{Ne^2}{\epsilon_0 m} \cdot \frac{1}{\omega_b^2 - \omega^2 + i\omega\gamma}$$

- In a metal there are no restoring forces $\Rightarrow \omega_b \rightarrow 0$
- γ is due to, electron-electron and electron-phonon interactions.

$$\Rightarrow \chi(\omega) = -\frac{Ne^2}{\epsilon_0 m} \cdot \frac{1}{\omega^2 - i\omega\gamma}$$

In metals it is convenient to separate the contributions from free electrons & from the ionic lattice to the permittivity

$$\Rightarrow \epsilon_r(\omega) = \epsilon_L [1 + \chi(\omega)]$$

↳ free electrons.
↳ lattice → constant.

$$\Rightarrow \epsilon_r(\omega) = \epsilon_L \left[1 - \frac{\omega_p^2}{\omega^2 - i\omega\gamma} \right]$$

where,

$$\boxed{\omega_p^2 = \frac{Ne^2}{\epsilon_0 m}}$$

is the plasma frequency.

use "effective mass" m^* for a specific metal.

$$\epsilon_r(\omega) = \epsilon'(\omega) + i\epsilon''(\omega)$$

$$\Rightarrow \epsilon'(\omega) = \epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2 + \gamma^2} \right)$$

$$\epsilon''(\omega) = -\epsilon_L \frac{\gamma}{\omega} \left(\frac{\omega_p^2}{\omega^2 + \gamma^2} \right)$$

} Drude model
 Valid when contributions
 from bound electrons is
 very small (Al is a good
 drude metal)

→ Plasma freq. of metals is typically in UV.

$$\text{Conductivity } \bar{\sigma} = \sigma \bar{E}$$

$$\epsilon_r(\omega) = \epsilon_L - \frac{i \sigma(\omega)}{\omega \epsilon_0}$$

(i) low freq. limit

$$\omega \ll \gamma \ll \omega_p$$

$$10^9 \quad 10^{13} \quad 10^{16}$$

$$\epsilon' = \epsilon_L \left(1 - \frac{\omega_p^2}{\gamma^2} \right) \quad \& \frac{\omega_p^2}{\gamma^2} \gg 1$$

$$\Rightarrow \epsilon' \approx -\epsilon_L \frac{\omega_p^2}{\gamma^2} \quad \& \epsilon'' = -\epsilon_L \frac{\gamma}{\omega} \cdot \frac{\omega_p^2}{\gamma^2} = -\epsilon_L \frac{\omega_p^2}{\omega \gamma}$$

$$\Rightarrow \frac{|\epsilon'|}{|\epsilon''|} = \frac{\epsilon_L \omega_p^2 / \gamma^2}{\epsilon_L \omega_p^2 / \omega \gamma} = \frac{\omega}{\gamma} \ll 1$$

$$\Rightarrow \epsilon_r \approx i\epsilon'' \text{ almost pure imaginary.}$$

$$\epsilon_r \approx -i\epsilon_L \frac{\omega_p^2}{\omega}$$

Conductivity, $\epsilon_r - \epsilon_L = -i \frac{\sigma}{\omega \epsilon_0}$

$$\epsilon_r \gg \epsilon_L$$

$$\Rightarrow -i \epsilon_L \frac{\omega_p^2}{\omega \gamma} = -i \frac{\sigma}{\omega \epsilon_0}$$

$$\sigma = \sigma' + i\sigma''$$

$\sigma \approx \sigma'$ almost purely real.

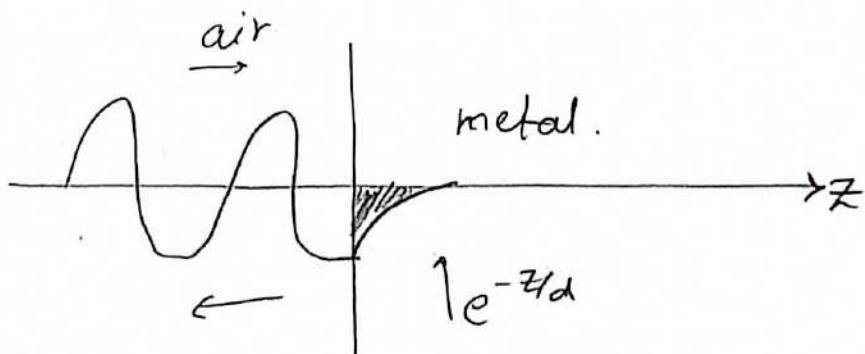
$$\sigma \approx \epsilon_0 \epsilon_L \frac{\omega_p^2}{\gamma} \quad \because \omega_p \gg \gamma \Rightarrow \sigma \text{ large.}$$

(ii) $\gamma \ll \omega \ll \omega_p$ (optical freq)

$$\epsilon' \approx \epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2} \right) \quad \epsilon'' \approx -\epsilon_L \frac{\omega_p^2 \gamma}{\omega^3}$$

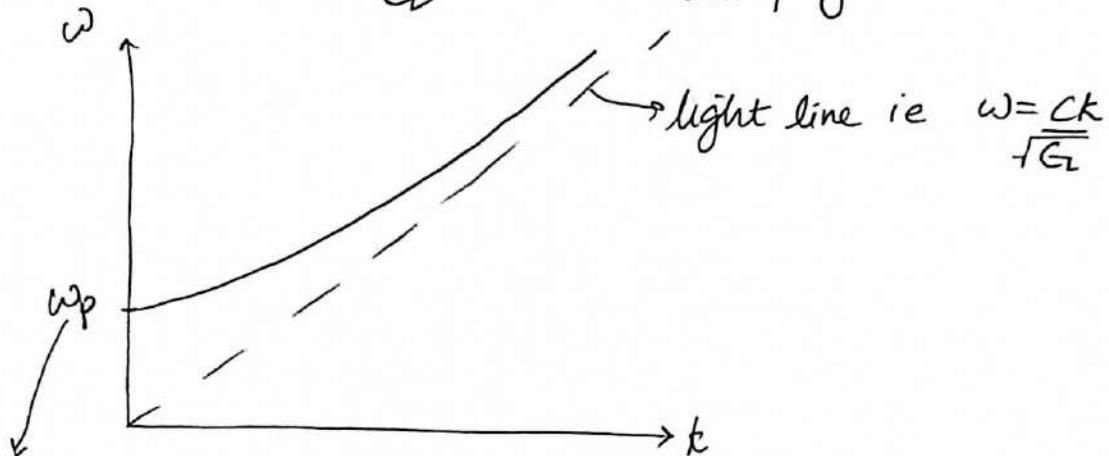
$$\frac{|\epsilon'|}{|\epsilon''|} = \frac{\omega}{\gamma} \gg 1 \quad \Rightarrow \quad \epsilon_{\text{r}}(\omega) \text{ mostly real & negative.}$$

with small imaginary part. $\rightarrow R \approx 90-95\%$ & $5-10\%$ absorption loss in metal.



$$\text{Recall, } k^2 = \frac{\omega^2}{c^2} \epsilon_r(\omega) = \frac{\epsilon_L}{c^2} (\omega^2 - \omega_p^2)$$

or $\omega^2 = \omega_p^2 + \frac{c^2}{\epsilon_L} k^2 \rightarrow \text{Dispersion relation. if we neglect damping.}$



At this frequency the oscillations have \propto wavelength \Rightarrow oscillations occur over the entire length. Here $k = \epsilon_r = 0$

> Plasmons

> Charge oscillations in a bulk metal are known as plasmons.

$$\epsilon' = \epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2} \right) \Big|_{\omega=\omega_p} = 0$$

$$\vec{D} = \epsilon_0 \vec{E} + \vec{P} = \epsilon_r \epsilon_0 \vec{E} = \underline{0} \quad \text{so} \quad \vec{P} = -\epsilon_0 \vec{E}$$

$$\& \nabla \times \vec{H} = -\frac{\partial \vec{B}}{\partial t} = 0.$$

The zeroes of the dielectric f.m. correspond to longitudinal excitations/oscillations of the electron plasma, which do not radiate or couple to propagating transverse EM fields.

Using \bar{J} instead of \bar{P} we have $\bar{J} = \sigma \bar{E}$

$$\nabla^2 \bar{E} = \mu_0 \epsilon_0 \epsilon_r \frac{\partial^2 \bar{E}}{\partial t^2} + \mu_0 \sigma \frac{\partial \bar{E}}{\partial t}$$

Time harmonic $\bar{E} = \bar{E}(r) e^{i\omega t}$

$$\Rightarrow \nabla^2 \bar{E}(r) + \frac{\omega^2}{c^2} \left(\epsilon_r - \frac{i\sigma}{\omega \epsilon_0} \right) \bar{E}(r) = 0$$

$$\Rightarrow \boxed{\epsilon_r = \epsilon_L - \frac{i\sigma}{\omega \epsilon_0}}$$

Surface Plasmons

- An EM wave propagating at the interface b/w a metal & a dielectric.
- Only TM exists

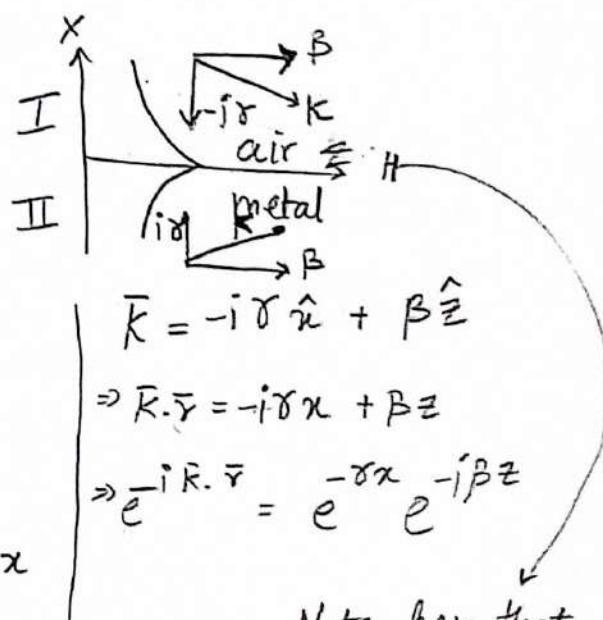
$$\bar{H}(x, z) = \int H(x) e^{-i\beta z}$$

$$\frac{d^2 H(x)}{dx^2} + \left[\epsilon_r(\omega) \frac{\omega^2}{c^2} - \beta^2 \right] H(x) = 0$$

$$H_I(x) = H_0 e^{-\gamma_I x} \quad H_{II}(x) = H_0 e^{\gamma_{II} x}$$

$$\Rightarrow \gamma_I^2 + \epsilon_r \frac{\omega^2}{c^2} - \beta^2 = 0 \quad \& \quad \gamma_{II}^2 + \epsilon_{II} \frac{\omega^2}{c^2} - \beta^2 = 0$$

$$\Rightarrow \gamma_I = \sqrt{\beta^2 - \epsilon_r \frac{\omega^2}{c^2}} \quad \& \quad \gamma_{II} = \sqrt{\beta^2 - \epsilon_{II} \frac{\omega^2}{c^2}}$$



Note here that β must be the same in both regions.

Applying B.C

→ Note $H_I(0) = H_{II}(0) = H_0$

→ Continuity of D_x

$$\nabla \times \vec{H} = \frac{\partial \vec{D}}{\partial t} = i\omega \vec{D} = i\omega (D_x \hat{x} + D_z \hat{z})$$

$$\Rightarrow \nabla \times \vec{H} \Big|_x = \frac{\partial H_z}{\partial y} - \frac{\partial H_y}{\partial z} = -\frac{\partial H_y}{\partial z} \text{ must be continuous.}$$

$$\frac{\partial H_i}{\partial z} = -i\beta H_i \quad (i=I, II) \Rightarrow \beta \text{ must be continuous which we saw from phase matching.}$$

$$\beta^2 = \gamma_I^2 + \epsilon_I \frac{\omega^2}{c^2} = \gamma_{II}^2 + \epsilon_{II} \frac{\omega^2}{c^2} \quad \text{--- (1)}$$

For $\epsilon_I = 1$ (air) & $\epsilon_{II} = \epsilon_r(\omega)$ for metal we have

$$\gamma_I^2 + \frac{\omega^2}{c^2} = \gamma_{II}^2 + \epsilon_r(\omega) \frac{\omega^2}{c^2} \quad \text{where } \epsilon_r(\omega) = \epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2}\right)$$

From continuity of E_z , use $\nabla \times \vec{H} = i\omega \vec{D} = i\omega \epsilon_r \vec{E}$ to see,

$$\frac{1}{\epsilon_I} \frac{\partial H_I}{\partial x} = \frac{1}{\epsilon_{II}} \frac{\partial H_{II}}{\partial x}$$

$$\Rightarrow -\frac{\gamma_I}{\epsilon_I} = \frac{\gamma_{II}}{\epsilon_{II}}$$

$$\Rightarrow \epsilon_r(\omega) = -\frac{\gamma_{II}}{\gamma_I} \quad \text{--- (2)}$$

must be negative!!

Dispersion relations

From ① & ②

$$\epsilon_r(\omega) = -\sqrt{\beta^2 - \epsilon_r(\omega) \frac{\omega^2}{c^2}}$$

$$\sqrt{\beta^2 - \frac{\omega^2}{c^2}}$$

$$\Rightarrow \omega = c\beta \sqrt{\frac{\epsilon_r(\omega) + 1}{\epsilon_r(\omega)}} \Leftrightarrow \boxed{\beta = \frac{\omega}{c} \sqrt{\frac{\epsilon_r(\omega)}{1 + \epsilon_r(\omega)}}}$$

In an ideal metal we can write ($\gamma=0$)

Need to
match to
this β .

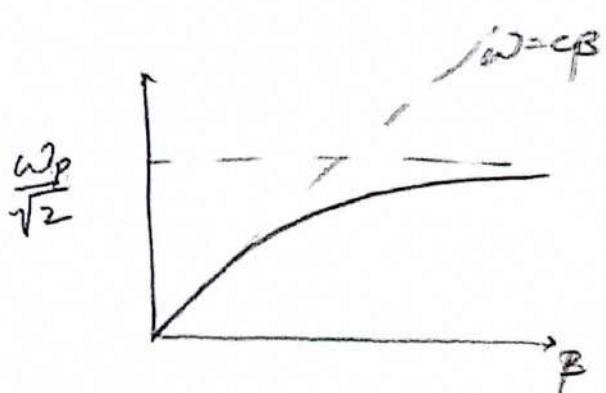
$$\epsilon_r(\omega) = \epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2}\right)$$

$$\Rightarrow \beta = \frac{\omega}{c} \sqrt{\frac{\epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2}\right)}{1 + \epsilon_L \left(1 - \frac{\omega_p^2}{\omega^2}\right)}}$$

> When $\omega \ll \omega_p$; $\beta \sim \frac{\omega}{c}$

> When $\omega = \frac{\omega_p}{\sqrt{1 + \frac{1}{\epsilon_L}}}$; β has a pole & if $\epsilon_L = 1$ which is true in many cases we get

$$\boxed{\omega = \frac{\omega_p}{\sqrt{2}}}$$



$$\boxed{\beta = \frac{\omega}{c} \sqrt{\frac{\epsilon_d \epsilon_r(\omega)}{\epsilon_d + \epsilon_r(\omega)}}}$$

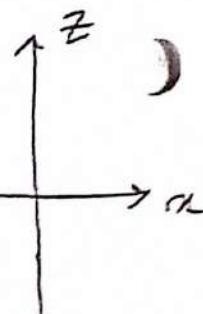
Match to this β to excite an SPP.

Surface Plasmon Polaritons (youtube - diff. notation of $i\kappa r$)

TE: $z > 0$

$$\bar{E}_1 = E \hat{y} e^{i(k_x x - \omega t)} e^{-\alpha_1 z}$$

$$\frac{\epsilon_1}{\epsilon_2} \text{ Air} \quad \text{Metal}$$



$z < 0$

$$\bar{E}_2 = E \hat{y} e^{i(k_x x - \omega t)} e^{\alpha_2 z} \quad (\alpha_1, \alpha_2 > 0)$$

$$\nabla \times \bar{E} = -\mu_0 \frac{\partial \bar{H}}{\partial t} \longrightarrow \mu_0 i \omega \bar{H}$$

$$\Rightarrow H_x = -\frac{1}{\mu_0 i \omega} \frac{\partial}{\partial z} (E_y) \quad \& \quad H_z = \frac{+1}{\mu_0 i \omega} \frac{\partial}{\partial x} (E_y)$$

$$\Rightarrow \underline{z > 0}: \quad \bar{H}_1 = \frac{1}{\mu_0 i \omega} [\alpha_1 \hat{x} + i k_x \hat{z}] E_1 e^{i(k_x x - \omega t)} e^{-\alpha_1 z}$$

$$\underline{z < 0}: \quad \bar{H}_2 = \frac{1}{\mu_0 i \omega} [-\alpha_2 \hat{x} + i k_x \hat{z}] E_2 e^{i(k_x x - \omega t)} e^{\alpha_2 z}$$

$\Theta z=0$ $H_{||}$ continuous. $X \Rightarrow$ cannot exist as TE since $\alpha_1 \neq \alpha_2$

TM:

$$\nabla \times \bar{H} = \mu_0 \frac{\partial \bar{E}}{\partial t} \quad \& \quad \bar{H}_1 = H_1 \hat{y} e^{i(k_x x - \omega t)} e^{-\alpha_1 z}$$

$$\bar{H}_2 = H_2 \hat{y} e^{i(k_x x - \omega t)} e^{-\alpha_2 z}$$

$$\Rightarrow E_x = \frac{1}{i \omega \epsilon} \frac{\partial H_y}{\partial z} \quad E_z = -\frac{1}{i \omega \epsilon} \frac{\partial H_y}{\partial x}$$

$$\underline{z > 0}: \bar{E}_1 = \frac{1}{\epsilon_1 i \omega} \left[-\alpha_1 \hat{x} - i k_x \hat{z} \right] H_1 e^{i(k_x x - \omega t)} e^{-\alpha_1 z}$$

$$\underline{z < 0}: \bar{E}_2 = \frac{1}{\epsilon_2 i \omega} \left[\alpha_2 \hat{x} - i k_x \hat{z} \right] H_1 e^{i(k_x x - \omega t)} e^{\alpha_2 z}$$

$$H_{11} @ z=0 \Rightarrow H_1 = H_2$$

$$E_{11} @ z=0 \Rightarrow \boxed{-\frac{\alpha_1}{\epsilon_1} = \frac{\alpha_2}{\epsilon_2}} \Rightarrow -\frac{\alpha_1}{\alpha_2} = \frac{\epsilon_1}{\epsilon_2}$$

Here α 's are the damping factor.

Similar to earlier we have

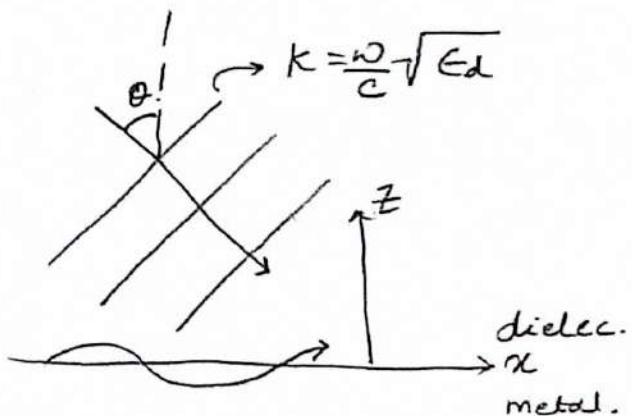
$$k_x = \frac{\omega}{c} \sqrt{\frac{\epsilon_1 \epsilon_2}{\epsilon_1 + \epsilon_2}}$$

Plane wave interaction

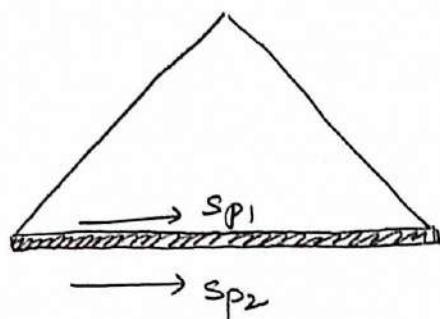
$$k_x = \frac{\omega}{c} \sqrt{\epsilon_d} \sin \theta$$

$$k_x = \frac{\omega}{c} \sqrt{\frac{\epsilon_d \epsilon_m}{\epsilon_d + \epsilon_m}}$$

$$\Rightarrow \boxed{\sin \theta = \sqrt{\frac{\epsilon_m}{\epsilon_d + \epsilon_m}}} \quad \begin{matrix} \\ \downarrow \\ [-1, 1] \end{matrix} \quad \begin{matrix} \\ \nearrow \\ > 1 \end{matrix} \quad \rightarrow \text{Wont work.}$$



Kretschmann Configuration.



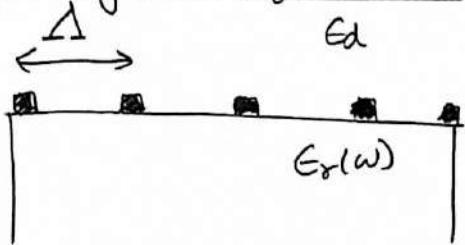
$$K_{\text{up}} = K_{\text{xsp}}$$

$$\frac{\omega}{c} \sqrt{\epsilon_d} \sin \theta = \frac{\omega}{c} \sqrt{\frac{\epsilon_m \epsilon_{air}}{\epsilon_m + \epsilon_{air}}} \quad \text{since it is excited in air}$$

$$\Rightarrow \sin \theta = \sqrt{\frac{\epsilon_m}{\epsilon_d (\epsilon_m + 1)}} = \sqrt{\frac{+8^2}{4 \times 1 + 8}} = \sqrt{\frac{64}{12}} = \sqrt{\frac{16}{3}} \quad \begin{cases} \epsilon_d \approx 4 \\ \epsilon_m = -8 \end{cases}$$

Possible to excite on the outside of the metal strip.

Grating configuration



$$\beta = k_z \text{incident} + \left(\frac{2\pi}{\lambda} \right) \rightarrow ik$$

$$\Rightarrow \frac{\omega}{c} \sqrt{\epsilon_d} \sqrt{\frac{\epsilon_r(\omega)}{\epsilon_d + \epsilon_r(\omega)}} = \sin \theta \frac{\omega}{c} \sqrt{\epsilon_d} + ik$$

$$\Rightarrow \frac{\omega}{c} \sqrt{\epsilon_d} \left[\sqrt{\frac{\epsilon_r(\omega)}{\epsilon_d + \epsilon_r(\omega)}} - \sin \theta \right] = ik.$$

$$\Rightarrow \omega = \frac{ik c}{\sqrt{\epsilon_d}} \left[\sqrt{\frac{\epsilon_r(\omega)}{\epsilon_d + \epsilon_r(\omega)}} - \sin \theta \right]^{-1}$$

$$\epsilon_r(\omega) \approx -i \frac{\sigma(\omega)}{\omega \epsilon_0} \xrightarrow{\text{im } M8} \approx -i(6.27 \times 10^6)$$

$\epsilon_d \approx 4.35$ for oxide U.

\Rightarrow Cannot use a real metal! If $\epsilon_r(\omega) \approx -8$ & $\epsilon_d \approx 4$ and $\theta = 45^\circ$

Linear Response Theory - Kramers-Kronig's Relations

Recall

$$X(\Delta) = -X_0'' \left[\frac{\Delta}{1+\Delta^2} + i \frac{1}{1+\Delta^2} \right] \quad \text{where } \Delta = \frac{\omega - \omega_0}{\delta\omega}$$

\hookrightarrow normalized detuning.

Features of a harmonic model

- i) Linearity $\bar{P} \propto \bar{E}$
- ii) Causality
- iii) Time shift invariance

• $\bar{P}(t) = \epsilon_0 \int_{-\infty}^{\infty} X(t, t') \bar{E}(t') dt' \longrightarrow$ most general form.
 \hookrightarrow response function. (Impulse Response)

Time shift invariance & causality

$\Rightarrow \bar{P}(t) = \epsilon_0 \int_{-\infty}^{\infty} X(t-t') \bar{E}(t') dt' \quad \star X(t-t') = 0 \text{ for } t' > t$

$\Rightarrow \bar{P}(t) = \epsilon_0 \int_{-\infty}^t X(t-t') \bar{E}(t') dt'$

→ For a real function $f(t)$ we have $f(t) = f^*(t)$ since no

imaginary part exists.

$$\Rightarrow F^*(\omega) = \int_{-\infty}^{\infty} f(t) e^{i\omega t} dt = \int_{-\infty}^{\infty} f(t) e^{-i\omega t} dt = F(-\omega)$$

Applying Fourier Transforms to Eq #

$$\bar{P}(\omega) = \mathcal{E} X(\omega) \bar{E}(\omega)$$

Impulse Response of the Harmonic Oscillator

$$X(\omega) = \frac{\omega_p^2}{\omega_0^2 - \omega^2 + i\omega\gamma} \quad \text{where } \omega_p^2 = \frac{Ne^2}{Gm}$$

$$\gamma = \gamma_r + \gamma_{nr} + \frac{2}{T_2'}$$

$$x(t) = \frac{1}{2\pi} \int_{-\infty}^{\infty} x(\omega) e^{i\omega t} d\omega$$

$$= \frac{\omega_p^2}{2\pi} \int_{-\infty}^{\infty} \frac{e^{i\omega t}}{\omega_0^2 - \omega^2 + i\omega\gamma} d\omega \quad \rightarrow \text{use complex integration}$$

Look at notes of 530/537 for complex integral evaluation.

$$x(t) = \frac{\omega_p^2}{\omega_a} \sin \omega_a t e^{-\frac{\gamma t}{2}} \theta(t)$$

$$\omega_a = \sqrt{\omega_0^2 - \gamma^2/4}$$

$$\gamma = \gamma_r + \gamma_{nr} + \frac{2}{T_2'},$$

$$\omega_p = \frac{Ne^2}{Gm}$$

$\theta(t) \rightarrow$ unit step fn.
↳ causality.

$\omega_a \rightarrow$ renormalized oscillation frequency

$\frac{\gamma}{2} \rightarrow$ damping rate

$x(t)$ is real. & is a damped sinusoid for $t \geq 0$.

Kramers-Kronig's Relations

- i) Causality $\Rightarrow X(t) = 0 \text{ for } t < 0$
- ii) Boundedness $\Rightarrow X(t)$ bounded for all t
- iii) Damping $\Rightarrow X(t) \rightarrow 0 \text{ as } t \rightarrow \infty$
- iv) Time shift invariance $\Rightarrow X(t, t') = X(t - t')$.

guarantee that $X(\omega)$ is analytic in the LHP \rightarrow by this convention.

Complex Analysis gives (Check 530/537 notes)

$$\begin{aligned} X'(\omega) &= -\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{X''(\omega')}{\omega' - \omega} d\omega' && \left. \begin{array}{l} \text{Hilbert Transforms} \\ \Rightarrow X' \text{ & } X'' \text{ form a} \\ \text{Hilbert transform pair.} \end{array} \right. \\ X''(\omega) &= \frac{1}{\pi} \int_{-\infty}^{\infty} \frac{X'(\omega')}{\omega' - \omega} d\omega' \end{aligned}$$

Also $X(\omega) = X^*(\omega)$

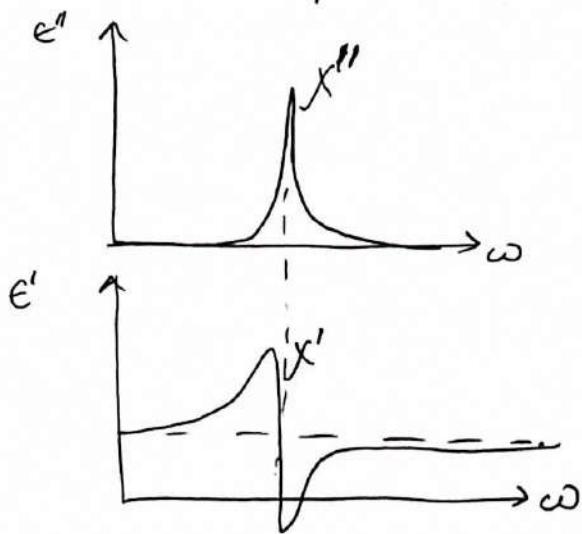
$\Rightarrow \operatorname{Re}\{X(\omega)\} = X'(\omega)$ is even in ω

$\operatorname{Im}\{X(\omega)\} = X''(\omega)$ is odd in ω .

$$\Rightarrow \boxed{\begin{aligned} X(\omega) &= -\frac{2}{\pi} \int_0^{\infty} \frac{\omega' X''(\omega')}{\omega'^2 - \omega^2} d\omega' \\ X''(\omega) &= \frac{2\omega}{\pi} \int_0^{\infty} \frac{X'(\omega')}{\omega'^2 - \omega^2} d\omega' \end{aligned}} \quad \begin{array}{l} \text{More useful} \\ \text{since } \omega \geq 0. \\ \text{Also } E(\omega) = 1 + X(\omega) \end{array}$$

Physical Significance

- i> Absorption \Leftrightarrow Dispersion.
- ii> Measure either Absorption or Dispersion is enough.
- iii> Sharp absorption peak \Rightarrow sharp change in ϵ or n .



Linear Response - Nonmonochromatic light

→ Phase velocity

$$V_p = \frac{\omega}{k(\omega)} = \frac{c}{n(\omega)} \rightarrow \text{velocity of a wavefront.}$$

→ A pure harmonic wave extends to $\pm\infty$ in space & time & has no information

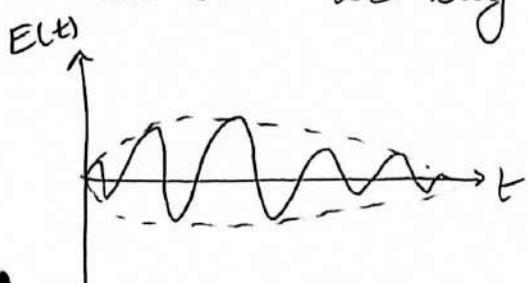
Wave packet

> Construct a wave packet around $\bar{\omega}$ with $BW = \Delta\omega$

$$\Rightarrow E(z, t) = \int_{\bar{\omega} - \frac{\Delta\omega}{2}}^{\bar{\omega} + \frac{\Delta\omega}{2}} E(\omega) e^{i(\omega t - \beta z)} d\omega$$

$$= \int_{-\infty}^{\infty} E(\omega) e^{i(\omega t - \beta z)} d\omega \quad \text{where } E(\omega) \text{ is defined only over } \bar{\omega} - \frac{\Delta\omega}{2} \text{ to } \bar{\omega} + \frac{\Delta\omega}{2}.$$

Since $\Delta\omega \ll \bar{\omega}$ we say the wave is quasimonochromatic

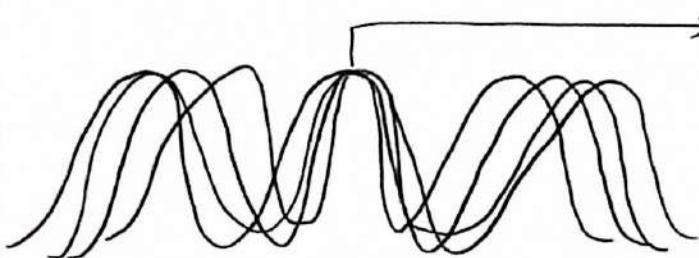


V_p is different for different frequencies.

Group Velocity.

> Velocity of the wave packet/envelope.

construction of envelope



At the peak of packet the waves are all in phase. \Rightarrow phase variation is zero w.r.t frequency.

$$\Rightarrow \phi = \omega t - \beta z \Rightarrow \frac{d\phi}{d\omega} = 0 \Rightarrow t = \frac{d\beta}{d\omega} z$$

Group delay $t_g = \frac{d\beta}{d\omega} z = \frac{z}{v_g}$

Group velocity

$$v_g = \left. \frac{d\omega}{d\beta} \right|_{\bar{\omega}}$$

$$v_g = \frac{c}{n + \omega \frac{dn}{d\omega}}$$

since $\frac{d\beta}{d\omega} = \frac{d}{d\omega} \left(\frac{n\omega}{c} \right) = \frac{n}{c} + \frac{\omega}{c} \frac{dn}{d\omega}$



> Close to resonance $\frac{dn}{d\omega} < 0$ (when $\omega \approx \omega_0$)

and $v_g > c$ but this does not violate relativity since absorption is very strong and the pulse does not propagate beyond one pulse width or so.

> If $\Delta\omega_p \gg \gamma$, the pulse shape is severely distorted and v_g can

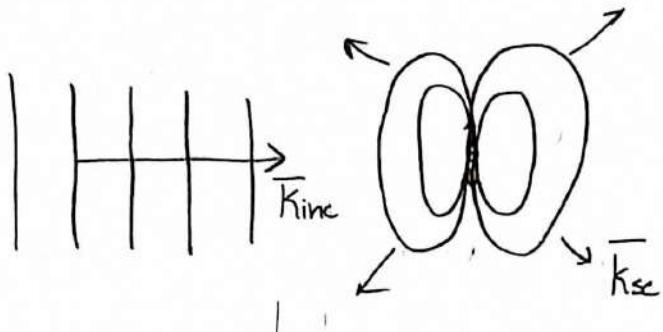
- > When light propagates through a gain medium (population inversion) we can have phase, group & energy velocities > 0 . $\chi \propto N \propto \epsilon$ & $N = N_1 - N_2$
 - > If $N_2 > N$, (pop. inv) χ flips sign &
-
- > This was resolved by R. Chiao (1993)
 - > Causality is not violated. This happens due to a pulse reshaping and there is no information at the peak that wasn't already in the tail of the pulse. Relativity is not violated

Pulse Propagation.

- 1) Given $E_i(z_i, t) \rightarrow$ find $E(z, \omega) = F\{E_i(z_i, t)\}$
 - 2) $E(z, \omega) = H(\omega) E(z, \omega)$
 ↳ Transfer function.
 $H(\omega) = |H(\omega)| e^{-i\phi(\omega)}$
 in a medium:- $H(\omega) = e^{-i\tilde{k}(\omega)z}$ with $\tilde{k}(\omega) = \frac{\tilde{n}(\omega)}{c}\omega$
 or $H(\omega) = e^{-\alpha_m(\omega)z - i\Delta\beta_m(\omega)z}$ with $\alpha_m = \frac{\beta}{2v} \chi''$ &
 $\Delta\beta_m = \frac{\beta}{2} \chi'$.
 - 3) $E(z, t) = F^{-1}\{E(z, \omega)\}$
- Or directly use time domain convolution.
- $E_{out}(t) = h(t) * E_{in}(t)$. → real impulse response.

Light Scattering

① Single dipole oscillator



Induced dipole moment

$$\vec{P}(\omega) = \alpha(\omega) \hat{e} E_{\text{inc}} \cos \omega t$$

horizontally: $\alpha(\omega) = \frac{e^2/m}{\omega_0^2 - \omega^2 + i\omega\gamma}$ polarizability

Power radiated by dipole

$$\langle w \rangle = \frac{p^2 \omega^4}{6\pi \epsilon_0 c^3}$$

$$= \frac{\omega^4 \alpha^2 E_{\text{inc}}^2}{6\pi \epsilon_0 c^3} \langle \cos^2 \omega t \rangle \quad \text{where}$$

$$\frac{1}{\epsilon_0 c} \langle \vec{S} \cdot \vec{S} \rangle = E_{\text{inc}}^2 \langle \cos^2 \omega t \rangle$$

is the avg Poynting vector.

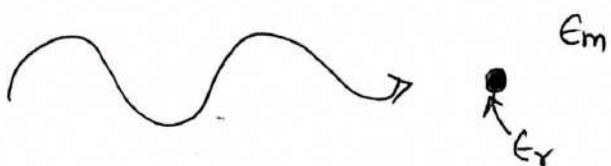
$$\Rightarrow \boxed{\langle w \rangle = \left[\frac{\omega^4 \alpha^2(\omega)}{12\pi \epsilon_0^2 c^4} \right] \langle \vec{S} \cdot \vec{S} \rangle} \quad \hookrightarrow I_{\text{inc}} \rightarrow \text{Intensity incident}$$

$$\Rightarrow \boxed{\langle w \rangle = \sigma(\omega) I_{\text{inc}}^{(\omega)}}$$

Scattering cross section.

② Small Spherical Particle. ($a \ll \lambda$)

↳ (Rayleigh scattering.)



Electrostatic approximation.

$$\Rightarrow \alpha = a^3 \frac{\epsilon_r - \epsilon_m}{\epsilon_r + 2\epsilon_m}$$

$$\Rightarrow \sigma = \frac{\omega^4 a^6}{12\pi \epsilon_0^2 c^4} \left(\frac{\epsilon_r - \epsilon_m}{\epsilon_r + 2\epsilon_m} \right)^2$$

For frequency ω far away from resonances σ goes as ω^4

$$\text{or } \sigma \propto \frac{1}{\lambda^4}$$

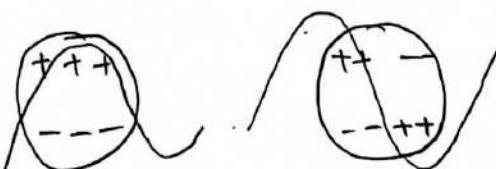
For metal particles when $\epsilon_r < 0$ we can find $\epsilon_r = -2\epsilon_m$

\Rightarrow "Fröhlich condition".

Which gives rise to strong colors in visible light from metal nanoparticles. (Stained glass in Churches that don't fade).

When $a \sim \lambda$, multiple resonance modes may be excited
↳ (quadrupoles, octupoles etc.)

which are analysed using Mie theory.



> $a \gg \lambda \Rightarrow \sigma$ is basically constant and approaches the physical dimensions. White light scatters into white light. (clouds, milk etc) \rightarrow most white stuff.

$$\sigma \sim \pi a^2$$

③ N molecules in a confined region $\ll \lambda$



$$\vec{P}_{\text{blob}} = N \vec{P}_i$$

$$\sigma_{\text{blob}} = \frac{\omega^4 N^2 \alpha^2(\omega)}{12 \pi \epsilon_0^2 c^4}$$

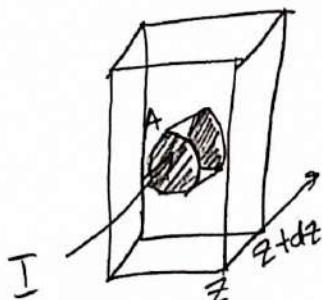
dependence on $N^2 \Rightarrow$ "Coherent scattering".
Coherent since dipoles are all in phase.

$$\sigma_N = N^2 \sigma_{\text{dipole}}$$

When "blob" size $\geq \lambda$ we get $\sigma_N = N \sigma_{\text{dipole}}$.

④ Scattering by a collection of N independent scatterers

→ Suppose there are N scatterers per unit volume.



$$P(z) = A I(z)$$

$$P(z+dz) = A I(z+dz)$$

Power scattered out of the beam = # atoms \times Scattered power per atom

$$P_{\text{scatt}} = (N \cdot A \cdot dz) \sigma I(z)$$

Conservation of energy $\Rightarrow P(z) - P(z+dz) = N A \sigma I dz$

or
$$\frac{dI}{dz} = -\sigma N I$$
 $\underbrace{- A dI(z)}_{\text{in } \lim_{dz \rightarrow 0}}$

$$\Rightarrow I(z) = I_{\text{inc}} e^{-\sigma N z}$$

Exponential attenuation from pure scattering (ignoring absorption)

More density \Rightarrow more scattering

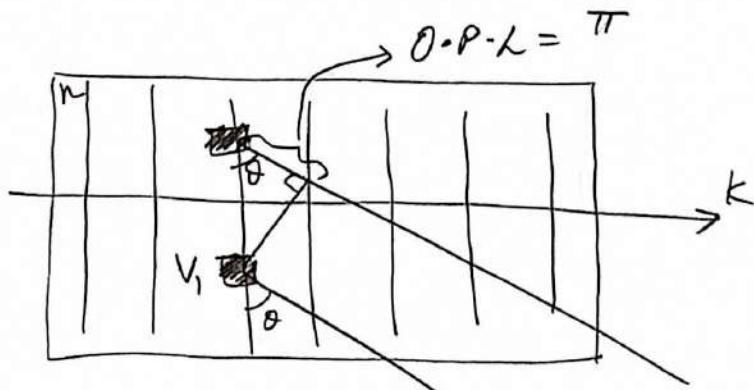
Here we have made a big assumption that each atom scatters independently of others. \Rightarrow each atom scatters power out of beam direction.

Accounting for absorption $\Rightarrow \sigma_{\text{extinction}} = \sigma_{\text{absorption}} + \sigma_{\text{scattering}}$.

#5) Scattering by a perfectly uniform distribution of scatterers

> There is a seeming contradiction b/w previous models (where a uniform dielectric that does not absorb light) that allow non attenuating propagation through a dielectric and #4 (which suggests atoms always cause attenuation by scattering).

Consider a plane wave inside a dielectric. For any direction θ , given a volume ΔV_1 that scatters along θ , there is always a volume ΔV_2 from where scattering along θ is out of phase with original scatterer. Therefore scattering occurs only along \bar{k} which is basically the wave propagation itself.



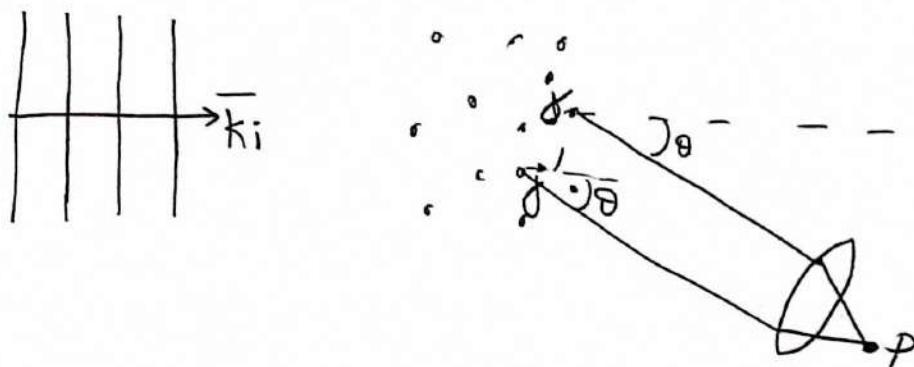
If $\Delta V_1, \Delta V_2 \ll \lambda^3$
they are identical
scatterers.

→ Fluctuations in n with space cause attenuation (ie side scattering)

#6

N point scatterers over region $\gg \lambda$

Consider plane wave incident on a collection of scatterers. Use a lens to focus all light scattered at an angle θ to a point P.



Field at P due to j^{th} dipole is

$$E_j = E_0 e^{-i\bar{k} \cdot \bar{r}_j} + i\theta_j \quad \text{where } E_0 \text{ is field amplitude from one dipole.}$$

Intensity at P is proportional to

$$I_p = \left(\sum_{j=1}^N E_j \right) \left(\sum_{j'=1}^N E_j^* \right) = |E_p|^2$$

$$= \sum_{jj'} |E_0|^2 e^{-i\bar{k} \cdot (\bar{r}_j - \bar{r}_{j'})} + i(\phi_j - \phi_{j'})$$

The incident wave sets up the relative phases as

$$\phi_j - \phi_{j'} = \bar{k}_i \cdot (\bar{r}_j - \bar{r}_{j'})$$

Thus

$$I_p = I_0 \sum_{j,j'} e^{-i(\bar{k} - \bar{k}_i) \cdot (\bar{r}_j - \bar{r}_{j'})}$$

i) Forward scattering: $\bar{k} = \bar{k}_i$

$$I_p = I_0 \left(\sum_{j=1}^N 1 \right) \left(\sum_{j'=1}^N 1 \right) = N^2 I_0$$

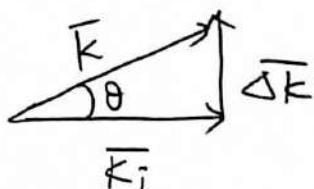
↳ goes as N^2 like coherent scattering.

ii) Side scattering $\bar{k} \neq \bar{k}_i$

Define $\Delta \bar{k} = \bar{k} - \bar{k}_i$ as the "scattering wave vector"

We've been assuming elastic scattering. (i.e. no absorption)
so that,

$$|\bar{k}| = |\bar{k}_i| = \frac{\omega}{c}$$



$$I_p = I_0 \sum_{j=1}^N \sum_{j'=1}^N e^{-i\Delta \bar{k} \cdot (\bar{r}_j - \bar{r}_{j'})}$$

We'll consider 2 cases - dipole scatterers arranged randomly or in a perfect crystal.

a) Random Scatterers

⇒ for any atom j , the sum

$$\sum_{j'} e^{-i \bar{\Delta k} \cdot (\bar{r}_j - \bar{r}_{j'})} = e^{-i \bar{\Delta k} \cdot \bar{r}_j} \underbrace{\sum_{j' \neq j} e^{i \bar{\Delta k} \cdot \bar{r}_{j'}}}_{+ 1}$$

This term oscillates wildly over j' & averages to 0.

Thus $I_p = I_0 \sum_{j=1}^N 1 = N I_0$

Thus we recover the case of N "independant" scatterers when they are randomly positioned.

⇒ Incoherent scattering $\propto N$. ie $I = N I_0$.

Side scattering occurs & $I(z) = I_{\text{inc}} e^{-\sigma N z}$

b) Scattering centers arranged on a perfect lattice

⇒ It is equivalent to the homogeneous case when $\lambda \gg a$.

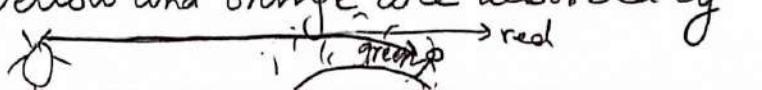
⇒ only coherent forward scattering occurs.

Real life: Thermal fluctuations & finite medium size along with defects always cause density fluctuations ⇒ exponential attenuation.

Wavelength dependence of scattered light

Why is the sky blue?

$$\text{Recall } \sigma(\omega) = \frac{\omega^4 \alpha^2(\omega)}{12\pi G_0 c^4}$$

- > When $\alpha(\omega)$ shows strong resonances we may see colourful scattering at those frequencies (metal nanoparticles)
- > Rayleigh Scattering
- > Nonresonant scattering in transparent media when $\alpha \ll \lambda$.
- > $\omega \ll \omega_0, \omega \gamma \ll \omega_0^2 \Rightarrow \alpha_{CEO}(\omega) \approx \frac{c^2}{m\omega_0^2} \Rightarrow$ freq. independant
↳ classical electric oscillator
- $\Rightarrow \sigma \propto \omega^4 \Rightarrow$ blue light (400nm) scatters 16 times more than red light (800nm). \Rightarrow Blue sky.
- > Red sunset \Rightarrow more atmosphere \Rightarrow blue light scatters away & red light scatters towards observer.
- > Clouds are large water molecules & scatter equally at all frequencies ($\alpha \geq \lambda$)
- > Gray clouds are due to attenuation of white light as it passes through the clouds.
- > Green flash right after sunset: Blue is scattered away & green bends more than red. Yellow and orange are absorbed by water vapour & O_2 molecules.


Geometrical Optics

From Waves to Rays

$$\nabla^2 \bar{E} + k^2 \bar{E} = 0 \quad \text{where } k^2 = \frac{n^2 \omega^2}{c^2}, \text{ let } k_0 = \frac{\omega}{c}$$

Consider the solution (scalar soln.)

$$E(\vec{r}, t) = A(\vec{r}) e^{-ik_0 S(\vec{r})} e^{i\omega t} \longrightarrow (\text{Wentzel, Kramers, Brillouin, WKB Solution.})$$

- 1) It is a scalar solution
- 2) Assume A varies slowly with \vec{r} .
- 3) $S(\vec{r})$ is called the Eikonal function.

Such a solution is useful when $n \rightarrow n(\vec{r})$.

$$\text{Ex: } n = n(z)$$

$$\Rightarrow E(z, t) = A(z) e^{i[\omega t - k_0 S(z)]}$$

or

$$E(z) = A(z) e^{-i k_0 S(z)}$$

Substituting into HH.

$$\Rightarrow \left[\frac{d^2 A}{dz^2} - A \left(k_0 \frac{ds}{dz} \right)^2 \right] - i \left[2 k_0 \frac{ds}{dz} \frac{dA}{dz} + A k_0 \frac{d^2 s}{dz^2} \right] + k^2 A = 0$$

$$\Rightarrow \frac{d^2 A}{dz^2} + \left[\frac{n(z)^2 \omega^2}{c^2} - (k_0 \frac{ds}{dz})^2 \right] A = 0 \quad \& \quad 2 k_0 \frac{ds}{dz} \frac{dA}{dz} + A k_0 \frac{d^2 s}{dz^2} = 0$$

If $n(z)$ varies slowly, then A should be slowly varying and $\frac{d^2 A}{dz^2} \rightarrow 0$ (aka WKB approximation)

$$\Rightarrow \frac{n(z)^2 \omega^2}{c^2} - k_0^2 \left(\frac{ds}{dz} \right)^2 = 0$$

$$\Rightarrow \frac{ds(z)}{dz} = n(z)$$

$$\Rightarrow \boxed{s(z) = \int_0^z n(z') dz'} \rightarrow \text{WKB solution for phase.}$$

$$\Rightarrow E(z) = A(z) e^{-ik_0 \int_0^z n(z') dz'}$$

slowly varying

WKB solution looks locally like a plane wave

Physical meaning: Assume $n(z) = n \Rightarrow s(z) = nz$

\Rightarrow optical path length = index \times physical length.

$\Rightarrow s(z)$ is the optical path length & $k_0 s = 2\pi \left(\frac{s}{\lambda} \right)$

$\Rightarrow \frac{s}{\lambda}$ is the "number of waves" in a given OPL.

WKB in 3D

$$n = n(\vec{r})$$

$$\Rightarrow E(\vec{r}) = A(\vec{r}) e^{-ik_0 S(\vec{r})}$$

$$\stackrel{\text{HH}}{\Rightarrow} \nabla^2 A + [n^2 - (\nabla S) \cdot (\nabla S)] k_0^2 A = 0$$

$$\& 2(\nabla S) \cdot (\nabla A) + (\nabla^2 S) A = 0$$

Neglect $\nabla^2 A$ assuming $n(\vec{r})$ varies slowly.

$$\Rightarrow \nabla S \cdot \nabla S = \boxed{|\nabla S|^2 = n^2(\vec{r})} \rightarrow \text{"Eikonal equation"}$$

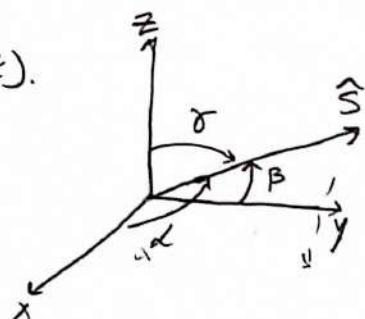
Note that when $S(\vec{r}) = \text{constant}$, this describes a surface of constant phase \Rightarrow the phase is $\phi = k_0 S$ everywhere on that surface.

Examples.

i) $n(\vec{r}) = n = \text{constant}$ (uniform medium)

Suppose initial solution is a plane wave with directional cosines α, β and γ . $\Rightarrow S = n(\alpha x + \beta y + \gamma z)$.

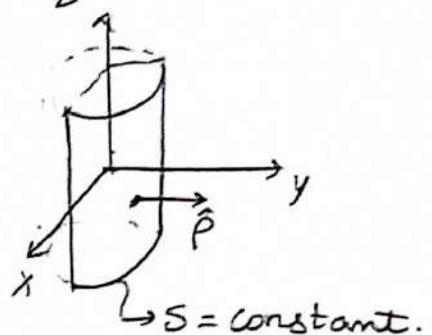
$$\& \nabla S = n \hat{S} \text{ where } \hat{S} = \alpha \hat{x} + \beta \hat{y} + \gamma \hat{z}$$



ii) $n = \text{constant}$, initial wave surface is a cylinder

$$\Rightarrow S = np \quad \& \quad \hat{S} = \hat{p}$$

$$\nabla S = n\hat{p} \rightarrow \nabla S = n\hat{S}$$



iii) $n = \text{constant}$, initial wave = sphere

$$S = nr ; \quad \nabla S = n\hat{r} ; \quad \nabla S = n\hat{S}$$

In general

$$\nabla S = n(\vec{r}) \hat{S}(\vec{r})$$

where $\hat{S}(\vec{r})$ is a unit vector normal to the surface at \vec{r} .

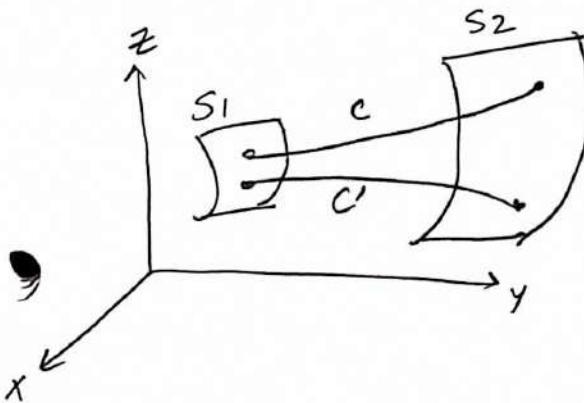
∇S is a vector perpendicular to the surface $S = \text{constant}$.

Def: A ray is the path taken by \hat{S} as the corresponding wave propagates through space.

When is this GO formalism valid? When $\lambda \rightarrow 0$ since then A is basically constant over a wavelength and $\nabla^2 A \rightarrow 0$.

Note, $\int_A^B |\nabla s| ds = s(\vec{r}_B) - s(\vec{r}_A) = \int_A^B n ds$.
 optical path length from $A \rightarrow B$.

\Rightarrow Two rays going from S_1 (one surface of constant phase) to S_2 (another surface of constant phase), the optical path length is the same.



$$\int_c n ds = \int_{c'} n ds = S_2 - S_1$$

Consider the rate of change of $n\hat{s}$ along a ray:

$$\begin{aligned} \frac{d}{ds}(n\hat{s}) &= \frac{d}{ds}(\nabla s) \\ &= \hat{s} \cdot \nabla(\nabla s) \\ &= \frac{\nabla s}{n} \cdot \nabla(\nabla s) \\ &= \frac{1}{2n} \nabla(\nabla s)^2 \\ &= \frac{1}{2n} \nabla n^2 = \nabla n \end{aligned}$$

$\Rightarrow \boxed{\frac{d}{ds} n\hat{s} = \nabla n}$ "ray equation"

Given $n(\vec{r})$ & an initial ray direction \hat{s}_0 , we can find the path taken by the ray.

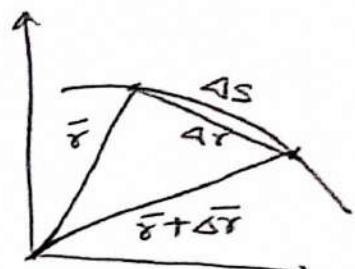
→ In words, the change in the optical path along the path is given by the gradient of the index of refraction.

→ s here is the parameter describing the ray.

Note:

$$\hat{s} = \frac{d\bar{r}}{ds} = \hat{x} \frac{dx}{ds} + \hat{y} \frac{dy}{ds} + \hat{z} \frac{dz}{ds}$$

ds = length element along curve



Examples.

i) $n = \text{constant}$; $\nabla n = 0$

$$\Rightarrow \frac{d}{ds} (n \hat{s}) = 0 \Rightarrow \hat{s} = \text{constant} \Rightarrow \text{straight line.}$$

ii) $n = n(y)$; $\nabla n = \frac{dn}{dy} \hat{j}$ & ray travelling along \hat{j} .

$$\Rightarrow \frac{d}{ds} = \frac{d}{dy} ; \quad \hat{s}_0 = \hat{j}$$

$$\Rightarrow \frac{d}{dy} (n \hat{s}) = \frac{dn}{dy} \hat{j} \Rightarrow \hat{s} = \hat{j} \Rightarrow \text{no change in direction.}$$

iii) $n = n(y)$ and $\hat{s}_0 = \hat{i}$, the ray bends towards higher n .

iv) In general, $n = n(y)$ $\hat{s}_0 = \hat{x} \sin \theta_0 + \hat{y} \cos \theta_0$

Vector ray equation can be split into 3 components.

$x: \frac{d}{ds} (n \sin\theta) = \frac{dn}{dx} = 0$

$$\hat{z} = \hat{x} \sin\theta + \hat{y} \cos\theta + \hat{z} \gamma.$$

$y: \frac{d}{ds} (n \cos\theta) = \frac{dn}{dy} \neq 0$

$z: \frac{d}{ds} (n \gamma) = 0 \quad \text{where } \gamma \text{ is direction cosine along } z.$

$\Rightarrow n\gamma = \text{constant}$ but $\gamma_0 = 0 \Rightarrow \gamma = 0 \Rightarrow \text{ray stays in } xy$ plane.

$x: n \sin\theta = \text{constant} = n_0 \sin\theta_0 \Rightarrow \text{Snell's Law is locally obeyed.}$

$y: \text{If } \frac{dn}{dy} > 0, \text{ the ray gradually turns towards } y \text{ axis.}$

In general, the ray bends towards higher refractive index.

Fermat's Principle

The loosest form of the statement is:

"Of all the paths light might take b/w 2 points, the actual path taken is the one that requires the least time".

Relationship b/w differential path element & time element is

$C dt = n ds \Rightarrow \int_A^B dt = \int_{A_i}^B n ds.$

$(V = \frac{C}{n})$

Eikonal equation gives,

$$\nabla S = n\hat{s}$$

$$\underbrace{\nabla \times (\nabla S)}_0 = \nabla \times n\hat{s} = 0$$

$$\Rightarrow \iint_A \nabla \times (n\hat{s}) \cdot d\bar{a} = 0 \quad \text{where } A \text{ is an open surface}$$

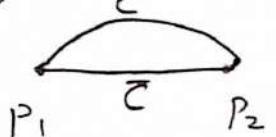
$$\Rightarrow \boxed{\oint_C n\hat{s} \cdot d\bar{r} = 0.} \rightarrow \text{Lagrange's integral invariant.}$$

This gives an easy proof of Fermat's Principle.

> Consider a ray passing through points P_1 and P_2 along \rightarrow a curve \bar{C} .

> Consider another curve \bar{C}' also passing through P_1 & P_2 .

Apply Lagrange's invariant to the loop,



$$\oint_C n\hat{s} \cdot d\bar{r} - \oint_{C'} n\hat{s} \cdot d\bar{r} = 0$$

$$\text{Along } \bar{C} \quad \hat{s} \parallel d\bar{r} \Rightarrow \int_C n ds = \int_{C'} n\hat{s} \cdot d\bar{r}$$

$$\text{Triangle inequality} \Rightarrow \hat{s} \cdot d\bar{r} \leq ds \Rightarrow \int_C n ds \leq \int_{C'} n ds$$

"Path of wavefront normal is also the path of front at time."

dr is along the boundary, whereas
 ds is along \bar{C} .

Generalized Fermat's Principle.

"The optical path traversed from P_1 to P_2 is that for which the integral $\int_{P_1}^{P_2} n ds$ is stationary."

⇒ The path integral is a local optimum.

Mathematically $\delta \int_{P_1}^{P_2} n ds = 0 \Rightarrow$ small change in path does not change the integral value.

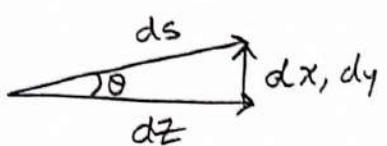
We can use the calculus of variations to derive the ray equation from this statement.

Graded Index (GRIN) optics

$$n = n(r) = n(\sqrt{x^2 + y^2}) \Rightarrow \text{Cylindrical symmetry.}$$

Use paraxial rays approximation ⇒ rays propagating at small angles with respect to Z-axis.

$$ds = dz \sqrt{1 + \left(\frac{dx}{dz}\right)^2 + \left(\frac{dy}{dz}\right)^2}$$



Taylor

$$\approx dz \left[1 + \frac{1}{2} \left(\frac{dx}{dz} \right)^2 + \frac{1}{2} \left(\frac{dy}{dz} \right)^2 \right]$$

$$\times \frac{dx}{dz} = \tan \theta = \sin \theta = \theta \quad \text{← paraxial approx.}$$

∴

$$\frac{dx}{dz}, \frac{dy}{dz} \ll 1$$

and

$$ds \approx dz$$

$$\hat{s} = \frac{dx}{dz} \hat{x} + \frac{dy}{dz} \hat{y} + \hat{z}$$

Ray equation,

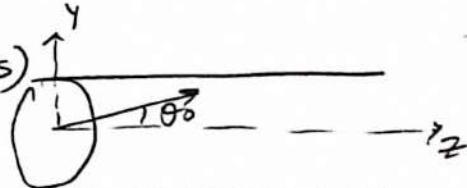
$$\frac{d}{dz}(n\hat{s}) = \frac{d}{dz}\left(n\frac{dx}{dz}\hat{x} + n\frac{dy}{dz}\hat{y} + n\hat{z}\right) = \nabla n$$

n does not vary with z so we have 2 equations,

$$\frac{d}{dz}\left(n\frac{dx}{dz}\right) = \frac{\partial n}{\partial x}$$

$$\frac{d}{dz}\left(n\frac{dy}{dz}\right) = \frac{\partial n}{\partial y}.$$

Meridional Rays



- > Rays that intersect the optical axis "z-axis".
- > Choose ray in yz plane WLOG.
- > Rays that do not intersect the optical axis are called skew rays.
- > Ray always remains in yz plane due to symmetry.

⇒ solve $\frac{d}{dz}\left(n\frac{dy}{dz}\right) = \frac{\partial n}{\partial y}$. subject to initial conditions

$$y(0) = y_0; \frac{dy}{dz}\Big|_0 = \tan\theta_0 \approx \theta_0$$

Assume $n^2(y) = n_0^2(1 - \alpha^2 y^2)$

Approximately,

$$n(y) \approx n_0 \left(1 - \frac{1}{2} \alpha^2 y^2\right)$$

$$\frac{\partial n}{\partial y} = -n_0 \alpha^2 y$$

Ray equation:

$$\frac{d}{dz} \left(n \frac{dy}{dz}\right) = n \frac{d^2 y}{dz^2} = -n_0 \alpha^2 y$$

$$\frac{d^2 y}{dz^2} = -\frac{n_0}{n} \alpha^2 y \approx -\alpha^2 y$$

Since $n \approx n_0$ if $\alpha^2 y^2 \ll 1$

$$\frac{d^2 y}{dz^2} + \alpha^2 y = 0$$

\Rightarrow Solution,

$$y(z) = A \cos \alpha z + B \sin \alpha z$$

$$y(0) = y_0 \Rightarrow A = y_0$$

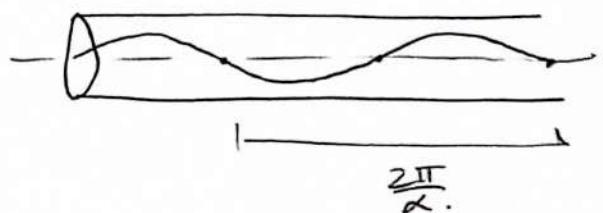
$$\left. \frac{dy}{dz} \right|_0 = \theta_0 \Rightarrow \alpha B = \theta_0 \Rightarrow B = \frac{\theta_0}{\alpha}$$

$$y(z) = y_0 \cos \alpha z + \frac{\theta_0}{\alpha} \sin \alpha z$$

$$\Rightarrow \frac{dy}{dz} = \theta(z) = -y_0 \alpha \sin \alpha z + \frac{\theta_0}{\alpha} \cos \alpha z$$

\Rightarrow Meridional rays are periodic with period or "pitch" = $\frac{2\pi}{\alpha}$

$$y_{\max} = \sqrt{y_0^2 + \left(\frac{\theta_0}{\alpha}\right)^2}$$



Skew rays

$$n^2 = n_0^2(1 - \alpha^2 p^2) = n_0^2(1 - \alpha^2(x^2 + y^2))$$

$$\Rightarrow n \approx n_0 [1 - \frac{1}{2} \alpha^2(x^2 + y^2)]$$

$$\frac{d^2x}{dz^2} = -\alpha^2 x \quad ; \quad \frac{d^2y}{dz^2} = -\alpha^2 y$$

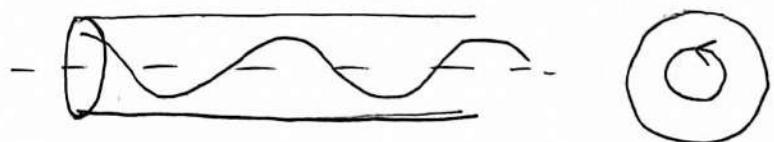
$\Rightarrow x(z)$, $y(z)$ are harmonic functions with period $\frac{2\pi}{\alpha}$

Initial conditions: $(x_0, y_0) \Leftarrow (\theta_{x_0}, \theta_{y_0})$. WLOG choose $x_0 = 0$

$$x(z) = \frac{\theta_{x_0}}{\alpha} \sin \alpha z \quad y(z) = y_0 \cos \alpha z + \frac{\theta_{y_0}}{\alpha} \sin \alpha z.$$

Special case: $\theta_{y_0} = 0$; $\theta_{x_0} = \alpha y_0$.

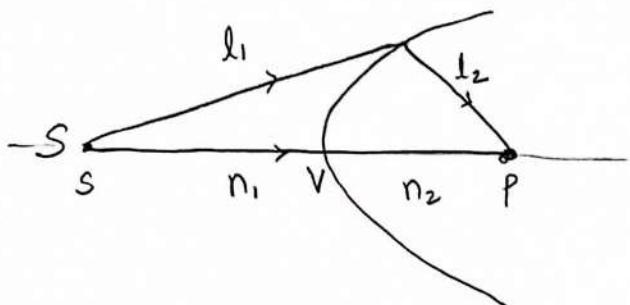
$$\left. \begin{array}{l} x(z) = y_0 \sin \alpha z \\ y(z) = y_0 \cos \alpha z \end{array} \right\} \text{helical ray along } p = y_0 = \text{constant.}$$



Lenses and Imaging Systems.

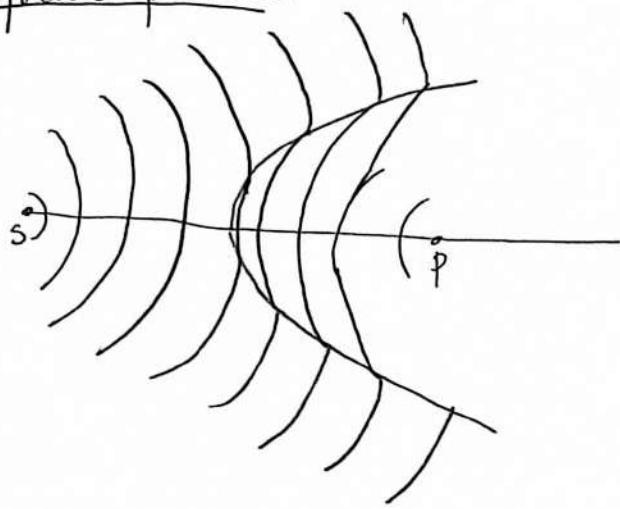
Review of basic GO

Image formation by a refracting surface.



Fermat's Principle $\Rightarrow n_1 l_1 + n_2 l_2 = \text{constant}$
 This is the equation of a "cartesian oval" of revolution.

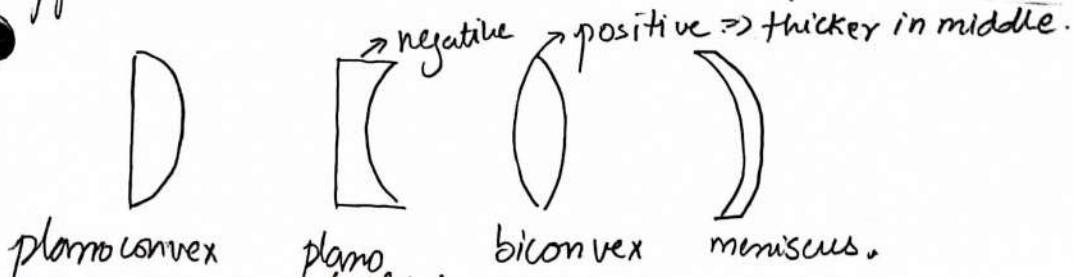
Wavefronts picture



- > When S is at infinity it can be shown that the Cartesian oval reduces to all ellipse -
- > Cartesian ovals are aspherical and cannot be efficiently manufactured so we use spherical lenses instead.

Practical lenses

- > For paraxial rays the cartesian oval can be approximated by a sphere. The imperfections caused due to this approximation are called spherical aberrations.

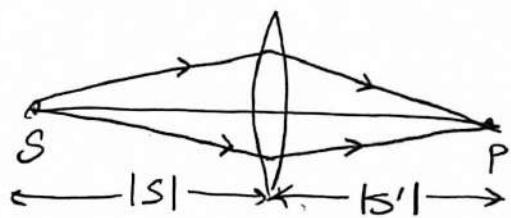


Sign convention.

- i) optical axis = \hat{z} axis
- ii) vertex of surface (V) is positioned at the origin.
- iii) left \Rightarrow -ve ; right \Rightarrow positive.

Thin lens equation.

$$\boxed{\frac{1}{s'} - \frac{1}{s} = \frac{1}{f}}$$

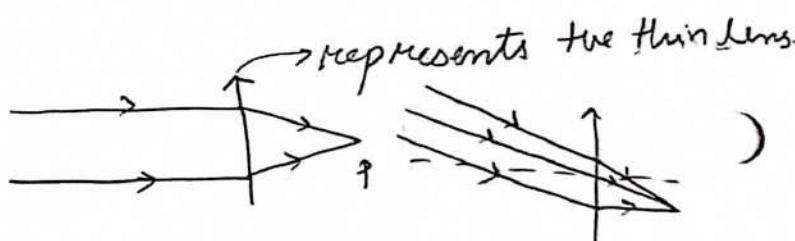


Special cases

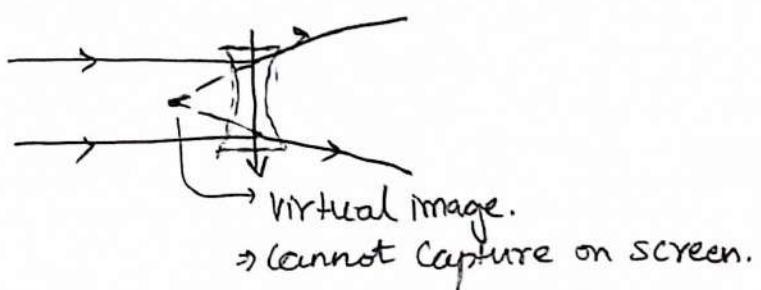
i) $s = -\infty \Rightarrow s' = f$

Similarly $s = -f \Rightarrow s' = \infty$

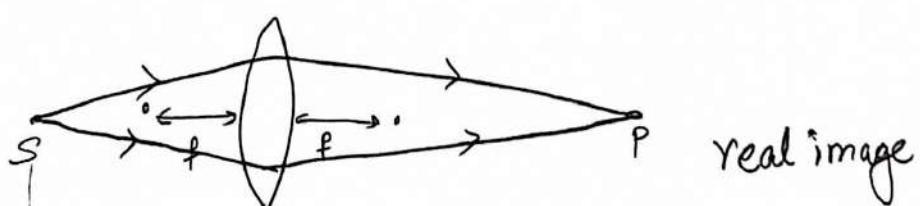
ii) $f < 0 \Rightarrow s' = f < 0$



iii) $s < -f$



a) $f > 0$



b) $f < 0$

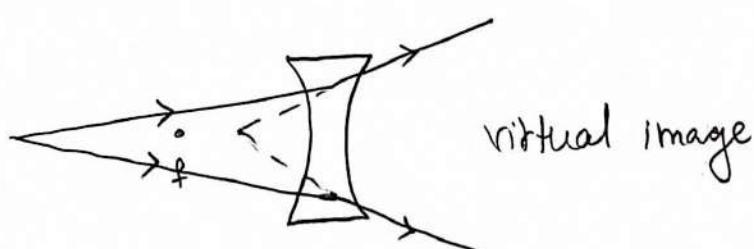
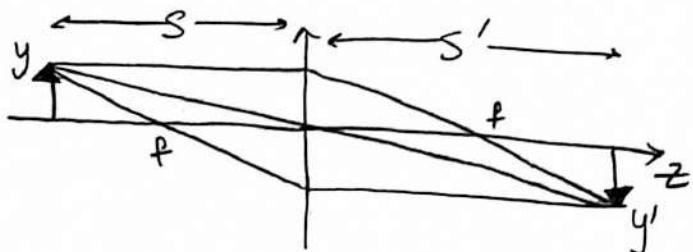


Image formation by ray tracing.

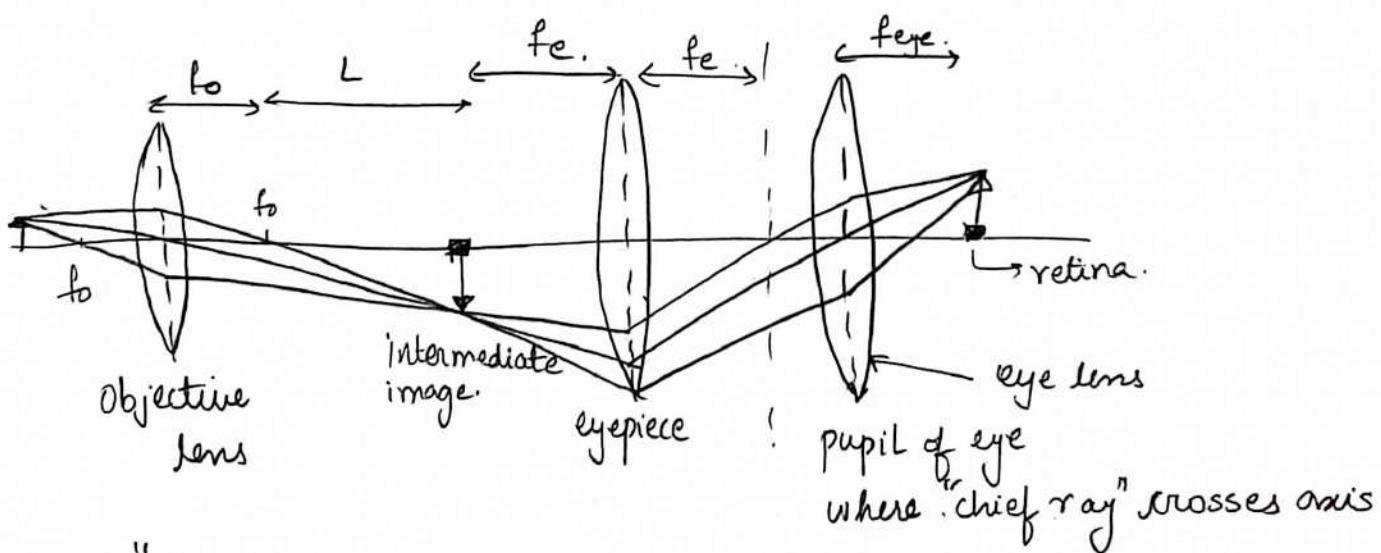
Rules

- Ray through centre of lens is undeviated.
- Collimated rays pass through focus.



$$\text{Magnification} = \frac{y'}{y} < 0 \Rightarrow \text{image inverted & real.}$$

Compound Microscope

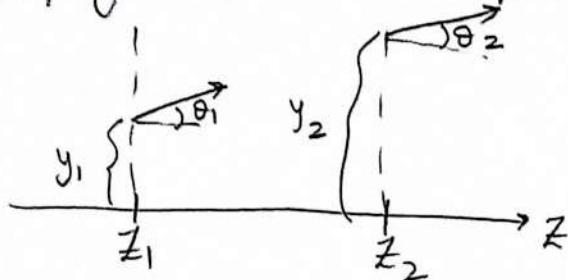


L = "tube length" $\approx 160\text{mm}$ usually

$$\frac{1}{f_{\text{eye}}} = 60 \text{ diopters} \quad (1 \text{ diopter} = 1 \text{ m}^{-1})$$

Paraxial ray tracing and Matrix (ABCD) optics.

1) Propagation in free space.



$$y_2 = y_1 + (z_2 - z_1) \tan \theta$$

$$\Rightarrow \boxed{\begin{aligned} y_2 &\approx y_1 + d\theta \\ \theta_2 &= \theta_1 \end{aligned}} \quad \text{where } d = \text{distance b/w reference planes.}$$

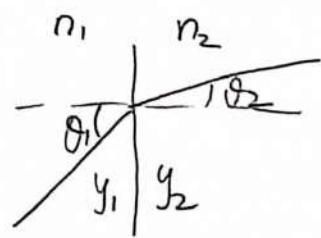
$$\Rightarrow \begin{bmatrix} y_2 \\ \theta_2 \end{bmatrix} = \begin{bmatrix} 1 & d \\ 0 & 1 \end{bmatrix} \begin{bmatrix} y_1 \\ \theta_1 \end{bmatrix} = \mathbf{T}$$

$\begin{bmatrix} y \\ \theta \end{bmatrix}$ is a "ray vector" $\begin{bmatrix} 1 & d \\ 0 & 1 \end{bmatrix}$ is the ray transfer matrix
of free space.

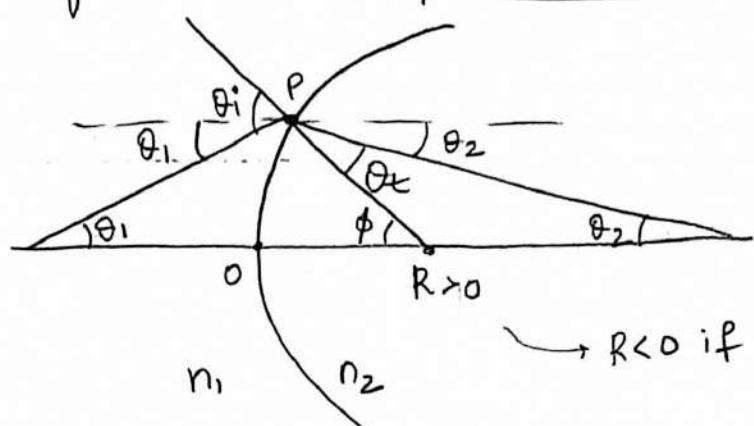
2) Refraction at planar interface

$$y_2 = y_1, \quad n_1 \sin \theta_1 = n_2 \sin \theta_2 \Rightarrow \theta_2 = \frac{n_1}{n_2} \theta_1$$

$$\Rightarrow \boxed{\begin{bmatrix} y_2 \\ \theta_2 \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ 0 & \frac{n_1}{n_2} \end{bmatrix} \begin{bmatrix} y_1 \\ \theta_1 \end{bmatrix}}$$



3. Refraction at spherical surface.



$$n_1 \sin \theta_i = n_2 \sin \theta_t \Rightarrow n_1 \theta_i = n_2 \theta_t$$

Geometry ...

$$\begin{bmatrix} A & B \\ C & D \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ -\frac{(n_2-n_1)}{n_2 R} & \frac{n_1}{n_2} \end{bmatrix}$$

$$\det \begin{bmatrix} AB \\ CD \end{bmatrix} = \frac{n_1}{n_2} \text{ always!}$$

for a paraxial system.

Some books use convention $\begin{bmatrix} y \\ r \theta \end{bmatrix}$ to make $\det = 1$

4. Thin lens.



$$\begin{bmatrix} AB \\ CD \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ -\frac{(n-1)}{R_2} & n \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -\frac{(n-1)}{n R_1} & \frac{1}{n} \end{bmatrix}$$

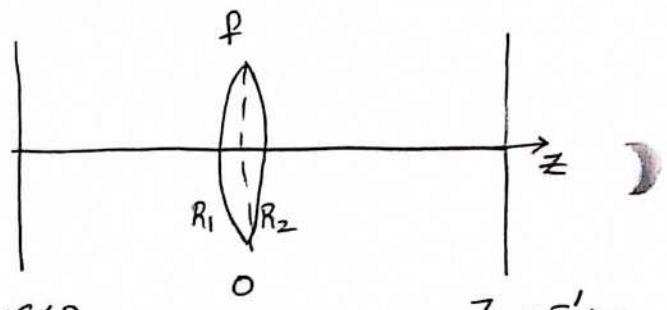
$\xrightarrow{R_2}$ $\xleftarrow{R_1}$

$$\Rightarrow \begin{bmatrix} A & B \\ C & D \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ -(n-1)\left(\frac{1}{R_1} - \frac{1}{R_2}\right) & 1 \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ -\frac{1}{f} & 1 \end{bmatrix} = \overleftrightarrow{S}$$

$R_1 > 0 \quad R_1 > 0, R_2 < 0 \quad R_1 < 0 \quad R_1 < 0, R_2 < 0$

Imaging.

$$\mathcal{T} = \overleftarrow{T_2} \cdot \overleftarrow{S} \cdot \overleftarrow{T_1}$$



$$z_2 = s' > 0$$

$$= \begin{bmatrix} 1 & s' \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -(n-1)\left(\frac{1}{R_1} - \frac{1}{R_2}\right) & 1 \end{bmatrix} \begin{bmatrix} 1 & -s \\ 0 & 1 \end{bmatrix} \quad z = s < 0$$

$$= \begin{bmatrix} 1 - s'(n-1)\left(\frac{1}{R_1} - \frac{1}{R_2}\right) & -s + s' \left\{ 1 + s(n-1)\left(\frac{1}{R_1} - \frac{1}{R_2}\right) \right\} \\ -(n-1)\left(\frac{1}{R_1} - \frac{1}{R_2}\right) & 1 + s(n-1)\left(\frac{1}{R_1} - \frac{1}{R_2}\right) \end{bmatrix}$$

$$= \begin{bmatrix} A & B \\ C & D \end{bmatrix}$$

> For image formation all rays coming from a point y_1 , must

end up at y_2 independant of $\theta_1 \Rightarrow B = 0$.

$$B = 0 \Rightarrow \frac{1}{s'} - \frac{1}{s} = (n-1) \left(\frac{1}{R_1} - \frac{1}{R_2} \right)$$

$$\text{Recall } \frac{1}{s'} - \frac{1}{s} = \frac{1}{f}$$

$$\Rightarrow \frac{1}{f} = (n-1) \left[\frac{1}{R_1} - \frac{1}{R_2} \right] \quad \text{"lens makers equation"}$$

∴ Thin lens

$$\begin{bmatrix} AB \\ CD \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ -\frac{1}{f} & 1 \end{bmatrix}$$

Imaging matrix is now,

$$\begin{bmatrix} AB \\ CD \end{bmatrix} = \begin{bmatrix} 1 - \frac{s'}{f} & 0 \\ -\frac{1}{f} & 1 + \frac{s}{f} \end{bmatrix}$$

$$y_2 = \left(1 - \frac{s'}{f}\right) y_1 \Rightarrow M = \frac{y_2}{y_1} = 1 - \frac{s'}{f}$$

$$\Rightarrow M = 1 - s' \left(\frac{1}{s'} - \frac{1}{s} \right) = \frac{s'}{s}$$

$$\Rightarrow M = \frac{s'}{s} \quad M < 0 \Rightarrow \text{inverted.}$$

$$D = 1 + \frac{s}{f} = \frac{s}{s'} = \frac{1}{M}$$

on axis $\Rightarrow y=0 \Rightarrow \theta_2 = \frac{1}{M} \theta_1 \rightsquigarrow \text{angular demagnification}$

$$\Rightarrow \begin{bmatrix} AB \\ CD \end{bmatrix} = \begin{bmatrix} M & 0 \\ -\frac{1}{f} & \frac{1}{M} \end{bmatrix}$$

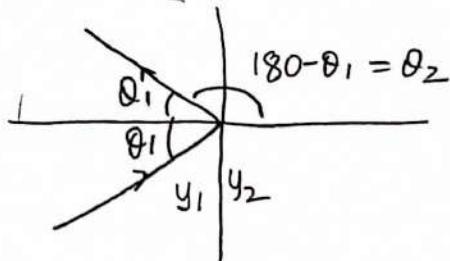
Planar mirror

$$\begin{bmatrix} AB \\ CD \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix} \Rightarrow \text{unchanged. Only direction of propagation changes.}$$

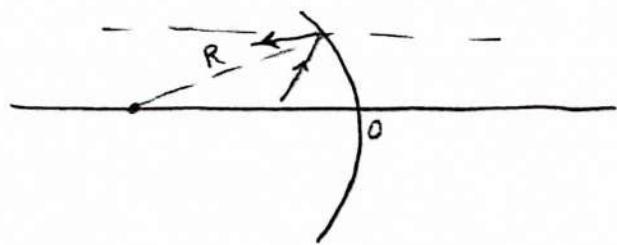
Not sure if this makes sense.

$$\theta_2 = \theta_1 \text{ or } \theta_2 = 180 - \theta_1 ?$$

\rightarrow If $\mathbb{Z} \geq 0$ is defined by direction of propagation, then this is correct!



Spherical Mirror



$$\vec{M} = \begin{bmatrix} 1 & 0 \\ \frac{2}{R} & 0 \end{bmatrix} \quad R < 0 \Rightarrow \text{Concave mirror} \\ R > 0 \Rightarrow \text{Convex mirror}$$

Quadratic GRIN medium.

$$\frac{d^2y}{dz^2} + \alpha^2 y = 0$$

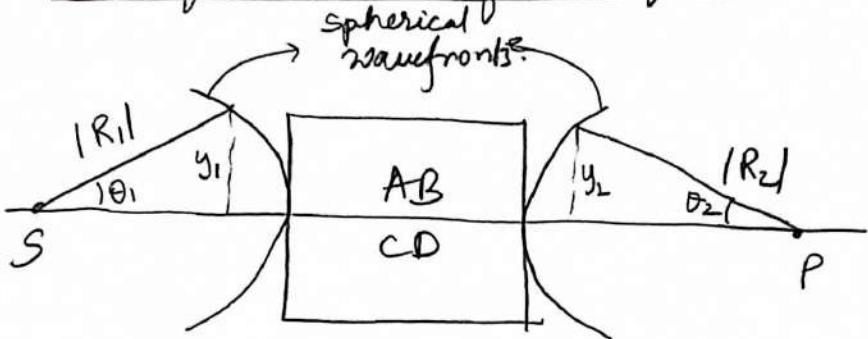
$$y(z) = y_0 \cos \alpha z + \frac{\theta_0}{\alpha} \sin \alpha z$$

$$\theta(z) = -y_0 \alpha \sin \alpha z + \theta_0 \cos \alpha z$$

$$\Rightarrow \begin{bmatrix} AB \\ CD \end{bmatrix} = \begin{bmatrix} \cos \alpha z & \frac{1}{\alpha} \sin \alpha z \\ -\alpha \sin \alpha z & \cos \alpha z \end{bmatrix}$$

(3.2)

Transformation of wavefronts

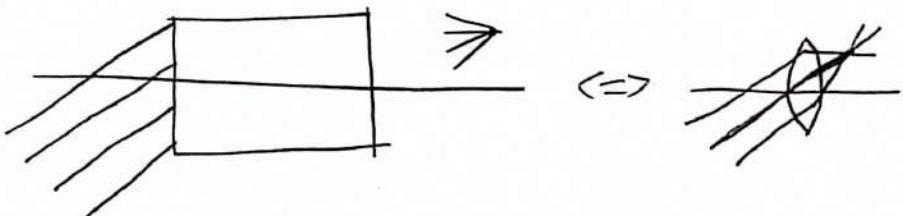


$$R_1 = \frac{y_1}{\theta_1} ; R_2 = \frac{y_2}{\theta_2} = \frac{Ay_1 + B\theta_1}{Cy_1 + D\theta_1}$$

$$\Rightarrow R_2 = \frac{AR_1 + B}{CR_1 + D} \quad \text{ABCO law!}$$

Other properties

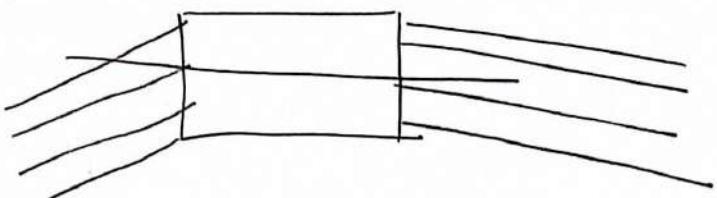
- $B=0 \Rightarrow$ reference planes S, P are conjugate planes
- $A=0 \Rightarrow y_2 = B\theta_1 \Rightarrow$ all rays entering at angle θ_1 end up at y_2



- $C=0 \Rightarrow \theta_2 = D\theta_1$

Telescopic system.

with angular magnification $\frac{\theta_2}{\theta_1}$



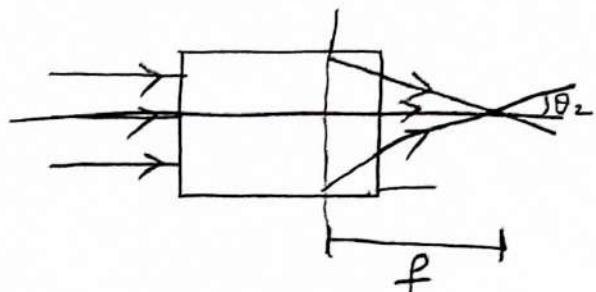
- $D=0$ is reciprocal of $A=0$ case since $\theta_2 = Cy_1$.

General meaning of C

$$C = -\frac{1}{f}$$

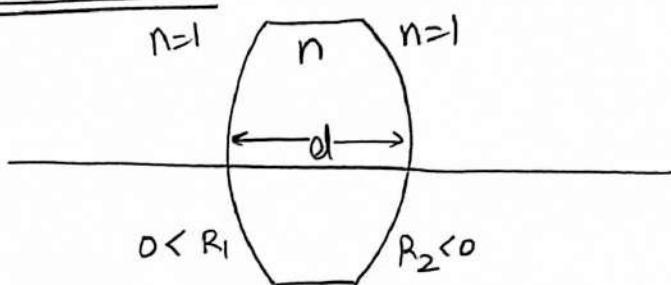
$$\theta_1 = 0$$

$$\theta_2 = Cy_1 \Rightarrow \theta_2 = -\frac{1}{f}y_1$$



⇒ For a complex system find ABCD matrix & $f = \frac{-1}{C}$

Thick lens

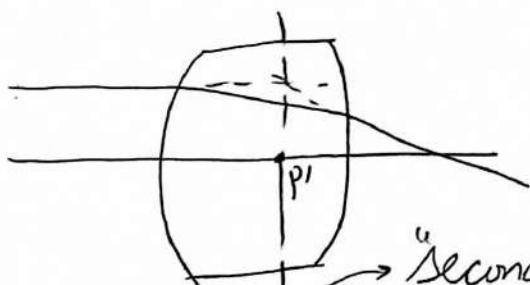


$$\begin{bmatrix} 1 & 0 \\ -\frac{1-n}{R_2} & n \end{bmatrix} \begin{bmatrix} 1 & d \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -\frac{n-1}{nR_1} & \frac{1}{n} \end{bmatrix} = \begin{bmatrix} 1 + d \frac{1-n}{nR_1} & \frac{d}{n} \\ -\frac{1-n}{R_2} \left(1 + d \frac{1-n}{nR_1} \right) + \frac{1-n}{R_1} & 1 - d \frac{1-n}{nR_2} \end{bmatrix}$$

$$C = -\frac{1}{f} = \frac{1-n}{R_1} - \frac{1-n}{R_2} - \frac{(1-n)^2 d}{n R_1 R_2}$$

$$\Rightarrow \frac{1}{f} = (n-1) \left[\left(\frac{1}{R_1} - \frac{1}{R_2} \right) + \frac{(n-1)d}{n R_1 R_2} \right]$$

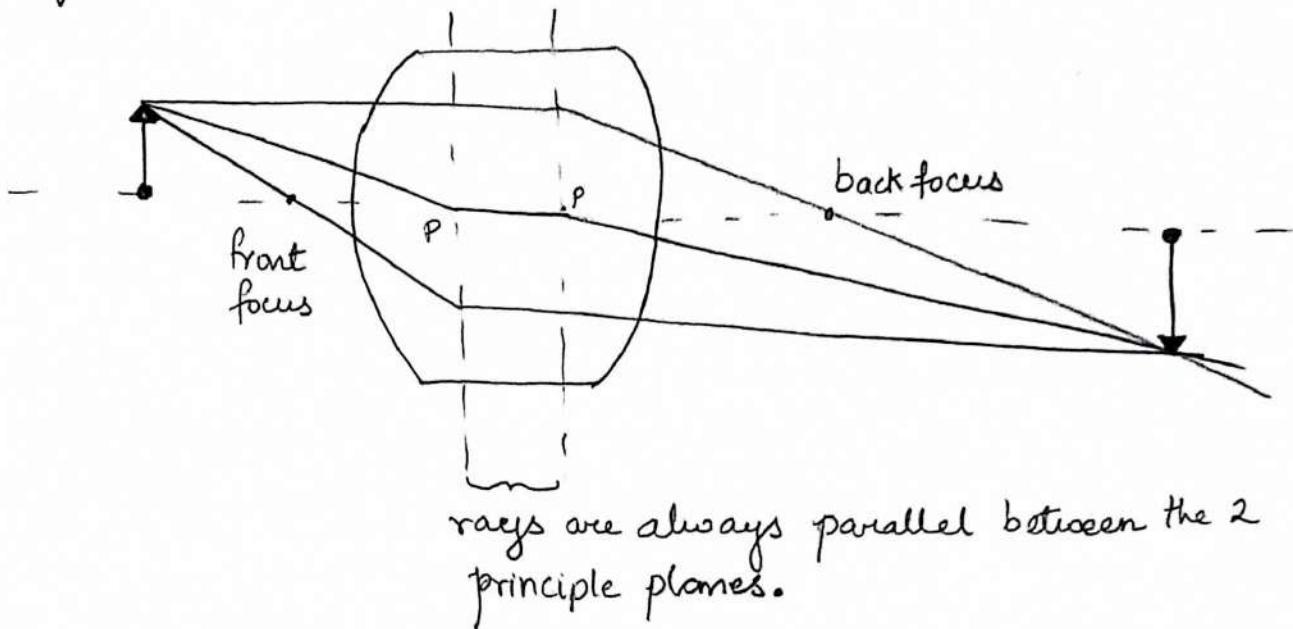
Lens Makers Eqn.



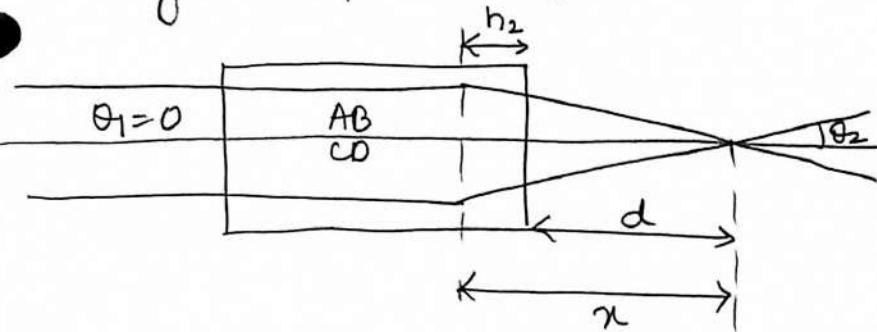
P' → "secondary principle point"

→ "secondary principle plane".
refraction appears to occur here.

> Similarly a primary principle point and plane can be defined.



Finding the principle planes.



$$\begin{bmatrix} 1 & d \\ 0 & 1 \end{bmatrix} \begin{bmatrix} A & B \\ C & D \end{bmatrix} = \begin{bmatrix} A+dc & B+dD \\ C & D \end{bmatrix}$$

$$\theta_1 = 0 \quad y_2 = (A+dc)y_1$$

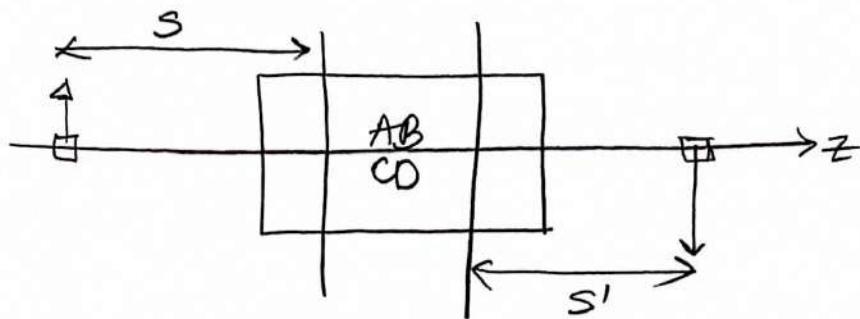
$$\text{at focus } y_2 = 0 \quad A+dc = 0 \quad d = -\frac{A}{C}$$

$$\theta_1 = 0 \quad \theta_2 = Cy_1 \quad ; \quad \tan \theta_2 = \theta_2 = -\frac{y_1}{x}$$

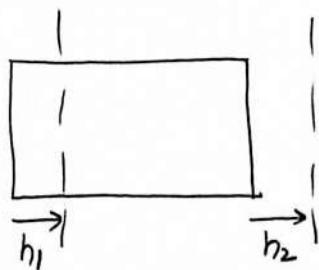
$$u = -\frac{y_1}{\theta_2} = -\frac{y_1}{Cy_1} = -\frac{1}{C} = f \quad ; \quad h_2 = d - u;$$

$$\boxed{h_2 = \frac{1-A}{C}} \quad \boxed{h_1 = \frac{D-1}{C}}$$

Once we have h_1 & h_2 , $\frac{1}{S'} - \frac{1}{S} = \frac{1}{f}$



$$\frac{1}{s'} - \frac{1}{s} = \frac{1}{f}$$



Here $h_1, h_2 > 0$ as drawn.

Propagation between the 2 PRs is given by.

$$\begin{bmatrix} 1 & h_2 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} A & B \\ C & D \end{bmatrix} \begin{bmatrix} 1 & -h_1 \\ 0 & 1 \end{bmatrix}$$

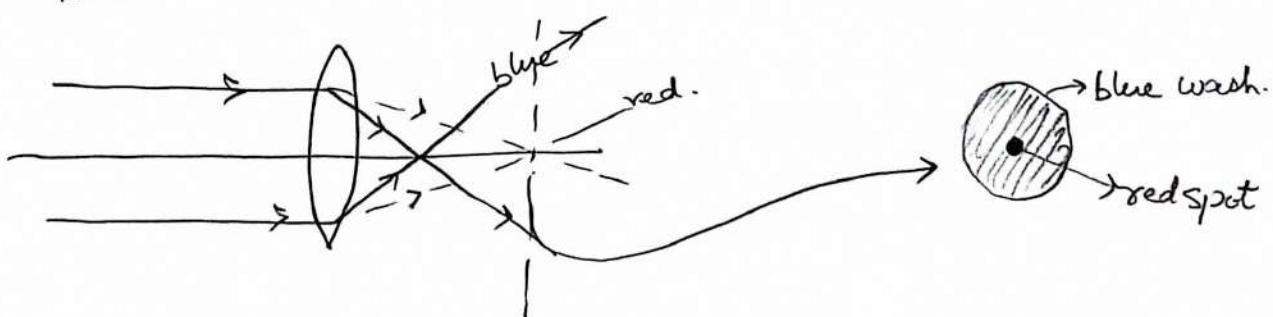
Lens Aberrations

- > Deviations from perfect point to point imaging are called aberrations
 > Chromatic & monochromatic aberrations.

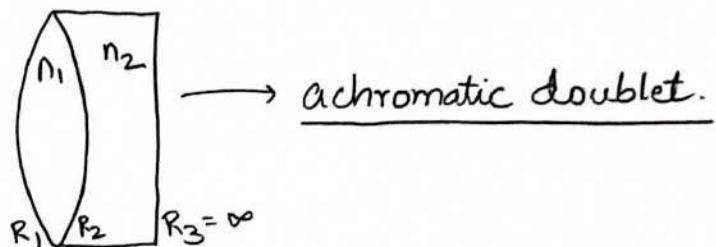
Chromatic aberrations

Recall $f(\lambda)^{-1} = [n(\lambda) - 1] \left[\frac{1}{R_1} - \frac{1}{R_2} \right]$ $\Rightarrow f$ is a fn. of frequency.

Blue bends more than red.



- > An achromat is a combination of two lenses that can achieve equal focal lengths at two wavelengths.



$$AIBCD = \begin{bmatrix} 1 & 0 \\ -\frac{1-n_2}{R_3} & n_2 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -\frac{n_2-n_1}{n_2 R_2} & \frac{n_1}{n_2} \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -\frac{n_1-1}{n_1 R_1} & \frac{1}{n_1} \end{bmatrix}$$

Calculating C,

$$\frac{1}{f} = \frac{n_1-1}{R_1} + \frac{n_2-n_1}{R_2} - \frac{n_2-1}{R_3}$$

Condition for achromaticity,

$$\frac{d}{d\lambda} \left(\frac{1}{f} \right) = 0 \Rightarrow \left(\frac{1}{R_1} - \frac{1}{R_2} \right) \frac{dn_1}{d\lambda} + \left(\frac{1}{R_2} - \frac{1}{R_3} \right) \frac{dn_2}{d\lambda} = 0$$

- > An "apoachromat" uses 3 elements for equal focus at 3 wavelengths.
- > Typically $R_3 \rightarrow \infty$ to reduce cost. R_1 & R_2 provide enough freedom to satisfy the above equation.

Monochromatic Aberrations.

- > Generally a result from breakdown of the paraxial approximation ie $\sin\theta \approx \theta$.
- > Use $\sin\theta \approx \theta - \frac{\theta^3}{3}$ aka. 3rd order aberrations, primary aberrations or Seidal aberrations.

Interferometry

> Coherent addition of waves!

$$\bar{E}_1 = \hat{\epsilon}_1 E_{01} e^{i(\omega_1 t - \vec{k}_1 \cdot \vec{r} + \phi_1)}$$

$$\bar{E}_2 = \hat{\epsilon}_2 E_{02} e^{i(\omega_2 t - \vec{k}_2 \cdot \vec{r} + \phi_2)}$$

$$\bar{E} = \bar{E}_1 + \bar{E}_2$$

Assume it is incident on a square law detector,

$$\langle |E(t)|^2 \rangle = \frac{1}{T} \int_t^{t+T} |E(t)|^2 dt \quad ; T \rightarrow \text{response time of detector.}$$

$$\Rightarrow \langle |E(t)|^2 \rangle = \langle (\bar{E}_1 + \bar{E}_2) \cdot (\bar{E}_1^* + \bar{E}_2^*) \rangle$$

$$= \langle |\bar{E}_1(t)|^2 \rangle + \langle |\bar{E}_2(t)|^2 \rangle$$

$$+ \langle \hat{\epsilon}_1 \cdot \hat{\epsilon}_2^* E_{01} E_{02}^* e^{i(\omega_1 - \omega_2)t - i(\vec{k}_1 - \vec{k}_2) \cdot \vec{r} + i(\phi_1 - \phi_2)} \rangle$$

+ l.c.

> Note that interference requires the 2 polarizations to be non orthogonal. ie. $\hat{\epsilon}_1 \cdot \hat{\epsilon}_2^* \neq 0$.

> If $T < \frac{2\pi}{|\omega_1 - \omega_2|}$ $\Rightarrow \langle I(t) \rangle$ will exhibit temporal beats at $\omega_1 - \omega_2$

> If $T > \frac{2\pi}{|\omega_1 - \omega_2|}$ $\Rightarrow \langle e^{i(\omega_1 - \omega_2)t} \rangle = 0$ unless $\omega_1 = \omega_2$

Therefore interference only occurs when $\omega_1 = \omega_2$ (if T is large). It is in this sense that we say that waves need to have the same frequency to interfere.)

Coherence: frequency & waveform are identical.
 ↳ (constant relative phase)

From here on, $\omega_1 = \omega_2 = \omega$.

$$\Rightarrow \langle |E(t)|^2 \rangle = \langle |E_{01}|^2 \rangle + \langle |E_{02}|^2 \rangle + \langle E_{01} E_{02}^* e^{-i(\bar{k}_1 - \bar{k}_2) \cdot \bar{r} + i(\phi_1 - \phi_2)} + cc \rangle$$

or

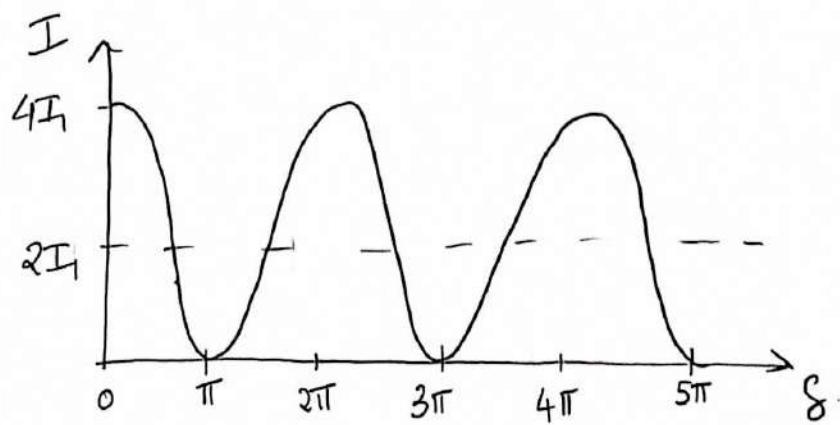
$$I(\bar{r}) = I_1 + I_2 + \sqrt{I_1 I_2} \left[e^{i(\bar{k}_1 - \bar{k}_2) \cdot \bar{r} + i(\phi_1 - \phi_2)} + cc \right]$$

Let $\delta = (\bar{k}_1 - \bar{k}_2) \cdot \bar{r} + (\phi_1 - \phi_2) \rightarrow$ phase diff. b/w the 2 waves
 (independant of t)

$$\Rightarrow \boxed{I(\bar{r}) = I_1 + I_2 + 2\sqrt{I_1 I_2} \cos \delta.}$$

$\delta = 2n\pi \Rightarrow$ constructive.
 $\delta = (2n+1)\pi \Rightarrow$ destructive.

$$\text{If } I_1 > I_2 \Rightarrow I = 4I_1 \cos^2 \frac{\delta}{2}$$



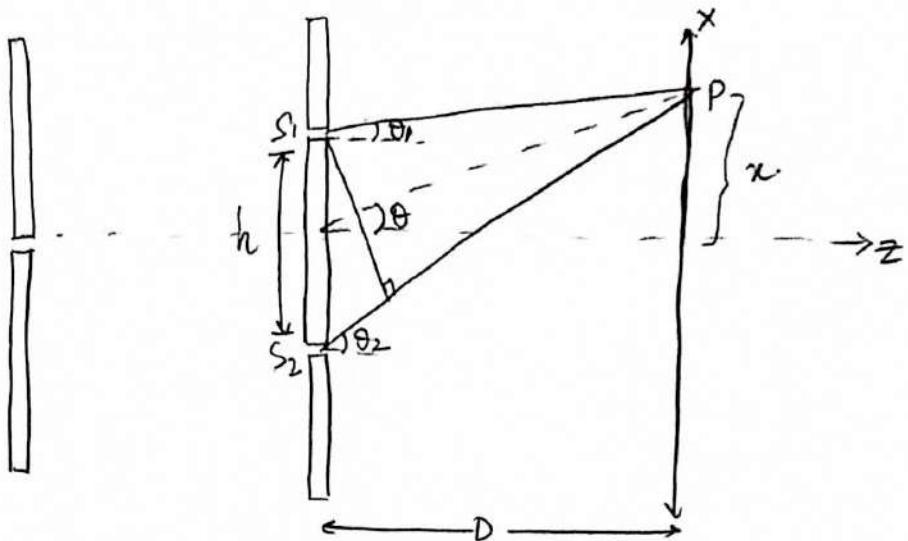
Usually $\phi_1 = \phi_2 \Rightarrow$ same initial phase.
 $\Rightarrow \delta = \bar{k}_1 \cdot \bar{r} - \bar{k}_2 \cdot \bar{r}$

In general if the 2 rays follow paths C_1 & C_2 we have.

$$\delta = \int_{C_1} \bar{k}_1(\bar{r}) \cdot d\bar{s} - \int_{C_2} \bar{k}_2(\bar{r}) \cdot d\bar{s}$$

Young's Interference

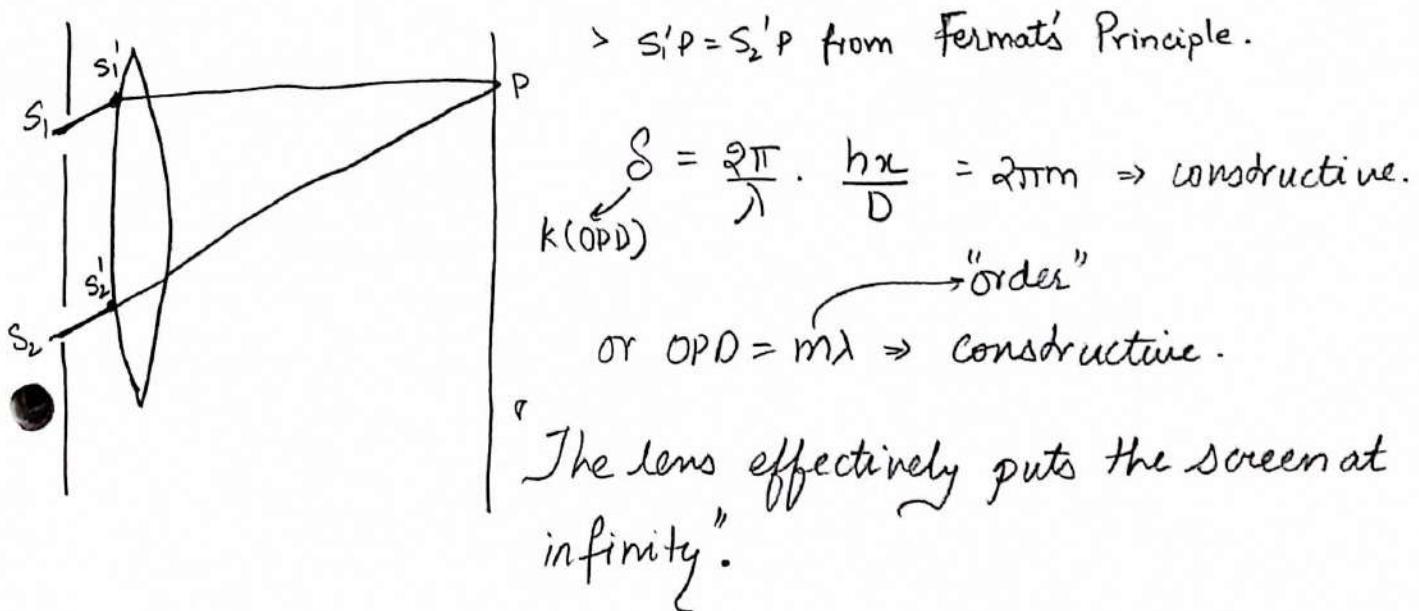
Double Slit Interference



If \$D \gg h\$, \$\theta_1 \approx \theta_2 \equiv \theta\$

$$OPD = \overline{S_2 P} - \overline{S_1 P} = h \sin \theta \approx h \theta. \quad ; \quad \theta = \tan \theta = \frac{x}{D}$$

Instead of assuming \$D \gg h\$, place a lens after the double slit & make \$D = f\$. Now \$OPD = h \sin \theta \propto \theta_1 = \theta_2 = \theta\$



Intensity Pattern.

> Some intensity and phase at both slits implies

$$I = 4I_1 \cos^2 \frac{\delta}{2} = 4I_1 \cos^2 \left(\pi \frac{Lx}{\lambda D} \right)$$

The maxima are spread out by an amount,

$$|\Delta x| = |m| \frac{D}{h} \Delta \lambda_0.$$

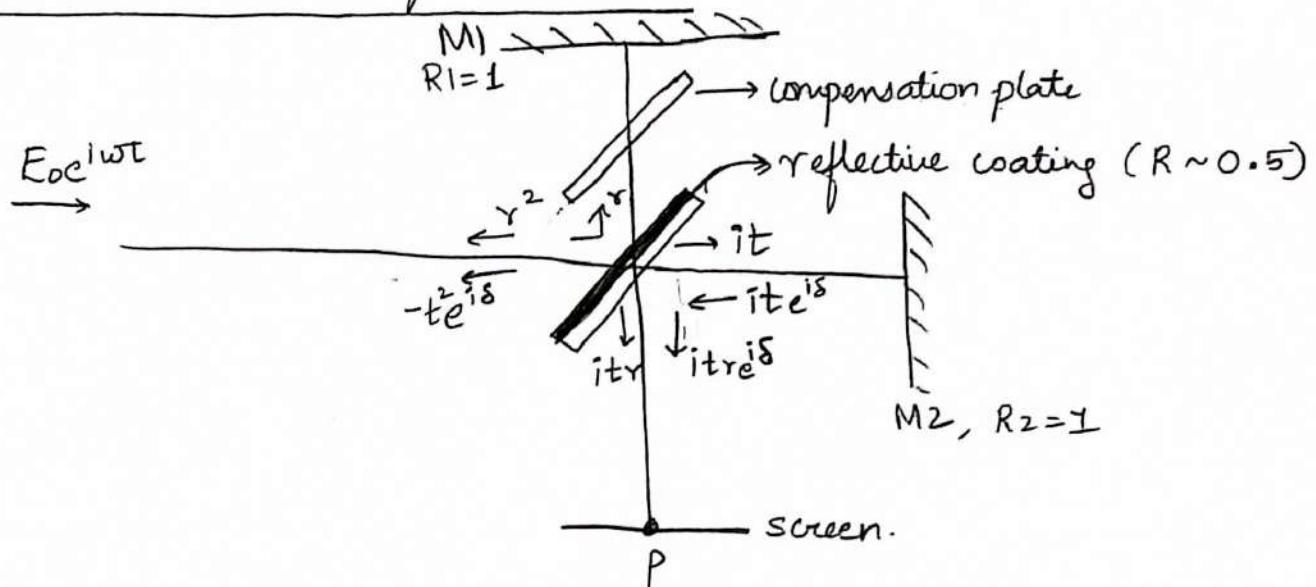
mean spacing b/w fringes $|\Delta x| = \frac{D}{h} \lambda_0.$

⇒ If $|m| < \frac{\lambda_0}{\Delta \lambda}$, the spread is negligible.

→ This experiment is an example of "interference by division of wavefronts".

Michelson Interferometer

4.5



- > For a lossless dielectric mirror, power conservation and time reversal symmetry give a scattering matrix as follows,

$$S = \begin{bmatrix} r & it \\ it & r \end{bmatrix} \quad t = \sqrt{1-r^2}$$

- > The reflected & transmitted waves are at 90° wrt each other.

$$> \delta = \frac{2\pi}{\lambda} 2d = \frac{2\omega}{c} d = \frac{2d}{c} \tau ; \text{ } d \text{ is excess distance of M2 wrt M1.}$$

- > $\tau = \frac{2d}{c}$ is the optical delay

$$\begin{aligned} E_R &= r(rE_0) + it(lite^{is}E_0) \\ &= [r^2 - (1-r^2)e^{is}]E_0 \\ &= 0 \text{ (if } \delta = 0 \text{ and } r^2 = 0.5) \end{aligned}$$

$$> E_T = r(itE_0) + it e^{i\delta} (rE_0) \\ = itr(1 + e^{i\delta}) E_0$$

If $r = t = \frac{1}{\sqrt{2}}$ & $\delta = 0$ we have $E_T = iE_0$.

$$> |E_T|^2 = |itr(1 + e^{i\delta})|^2 |E_0|^2$$

$$\Rightarrow I_t = |r|^2 |t|^2 |1 + e^{i\delta}|^2 I_0 \\ = RT (1 + e^{i\delta})(1 + e^{-i\delta}) I_0$$

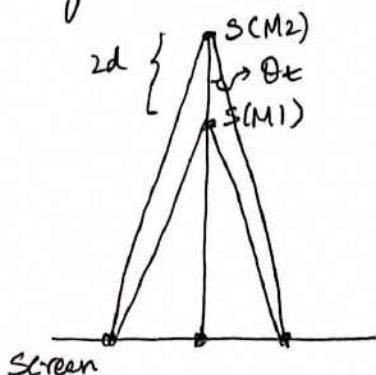
$$I_t = 2RT (1 + \cos \omega T) I_0$$

$$R = |r|^2; T = |t|^2$$

If $R = T = 0.5$;

$$I_t = \frac{1}{2} (1 + \cos \omega T) I_0$$

> Intensity pattern on screen is of concentric rings since source gives spherical waves. The image of S from M1 & M2 is as follows.



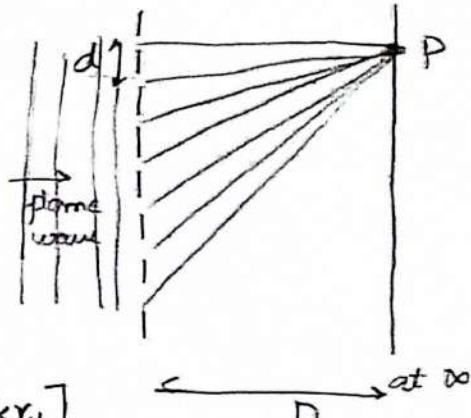
$$\Rightarrow \delta = \frac{2\pi}{\lambda} \cdot 2d \cos \theta_L = 2\pi m \Rightarrow \text{constructive.}$$

Interference of N-waves

Division of wavefront $\Leftrightarrow N$ slits

Total field at P,

$$E = E_0(r_1) e^{i(\omega t - kr_1)} + E_0(r_2) e^{i(\omega t - kr_2)} + \dots + E_0(r_N) e^{i(\omega t - kr_N)}$$



$$\simeq E_0(D) e^{i\omega t} \left[e^{-ikr_1} + e^{-ikr_2} + \dots + e^{-ikr_N} \right]$$

(Assume $E_0(r_1) = E_0(r_2) = \dots = E_0(D)$)

$$\Rightarrow E = E_0(D) e^{i(\omega t - kr_1)} \left[1 + e^{iK(r_2 - r_1)} + \dots + e^{-iK(r_N - r_1)} \right]$$

$$r_3 - r_1 = 2(r_2 - r_1); \quad r_4 - r_1 = 3(r_2 - r_1); \dots; \quad r_N - r_1 = (N-1)(r_2 - r_1)$$

We know $r_2 - r_1 = ds \sin \theta = \frac{s}{K}$ \rightarrow phase diff. b/w adjacent slits

$$\Rightarrow E = E_0 e^{i(\omega t - kr_1)} \left[1 + e^{-is} + (e^{-is})^2 + \dots + (e^{-is})^{N-1} \right]$$

$$= E_0 e^{i(\omega t - kr_1)} \begin{bmatrix} e^{-iNs} & -1 \\ e^{-is} & -1 \end{bmatrix} = E_0 e^{i(\omega t - kr_1)} \frac{\begin{bmatrix} e^{-iNs} & (e^{-iNs} - e^{iNs}) \\ e^{-is} & (e^{-is} - e^{is}) \end{bmatrix}}{\begin{bmatrix} e^{-is} & (e^{-is} - e^{is}) \\ e^{-is} & (e^{-is} - e^{is}) \end{bmatrix}}$$

$$= E_0 e^{i[\omega t - kr_1 - (N-1)s/2]} \begin{bmatrix} \sin \frac{Ns}{2} \\ \sin \frac{s}{2} \end{bmatrix}$$



Note, $D = r_1 + \frac{1}{2} (N-1) d \sin\theta$.

$$\Rightarrow E = E_0(D) e^{i(\omega t - kD)} \left[\frac{\sin \frac{N\delta}{2}}{\sin \frac{\delta}{2}} \right]$$

plane wave modulated by $\left[\quad \right]$ term.
(Similar to array factor from Antennas)

$$I = I_0 \frac{\sin^2 \left[\frac{Nkds\sin\theta}{2} \right]}{\sin^2 \left[\frac{kds\sin\theta}{2} \right]}$$

or

$$I = I_0 \left[\frac{\sin^2 \left(N\pi \frac{ds\sin\theta}{\lambda} \right)}{\sin^2 \left(\pi \frac{ds\sin\theta}{\lambda} \right)} \right]$$

oscillates much faster than the denominator.

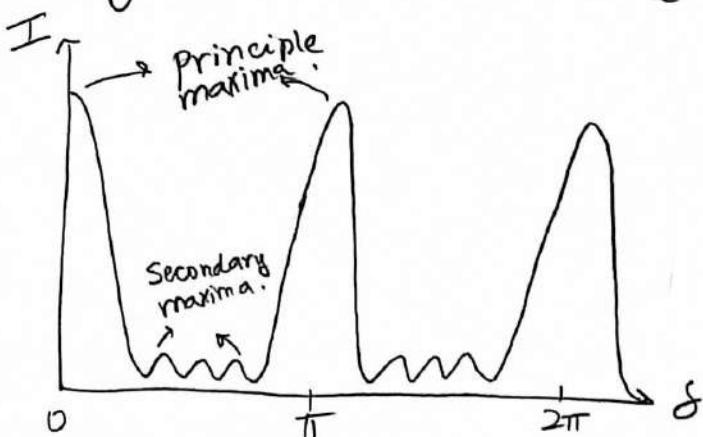
$$\delta = kds\sin\theta$$

As $\theta \rightarrow 0$ (ie $m=0$) $\Rightarrow \sin\theta = \theta$,

$$\Rightarrow I = I_0 N^2$$

\Rightarrow When $ds\sin\theta = m\lambda$ we have principle maxima.
grating equation with order m .

Secondary maxima occur at $\frac{N\delta}{2} = \frac{3\pi}{2}, \frac{5\pi}{2}, \dots$



$N \gg 1 \Rightarrow$ secondary maxima are small & principle maxima are very narrow.

$$\Delta\theta = \frac{\lambda}{Nd\cos\theta}$$

Angular width.

Fabry Perot Etalons and Interferometers.

Division of Amplitude

FP Etalon: A dielectric slab with parallel faces (may be coated)

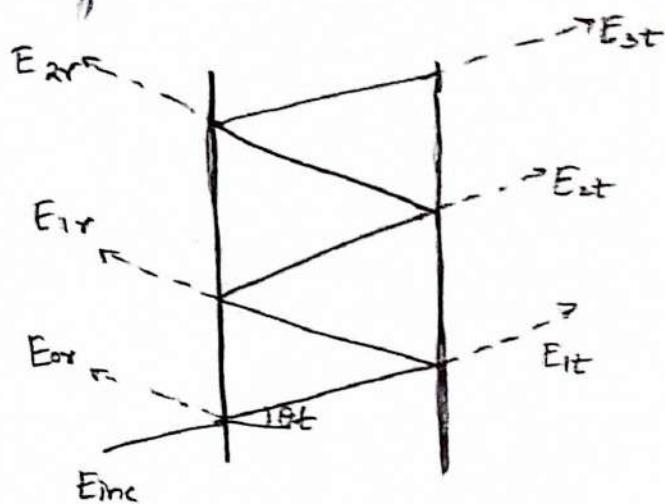
FP Interferometer: Two highly reflecting parallel mirrors.

Approach I: Follow the partial reflections.

$$t_{1,2} = \sqrt{1 - r_{1,2}^2}$$

> Even with 99.99% reflectivity,
at some "resonant frequencies"
the reflections inside build
up in amplitude to a large
steady state field & then

E_{1r}, E_{2r}, \dots add up to cancel E_{0r} . Which gives perfect transmission.



1) incident wave $\rightarrow E_{inc}$

2) one transmission $\rightarrow i_1, E_{inc}$.

3) propagation b/w mirrors $\rightarrow i_1, e^{-i\delta_1} E_{inc}$.

4) transmission through second mirror $\rightarrow (i_1)(i_2) e^{-i\delta_2} E_{inc}$.
 $= -(t_1 t_2) e^{-i\delta_2} E_{inc}$.

$$\text{Where } \frac{\delta}{2} = \frac{n \omega d}{c} \cos \theta_t$$

$$\therefore E_{2t} = (r_1 r_2 e^{i\delta}) (-t_1 t_2 e^{-i\delta_2}) E_{inc}$$

$$E_{3t} = (r_1 r_2 e^{i\delta})^2 (-t_1 t_2 e^{-i\delta_2}) E_{inc} \quad \& \text{ so on.}$$

$$\Rightarrow E_{\text{trans}} = \left[\sum_{l=0}^{\infty} (\gamma_1 \gamma_2 e^{i\delta})^l \right] (-t_1 t_2 e^{-i\delta_1}) E_{\text{inc.}}$$

$\gamma_1 < 1$ & $\gamma_2 < 1$ so we set $x = \gamma_1 \gamma_2 e^{-i\delta}$ \approx note $|x| < 1$

$$\Rightarrow \sum_{l=0}^{\infty} x^l = \frac{1}{1-x}$$

$$\Rightarrow E_{\text{trans}} = \frac{-t_1 t_2 e^{-i\delta_1}}{1 - \gamma_1 \gamma_2 e^{i\delta}} E_{\text{inc.}}$$

Similarly,

$$E_{\text{or}} = \gamma_1 E_{\text{inc}}$$

$$E_{1r} = (it_1) \gamma_2 (it_1) e^{-i\delta} E_{\text{inc.}}$$

$$E_{2r} = \gamma_1 \gamma_2 e^{-i\delta} E_{1r}$$

⋮

$$\Rightarrow E_{\text{refl}} = \left[\sum_{l=0}^{\infty} (\gamma_1 \gamma_2 e^{-i\delta})^l \right] (-t_1^2 \gamma_2 e^{-i\delta}) E_{\text{inc.}} + \gamma_1 E_{\text{inc.}}$$

$$\Rightarrow E_{\text{refl}} = \left[\gamma_1 - \frac{t_1^2 \gamma_2 e^{-i\delta}}{1 - \gamma_1 \gamma_2 e^{-i\delta}} \right] E_{\text{inc.}}$$

Approach II: Scattering & Transfer matrices.

$$S_{1,2} = \begin{bmatrix} \gamma_{1,2} & it_{1,2} \\ it_{1,2} & \gamma_{1,2} \end{bmatrix} ; M_{1,2} = \frac{1}{it_{1,2}} \begin{bmatrix} -1 & \gamma_{1,2} \\ -\gamma_{1,2} & 1 \end{bmatrix}$$

$$\Rightarrow M = \left\{ \left(\frac{1}{it_2} \right) \begin{bmatrix} -1 & \gamma_2 \\ -\gamma_2 & 1 \end{bmatrix} \right\} \begin{bmatrix} e^{-i\delta_2} & 0 \\ 0 & e^{i\delta_2} \end{bmatrix} \left\{ \left(\frac{1}{it_2} \right) \begin{bmatrix} -1 & \gamma_1 \\ -\gamma_1 & 1 \end{bmatrix} \right\}$$

$$\Rightarrow M_{22} = \frac{-1}{t_1 t_2} \left[-\gamma_1 \gamma_2 e^{-i\delta_2} + e^{i\delta_2} \right]$$

$S_{21} = S_{12} = t_{21} = \frac{1}{M_{22}}$ which gives the same results as Approach I

Intensity

$$I_{\text{trans}} = \left| \frac{-t_1 t_2 e^{-i\delta_2}}{1 - \gamma_1 \gamma_2 e^{-i\delta}} \right|^2 I_{\text{inc.}}$$

Consider $\gamma_1 = \gamma_2 = \gamma$; $t_1 = t_2 = t = \sqrt{1 - \gamma^2}$

$$\Rightarrow I_{\text{trans}} = \frac{t^4}{|1 - \gamma^2 e^{-i\delta}|^2} I_{\text{inc.}}$$

$$|1 - \gamma^2 e^{-i\delta}|^2 = (1 - R e^{-i\delta})(1 - R e^{i\delta}) = 1 - 2R \cos \delta + R^2$$

$$= 1 - 2R(1 - 2 \sin^2 \frac{\delta}{2}) + R^2$$

$$= (1 - R)^2 + 4R \sin^2 \frac{\delta}{2}$$

where $R = \gamma^2$

$$\Rightarrow I_{\text{trans}} = \frac{T^2}{(1-R)^2 \left[1 + \frac{4R}{(1-R)^2} \sin^2 \frac{\delta}{2} \right]} I_{\text{inc.}}$$

Define "contrast" as

$$F = \frac{4R}{(1-R)^2}$$

; note $\frac{T^2}{1-R^2} = 1$
 $\Rightarrow R+T=1$

$$\Rightarrow \frac{I_{\text{trans}}}{I_{\text{inc}}} = \frac{1}{1+F \sin^2 \frac{\delta}{2}} \rightarrow \text{Airy function.}$$

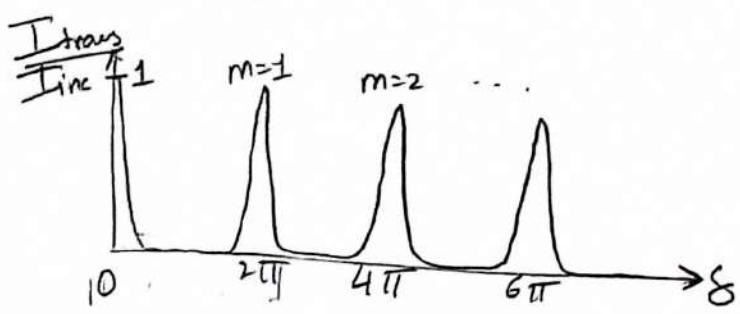
F is also called "coefficient of finesse."

> When $\delta = \frac{2\pi d w}{c} \cos \theta_t = 2n\pi$ we have $\sin^2 \frac{\delta}{2} = 0$

$$\Rightarrow I_{\text{trans}} = I_{\text{inc}} !$$

$$> \text{When } \delta = (2n+1)\pi ; \frac{I_{\text{trans}}}{I_{\text{inc}}} = \frac{1}{1+F}$$

> If R is very close to 1, F is very large. $\Rightarrow I_{\text{trans}} \ll I_{\text{inc.}}$



$$\text{Define } L = nd \cos \theta_t$$

\Rightarrow Separation b/w 2 orders ($\Delta m=1$) is called the "free spectral range".

\Rightarrow aka axial mode spacing.

$$\Delta \omega_{\text{FSR}} = 2\pi \frac{C}{2L}$$

$$\frac{2 \Delta \omega_{\text{FSR}}}{C} L = 2\pi$$

$$\Delta \nu_{\text{FSR}} = \frac{C}{2L}$$

> Full width at half maximum,

- ie. $\frac{1}{1+F \sin^2 \frac{\delta}{2}} = \frac{1}{2}$

$$\Rightarrow \boxed{\delta_{1/2} = \pm 2 \sin^{-1} \left(\frac{1}{\sqrt{F}} \right)}$$

When $F \gg 1$ $\sin \frac{\delta}{2} = \frac{\delta}{2} = \pm \frac{1}{\sqrt{F}}$

$$\Rightarrow \boxed{\Delta \delta_{1/2} \approx \frac{4}{\sqrt{F}}}$$

Recall $\delta = \frac{2\omega}{c} n d \cos \theta_t = \frac{2\omega}{c} L = 2\pi \nu \frac{2L}{c}$

- $\Delta \delta = 2\pi \Delta \nu \cdot \frac{2L}{c}$

$$\Rightarrow \Delta \nu_{1/2} = \frac{\Delta \delta_{1/2}}{2\pi} \cdot \frac{c}{2L} = \frac{2}{\pi \sqrt{F}} \Delta \nu_{\text{FSR}}$$

Define Finesse \mathcal{F} by $\boxed{\mathcal{F} = \frac{\pi}{2} \sqrt{F}}$

$$\Rightarrow \boxed{\frac{\Delta \nu_{1/2}}{\Delta \nu_{\text{FSR}}} = \mathcal{F}^{-1}}$$

→ With large F (\mathcal{F}) we have $\Delta \nu_{\text{FSR}} \gg \Delta \nu_{1/2}$ & this is useful in doing spectroscopy.

Diffraction

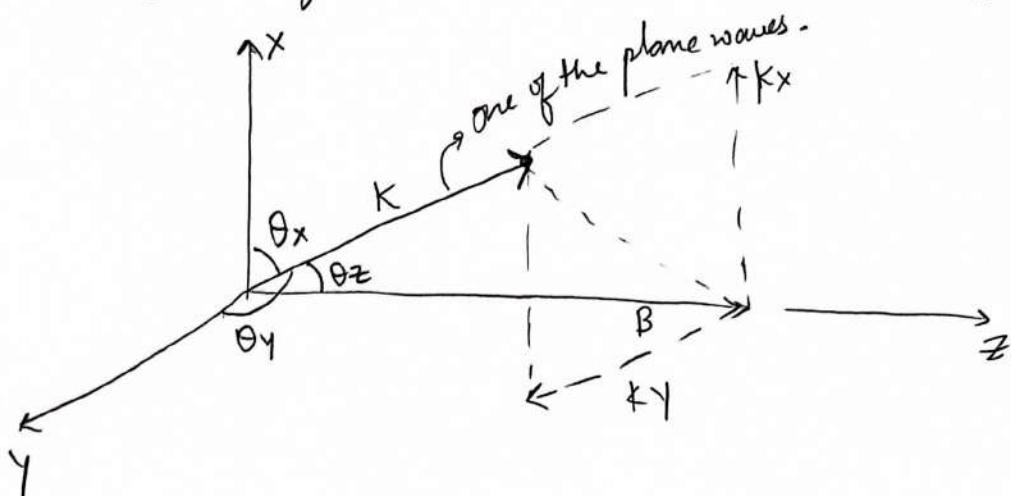
- > Confined light (or any wave) eventually spreads.
- > We consider only scalar diffraction theory \Rightarrow scalar fields.

Recall Angular spectrum

$$E_A(\vec{r}) = \iint_{-\infty}^{\infty} A(k_x, k_y) e^{-i(k_x x + k_y y + kz)} dk_x dk_y$$

- We break down the field into plane waves propagating at different angles. Each plane wave has a value for a given k_x, k_y called the angular spectrum (ie. $A(k_x, k_y)$)
- $\vec{k} = k_x \hat{x} + k_y \hat{y} + k_z \hat{z}$.

- Therefore for a given \vec{r} , the value of E can be written as a sum of plane waves in different directions (with some \vec{k}). The "weight" of each direction is given by $A(k_x, k_y)$



$$\beta = \sqrt{k^2 - k_x^2 - k_y^2}$$

$k_x \rightarrow$ spatial frequency along x direction.

$\Rightarrow \theta_x = \cos^{-1}\left(\frac{k_x}{k}\right)$ gives prop. angle wrt x axis

If $k_x^2 + k_y^2 > \left(\frac{2\pi}{\lambda}\right)^2 \Rightarrow \beta$ is imaginary \Rightarrow evanescent mode.

> Therefore there is an upper limit on the spatial frequency that can propagate. Higher frequencies (spatial) that capture variations that are smaller are evanescent!

> Such an angular spectrum decomposition directly hints at diffraction since we have plane waves going in all directions. (Unless we started with a plane wave).

> Once we know $A(k_x, k_y)$ at z_1 , ie $A(k_x, k_y; z_1)$ it is trivial to find $A(k_x, k_y; z_2)$. Let $z_1=0$, $z_2=z$. Then we have

$$A(k_x, k_y; z) = A(k_x, k_y; 0) e^{-i\beta z}$$

> Let $H(k_x, k_y; z) = e^{-i\beta z}$ be the transfer fn. of the angular spectrum.

(5.3)

Therefore, propagation in free space amounts to multiplication of angular spectrum by the transfer function

$$A(k_x, k_y; o) = \frac{1}{4\pi^2} \iint_{-\infty}^{\infty} E(x', y'; o) e^{i(k_x x' + k_y y')} dx' dy'$$

$$\begin{aligned} E(x, y; z) &= \iint_{-\infty}^{\infty} A(k_x, k_y; o) H(k_x, k_y; z) e^{-i(k_x x + k_y y)} dk_x dk_y \\ &= \frac{1}{4\pi^2} \iint_{-\infty}^{\infty} E(x', y'; o) \times \left[\iint_{-\infty}^{\infty} H(k_x, k_y; z) e^{-i[k_x(x-x') + k_y(y-y')]} dk_x dk_y \right] dx' dy' \end{aligned}$$

The term in brackets looks like an inverse Fourier Transform & indeed it is the "spatial impulse response"

$$h(x, y; z) = \frac{1}{4\pi^2} \iint_{-\infty}^{\infty} H(k_x, k_y; z) e^{-i(k_x x + k_y y)} dk_x dk_y$$

The fields then are the convolution integral.

$$E(x, y; z) = \iint_{-\infty}^{\infty} E(x', y'; o) h(x-x'; y-y'; z) dx' dy'$$

$E(x, y; z) = E(x', y'; o) \otimes h(x, y; z)$

Fresnel Approximation : Propagation of Paraxial Waves.

$$> H(k_x, k_y; z) = e^{-i\beta z} \quad \text{where } \beta = \sqrt{k_x^2 + k_y^2 - k^2}$$

Assuming waves propagate close to z axis we have,

$$\beta = k \sqrt{1 - \frac{k_x^2}{k^2} - \frac{k_y^2}{k^2}} \approx k \left[1 - \frac{1}{2} \left(\frac{k_x^2 + k_y^2}{k^2} \right) \right]$$

So,

$$\beta \approx k - \frac{k_x^2 + k_y^2}{2k}$$

$\sin\theta \approx \theta$ comes from the fact that
 $\frac{k_x}{k} \ll 1 \Rightarrow \frac{k_x}{k} = \sin\theta \ll 1$!

$$\Rightarrow H(k_x, k_y; z) = e^{-i\beta z} \approx e^{-ikz} e^{i(K_x^2 + K_y^2)z/2k}$$

Impulse response: $h(x, y; z) \approx \frac{e^{-ikz}}{i\lambda z} e^{-ik(x^2 + y^2)/2z}$

Conclusion

i) Each spatial frequency component acquires a quadratic phase $(K_x^2 + K_y^2)z/2k$ on propagating over a distance z .

ii) The initial field is convolved by a quadratic phase $k(x^2 + y^2)/2z$.

In either case, the phase accumulation is quadratic

Gaussian Beams

Assume $E(x, y; 0) = E_0 e^{-\frac{(x^2+y^2)}{W_0^2}}$ where $W_0 \rightarrow$ beam waist
ie the radius where amplitude drops by $1/e$.

$$A(k_x, k_y; 0) = -E_0 \iint_{-\infty}^{\infty} e^{-(x^2+y^2)/W_0^2} e^{i(k_x x + k_y y)} dx dy$$

$$A(k_x, k_y; 0) = E_0 \pi W_0^2 e^{-W_0^2 (k_x^2 + k_y^2)/4}$$

From Fresnel approximation, at a distance z we have

$$A(k_x, k_y; z) = E_0 \pi W_0^2 e^{-W_0^2 (k_x^2 + k_y^2)/4} e^{-ikz} e^{i(k_x^2 + k_y^2) \frac{z}{2k}}$$

$$\Rightarrow E(x, y; z) = \frac{E_0}{4\pi} W_0^2 e^{-ikz} \iint_{-\infty}^{\infty} e^{-(k_x^2 + k_y^2)(\frac{W_0^2}{4} - \frac{i\frac{z}{2k}})} \times e^{-i(k_x x + k_y y)} dk_x dk_y$$

$$= \frac{E_0}{1 - i \frac{2z}{K W_0^2}} e^{-(x^2 + y^2)/W_0^2(1 - i \frac{2z}{K W_0^2})}$$

Define, the Rayleigh range $Z_R = \frac{k W_0^2}{2} = \frac{2\pi}{\lambda} \frac{W_0^2}{2}$

$$\Rightarrow Z_R = \frac{\pi W_0^2}{\lambda}$$

> The factor $1 - i \frac{2z}{kW_0^2} = 1 - i \frac{z}{Z_R} = \sqrt{1 + \left(\frac{z}{Z_R}\right)^2} e^{-i \tan^{-1}\left(\frac{z}{Z_R}\right)}$

polar.

$$= \sqrt{1 + \left(\frac{z}{Z_R}\right)^2} e^{-i\phi(z)}$$

$$= \frac{w(z)}{W_0} e^{-i\phi(z)}$$

> Alternatively,

$$\frac{1}{1 - i\left(\frac{z}{Z_R}\right)} = \frac{Z_R}{Z_R - iz} = \frac{iZ_R}{z + iZ_R} = \frac{iZ_R}{\underline{q(z)}}$$

where $\boxed{q(z) = z + iZ_R}$ is the 'q-parameter'

Define : Wavefront radius as

$$\boxed{R(z) = z \left[1 + \left(\frac{Z_R}{z} \right)^2 \right]}$$

Define : Beam radius as

$$\boxed{W(z) = W_0 \left[1 + \left(\frac{z}{Z_R} \right)^2 \right]} ; W_0 \text{ is the beam waist.}$$

> The Gaussian beam envelope can then be written as,

$$\mathcal{E}(x, y, z) = E_0 \frac{W_0}{W(z)} e^{-i(-\phi(z) + k(x^2 + y^2)/2R(z))} e^{-(x^2 + y^2)/W^2(z)}$$

Also,

$$\boxed{E(x, y, z) = E_0 \left[\frac{iz_R}{q(z)} \right] e^{-ik(x^2+y^2)/2q(z)}}$$

Note,

$$\frac{1}{q(z)} = \frac{1}{R(z)} - i \frac{\lambda}{\pi w^2(z)}$$

The full field $E(x, y, z) = E(x, y, z) e^{-i\beta z}$.

Standard Form of the Gaussian Beam.

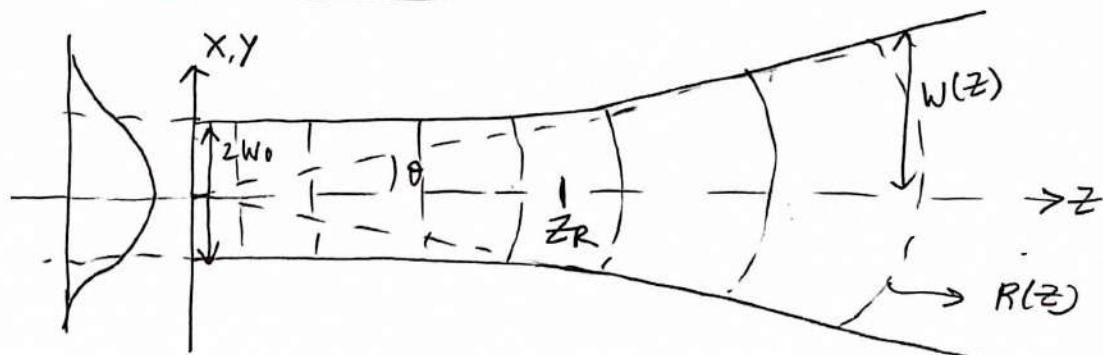
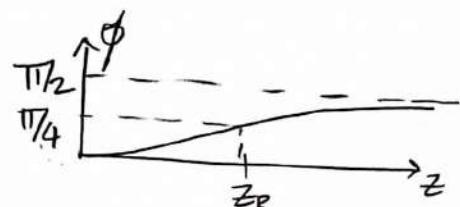
Define origin where beam radius is minimum $\Rightarrow w(z=0) = w_0$, and the wavefront radius $R(z=0) = \infty$. Then,

$$\frac{1}{q(z)} = \frac{1}{q_0} = \frac{1}{R(z)} - i \frac{\lambda}{\pi w^2(z)} = 0 - i \frac{\lambda}{\pi w_0^2}$$

$$\Rightarrow q_0 = i \frac{\pi w_0^2}{\lambda} = i \underline{\underline{z_R}}$$

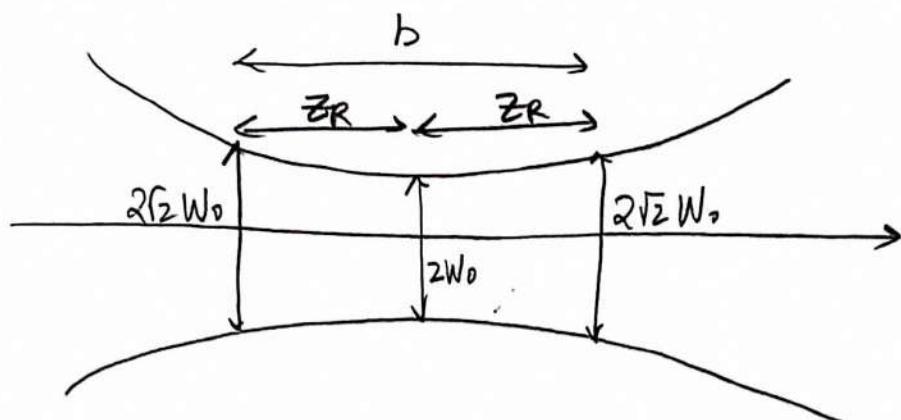
$$\Rightarrow w(z) = w_0 \sqrt{1 + \left(\frac{z}{z_R}\right)^2}; R(z) = z \left[1 + \left(\frac{z_R}{z}\right)^2 \right];$$

$$\boxed{\phi(z) = \tan^{-1} \left(\frac{z}{z_R} \right)} \rightarrow \text{Gouy phase}$$



At $z = z_R$; $w(z) = \sqrt{2} w_0$. Beyond z_R , the beam is not collimated and the divergence due to diffraction becomes significant.

Define $b = 2z_R$ → confocal parameter.



→ A small beam waist corresponds to a small confocal parameter. Since $b = 2z_R = \frac{2\pi w_0^2}{\lambda}$

$$\theta = \tan \theta = \frac{w(z)}{z} \approx \frac{w_0}{z_R} = \frac{\lambda}{\pi w_0}$$

$$\Rightarrow \boxed{\theta = \frac{\lambda}{\pi w_0}} \text{ Angular spread of Gaussian beam.}$$

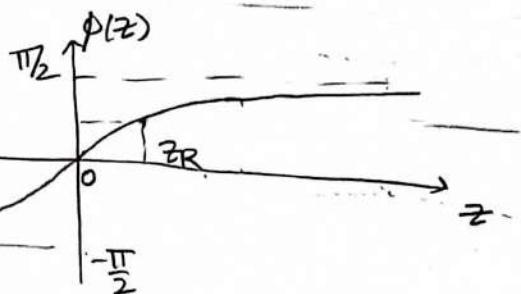
> Narrow beam \Rightarrow more divergence.

Gruy Phase Shift

$e^{-ikz+i\phi(z)}$ is the phase shift of a Gaussian beam.

$$\text{Where } \phi(z) = \tan^{-1}\left(\frac{z}{z_R}\right)$$

> This is an additional (axial) phase shift beyond the normal phase advance, called the Gruy phase shift.



From focus to $\infty \rightarrow \frac{\pi}{2}$

From $-\infty$ to $\infty \rightarrow \pi$.

From $-z_R$ to $z_R \rightarrow \frac{\pi}{2}$

Beam Power

In cylindrical coordinates,

$$I(p) = \frac{2P}{\pi w^2} e^{-2p^2/w^2} \quad \text{where } P \text{ the total power in the beam.}$$

The power transmitted through a circular aperture of diameter D is given by

$$P_{\text{trans}} = \frac{2P}{\pi w^2} \int_0^{D/2} e^{-2p^2/w^2} 2\pi p dp = P \left(1 - e^{-D^2/2w^2}\right).$$

> In reality diffraction from the edges will introduce a "ripple" that can only be reduced by increasing D.

Gaussian Beam Propagation & ABCD matrices.

$R(z)$ and $W(z)$ are connected by

$$\frac{1}{q(z)} = \frac{1}{R(z)} - \frac{i\lambda}{\pi W^2(z)} \quad \text{--- ①}$$

$$\textcircled{①} q(z) = z - iz_R \quad \text{--- ②}$$

If $q(z_1) = q_1 \Rightarrow q(z_2) = q_2 = q_1 + d$ from ②

$$\Rightarrow q_2 = \frac{Aq_1 + B}{Cq_1 + D} \text{ where } \begin{pmatrix} A & B \\ C & D \end{pmatrix} = \begin{pmatrix} 1 & d \\ 0 & 1 \end{pmatrix}$$

- A thin lens changes R but not W .



From ①, $\frac{1}{q_2} - \frac{1}{q_1} = \frac{1}{R_2} - \frac{1}{R_1}$

From GO we know how a thin lens changes R , ie

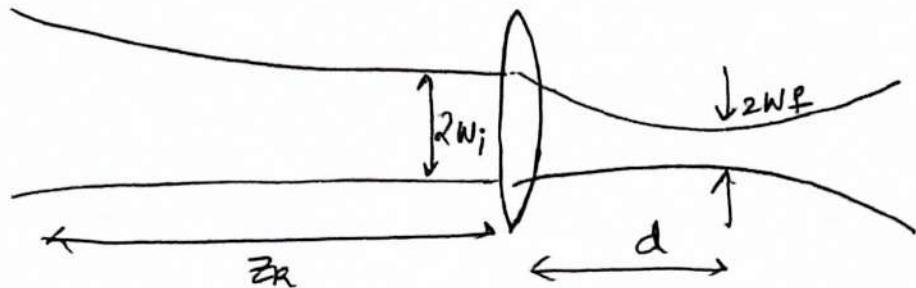
$$R_2 = \frac{AR_1 + B}{CR_1 + D} \text{ where } \begin{pmatrix} A & B \\ C & D \end{pmatrix} = \begin{pmatrix} \gamma & 0 \\ -1/f & 1 \end{pmatrix}$$

Combining these results gives,

$$q_2 = \frac{Aq_1 + B}{Cq_1 + D}$$

→ ABCD law. This law is more general than Gaussian beams. It says that the complex radius of curvature q is changed by a system exactly as R is changed.

Focussing a Gaussian Beam.



Assume
 $z_R \gg f$

$$\frac{1}{q_1} = \frac{1}{\infty} - i \frac{\lambda}{\pi W_i^2} \Rightarrow q_1 = i \frac{\pi W_i^2}{\lambda} \text{ where } z_{Ri} = \frac{\pi W_i^2}{\lambda}$$

The ABCD matrix from lens to focus is

$$\begin{bmatrix} A & B \\ C & D \end{bmatrix} = \begin{bmatrix} 1 & d \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -\frac{1}{f} & 1 \end{bmatrix} = \begin{bmatrix} 1 - \frac{d}{f} & d \\ -\frac{1}{f} & 1 \end{bmatrix}$$

So the ABCD law gives

$$q_2 = \frac{(1 - \frac{d}{f})(i \frac{\pi W_i^2}{\lambda}) + d}{(-\frac{1}{f})(i \frac{\pi W_i^2}{\lambda}) + 1} = \left[\frac{1}{R_2} - i \frac{\lambda}{\pi W_2^2} \right]^{-1}$$

Equating real & imaginary parts gives,

$$R_2(d) = \frac{\left(\frac{d}{z_{Ri}}\right)^2 + \left(1 - \frac{d}{f}\right)^2}{\left(\frac{d}{z_{Ri}}\right)^2 - \frac{1}{f}(1 - \frac{d}{f})}$$

$$\therefore W_2^2(d) = W_i^2 \left(1 - \frac{d}{f}\right)^2 + W_i^2 \left(\frac{d}{z_{Ri}}\right)^2$$

The waist of the focussed beam occurs when $R_2(d) = \infty$

Therefore,

$$\frac{d}{z_{R_i}^2} - \frac{1}{f} \left(1 - \frac{d}{f}\right) = 0 \Rightarrow d \left(\frac{1}{z_{R_i}^2} + \frac{1}{f^2}\right) = \frac{1}{f},$$

$$\Rightarrow d = \frac{f}{1 + \frac{f^2}{z_{R_i}^2}}$$

When the incident beam is well collimated ie $z_{R_i} \gg f$ then

$$d = f \text{ as in GO}$$

Insert d into expression for $W_2^2(cd)$,

$$W_f = \frac{\lambda f}{\pi W_i} \frac{1}{\sqrt{1 + f^2/z_{R_i}^2}} \rightarrow \text{Beam waist at focus.}$$

Again when $z_{R_i} \gg f$,

$$W_f = \frac{\lambda f}{\pi W_i}$$

Define beam diameters as $D_i = 2W_i$, $D_f = 2W_f$, then

$$D_f = \frac{4}{\pi} \lambda \left(\frac{f}{D_i}\right)$$

$\frac{f}{D_i}$ is the f-number ie $f'^{\#} \approx \frac{4}{\pi} \approx 1.27$

$$\Rightarrow D_f = 1.27 \lambda f'^{\#}$$

$f'^{\#} \uparrow \Rightarrow$ fast lens since it focuses at a very close range

The fastest practical lenses have $f'^{\#} \approx 1 \Rightarrow D_{f\min} \approx \lambda \Rightarrow$ Diffraction limit ||

Fraunhofer Diffraction

We start with the Fresnel Diffraction integral.

$$E(x, y, z) = \frac{i}{\lambda} \cdot \frac{e^{-ikz}}{z} \iint_{-\infty}^{\infty} E(x', y'; 0) e^{-ik[(x-x')^2 + (y-y')^2]/2z} dx' dy'$$

Expanding the phase term,

$$k \frac{(x-x')^2 + (y-y')^2}{2z} = k \left[\frac{x^2+y^2}{2z} - \frac{xx'+yy'}{z} + \frac{x'^2+y'^2}{2z} \right] \quad \text{ignore this}$$

Fraunhofer ("far field") approximation $\Rightarrow \frac{z}{k} \gg \frac{x'^2+y'^2}{2}$

$$\Rightarrow \text{Let } W = \max(x', y'), \text{ we need } z \gg \frac{kW^2}{2} = \frac{\pi W^2}{\lambda} \rightarrow \text{looks like Rayleigh range!}$$

Hence we need to be many Rayleigh ranges away.

$$\Rightarrow E(x', y'; z) = \frac{ie^{-ikz}}{\lambda z} \cdot e^{-ik(x^2+y^2)/2z} \iint_A E(x', y'; 0) e^{ik(xx'+yy')/z} dx' dy'$$

Fraunhofer Diffraction integral; $k_x = k \sin \theta_x = k \tan \theta_x = \frac{kx}{z}$

$$\Rightarrow E(x', y'; z) = \frac{ie^{-ikz}}{\lambda z} \cdot e^{-ik(x^2+y^2)/2z} \iint_A E(x', y'; 0) e^{i(k_x x' + k_y y')} dx' dy'$$

High Q tunable filters @ high frequencies.

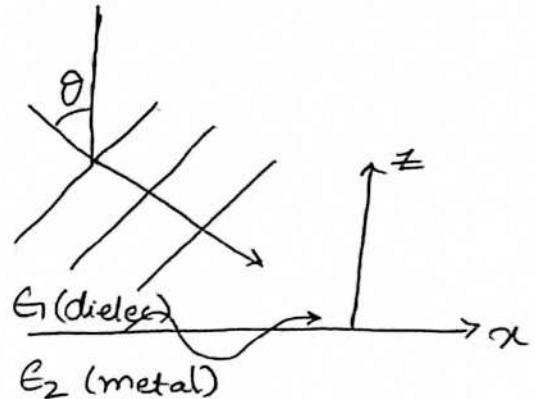
① Surface Plasmon Polaritons excited by a grating.

SPP basics - Plane wave interaction.

$$k_{x_1} = k_{x_2}$$

$$\beta = k_{x_2} = \frac{\omega}{c} \sqrt{\frac{\epsilon_1 \epsilon_2}{\epsilon_1 + \epsilon_2}} \rightarrow \text{dispersion relation.}$$

$$k_{x_1} = \frac{\omega}{c} \sqrt{\epsilon_1} \sin \theta$$



$$\Rightarrow \sin \theta = \sqrt{\frac{\epsilon_m}{\epsilon_d + \epsilon_m}}$$

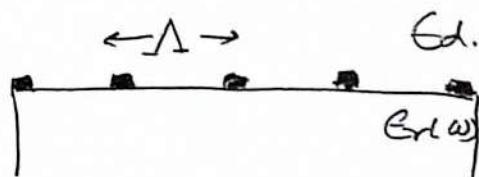
Also need

$$\frac{\epsilon_1}{\epsilon_2} = -\frac{\alpha_1}{\alpha_2} \rightarrow \text{damping factor.}$$

> 1 since $\epsilon_m < 0$.

To match the momentum along x use grating!

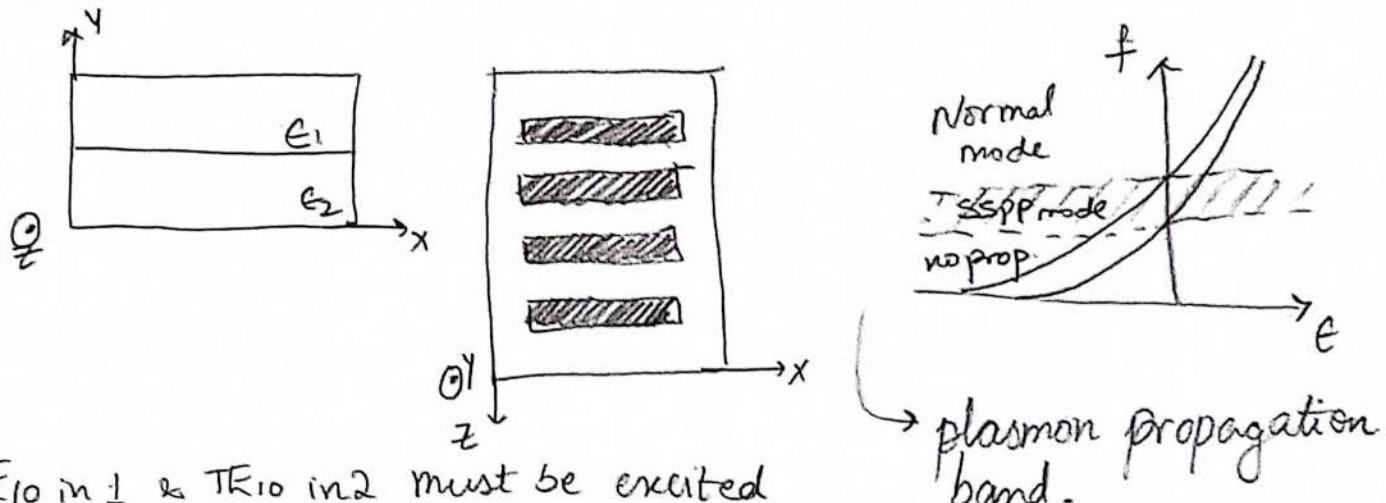
$$\beta = k_{x_1} + \frac{2\pi}{\lambda}$$



$$\Rightarrow \frac{\omega}{c} \sqrt{\epsilon_d} \left[\sqrt{\frac{\epsilon_r(\omega)}{\epsilon_d + \epsilon_r(\omega)}} - \sin \theta \right] = 1/\lambda.$$

- > Design ϵ_d & $\epsilon_r(\omega)$ to get a realisable $1/\lambda$ value at THz!
- > Cannot use a real metal since $\epsilon_r(\omega)$ is purely imaginary at subTHz
- > How can a grating be implemented if ϵ_1 & ϵ_2 are 2D lattices
- > What would the dispersion relation look like & can plasmons even be excited in a 2D lattice?
- > Are the E & H fields in a lattice confined to the lumped elements?

II Spoof SPP in SIW for band pass response.



- > TE₁₀ in 1 & TM₁₀ in 2 must be excited
& ε₁ & ε₂ must have opposite signs.
- > ε₁ & ε₂ in reality are positive by due to structural dispersion
- > Can we achieve such a structural dispersion to make ε₁ & ε₂ of opposite signs in a 2D lattice? This would probably allow plasmon propagation at the interface.
- > PPW : structural dispersion → $\epsilon_{eff} = \epsilon_b - \frac{\lambda_0^2}{4a^2}$ —①
⇒ ε_{eff} can be negative!
- > Mode coupling should be negligible.
The rods are basically killing the TM₁₀ mode since it has an E_x-component.
- > The waveguide structural dispersion then allows ε₁ & ε₂ to be of opposite signs as per Eq① & a dipole along E_y can be used to excite the SPP.
- > 2D lattice to control ε₁ & ε₂ — anal?

↳ comes from

$$\lambda g_{mn} = \left[\sqrt{\frac{1}{\lambda^2} - \left(\frac{m}{2a} \right)^2 - \left(\frac{n}{2b} \right)^2} \right]^{-1}$$

* → Combine ideas I & IV using Lattice dispersion to design tunable filters?

