

To Grid or Not To Grid

Atomic Methods for Sparse Inverse Problems

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under the direction of
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November 7th, 2025



Part I: What is it about?

1. Linear Inverse Problems
2. Sparse Signals
3. Penalized Optimization
4. Atomic Reconstruction Methods



“Méthodes atomiques
pour la résolution de
problèmes inverses
parcimonieux”



“En science, un **problème inverse** est une situation dans laquelle on tente de déterminer les causes d'un phénomène à partir des observations expérimentales de ses effets.”

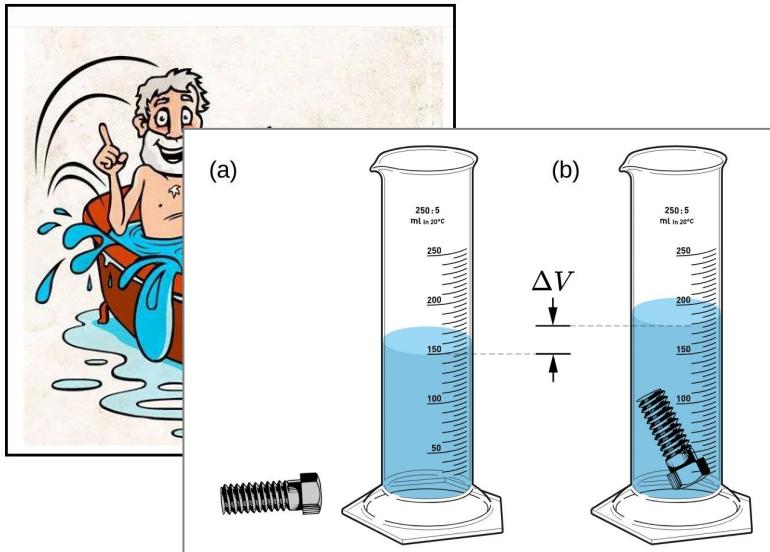
Wikipedia, FR

“An **inverse problem** in science is the process of calculating from a set of observations the causal factors that produced them [...]”

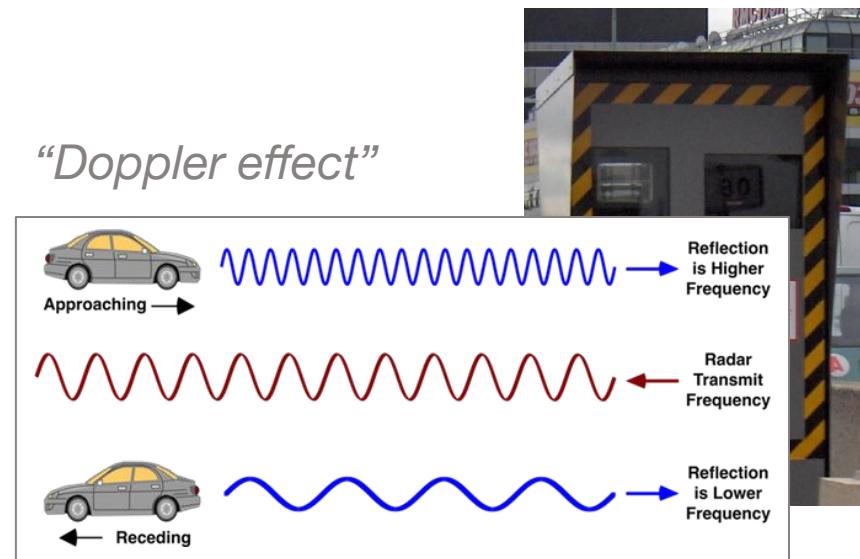


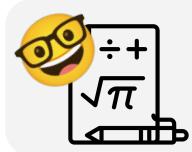
Wikipedia, EN

Inverse Problems ?



“Doppler effect”

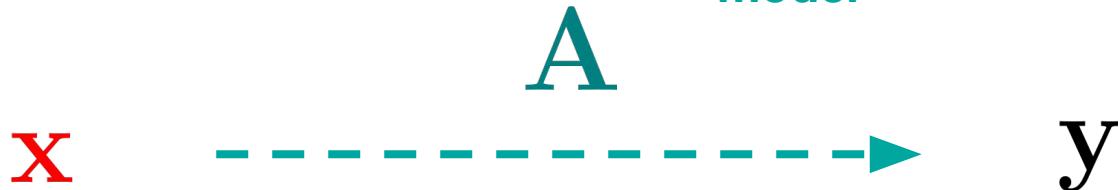




Unknown signal

- volume
- speed
- image
- time series
- ...

Measurement model



Physical observations

- distance
- frequency
- image
- interferences
- ...

“Forward/design equation”

$$\textcolor{yellow}{y} = \textcolor{teal}{A} \textcolor{red}{x}$$

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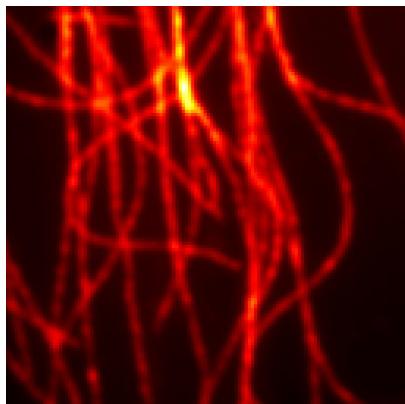
y

y

A



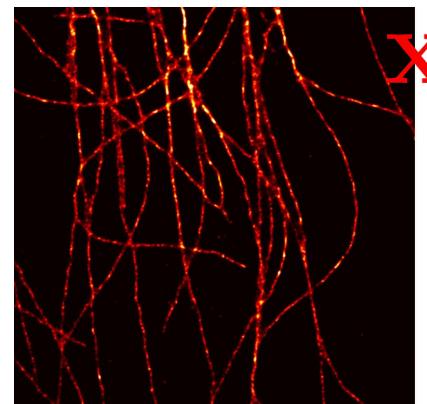
y



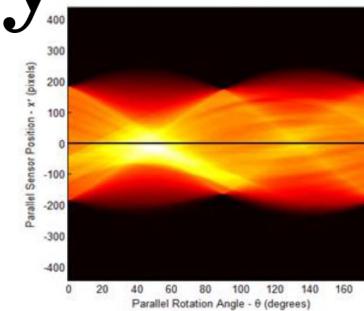
A



x

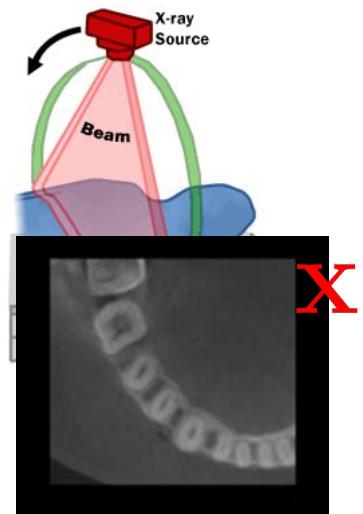


y



“Deconvolution”

EPFL



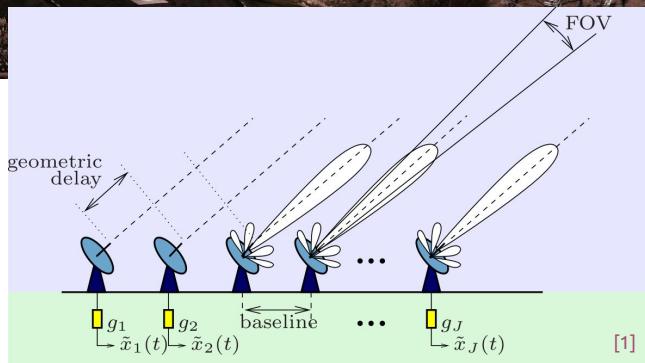
“SMLM”

[Credits: Bastien Laville]

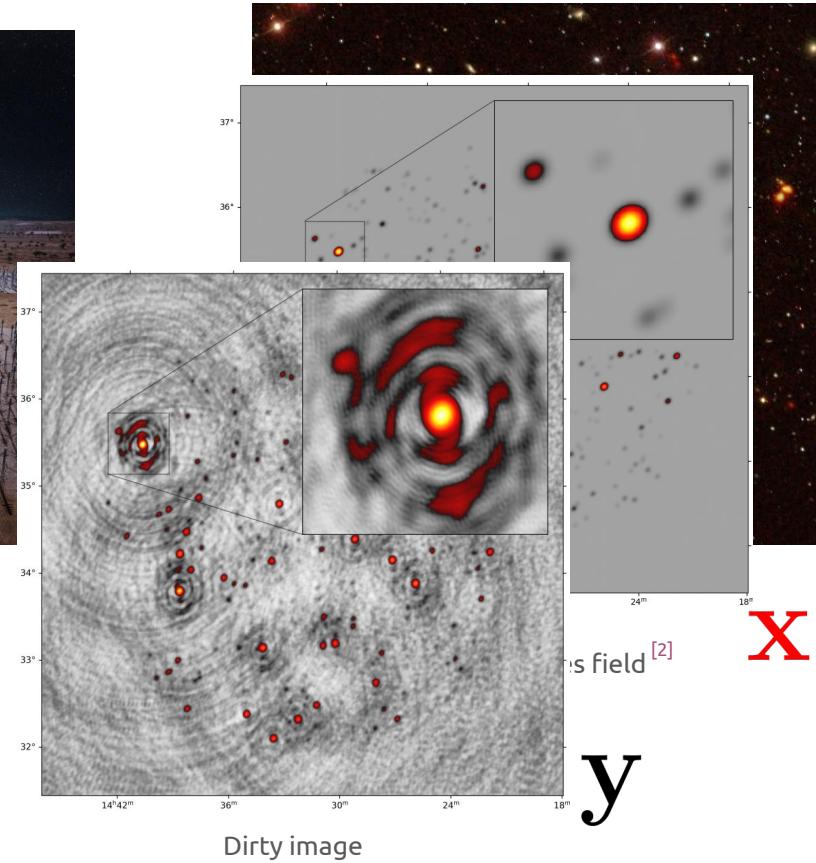


[Credits: SKAO]

A



[1]



y

X

Dirty image

[1] Van der Veen et al. "Signal Processing for Radio Astronomy", 2019

[2] Williams WL et al., "LOFAR 150-MHz observations of the Boötes field: catalogue and source counts.", *Monthly Notices of the Royal Astronomical Society*, 2016



$$\mathbf{x} = \mathbf{A}^{-1} \mathbf{y}$$

~~A⁻¹~~

- Measurement noise
- Instability
- Non-uniqueness
- Missing data



$$\mathbf{y} = \mathbf{Ax} + \mathbf{n}$$

$$\mathbf{Ax}_1 = \mathbf{Ax}_2$$

~~A⁻¹~~



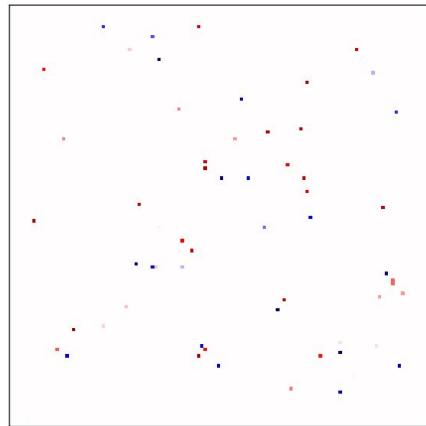
Part I: What is it about?

1. ~~Linear Inverse Problems~~ $y = \mathbf{A}\mathbf{x}$
2. Sparse Signals
3. Penalized Optimization
4. Atomic Reconstruction Methods

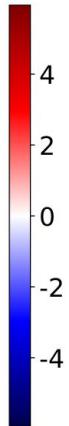


“Méthodes atomiques
pour la résolution de
problèmes inverses
parcimonieux”

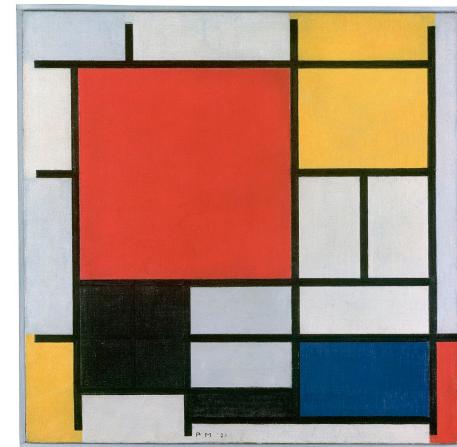
2. Parcimonie ■ / Sparsity



Simulated sparse signal

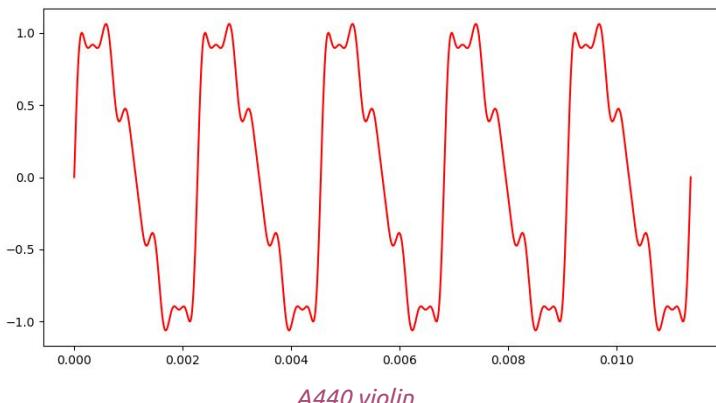


GLEAM observations of the sky

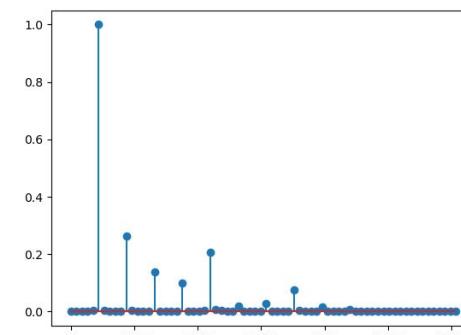


*Composition in Red, Yellow, Blue and Black -
Piet Mondrian, 1921*

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A440 violin



$$\mathbf{x} = \alpha_1 \mathbf{a}_1 + \alpha_2 \mathbf{a}_2 + \dots + \alpha_K \mathbf{a}_K$$

EPFL 2. (Bonus) All signals can be approx.-sparse



1% information

2. (Bonus) Reconstruction example with sparsity

Something like inpainting or deblurring of an image, with simple sparsity operator

Or Radio Interferometry

Results are maybe not as impressive as what we can see nowadays with AI, but it is controlled, stable and reproducible -> another debate

2. (Bonus) Historical moment

Consequence of this approach: change of paradigm, birth of compressed sensing, reduction of the number of observations

Part I: What is it about?

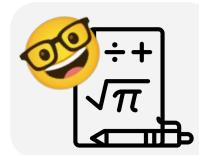
1. ~~Linear Inverse Problems~~
2. ~~Sparse Signals~~
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$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
$$\mathbf{x} = \sum \alpha_k \mathbf{a}_k$$



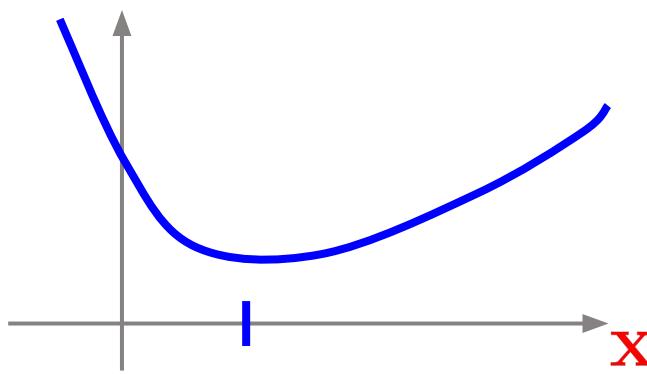
“Méthodes atomiques
pour la **résolution** de
problèmes inverses
parcimonieux”

3. Penalized Optimization



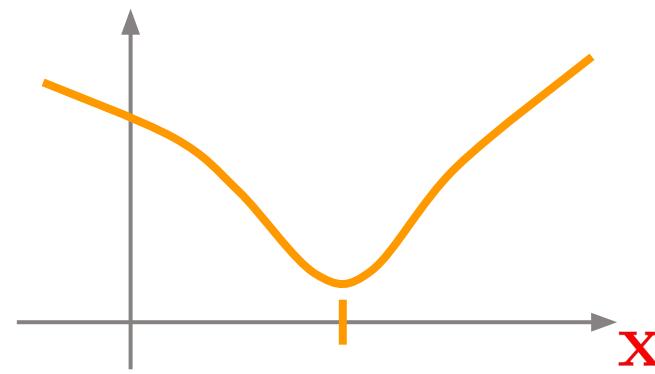
$$\mathbf{y}^T \mathbf{y} - \mathbf{A}^T \mathbf{A} \mathbf{x} = 0$$

$$E(\mathbf{y} - \mathbf{A}\mathbf{x})$$



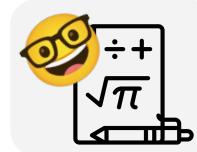
“Data-fidelity”

$$R(\mathbf{x})$$

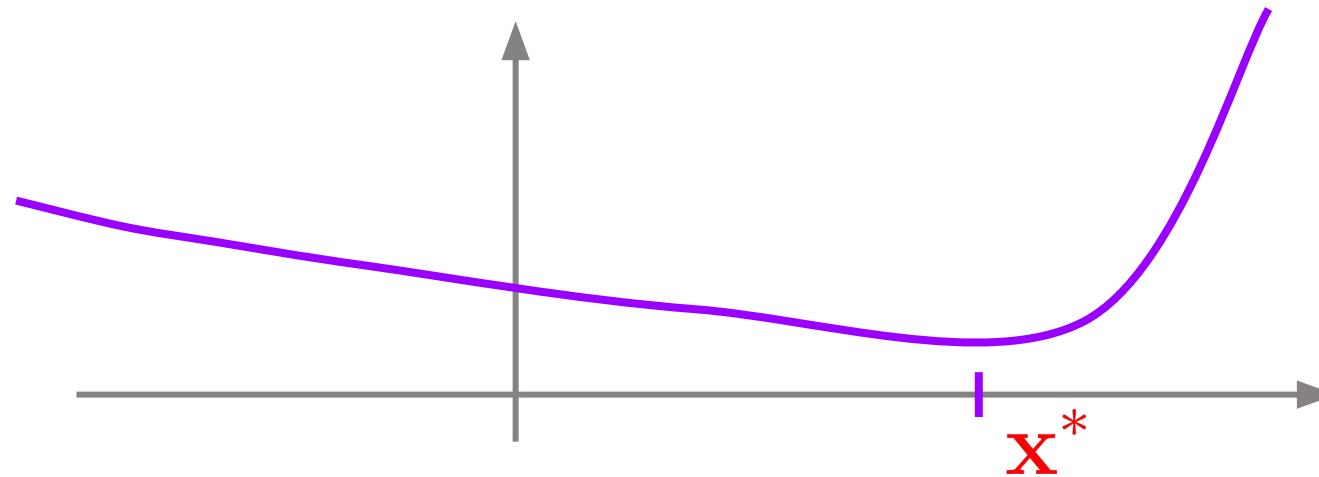


“Penalty/Regularization”

3. Penalized Optimization



$$\min_{\mathbf{x}} E(\mathbf{y} - \mathbf{Ax}) + \lambda R(\mathbf{x})$$



LASSO problem^[3]:

$$\min_{\mathbf{x}} \frac{1}{2} \|\mathbf{y} - \mathbf{Ax}\|_2^2 + \lambda \|\mathbf{x}\|_1$$

[3] Tibshirani R. "Regression Shrinkage and Selection via the Lasso", *Journal of the Royal Statistical Society Series B (Methodological)*, 1996.

Part I:

What is it about?

1. Linear Inverse Problems
2. Sparse Signals
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4. Atomic Reconstruction Methods

$$\begin{aligned} \mathbf{y} &= \mathbf{A}\mathbf{x} \\ \mathbf{x} &= \sum_{\mathbf{k}} \alpha_k \mathbf{a}_k \\ &\min_{\mathbf{x}} \mathcal{E} \end{aligned}$$

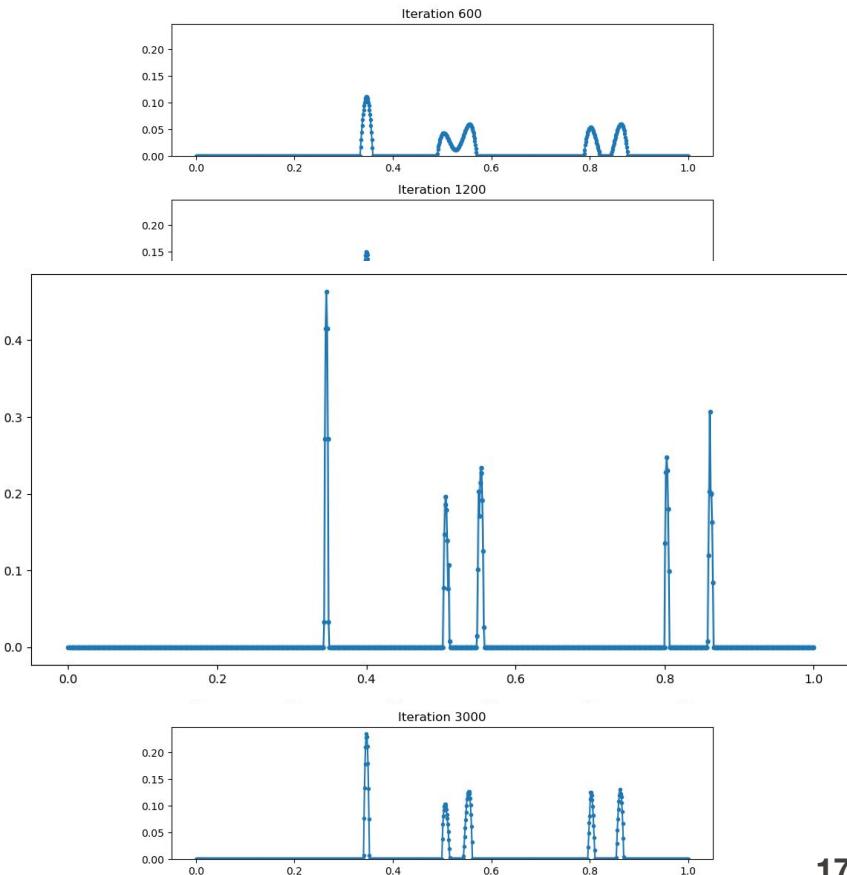
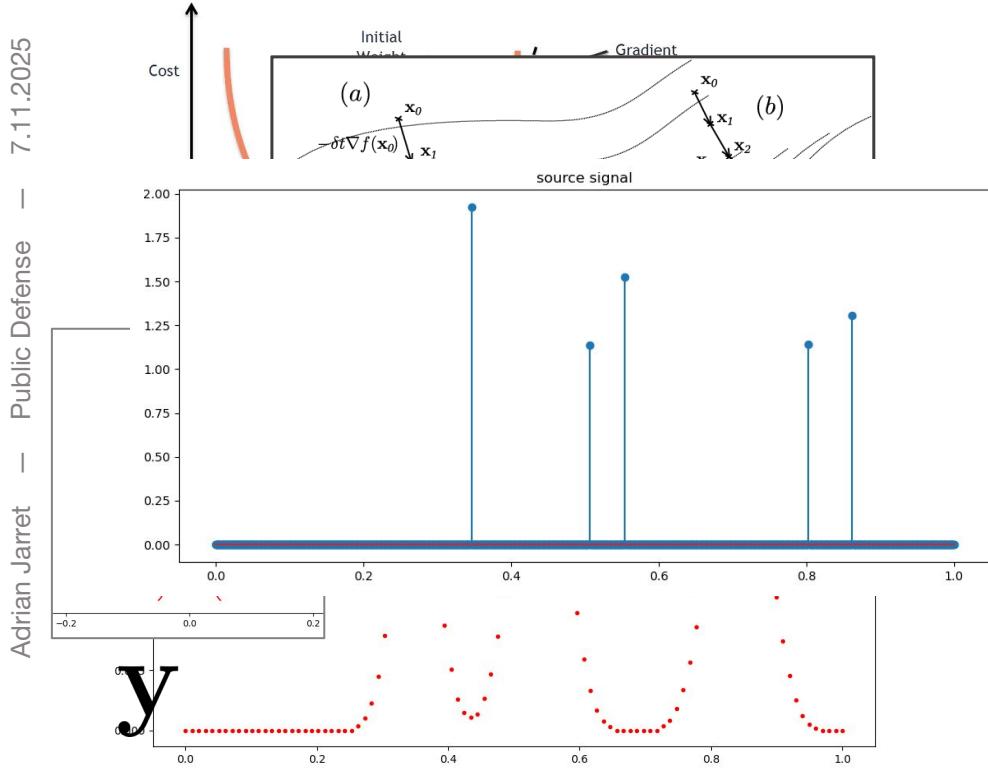
“**Méthodes atomiques**
pour la **résolution** de
problèmes inverses
parcimonieux”



4. Numerical reconstruction methods

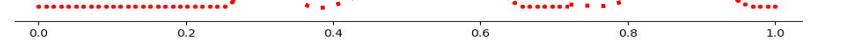
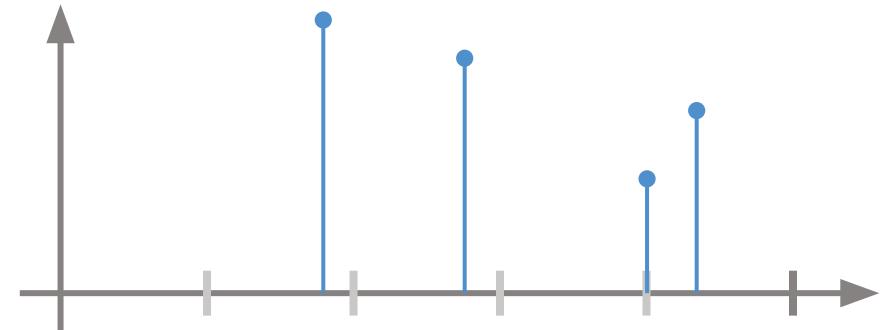
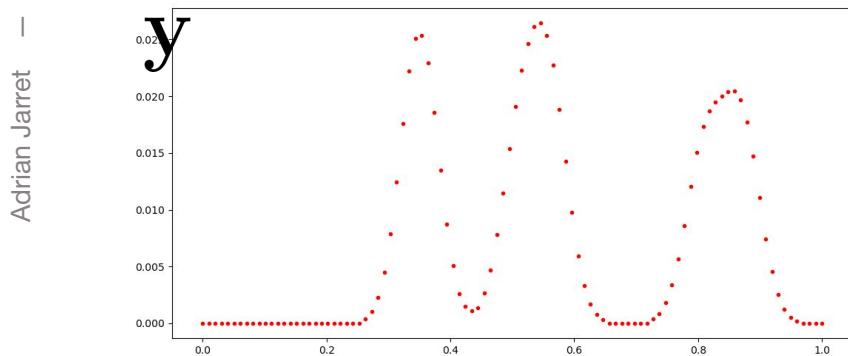
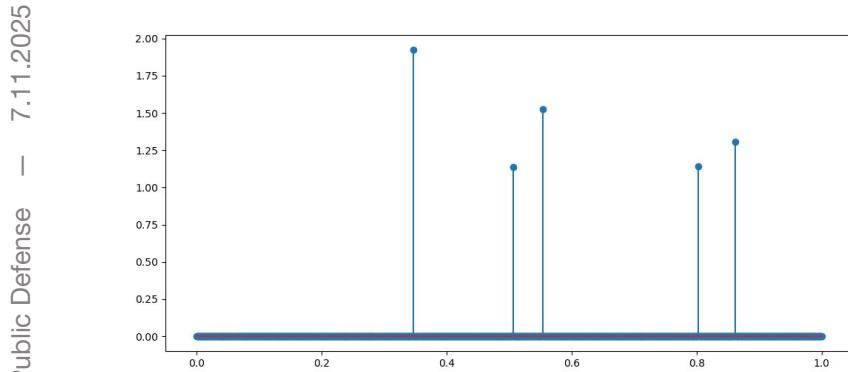
$$\min_{\mathbf{x}} E(\mathbf{y} - \mathbf{Ax}) + \lambda R(\mathbf{x})$$

“Classical” gradient descent:



EPFL 4. Numerical reconstruction methods

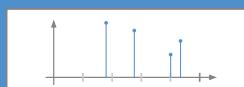
Atomic method:



Part I: What is it about?

1. ~~Linear Inverse Problems~~
2. ~~Sparse Signals~~
3. ~~Penalized Optimization~~
4. ~~Atomic Reconstruction Methods~~
5. **To Grid or Not To Grid ?**

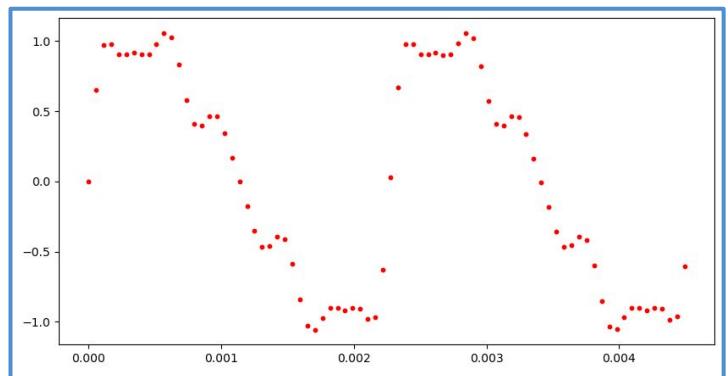
$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
$$\mathbf{x} = \sum \alpha_k \mathbf{a}_k$$
$$\min_{\mathbf{x}} \mathcal{E}$$



“Méthodes atomiques
pour la résolution de
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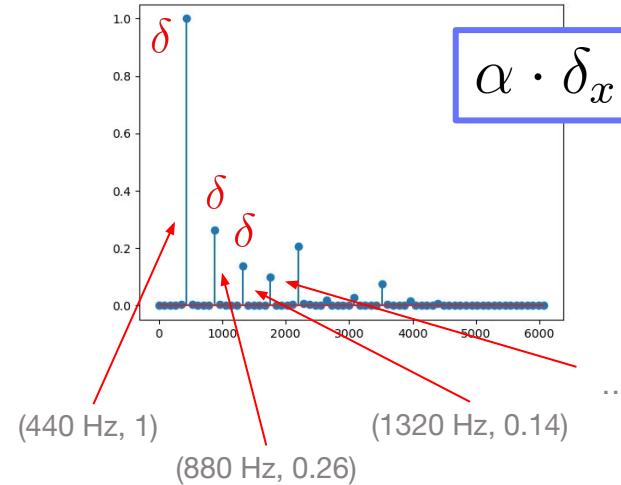
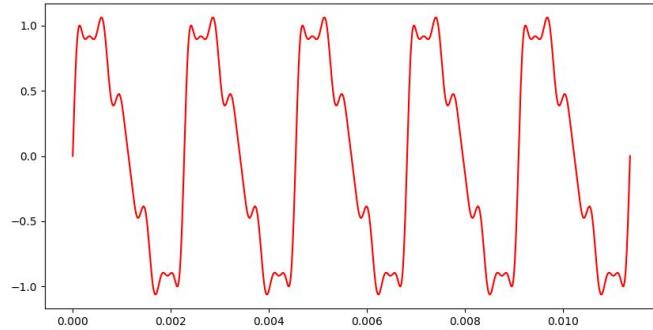
5. Continuous-domain reconstruction



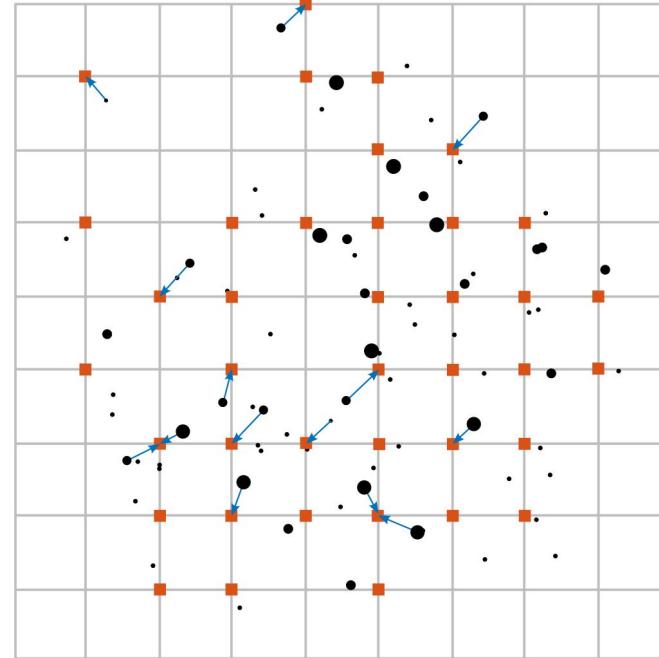
5. Continuous-domain reconstruction?

 $\in \mathcal{M}(\mathbb{R}^d)$

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$$\mathbf{x} \in \mathbb{R}^N \quad m = \sum \alpha_k \delta_{x_k}$$



[Credits: H. Pan et al., "LEAP", A&A, 2017.]

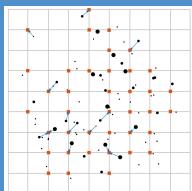
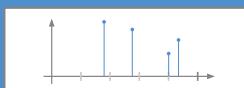
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+ Radio Interferometry

$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
$$\mathbf{x} = \sum \alpha_k \mathbf{a}_k$$
$$\min_{\mathbf{x}} \mathcal{E}$$



Challenges:

- Time efficiency
- Accuracy
- Scalability



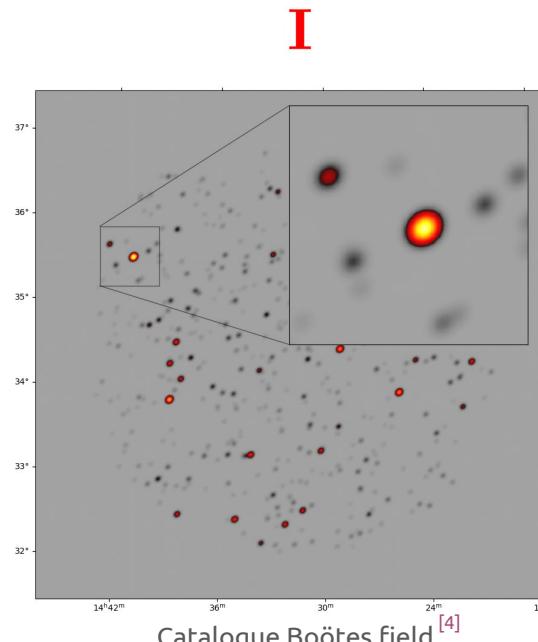
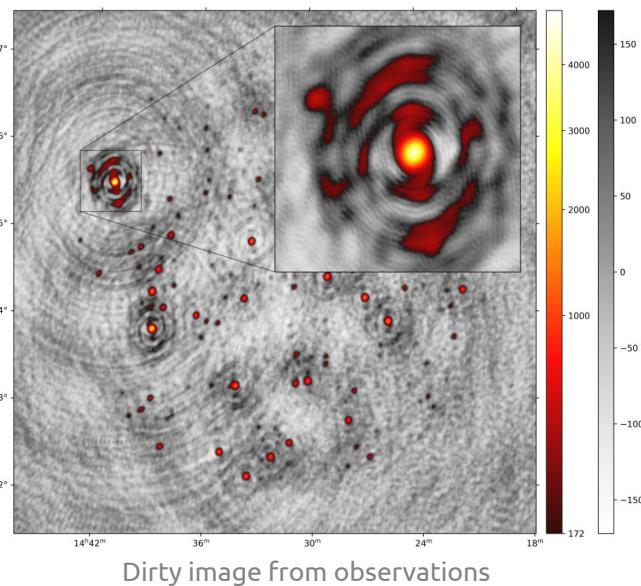
Part II: Focus on key contributions

1. PolyCLEAN:
A Polyatomic Method for
Radio Interferometry
2. Decoupling of Composite
Problems

EPFL Radio Interferometric Imaging - Dirty Image

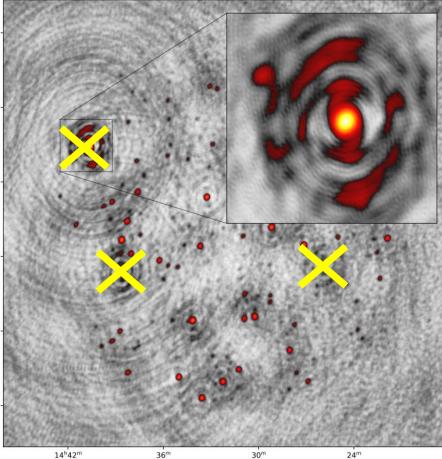
$$\mathbf{V} = \Phi \mathbf{I} + \epsilon$$

$$\mathbf{I}_D = \Phi^* \mathbf{V}$$



EPFL Classical Approaches

The CLEAN family ^[3]

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- 
- ✓ Intuitive and simple method, long-developed and fast
 - ✗ Sensitive to stop, objective function unclear, physically impossible artefacts

[3] Högbom JA., "Aperture Synthesis with a Non-Regular Distribution of Interferometer Baselines", *Astronomy and Astrophysics Supplement Series*, 1974.

Optimization-based Methods ^[4]

$$\arg \min_{\mathbf{I} \in \mathbb{R}^N} \frac{1}{2} \|\mathbf{y} - \Phi \mathbf{I}\|_2^2 + \mathcal{R}(\mathbf{I})$$

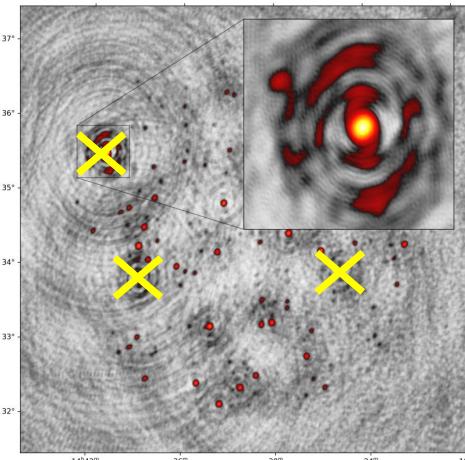


Controlled solutions, versatile priors, excellent results, additional tools

[4] Wiaux Y. et al., "Compressed sensing imaging techniques for radio interferometry", *Monthly Notices of the Royal Astronomical Society*, 2009.

EPFL Classical Approaches

The CLEAN family ^[3]



I_D

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Intuitive and simple method,
long-developed and fast
Sensitive to stop, objective function
unclear, physically impossible artefacts



Optimization-based Methods ^[4]

$$\arg \min_{\mathbf{I} \in \mathbb{R}^N} \frac{1}{2} \|\mathbf{y} - \Phi \mathbf{I}\|_2^2 + \lambda \|\mathbf{I}\|_1$$

LASSO



Controlled solutions, versatile priors,
excellent results, additional tools



Numerically heavy, little adoption in the
field

[3] Högbom JA., "Aperture Synthesis with a Non-Regular Distribution of Interferometer Baselines", *Astronomy and Astrophysics Supplement Series*, 1974.

[4] Wiaux Y. et al., "Compressed sensing imaging techniques for radio interferometry", *Monthly Notices of the Royal Astronomical Society*, 2009.

Our Polyatomic Frank-Wolfe Algorithm

$$\arg \min_{\mathbf{x} \in \mathbb{R}^N} \mathcal{J}(\mathbf{x}) := \frac{1}{2} \|\mathbf{y} - \mathbf{A}\mathbf{x}\|_2^2 + \lambda \|\mathbf{x}\|_1$$

Algorithm 3: Polyatomic Frank-Wolfe algorithm of quality $\delta^{[9]}$

Initialize $\mathbf{x}_0 \in \mathcal{D}$, $\mathcal{S}_0 = \emptyset$

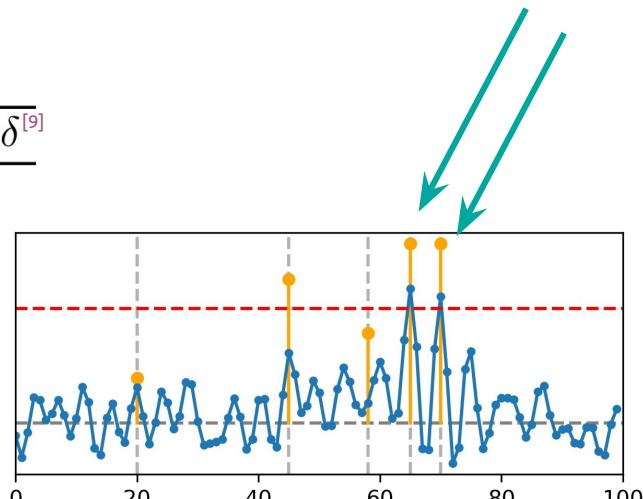
for $k = 1, 2, \dots$ **do**

1) Find update directions:

$$\begin{aligned} \mathcal{I}_k &\leftarrow \{1 \leq j \leq N : |\boldsymbol{\eta}_k[j]| \geq \|\boldsymbol{\eta}_k\|_\infty - \delta/k\} \\ \mathcal{S}_k &\leftarrow \mathcal{S}_{k-1} \cup \mathcal{I}_k \end{aligned}$$

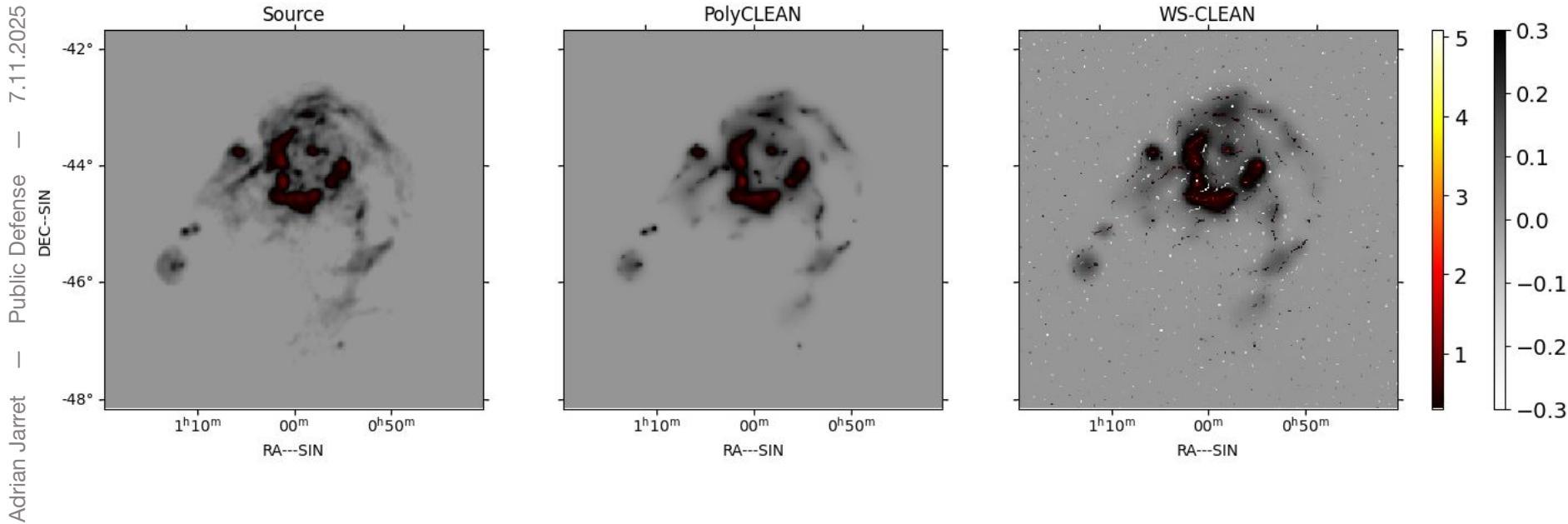
2) Reweighting:

$$\mathbf{x}_k \leftarrow \underset{\text{Supp}(\mathbf{x}) \subset \mathcal{S}_k}{\arg \min} \frac{1}{2} \|\mathbf{y} - \mathbf{A}\mathbf{x}\|_2^2 + \lambda \|\mathbf{x}\|_1$$

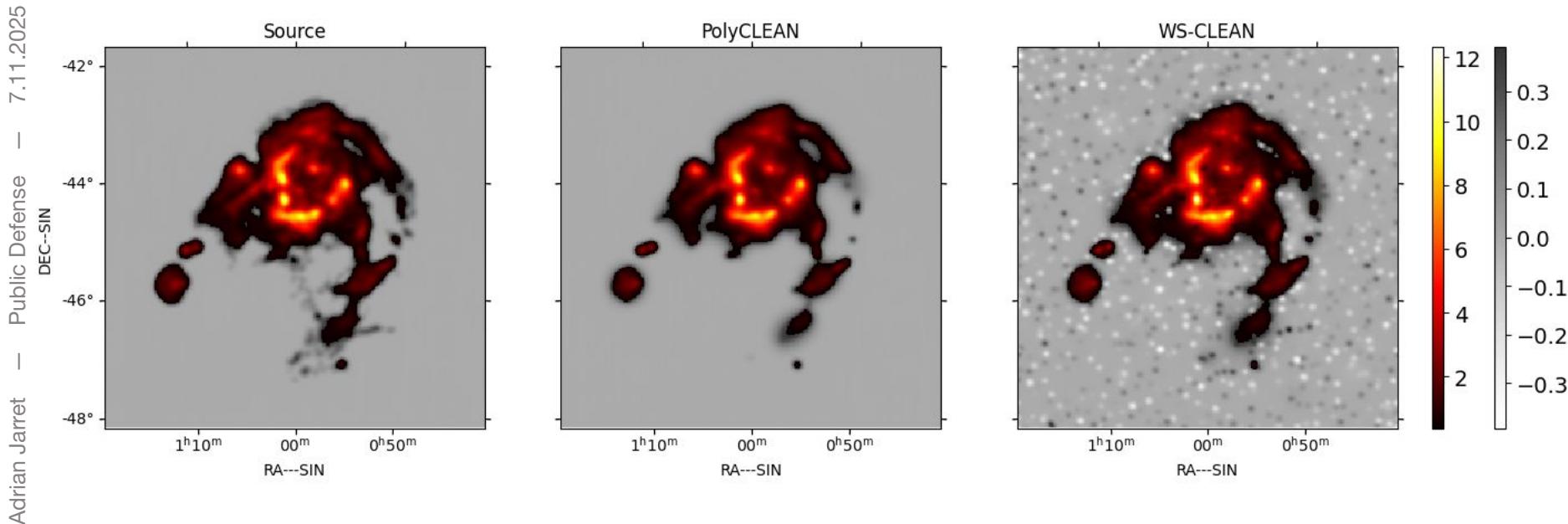


$|\mathcal{S}_k| \ll N$

Diffuse emission reconstruction



Diffuse emission reconstruction



EPFL Summary and Benefits



Design of optimization algorithm



Real world application



Best of both worlds:

- Benefits of CLEAN → **Atomic, fast**
- Benefits of convex optimization → **Accurate**
- Sparsity-aware processing (HVOX) → **Numerical efficiency**
- Question of resolution
 - CLEAN beam is too coarse, certificate beam is **data-inspired**

Part II: Focus on key contributions

1. PolyCLEAN:
~~A Polyatomic Method for
Radio Interferometry~~
2. Decoupling of Composite
Problems

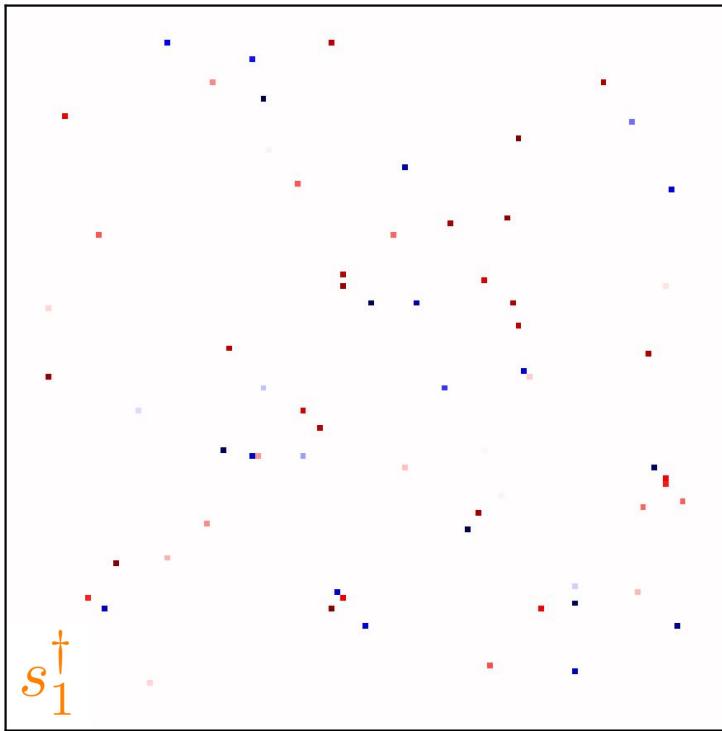
Sparse-plus-Smooth Problems



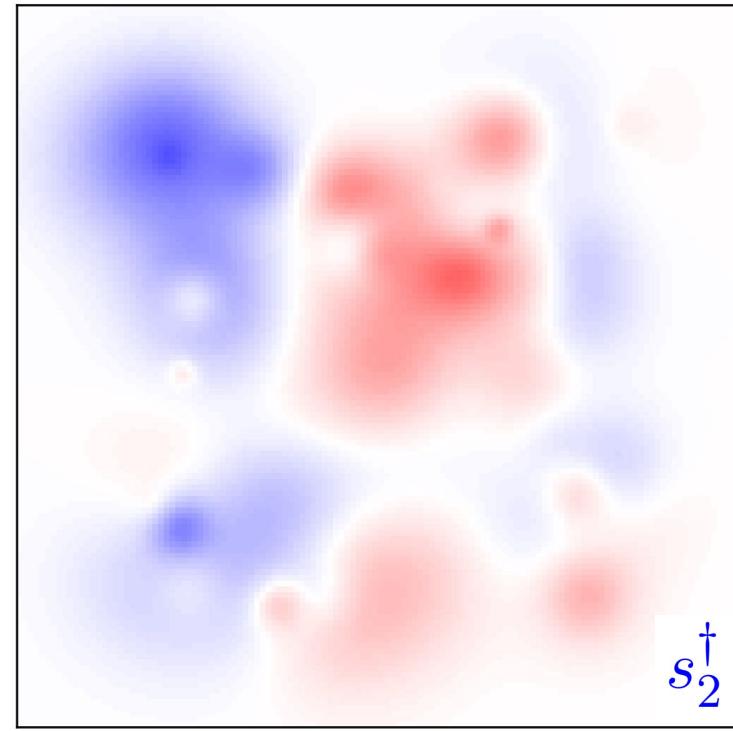
GLEAM survey of the radio sky, J2000 coordinates (9h37min15.21s, 50°25'03.1")

$$s_1^\dagger + s_2^\dagger$$

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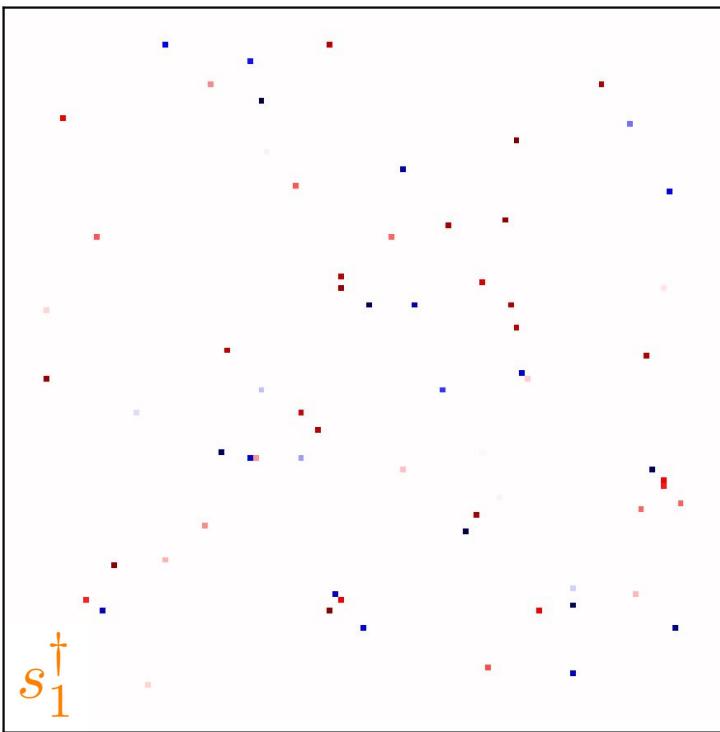


Sparse foreground

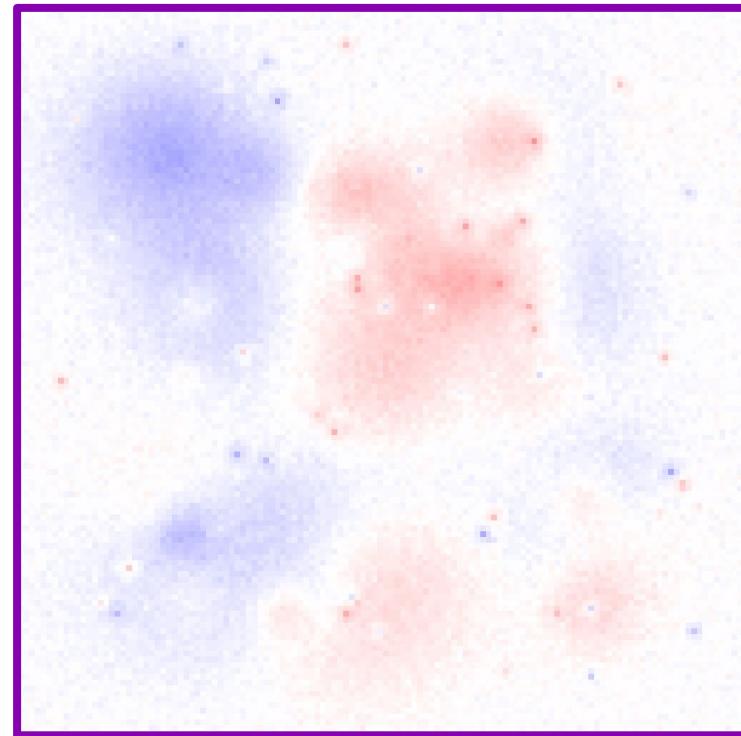


Smooth background

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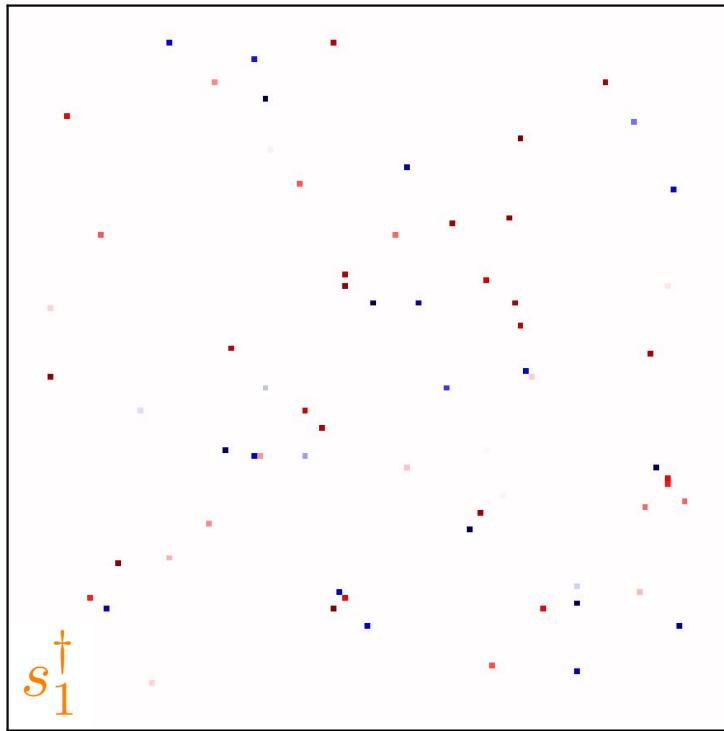


Sparse foreground

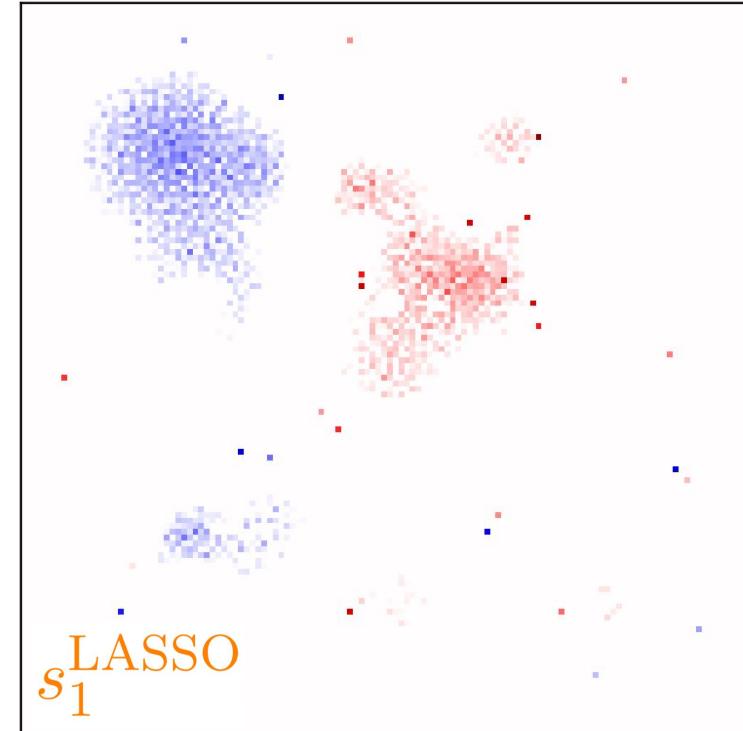


Dirty Image

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Sparse foreground



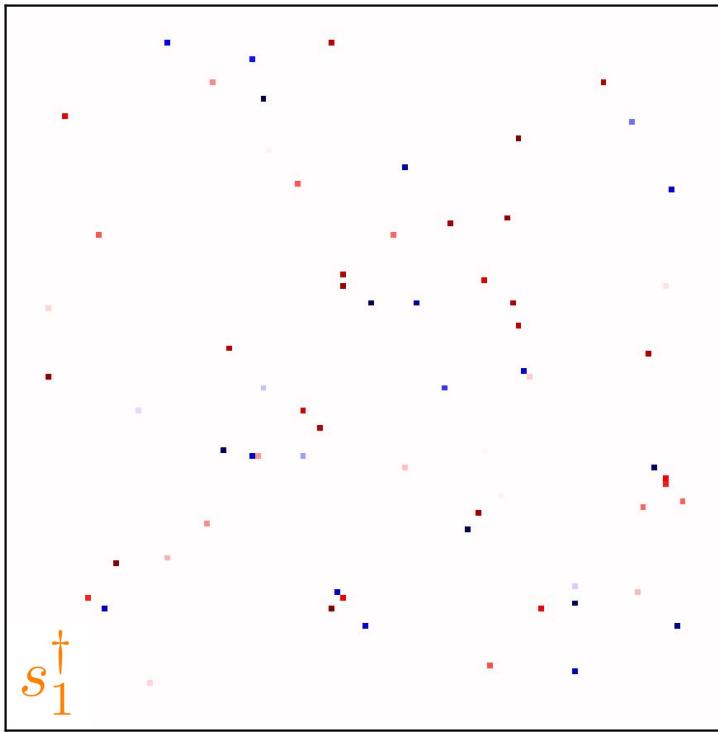
s_1^{\dagger}
LASSO

LASSO reconstruction

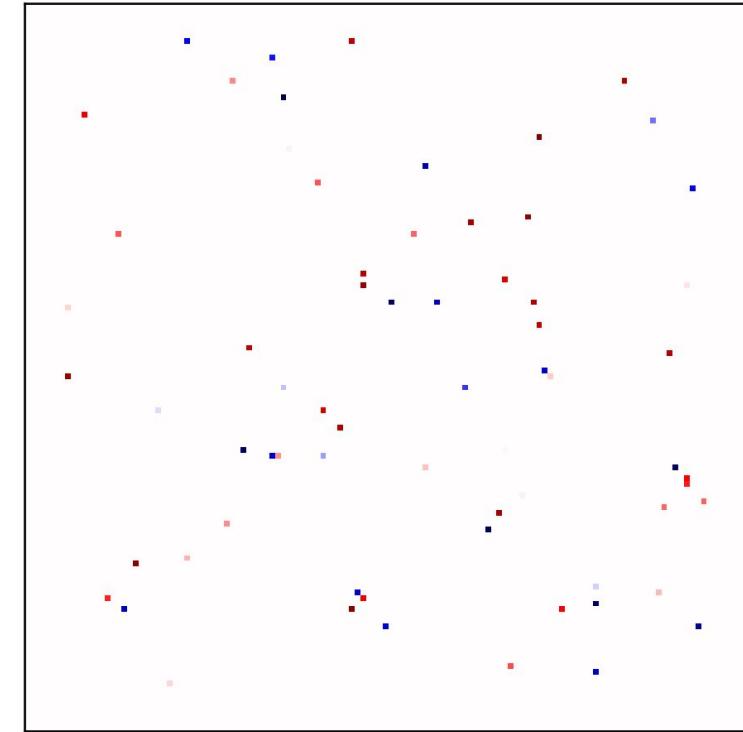
EPFL Sparse-plus-Smooth Problems

$$\arg \min_{s_1, s_2} \mathcal{J}(s_1, s_2)$$

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Sparse foreground



Composite model

$$\arg \min_{s_1, s_2 \in \mathcal{M}(\mathcal{X}) \times \mathcal{L}_2(\mathcal{X})} \frac{1}{2} \|\mathbf{y} - \Phi(s_1 + s_2)\|_2^2 + \lambda_1 \|s_1\|_{\mathcal{M}} + \frac{\lambda_2}{2} \|s_2\|_{\mathcal{L}_2}^2$$

Our Representer Theorem
[Theorem 6.1]

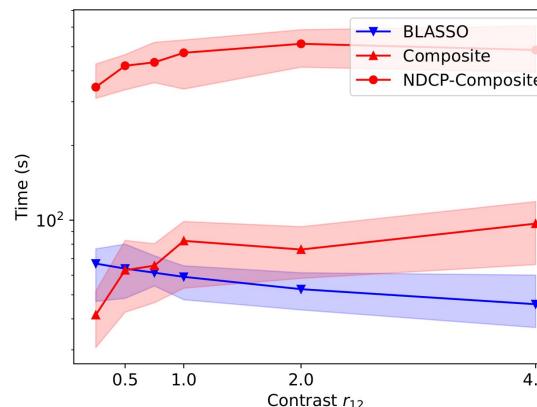
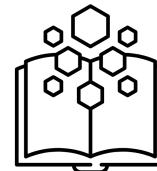
$$\left\{ \begin{array}{l} \widehat{s}_1 \in \arg \min_{s_1 \in \mathcal{B}} \frac{1}{2} \|\mathbf{M}_{\lambda_2}^{-\frac{1}{2}} (\mathbf{y} - \Phi(s_1))\|_2^2 + \lambda_1 \|s_1\|_{\mathcal{B}} \\ \widehat{s}_2 = \frac{1}{\lambda_2} \Phi^* \mathbf{M}_{\lambda_2}^{-1} (\mathbf{y} - \mathbf{w}) \end{array} \right.$$

$$\mathbf{M}_{\lambda_2} := \frac{1}{\lambda_2} (\Phi \Phi^* + \lambda_2 \mathbf{I}_L)$$

$$\mathbf{w} = \Phi(\widehat{s}_1)$$

EPFL Summary and Benefits

- Better understanding of the composite optimization problem
- Nice mathematical framework
- More efficient optimization algorithms



A Short Summary of this Thesis

1. Polyatomic Frank-Wolfe
2. PolyCLEAN
3. Continuous-Domain PFW
4. Composite Reconstruction

Better understanding
of the
problems



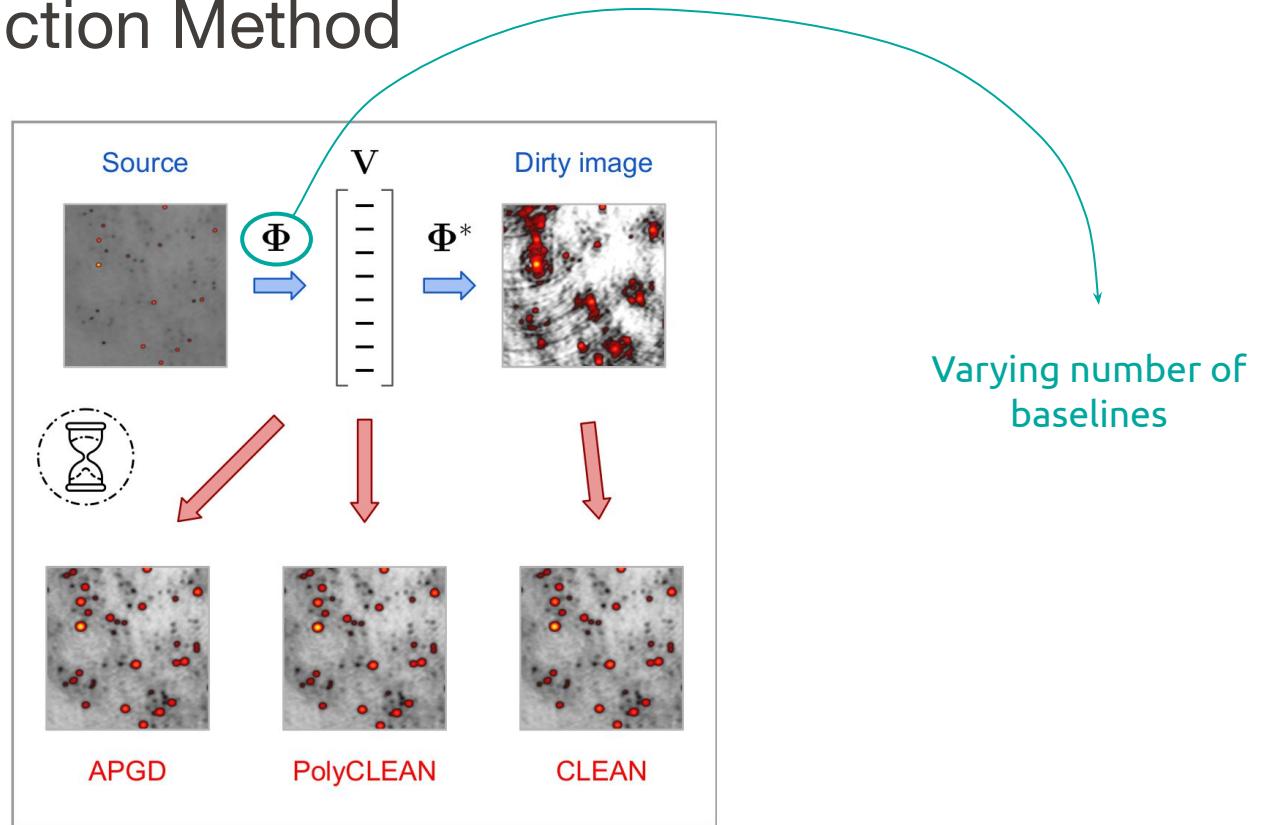
Develop more
efficient
numerical methods



Thank you !

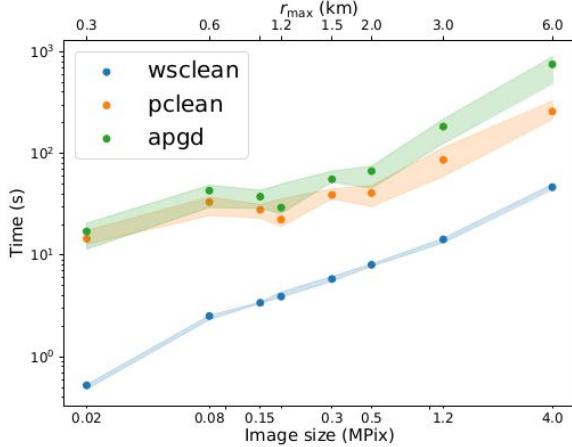
Supplementary slides

A Fast Reconstruction Method

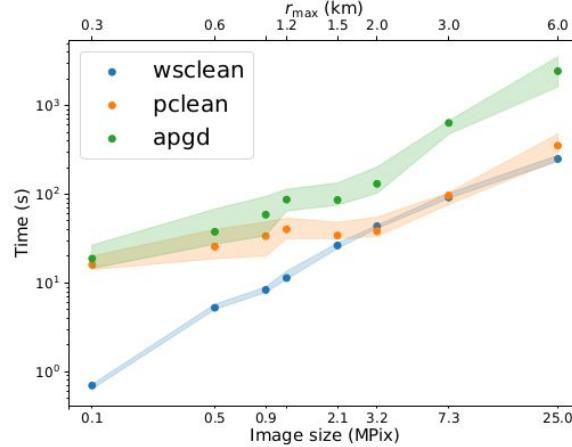


[11] Jarret A et al., "PolyCLEAN: Atomic optimization for super-resolution imaging and uncertainty estimation in radio interferometry", *Astronomy & Astrophysics*, 2025.

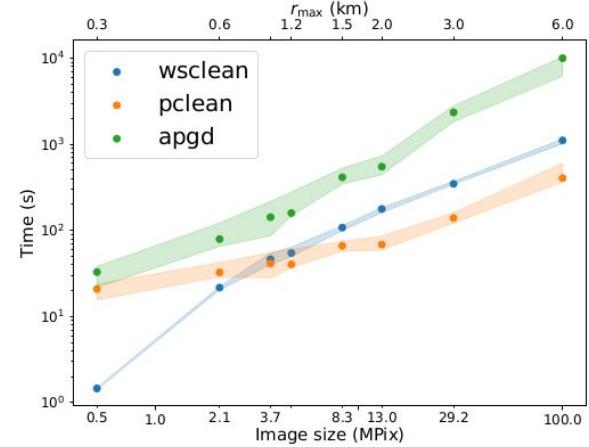
A Fast Reconstruction Method



SRF 2



SRF 5



SRF 10

[11] Jarret A et al., "PolyCLEAN: Atomic optimization for super-resolution imaging and uncertainty estimation in radio interferometry", *Astronomy & Astrophysics*, 2025.

Benefits of Polyatomic Frank-Wolfe

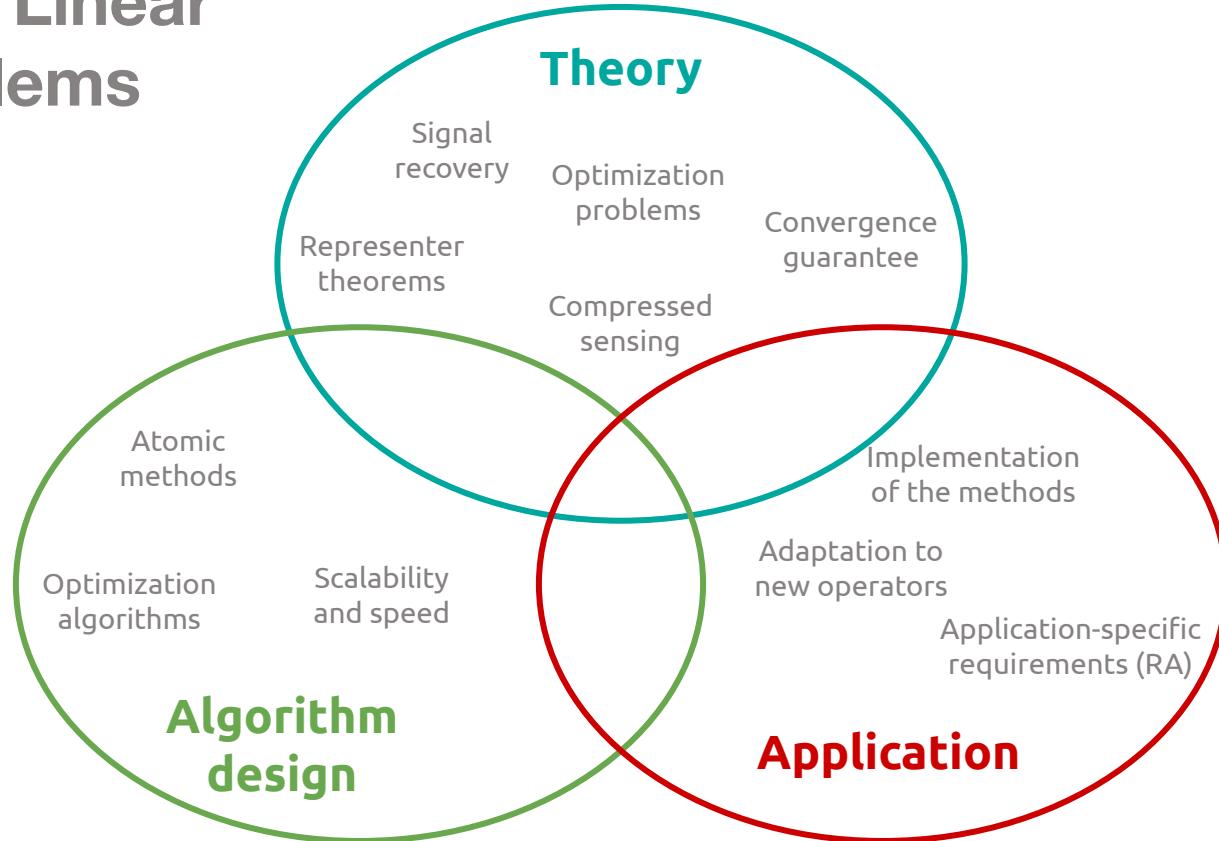
- Polyatomic → **Fast** $\mathcal{S}_k \leftarrow \mathcal{S}_{k-1} \cup \mathcal{I}_k$
- Sparse iterates → **Scalable** $\mathbf{x}_k = \sum_{i=1}^{N_k} \alpha_i^{[k]} \mathbf{e}_i^{[k]}$
- Convergence → **Principled**

Theorem 3.2 (Convergence of Polyatomic FW).^[9]

$$\mathcal{J}(\mathbf{x}_k) - \mathcal{J}^* \leq \frac{2}{k+2} (C_f + 2\delta)$$

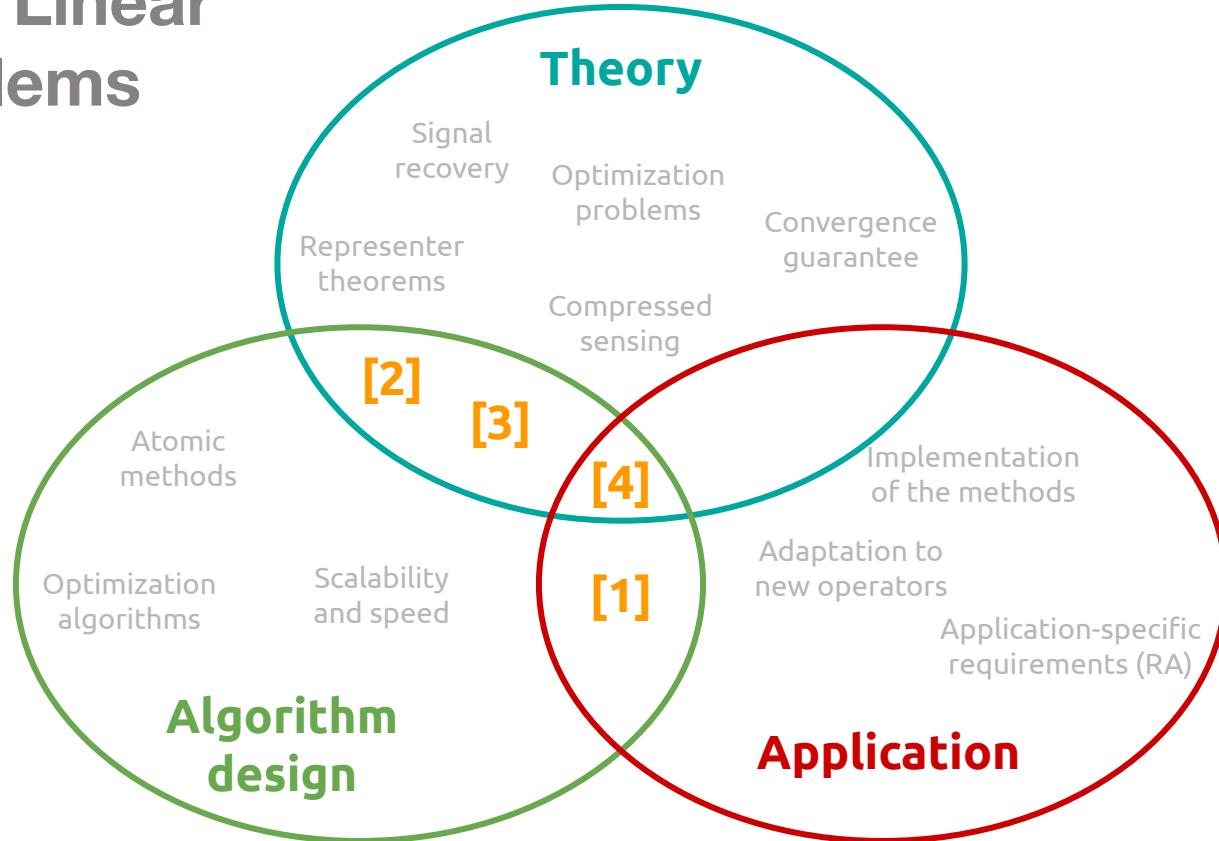
[9] Jarret A, Fageot J, Simeoni M. "A Fast and Scalable Polyatomic Frank-Wolfe Algorithm for the LASSO", *IEEE Signal Processing Letters*, 2022.

Landscape of Linear Inverse Problems



Landscape of Linear Inverse Problems

- [1] PFW
- [2] CD-PFW
- [3] Composite
- [4] PolyCLEAN



Conclusion and perspectives

Mathematics-aware numerical solvers:

- Principled (poly)atomic methods
- Decoupled algorithms
- Sparsity-aware processing

Chapters 3, 5

Chapter 6

Chapter 8

Open questions:

- Resolution of the reconstruction and quantitative imaging
- More advanced *traditional* methods
- Mixed learning-based approaches

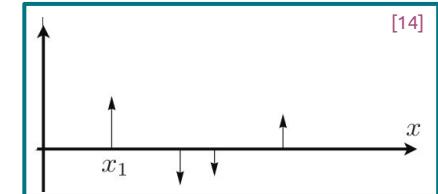
Composite Representer Theorem (in the literature)

$$\arg \min_{s_1, s_2 \in \mathcal{M}(\mathcal{X}) \times L_2(\mathcal{X})} \frac{1}{2} \|y - \Phi(s_1 + s_2)\|_2^2 + \lambda_1 \|s_1\|_{\mathcal{M}} + \frac{\lambda_2}{2} \|s_2\|_{L_2}^2$$

[13]

Representer theorem^[13]:

- $s_1^* \rightarrow \text{Sparse measure}$
 $s_1^* = m[\mathbf{a}^*, \mathbf{x}^*]$
- $s_2^* \rightarrow \text{Smooth quadratic solution}$
 $s_2^* = \Phi^* \mathbf{u}^*$



[13] Debarre T et al. "Continuous-Domain Formulation of Inverse Problems for Composite Sparse-Plus-Smooth Signals", *IEEE Open Journal of Signal Processing*, 2021.

[14] Unser M et al., "Splines Are Universal Solutions of Linear Inverse Problems with Generalized TV Regularization", *SIAM Review*, 2017.