

# Part IA - Dynamics and Relativity

## Theorems with Proof

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### Basic concepts

Space and time, frames of reference, Galilean transformations. Newton's laws. Dimensional analysis. Examples of forces, including gravity, friction and Lorentz. [4]

### Newtonian dynamics of a single particle

Equation of motion in Cartesian and plane polar coordinates. Work, conservative forces and potential energy, motion and the shape of the potential energy function; stable equilibria and small oscillations; effect of damping.

Angular velocity, angular momentum, torque.

Orbits: the  $u(\theta)$  equation; escape velocity; Kepler's laws; stability of orbits; motion in a repulsive potential (Rutherford scattering). Rotating frames: centrifugal and coriolis forces. \*Brief discussion of Foucault pendulum.\* [8]

### Newtonian dynamics of systems of particles

Momentum, angular momentum, energy. Motion relative to the centre of mass; the two body problem. Variable mass problems; the rocket equation. [2]

### Rigid bodies

Moments of inertia, angular momentum and energy of a rigid body. Parallel axes theorem. Simple examples of motion involving both rotation and translation (e.g. rolling). [3]

### Special relativity

The principle of relativity. Relativity and simultaneity. The invariant interval. Lorentz transformations in  $(1+1)$ -dimensional spacetime. Time dilation and length contraction. The Minkowski metric for  $(1+1)$ -dimensional spacetime. Lorentz transformations in  $(3+1)$  dimensions. 4-vectors and Lorentz invariants. Proper time. 4-velocity and 4-momentum. Conservation of 4-momentum in particle decay. Collisions. The Newtonian limit. [7]

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# 1 Newtonian dynamics of particles

## 1.1 Newton's laws of motion

**Law** (Newton's First Law of Motion). A body remains at rest, or moves uniformly in a straight line, unless acted on by a force. (This is in fact Galileo's Law of Inertia)

**Law** (Newton's Second Law of Motion). The rate of change of momentum of a body is equal to the force acting on it (in both magnitude and direction).

**Law** (Newton's Third Law of Motion). To every action there is an equal and opposite reaction: the forces of two bodies on each other are equal and in opposite directions.

## 1.2 Galilean transformations

**Law** (Galilean relativity). The *principle of relativity* asserts that the laws of physics are the same in inertial frames.

## 1.3 Newton's Second Law

**Law.** The *equation of motion* for a particle subject to a force  $F$  is

$$\frac{d\mathbf{p}}{dt} = \mathbf{F},$$

where  $\mathbf{p} = m\mathbf{v} = m\dot{\mathbf{r}}$  is the (linear) momentum of the particle. We say  $m$  is the (inertial) mass of the particle, which is a measure of its reluctance to accelerate.

## 2 Dimensional Analysis

### 2.1 Units

### 2.2 Scaling

## 3 Forces

### 3.1 Force and potential energy in one dimension

**Proposition.** Suppose the equation of a particle satisfies

$$m\ddot{x} = -\frac{dV}{dx}. \quad (*)$$

Then the total energy

$$E = T + V = \frac{1}{2}m\dot{x}^2 + V(x)$$

is conserved, i.e.  $\dot{E} = 0$ .

*Proof.*

$$\begin{aligned} \frac{dE}{dt} &= m\dot{x}\ddot{x} + \frac{dV}{dx}\dot{x} \\ &= \dot{x} \left( m\ddot{x} + \frac{dV}{dx} \right) \\ &= 0 \end{aligned}$$

□

### 3.2 Motion in a potential

### 3.3 Equilibrium points

### 3.4 Force and potential energy in three dimensions

**Proposition.** If  $\mathbf{F}$  is conservative, then the energy

$$\begin{aligned} E &= T + V \\ &= \frac{1}{2}m|\mathbf{v}|^2 + V(\mathbf{r}) \end{aligned}$$

is conserved. Then the work done is equal to the change in potential energy, and is independent of the path taken between the end points.

In particular, if we travelled around a closed loop, no work is done.

*Proof.*

$$\begin{aligned} \frac{dE}{dt} &= m\ddot{\mathbf{r}} \cdot \dot{\mathbf{r}} + \frac{\partial V}{\partial x_i} \frac{dx_i}{dt} \\ &= (m\ddot{\mathbf{r}} + \nabla V) \cdot \dot{\mathbf{r}} \\ &= (m\ddot{\mathbf{r}} - \mathbf{F}) \cdot \dot{\mathbf{r}} \\ &= 0 \end{aligned}$$

In this case, the work done is

$$W = \int_C \mathbf{F} \cdot d\mathbf{r} = - \int_C (\nabla V) \cdot d\mathbf{r} = V(\mathbf{r}_1) - V(\mathbf{r}_2).$$

□

### 3.5 Central forces

**Proposition.**  $\nabla r = \hat{\mathbf{r}}$ .

*Proof.* We know that

$$r^2 = x_1^2 + x_2^2 + x_3^2.$$

Then

$$2r \frac{\partial r}{\partial x_i} = 2x_i.$$

So

$$\frac{\partial r}{\partial x_i} = \frac{x_i}{r} = (\hat{\mathbf{r}})_i.$$

□

**Proposition.** Let  $\mathbf{F} = -\nabla V(r)$  be a central force. Then

$$\mathbf{F} = -\nabla V = -\frac{dV}{dr} \hat{\mathbf{r}},$$

where  $\hat{\mathbf{r}} = \mathbf{r}/r$  is the unit vector in the radial direction pointing away from the origin.

*Proof.* Continuing the proof above,

$$(\nabla V)_u = \frac{\partial V}{\partial x_i} = \frac{dV}{dr} \frac{\partial r}{\partial x_i} = \frac{dV}{dr} (\hat{\mathbf{r}})_i$$

□

**Proposition.** Angular momentum is conserved by a central force.

*Proof.*

$$\frac{d\mathbf{L}}{dt} = m\dot{\mathbf{r}} \times \dot{\mathbf{r}} + m\mathbf{r} \times \ddot{\mathbf{r}} = \mathbf{0} + \mathbf{r} \times \mathbf{F} = \mathbf{0}.$$

where the last equality comes from the fact that  $\mathbf{F}$  is parallel to  $\mathbf{r}$  for a central force. □

### 3.6 Gravity

**Law.** If a particle of mass  $M$  is fixed at a origin, then a second particle of mass  $m$  experiences a potential energy

$$V(r) = -\frac{GMm}{r},$$

where  $G \approx 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$  is the *gravitational constant*.

The gravitational force experienced is then

$$\mathbf{F} = -\nabla V = -\frac{GMm}{r^2} \hat{\mathbf{r}}.$$

**Proposition.** The gravitational potential due to many fixed masses  $M_i$  at points  $\mathbf{r}_i$  is

$$\Phi_g(\mathbf{r}) = -\sum_i \frac{GM_i}{|\mathbf{r} - \mathbf{r}_i|}.$$

Again,  $V = m\Phi_g$  for a particle of mass  $m$ .

**Proposition.** The external gravitational potential of a spherically symmetric object of mass  $M$  is the same as that of a point particle with the same mass at the center of the object, i.e.

$$\Phi_g(r) = -\frac{GM}{r}.$$

*Proof.* c.f. Vector Calculus

□