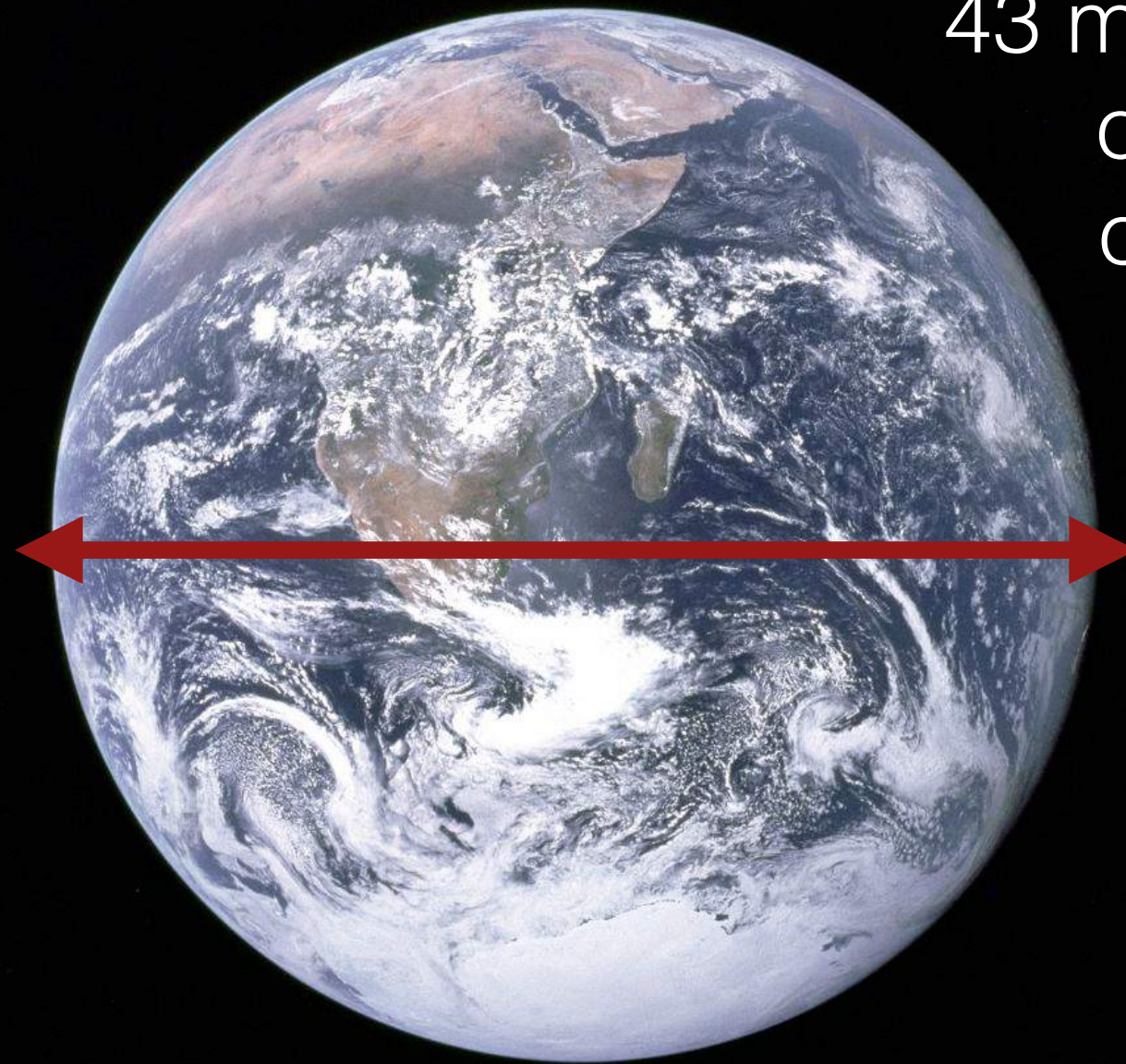


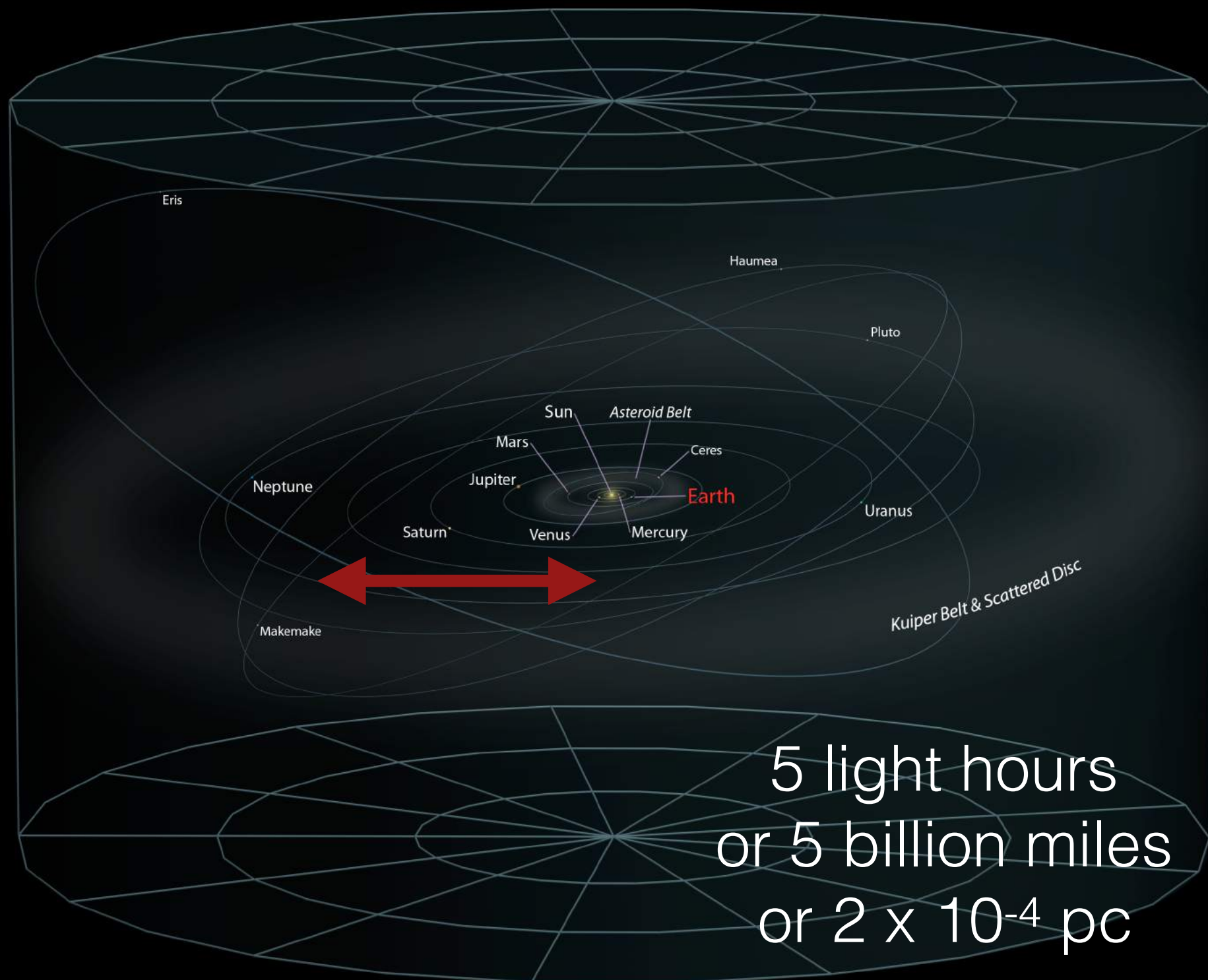
How big is our
Universe?

The Earth

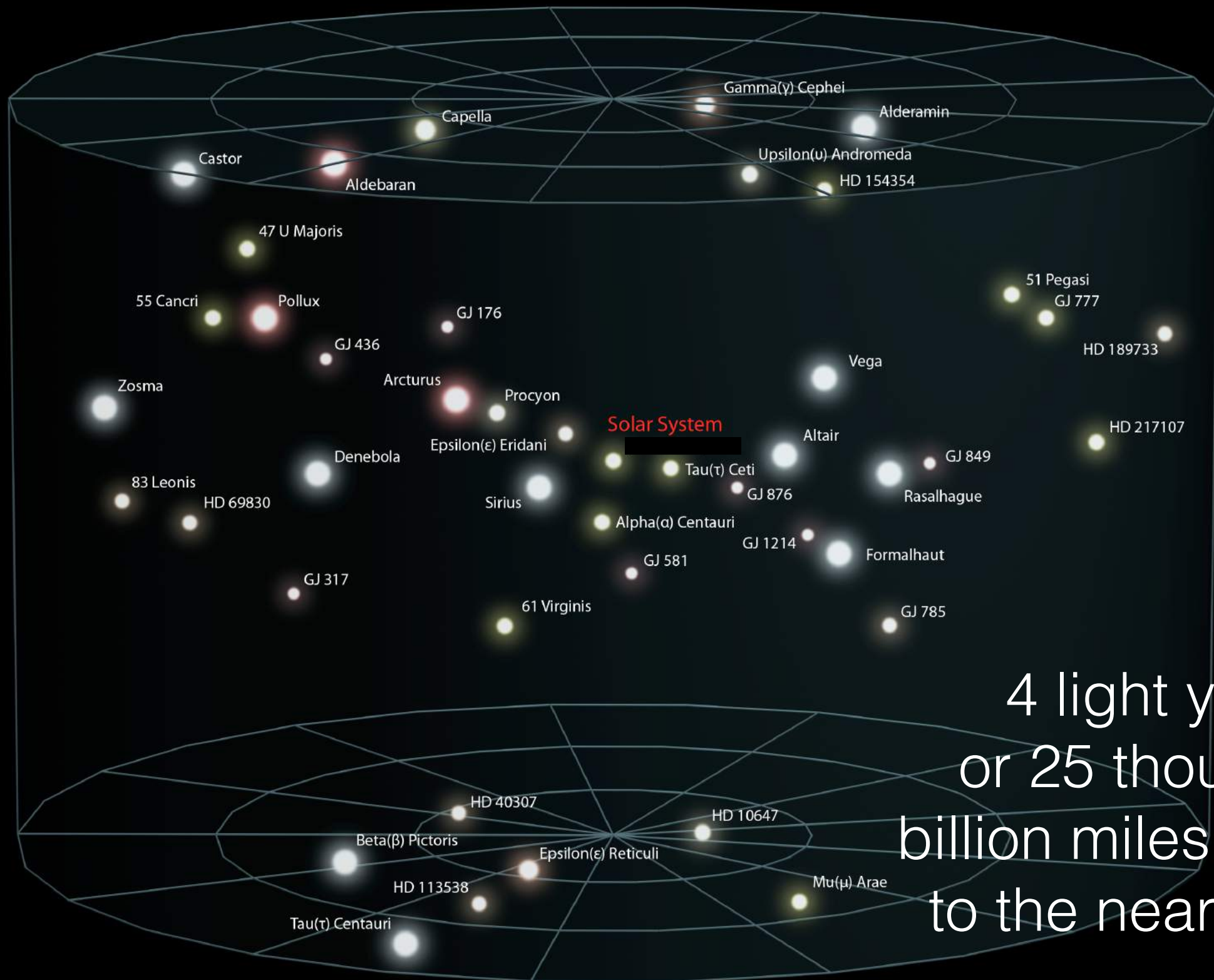


43 milli-light seconds
or 8000 miles
or 4×10^{-7} pc

The Solar System

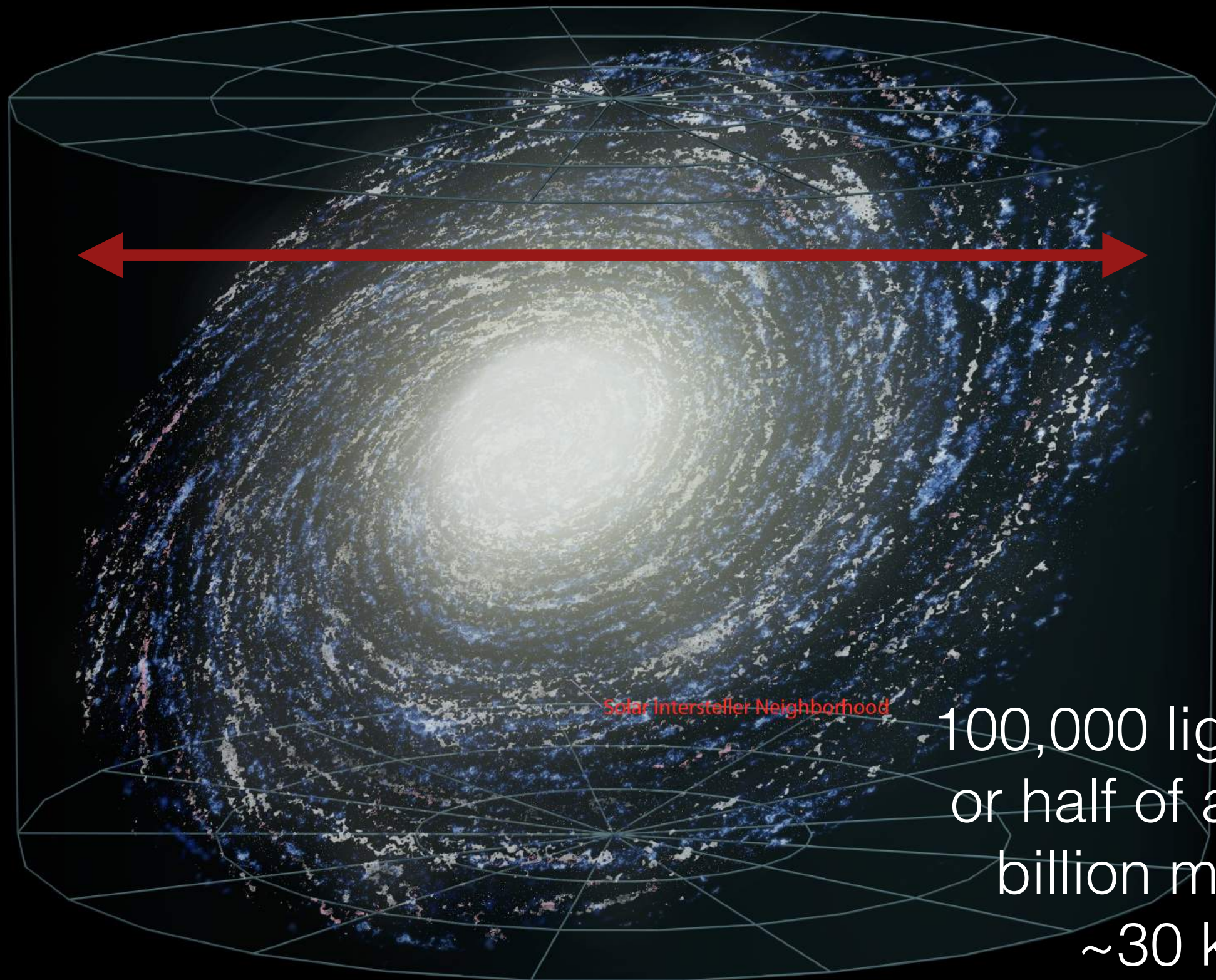


Nearby Stars



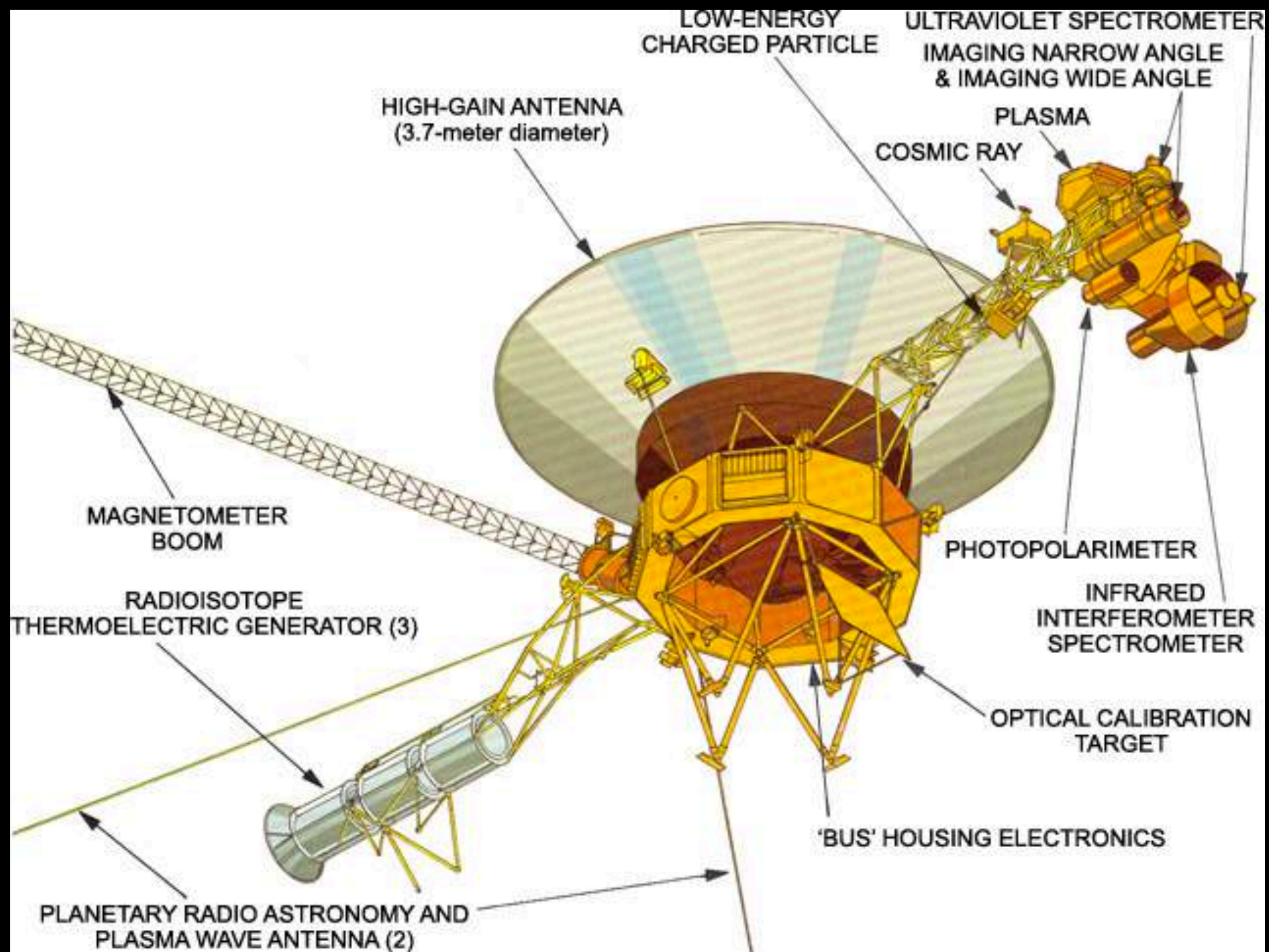
4 light years
or 25 thousand,
billion miles or ~1pc
to the nearest star

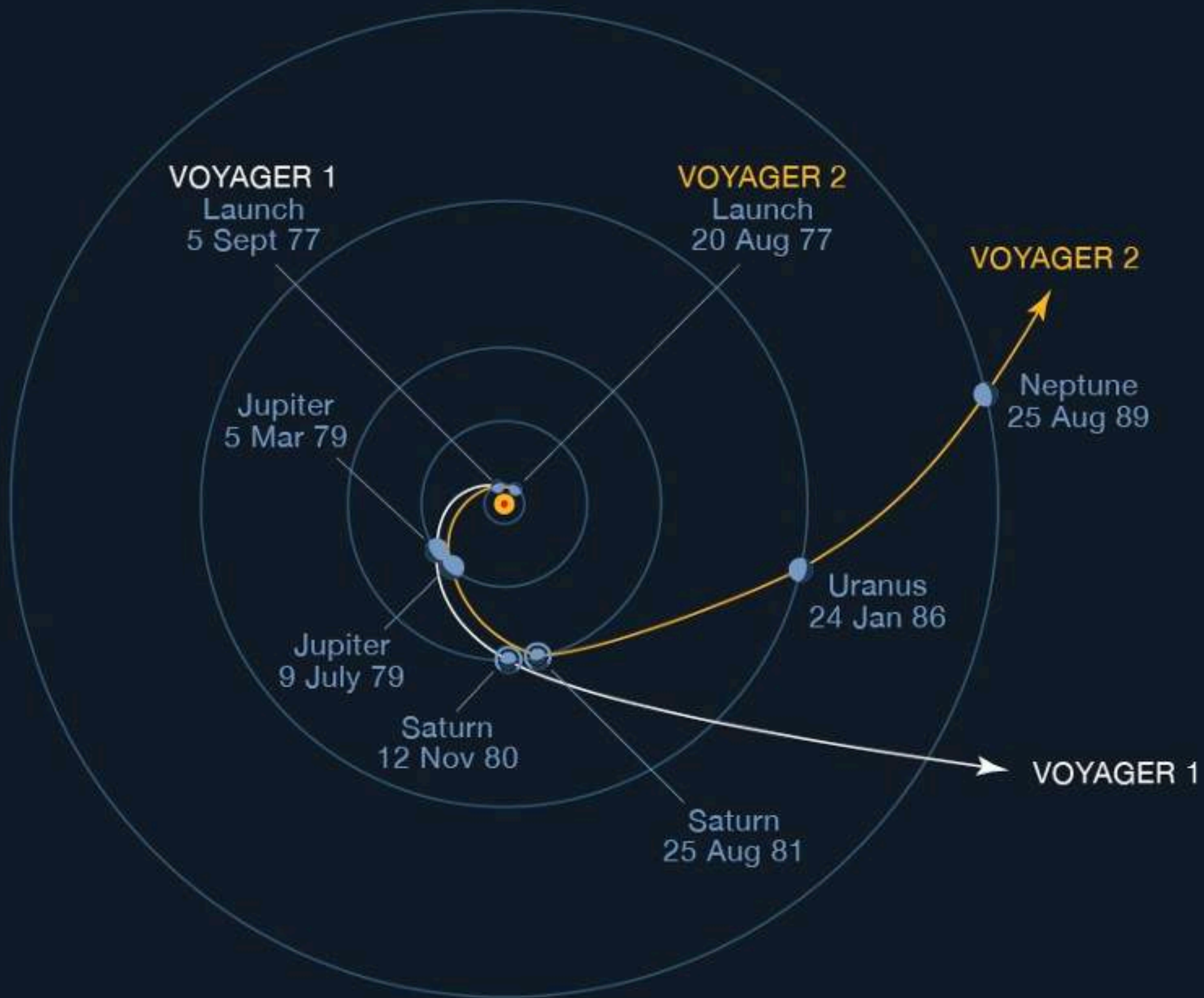
The Milky Way

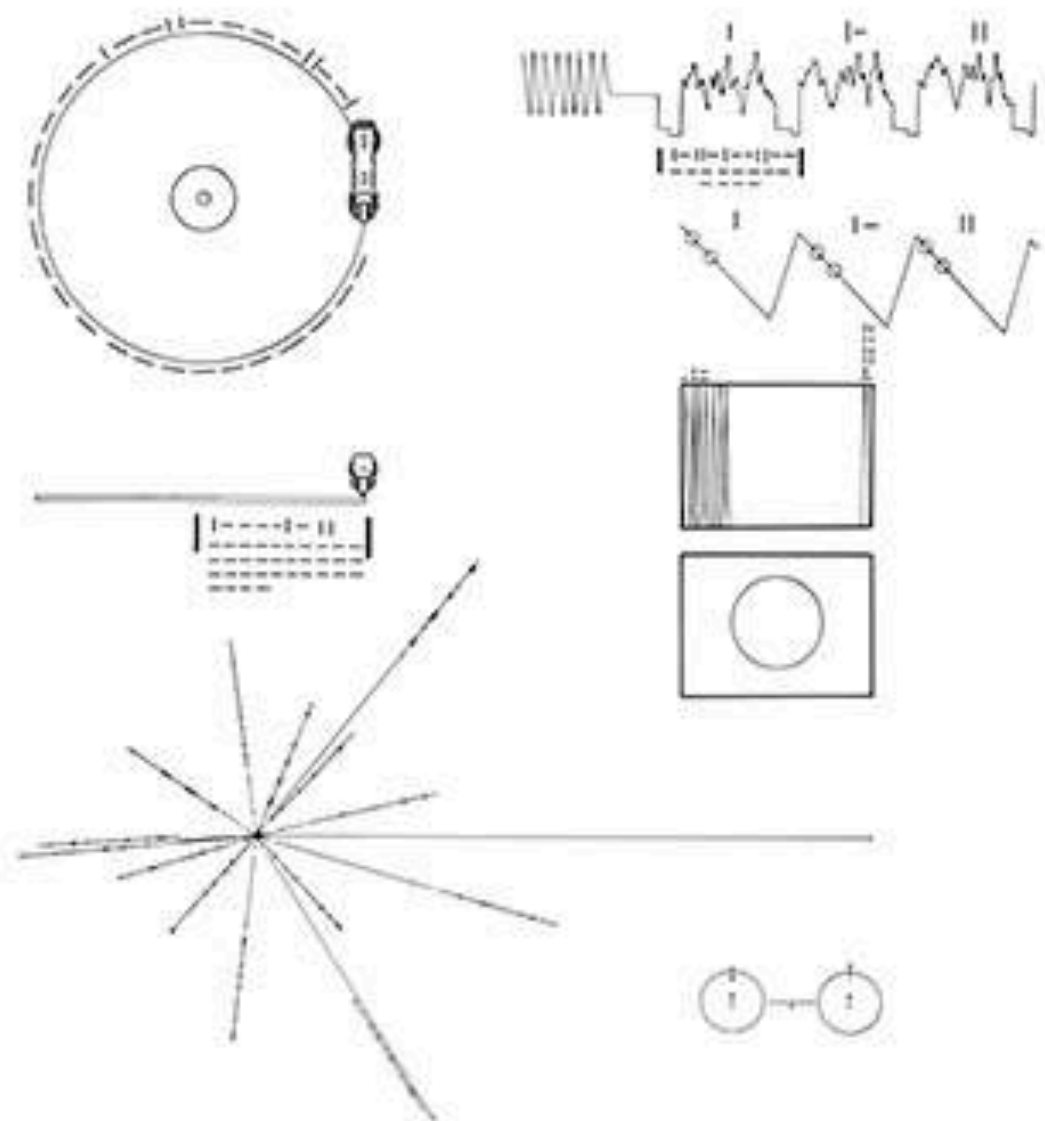
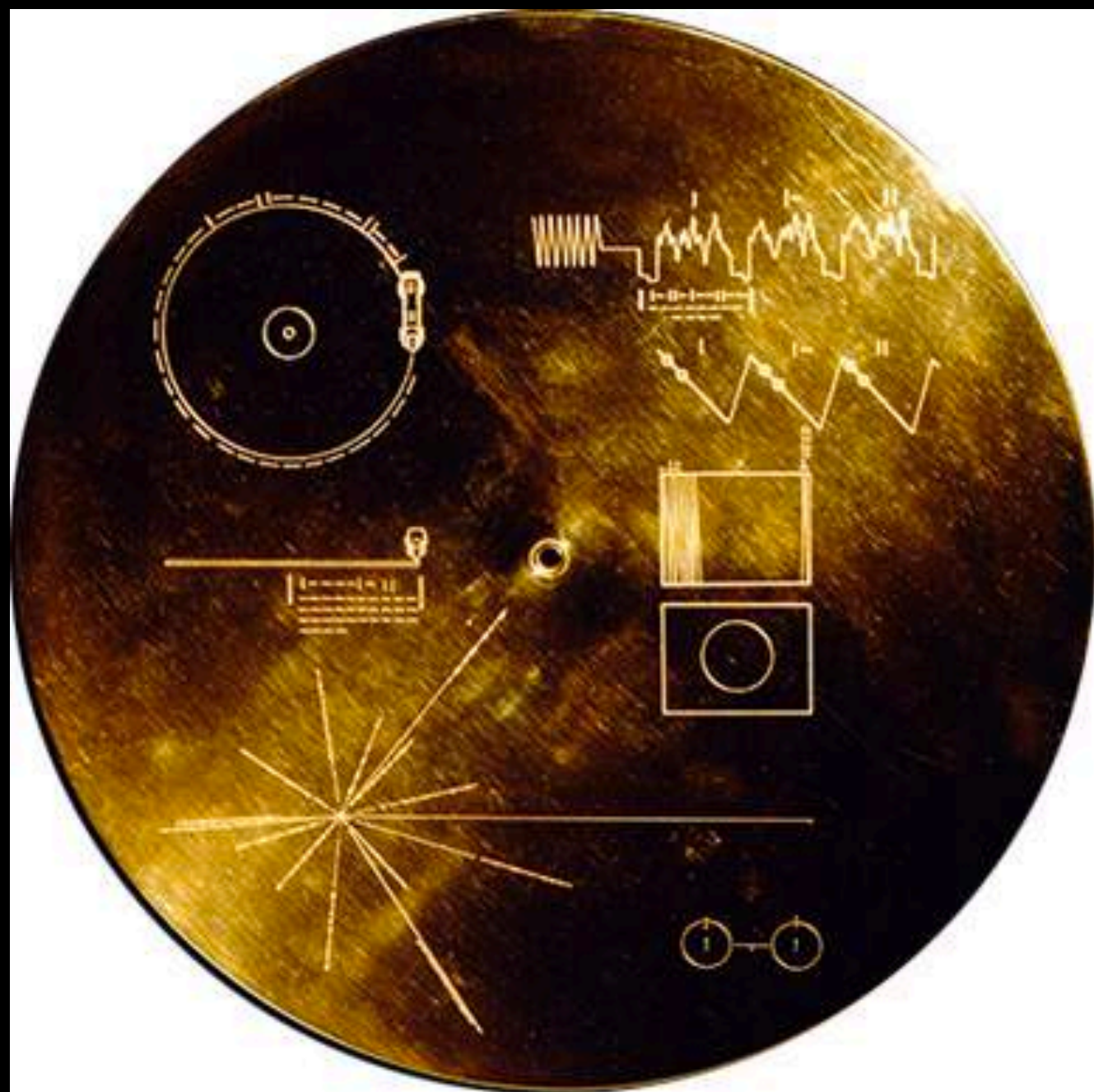


Solar Interstellar Neighborhood

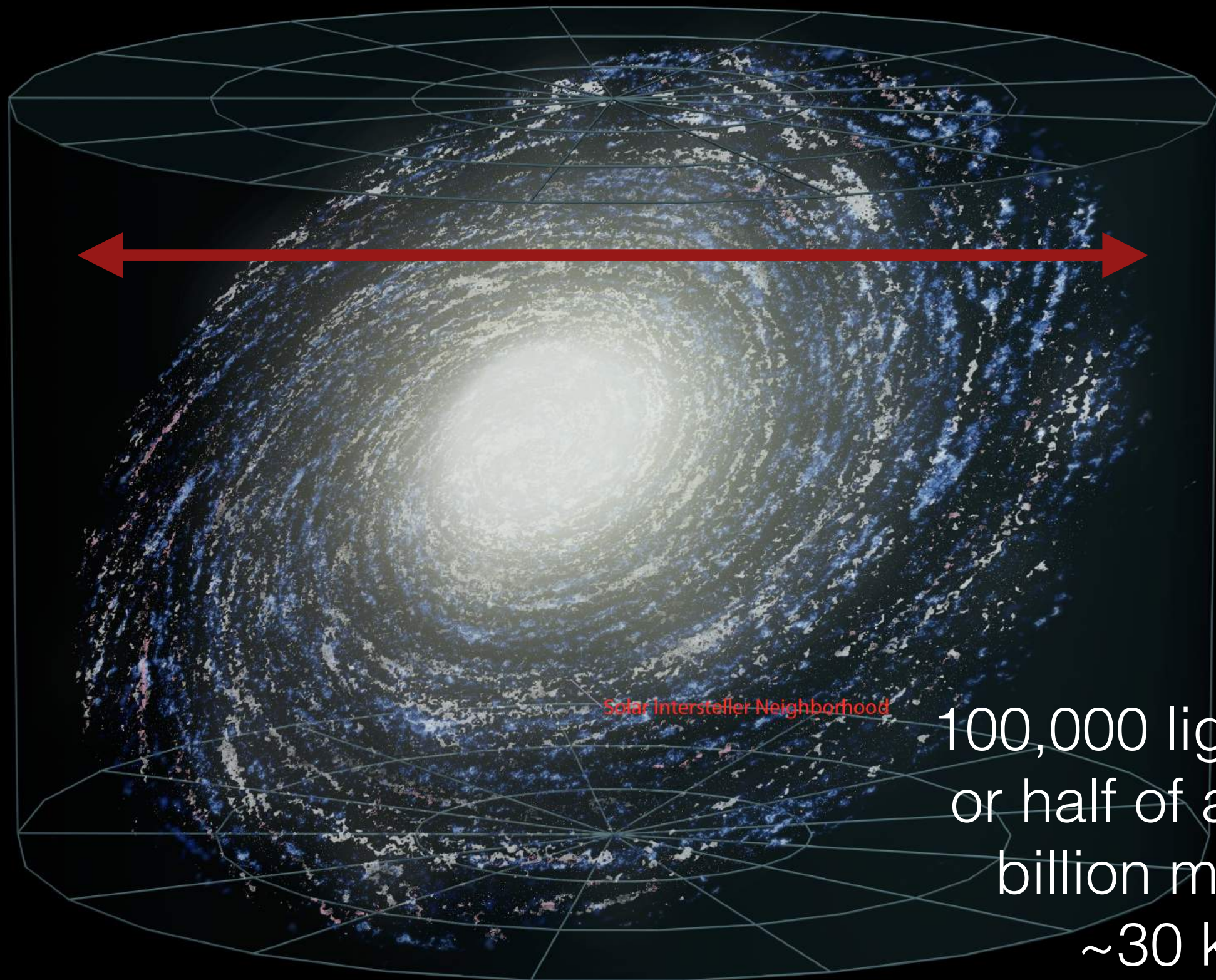
100,000 light years
or half of a billion,
billion miles or
~30 kpc







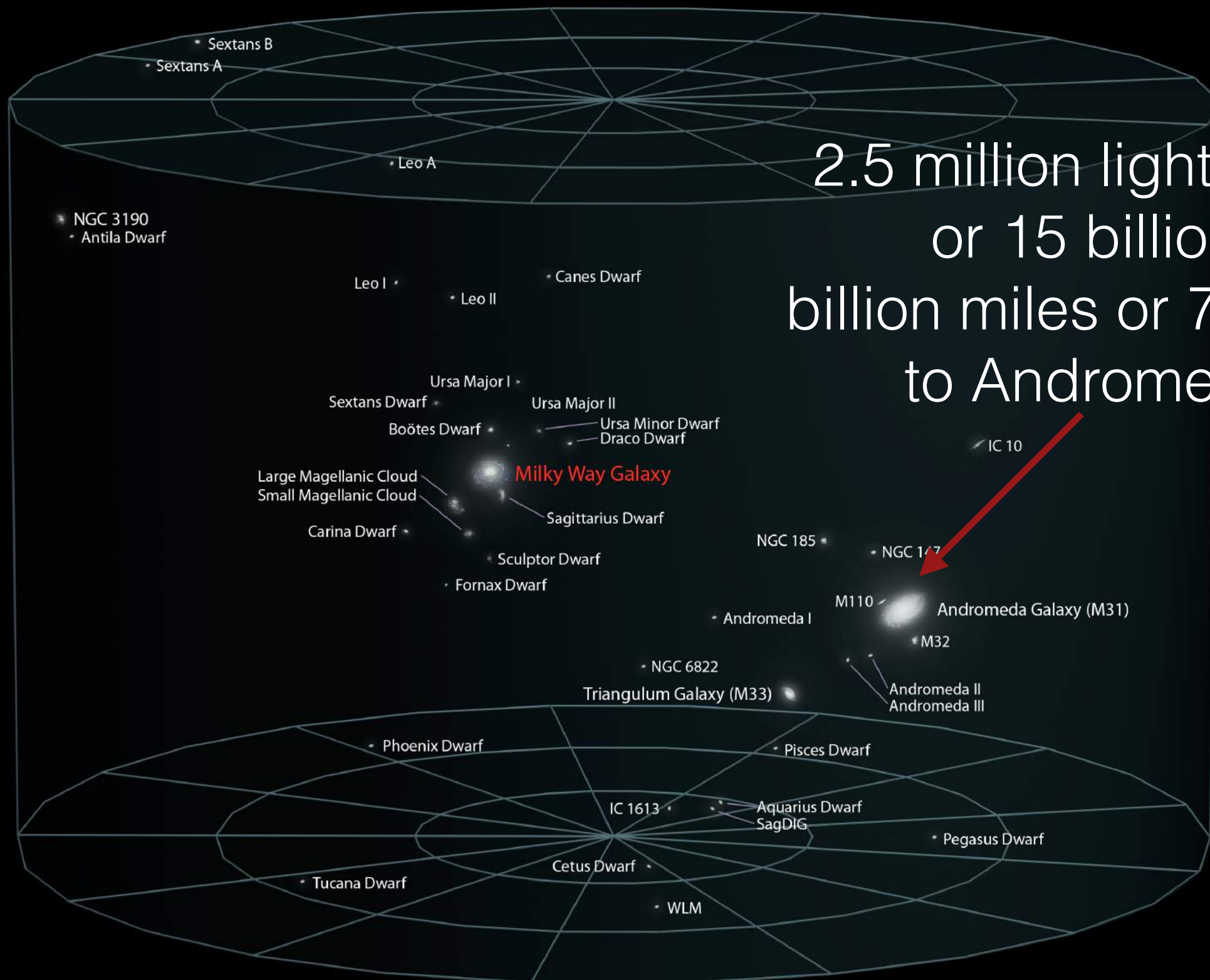
The Milky Way



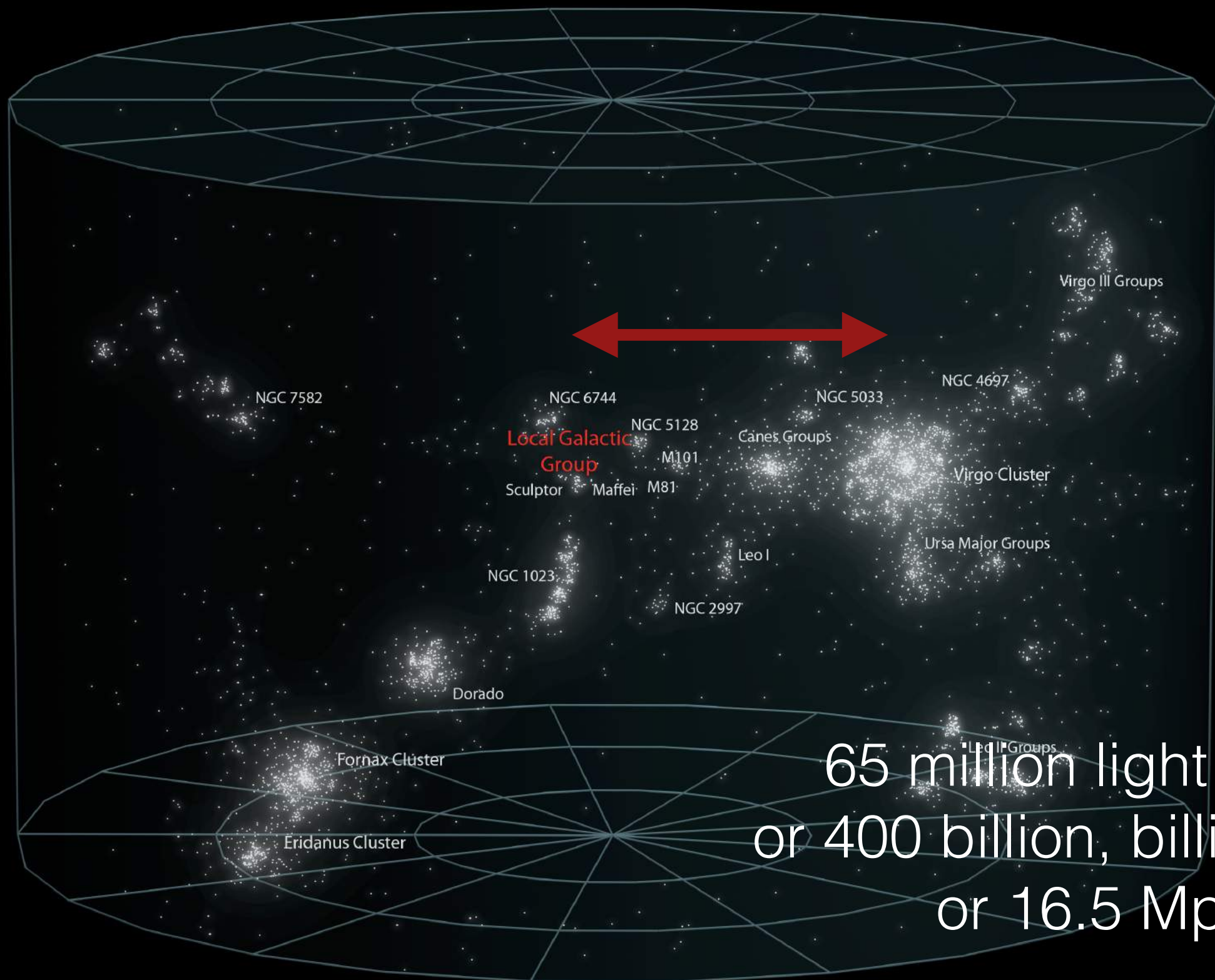
Solar Interstellar Neighborhood

100,000 light years
or half of a billion,
billion miles or
~30 kpc

The Local Group

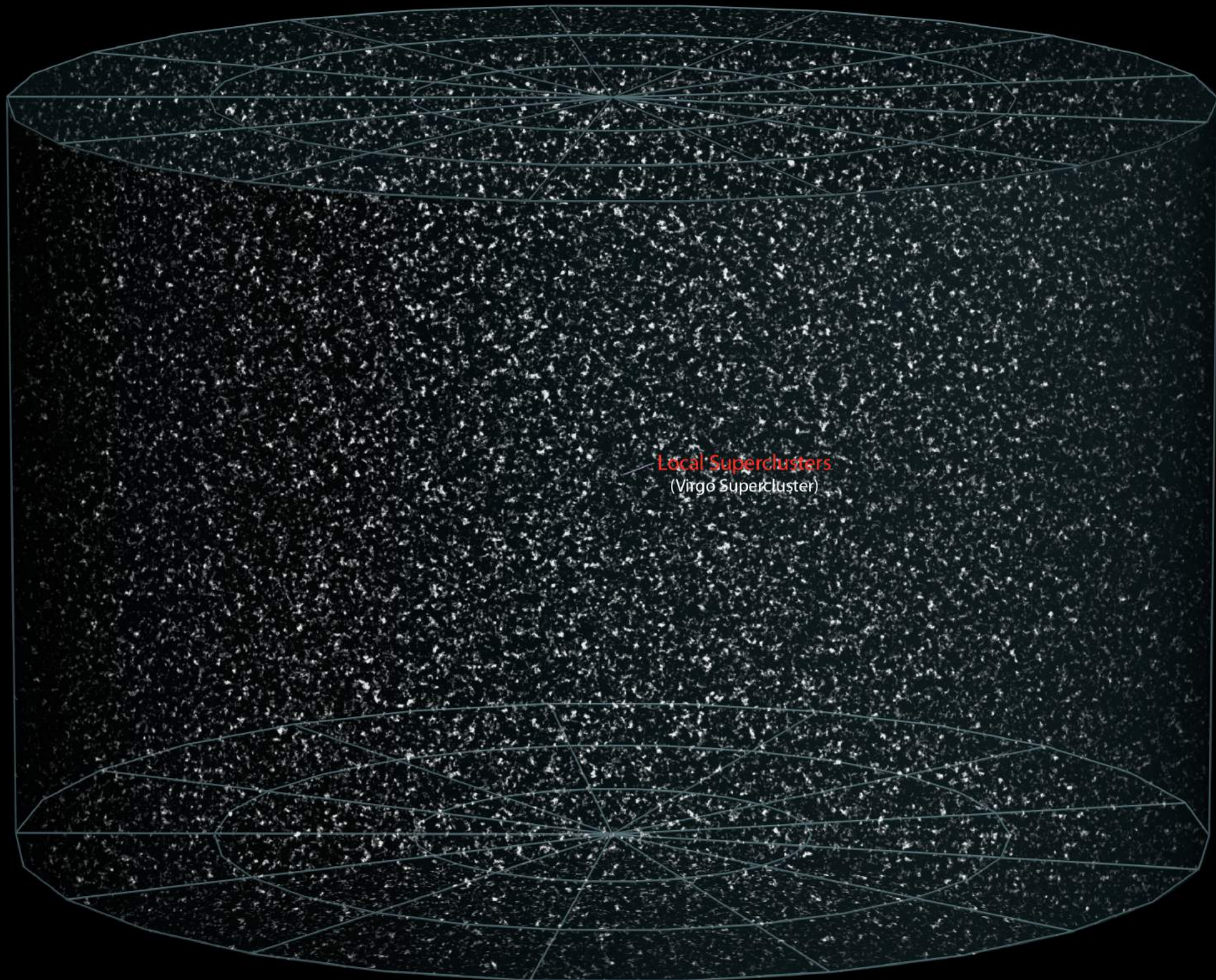


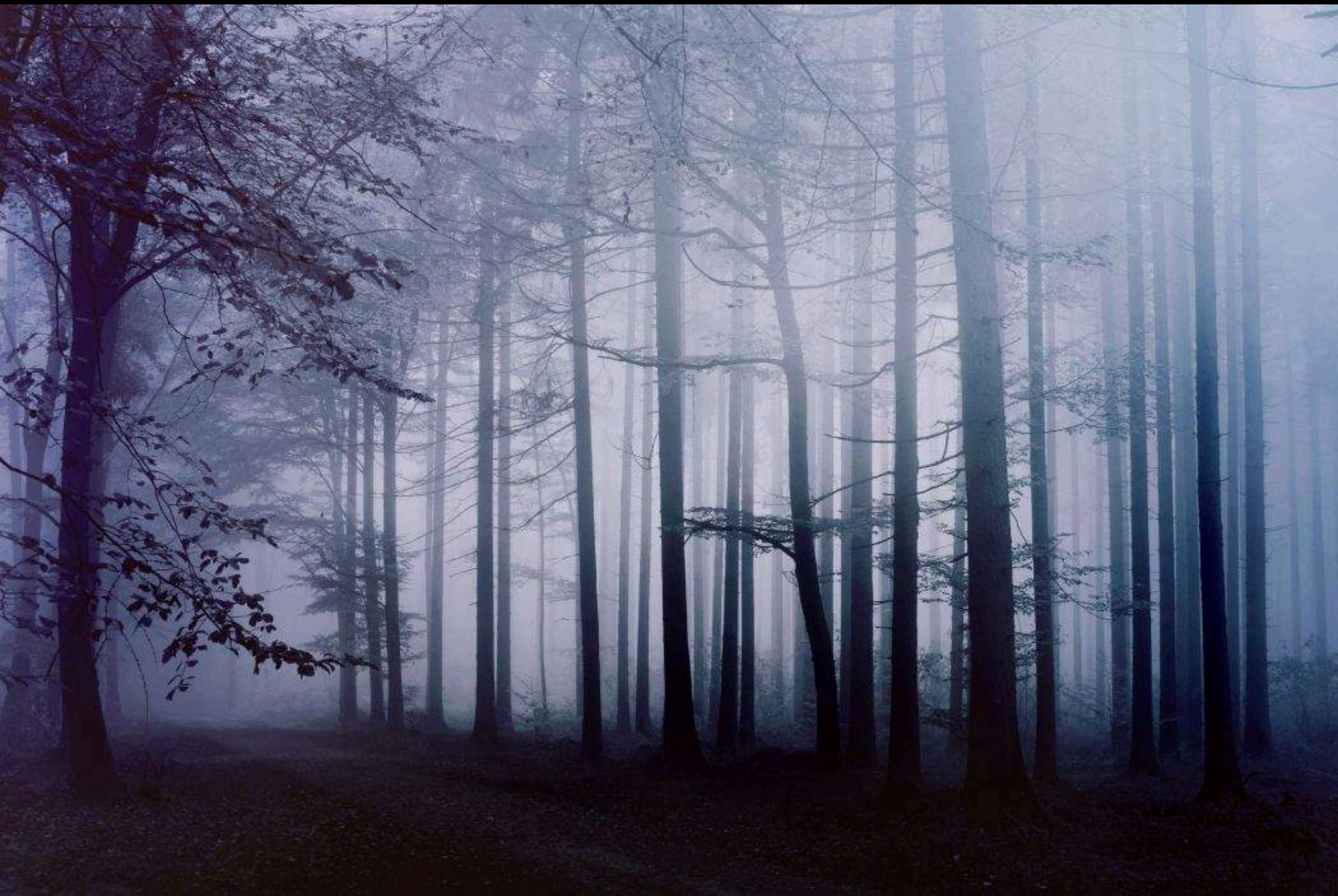
Virgo Supercluster



65 million light years
or 400 billion, billion miles
or 16.5 Mpc

We are just one of many









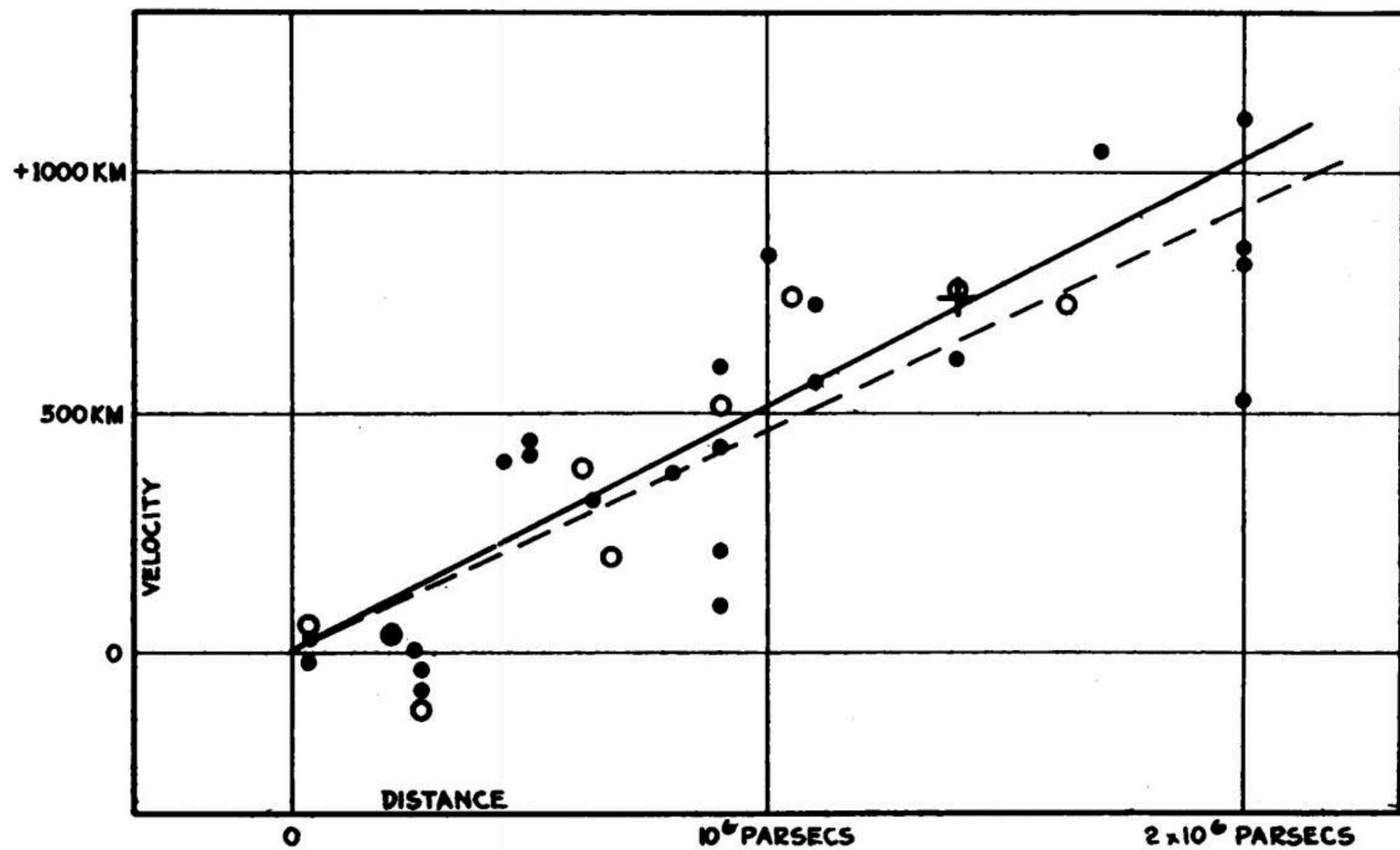
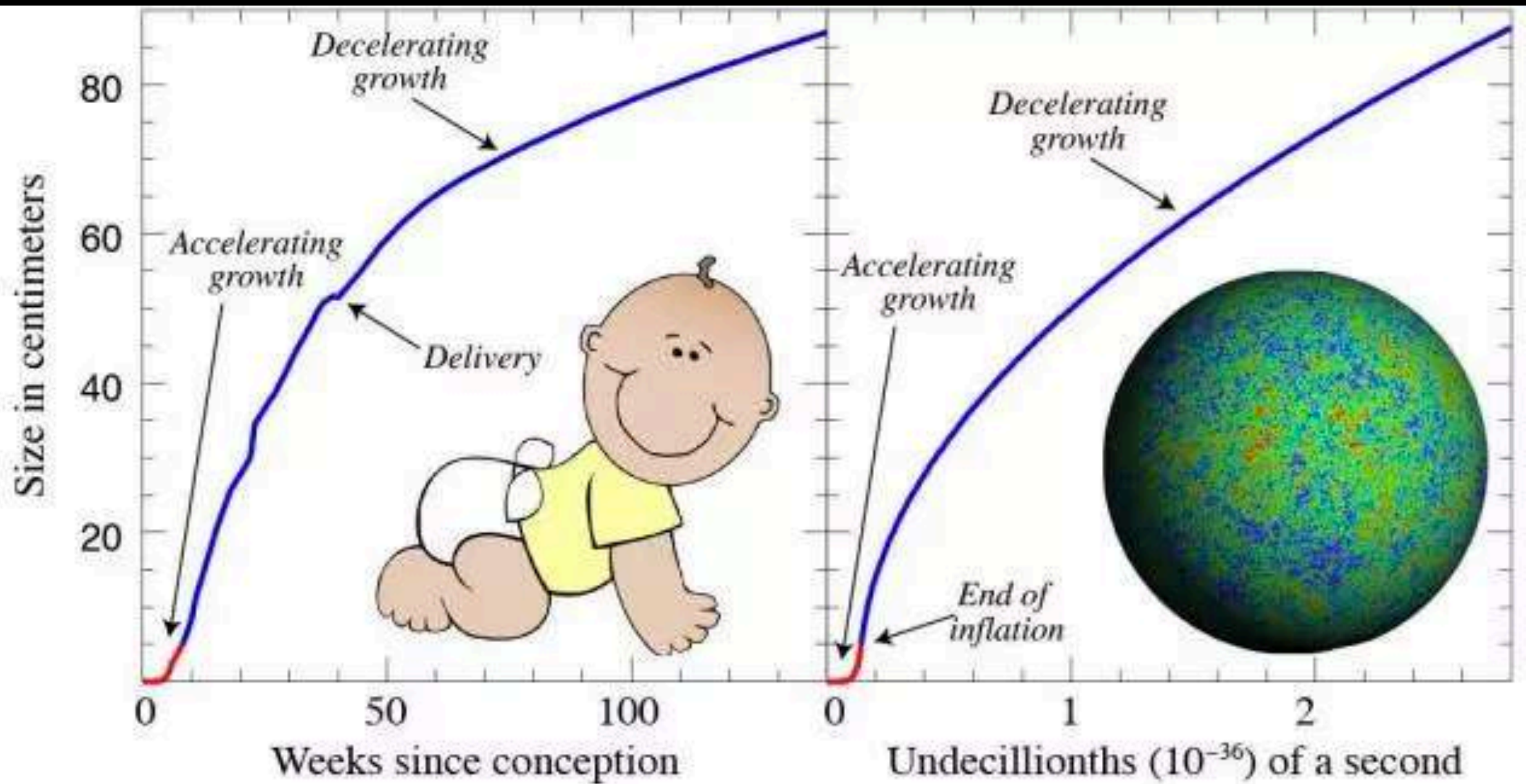


FIGURE 1

Velocity-Distance Relation among Extra-Galactic Nebulae.

Hubble (1929)



Our Mathematical Universe by Max Tegmark

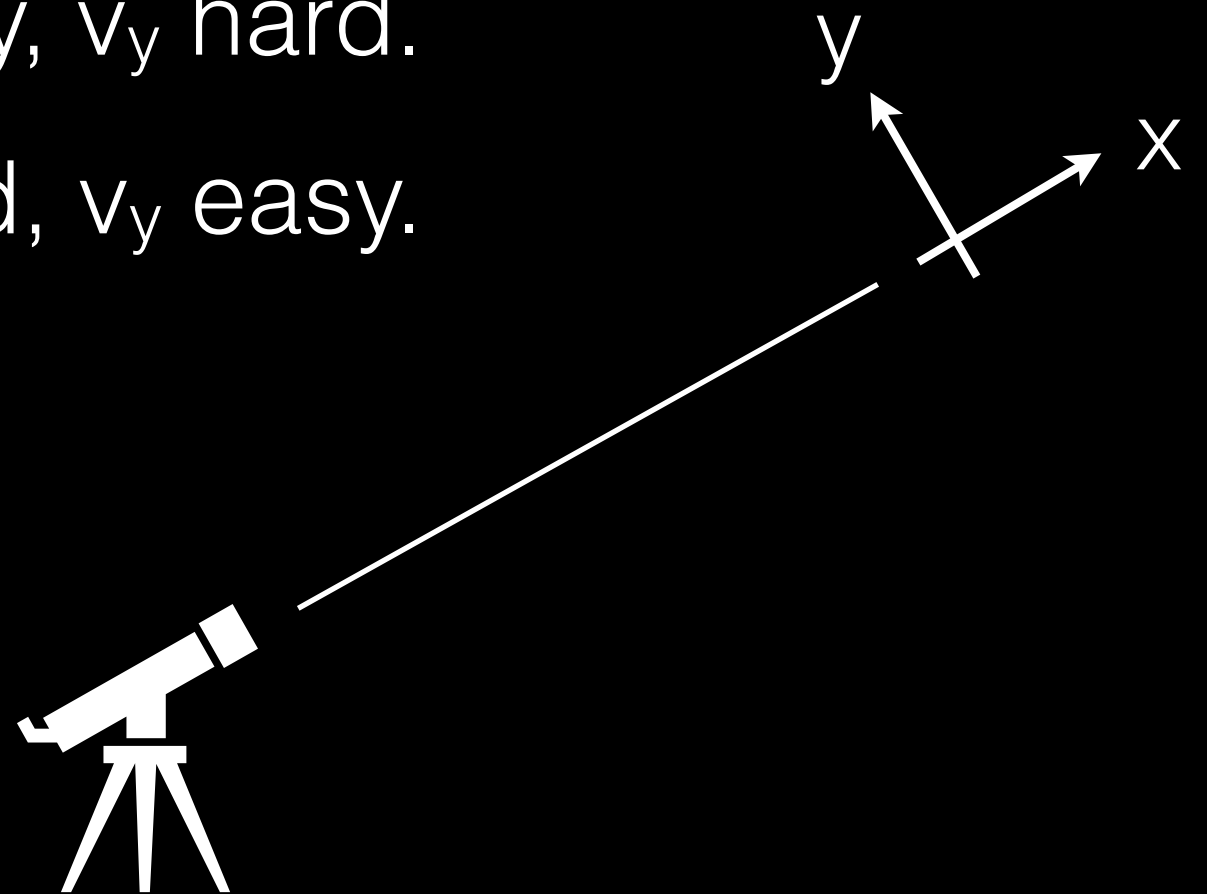
Which of components of a distant galaxy's position and velocity are easy/hard to observe?

A. d_x easy, d_y hard; v_x easy, v_y hard.

B. d_x easy, d_y hard; v_x hard, v_y easy.

C. d_x hard, d_y easy; v_x easy, v_y hard.

D. d_x hard, d_y easy; v_x hard, v_y easy.



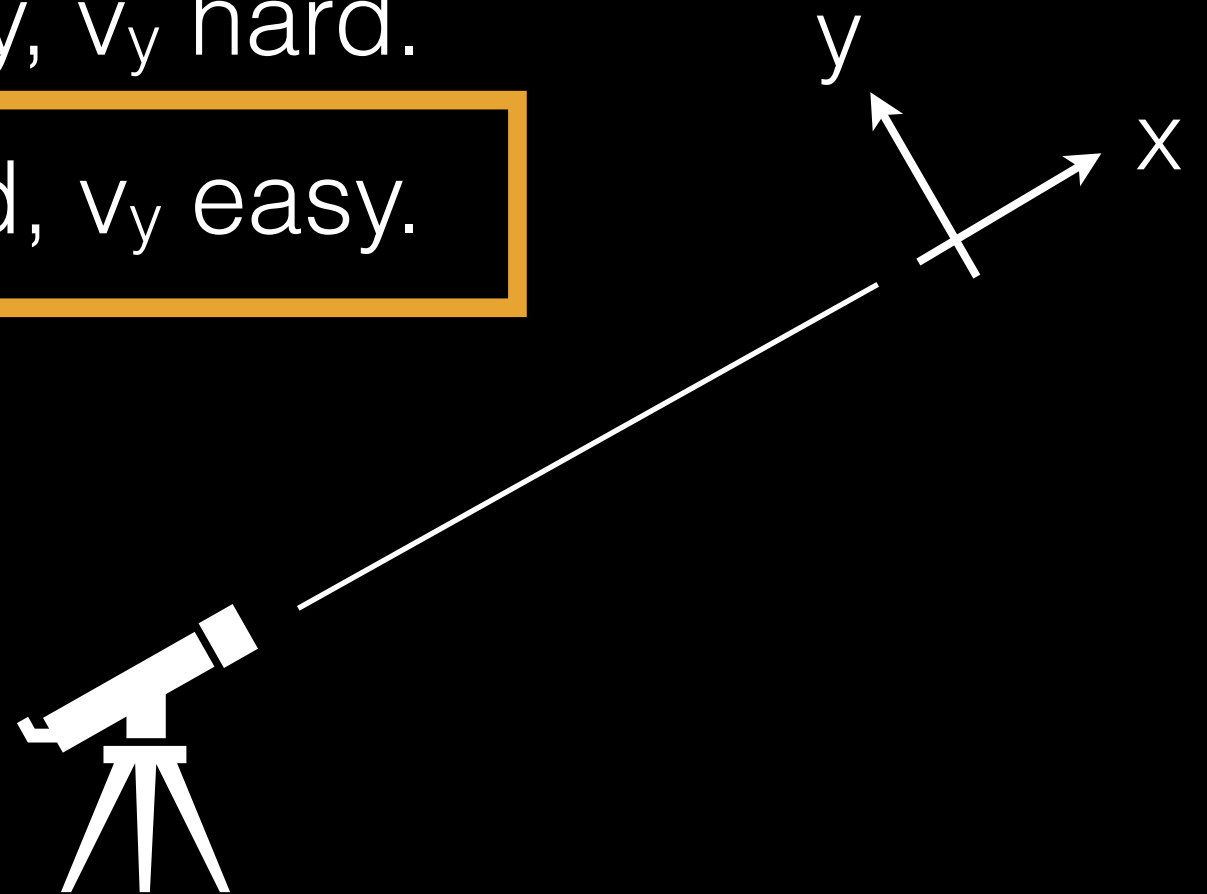
Which of components of a distant galaxy's position and velocity are easy/hard to observe?

A. d_x easy, d_y hard; v_x easy, v_y hard.

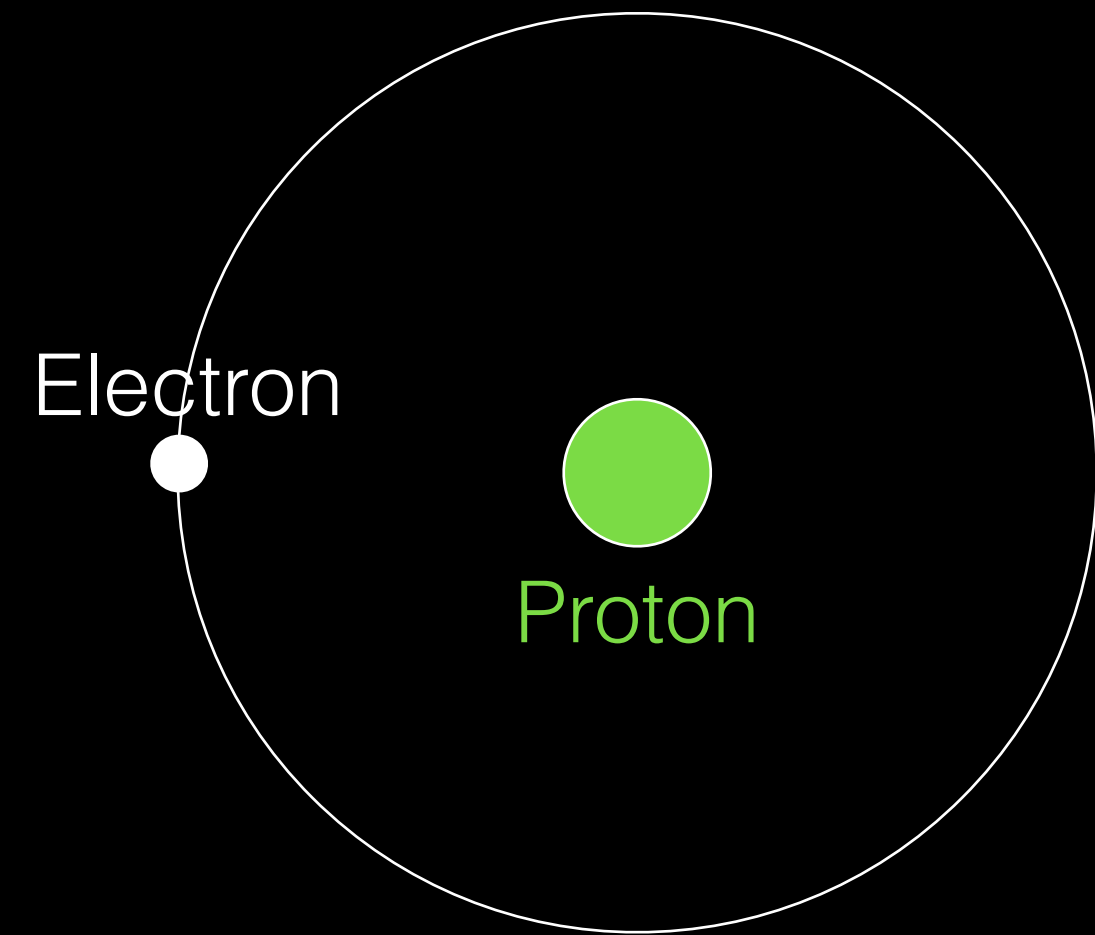
B. d_x easy, d_y hard; v_x hard, v_y easy.

C. d_x hard, d_y easy; v_x easy, v_y hard.

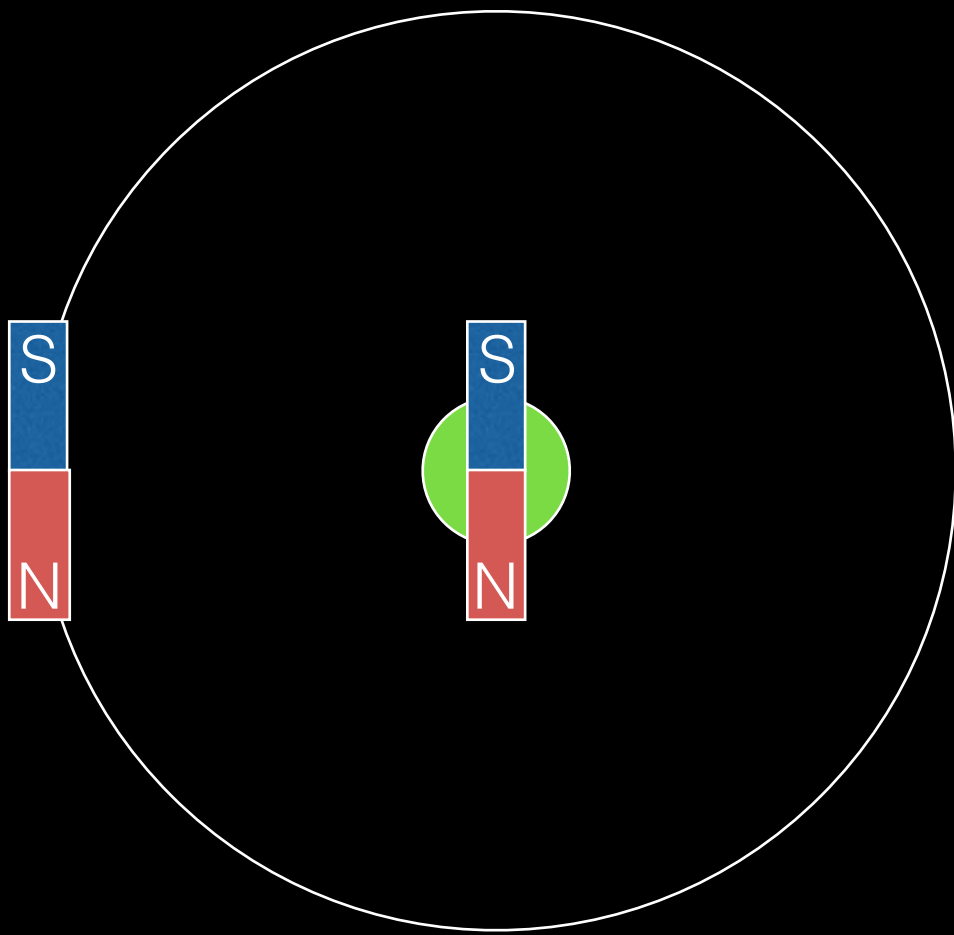
D. d_x hard, d_y easy; v_x hard, v_y easy.



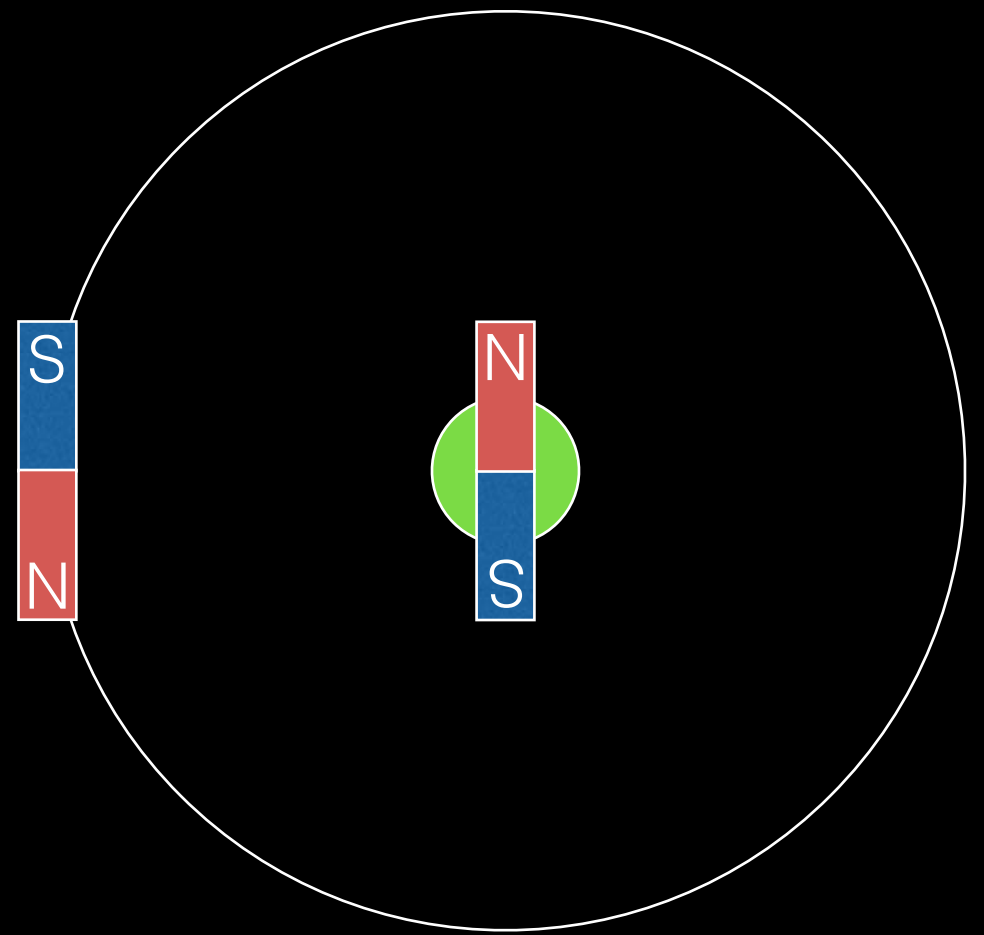
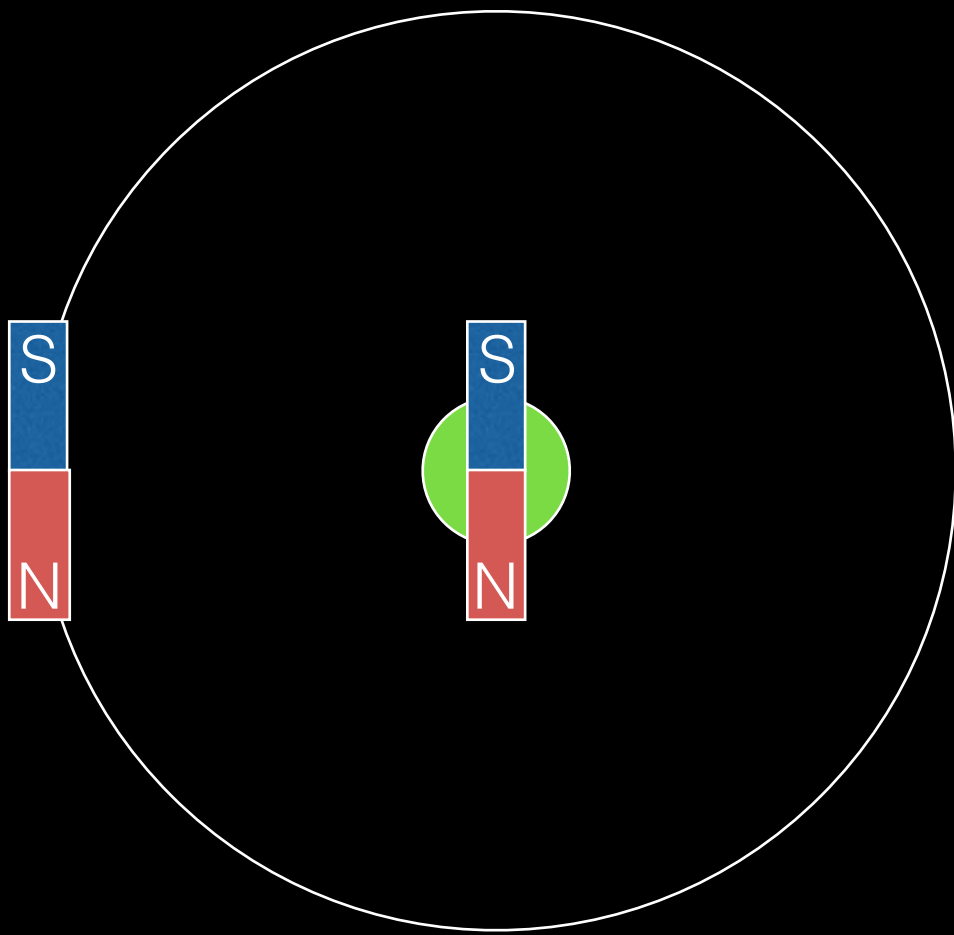
Hydrogen atoms emit radio waves with 21cm wavelengths



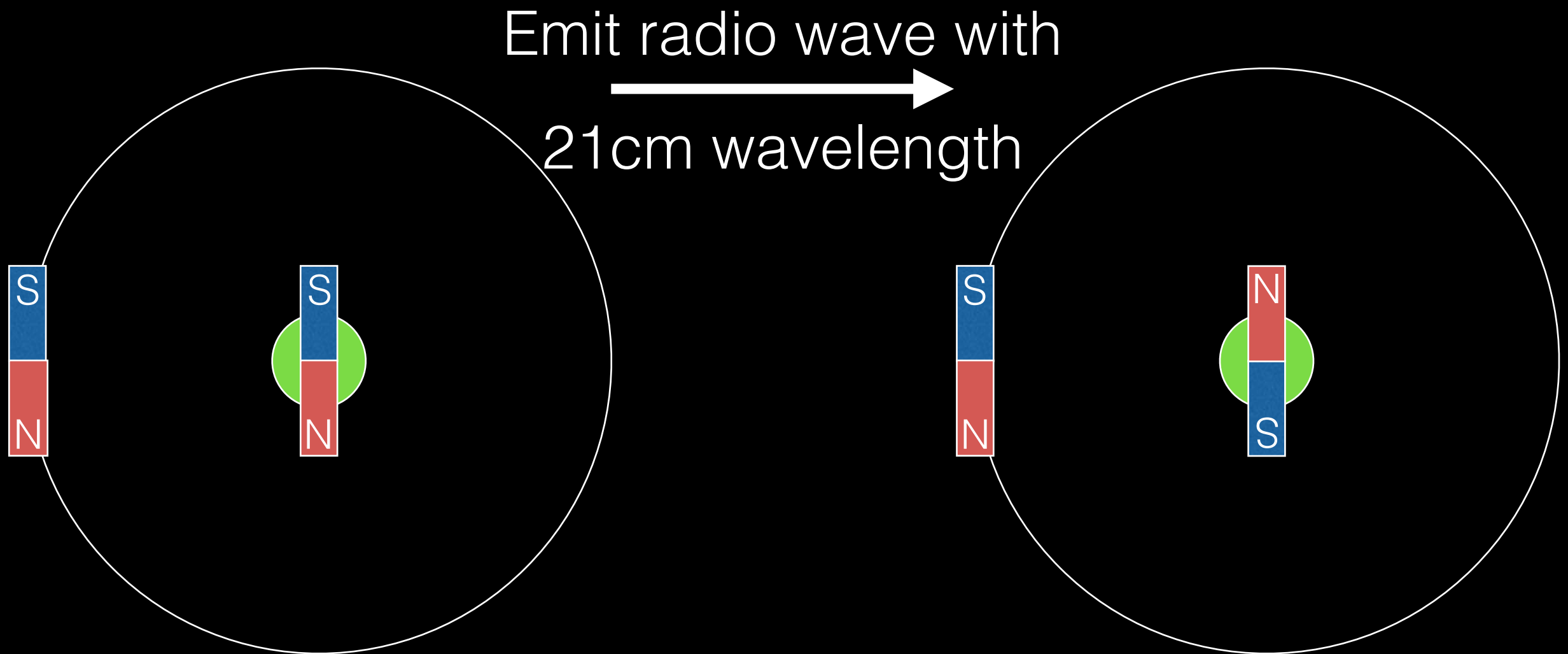
Hydrogen atoms emit radio waves with 21cm wavelengths



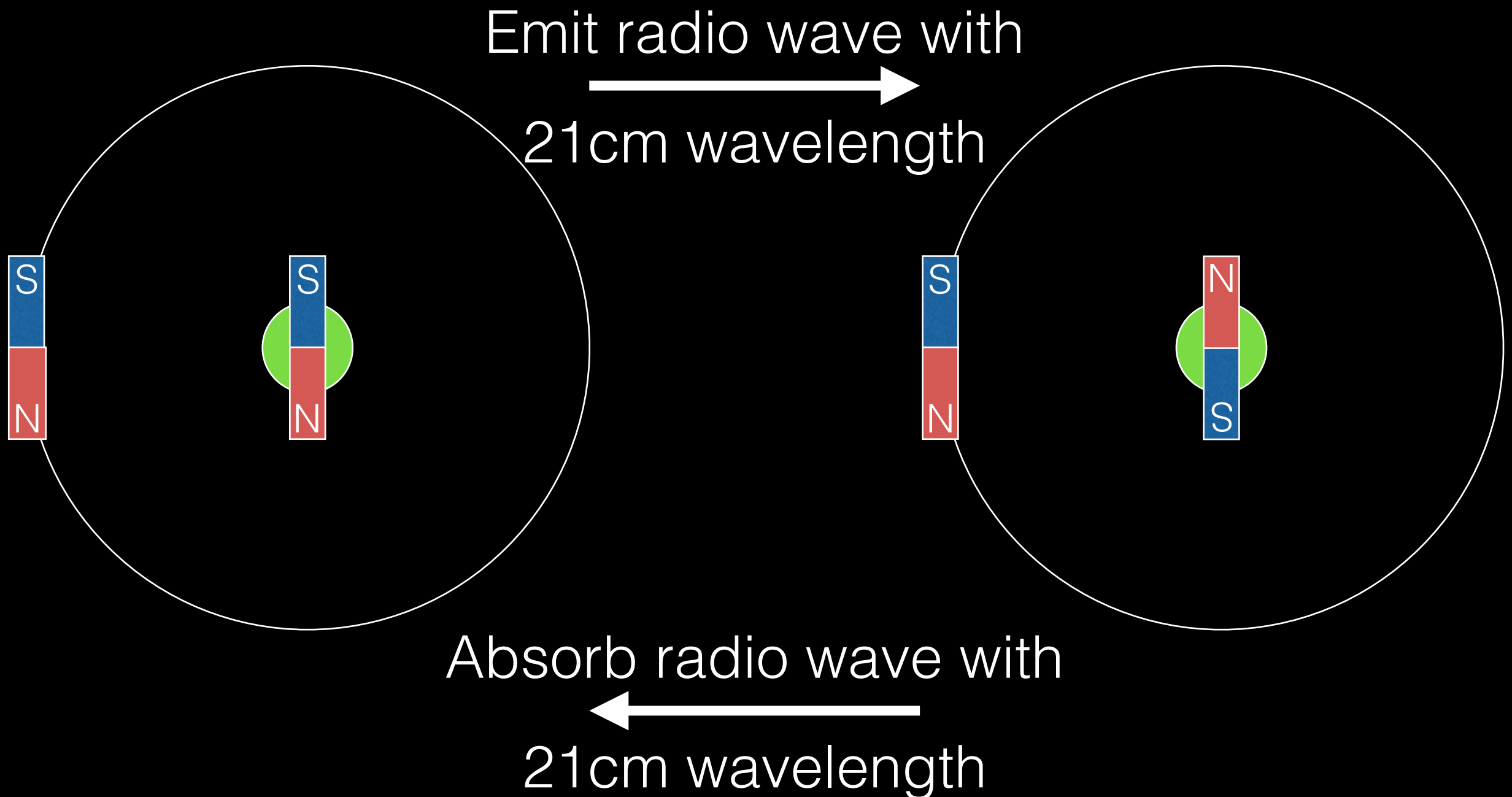
Hydrogen atoms emit radio waves with 21cm wavelengths

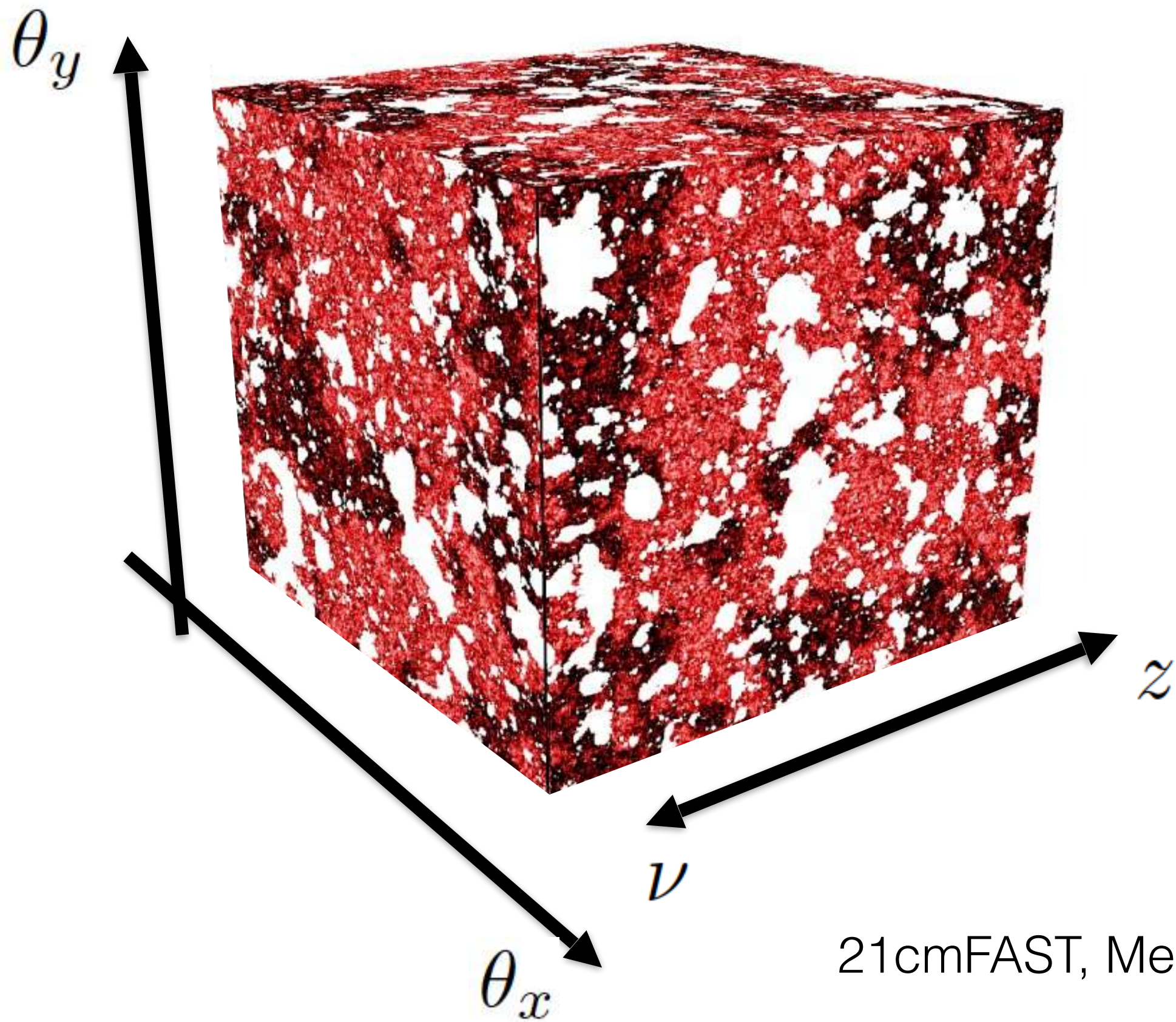


Hydrogen atoms emit radio waves with 21cm wavelengths



Hydrogen atoms emit radio waves with 21cm wavelengths





21cmFAST, Mesinger et al.

The night sky is a time
machine

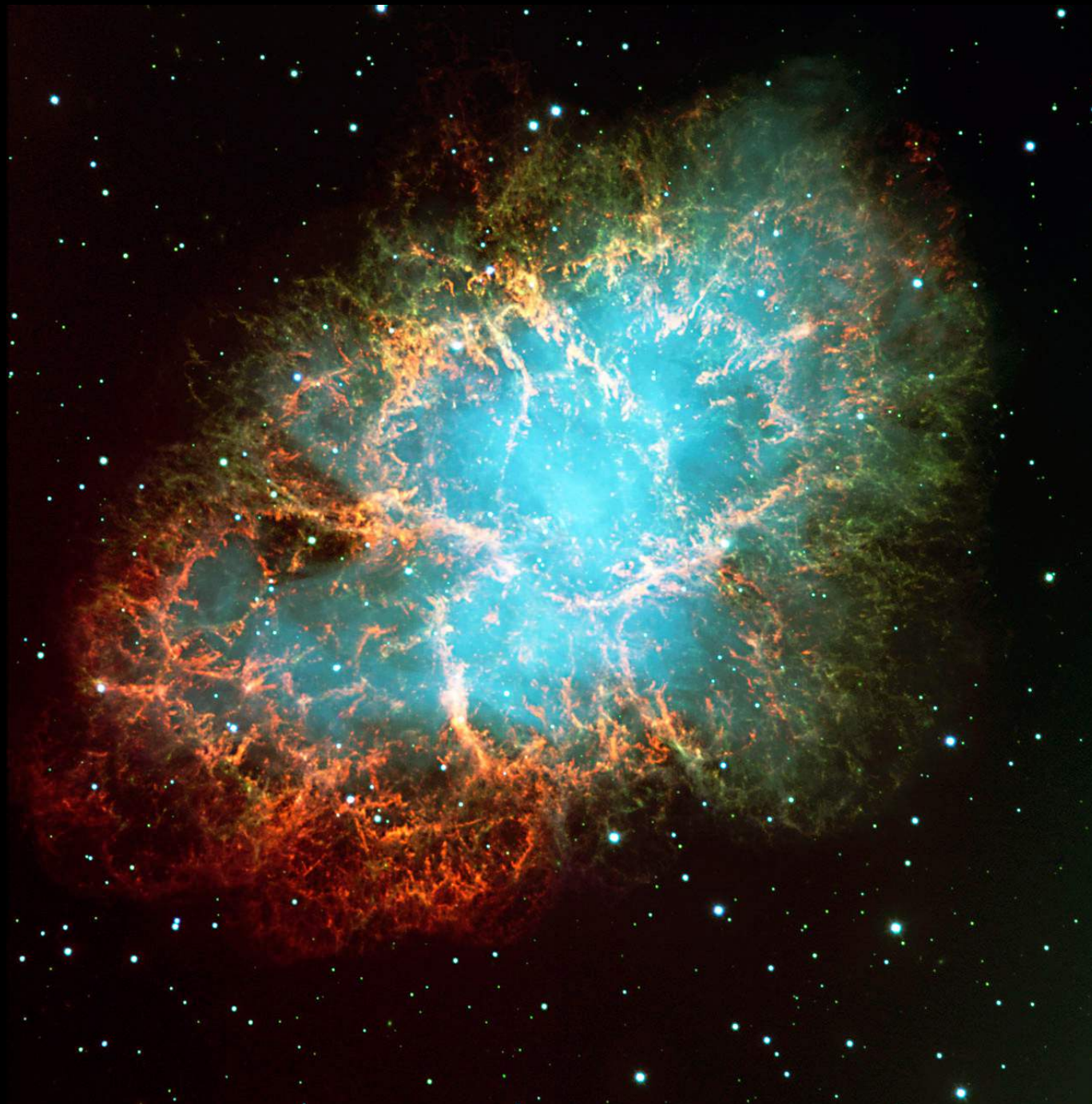
The night sky is a time machine

- Light coming to us from 4 light years away started traveling towards us 4 years ago.

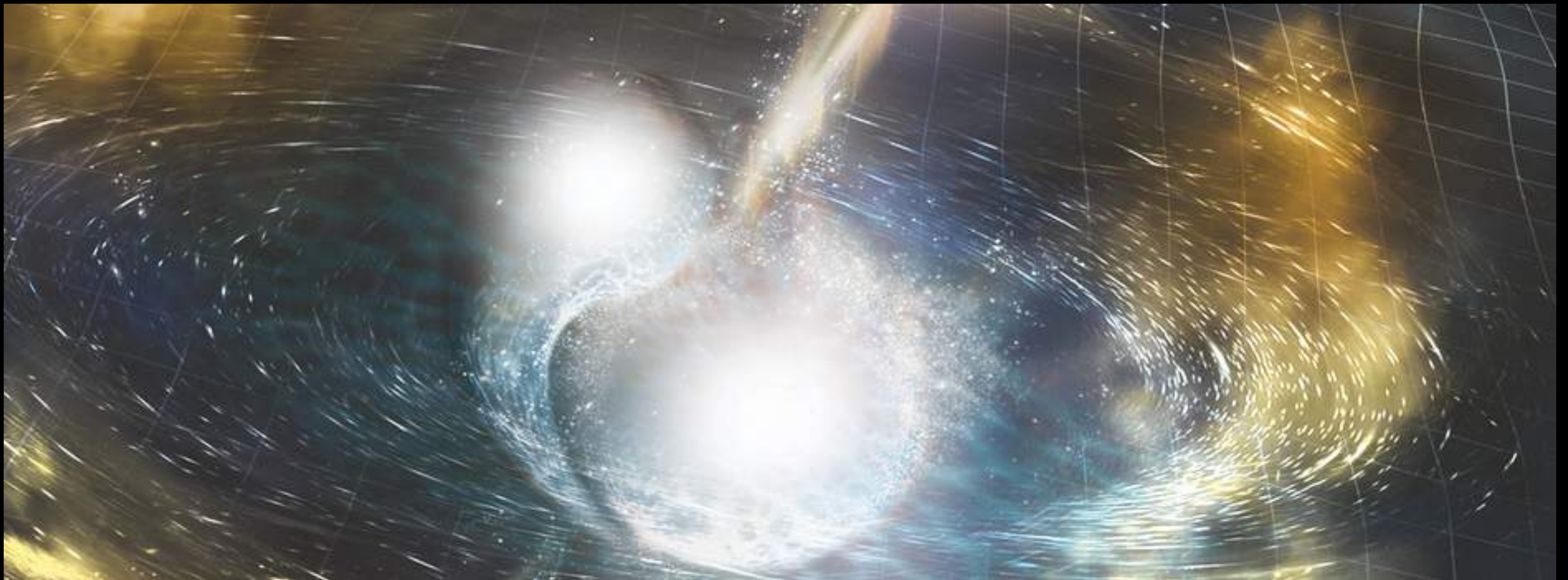
The night sky is a time machine

- Light coming to us from 4 light years away started traveling towards us 4 years ago.
- Light coming to us from 4000 light years away started traveling towards us 4000 years ago.

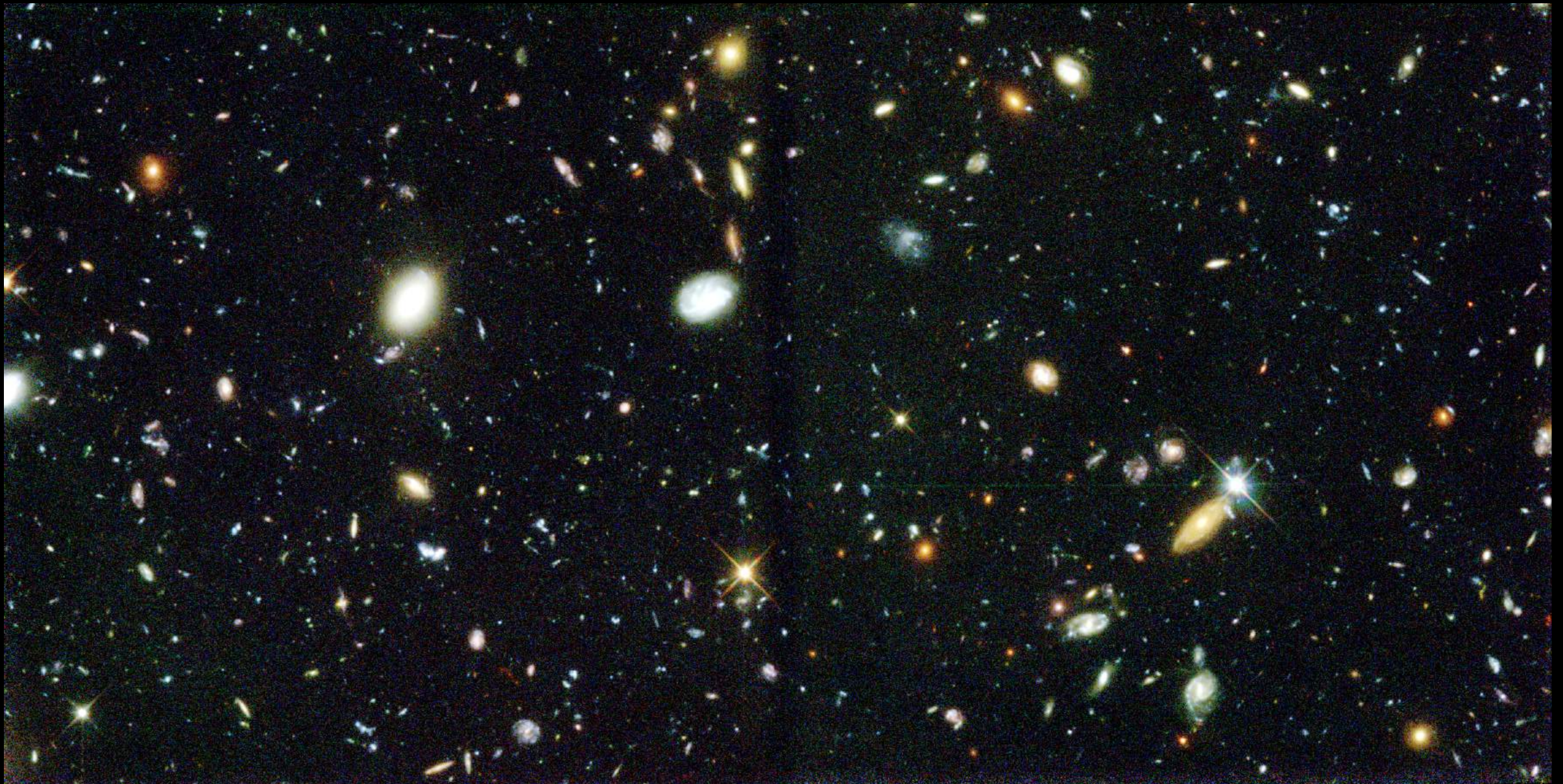
We see the Crab Nebula not as it is today, but as it was 6500 years ago



The neutron star merger that was recently detected happened about 120 million years ago!



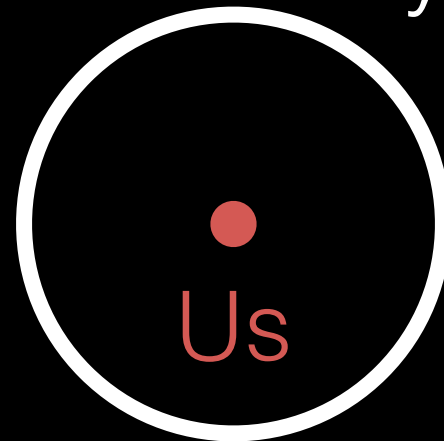
We see distant galaxies not as they are today, but they were billions of years ago



The farther away we look,
the farther back in time
we see

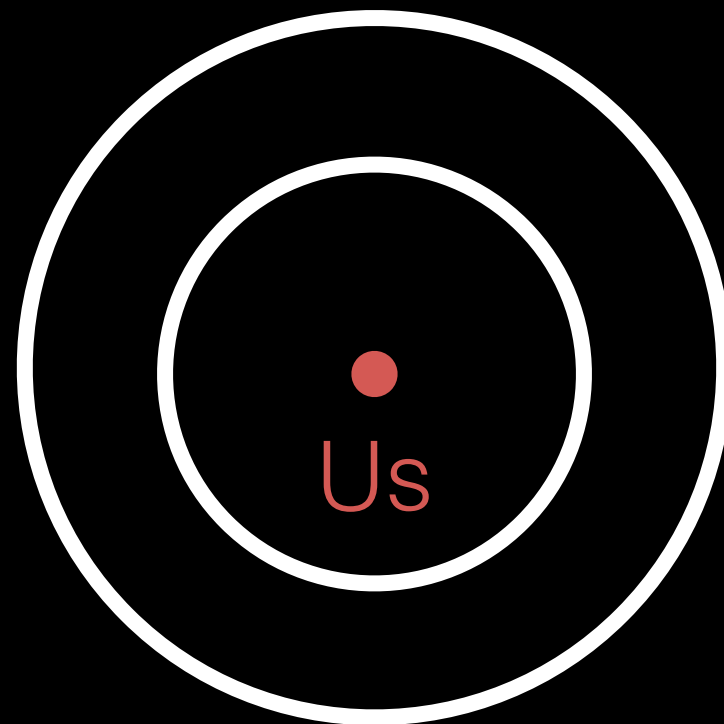
•
Us

10 years ago



Us

20 years ago



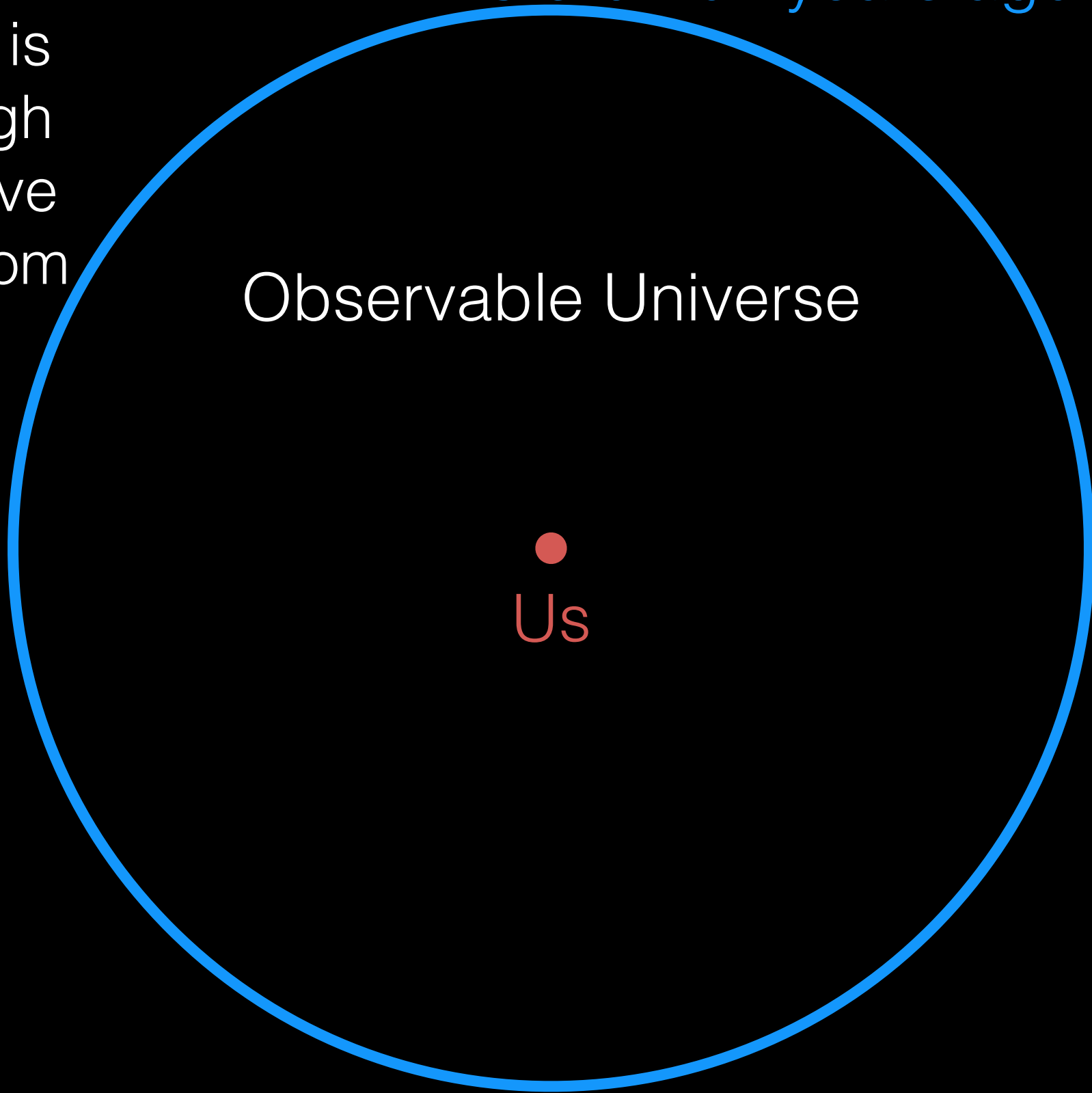
How far back have we
looked? How far back **can**
we look?

Unobservable
Universe:
our Universe is
not old enough
for light to have
reached us from
out here

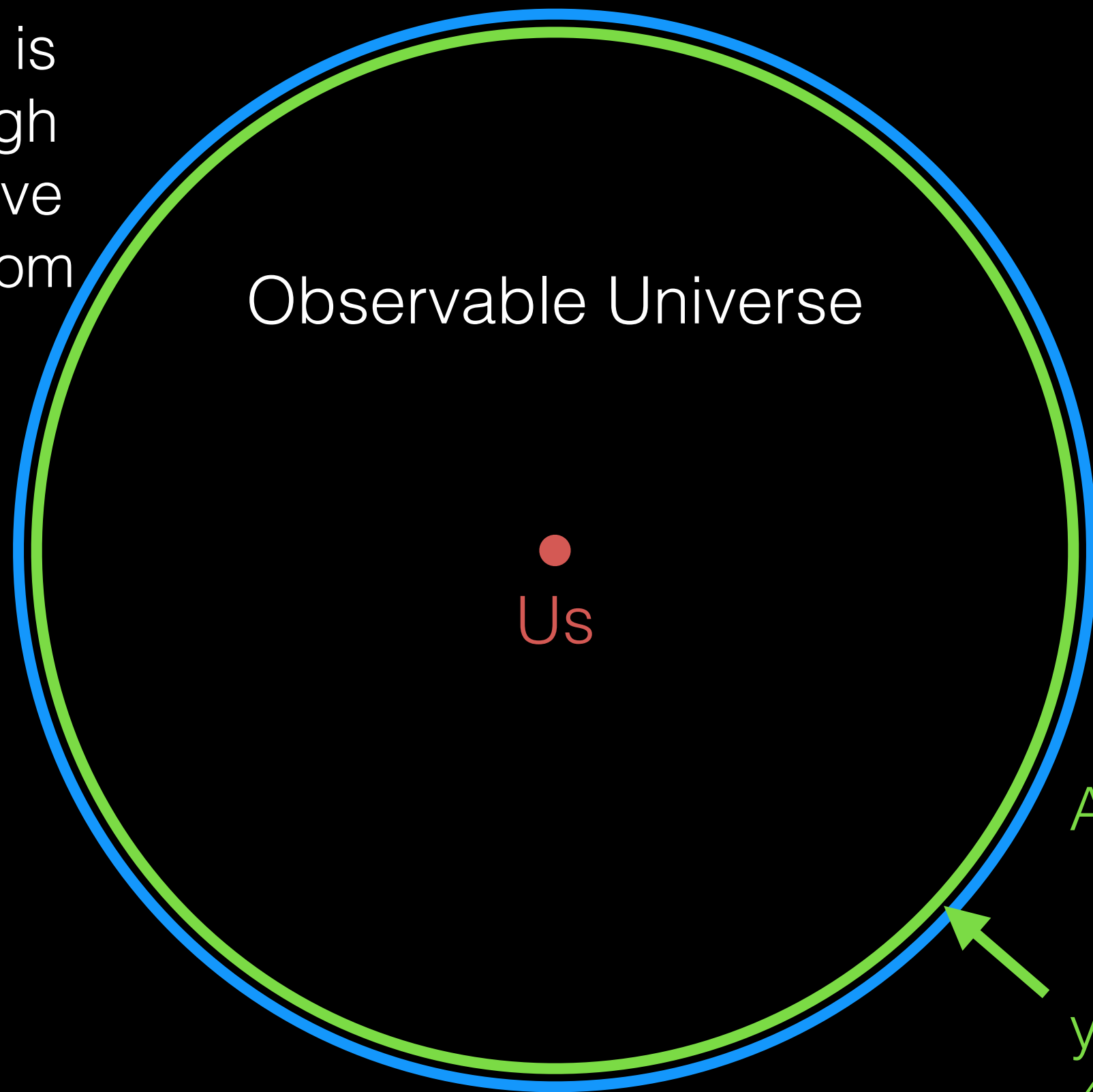
13.8 billion years ago

Observable Universe

●
Us

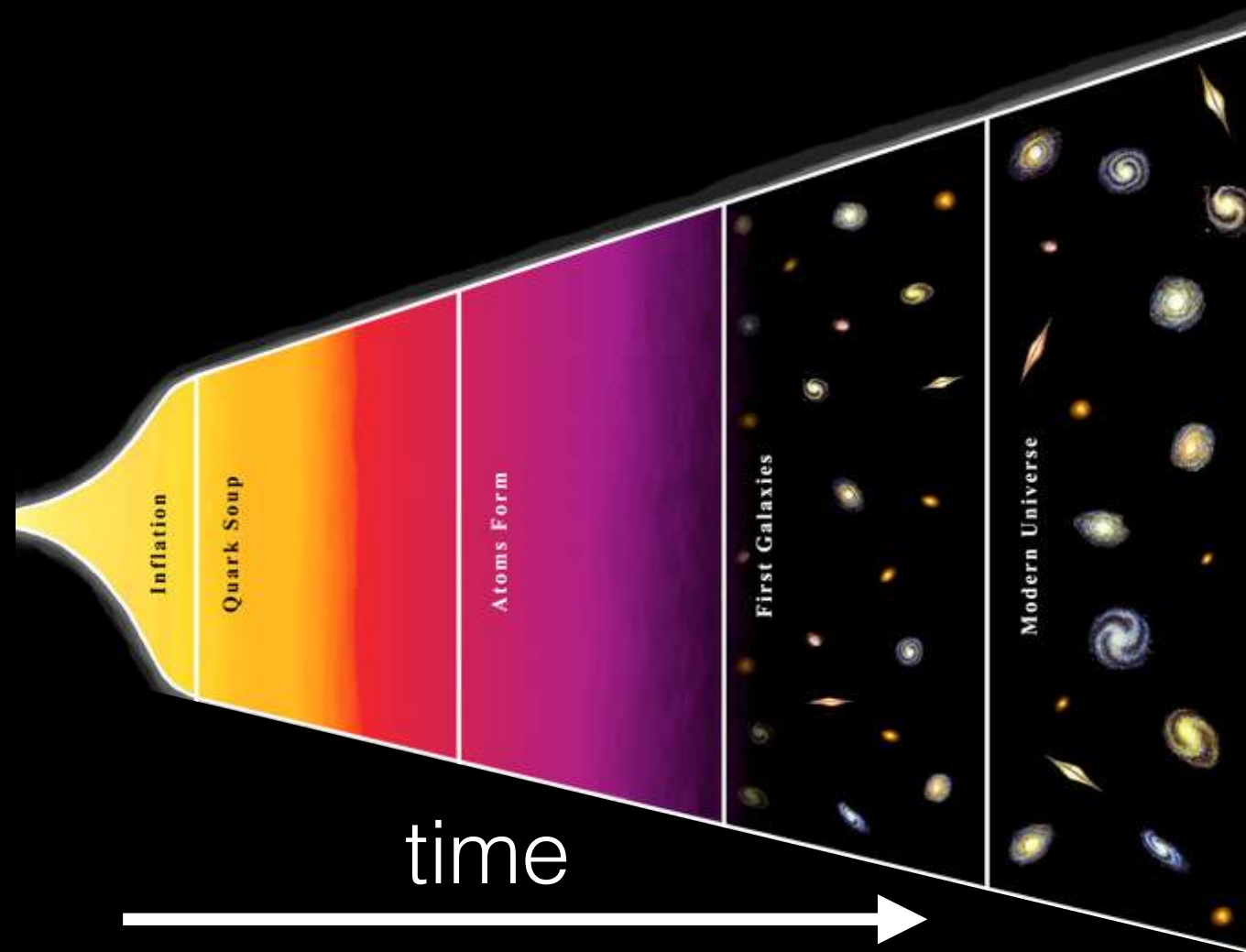


Unobservable
Universe:
our Universe is
not old enough
for light to have
reached us from
out here



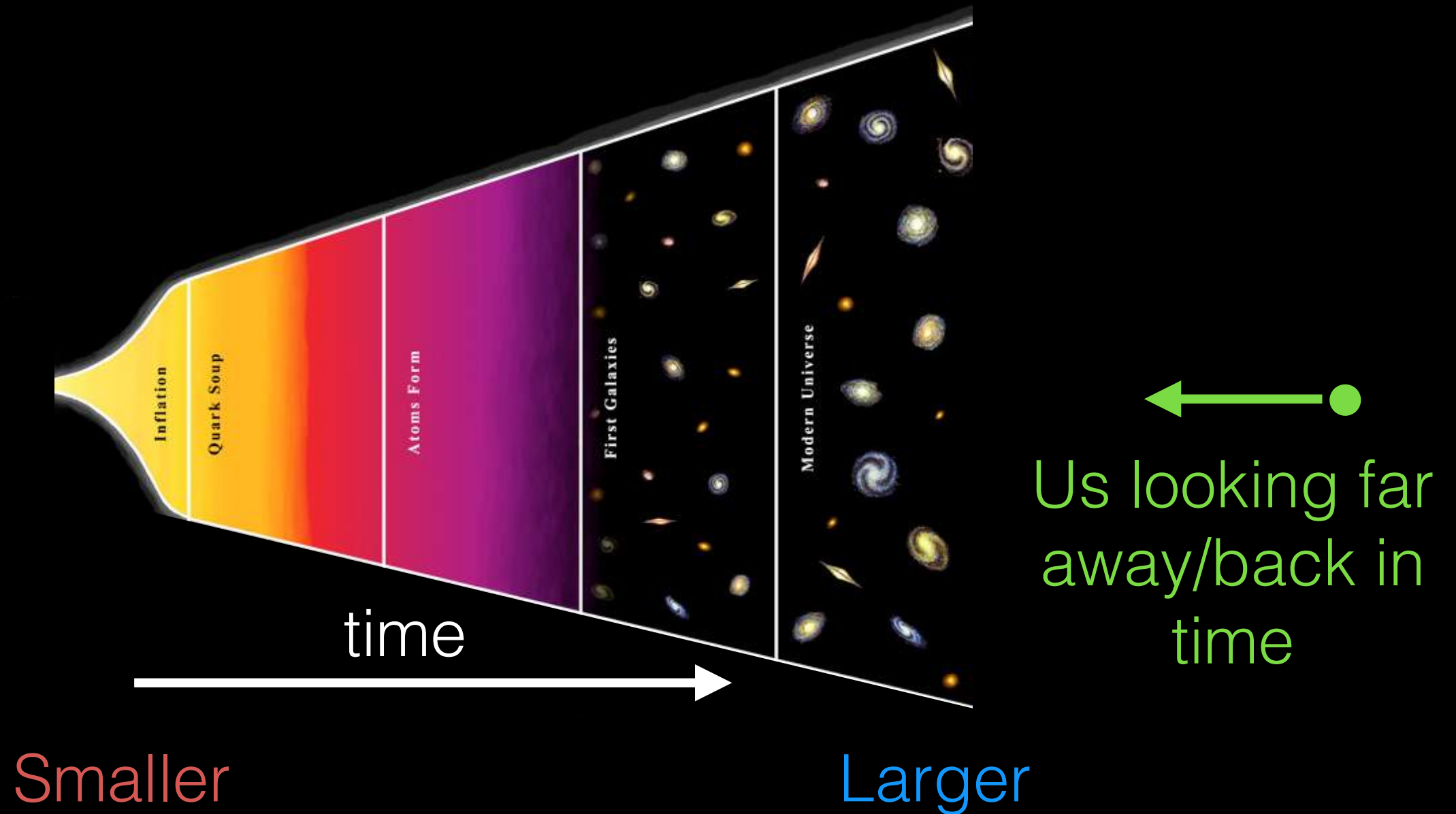
Actual limit we
can see. Not
13.8 billion
years ago, but
400,000 years
after that

Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe

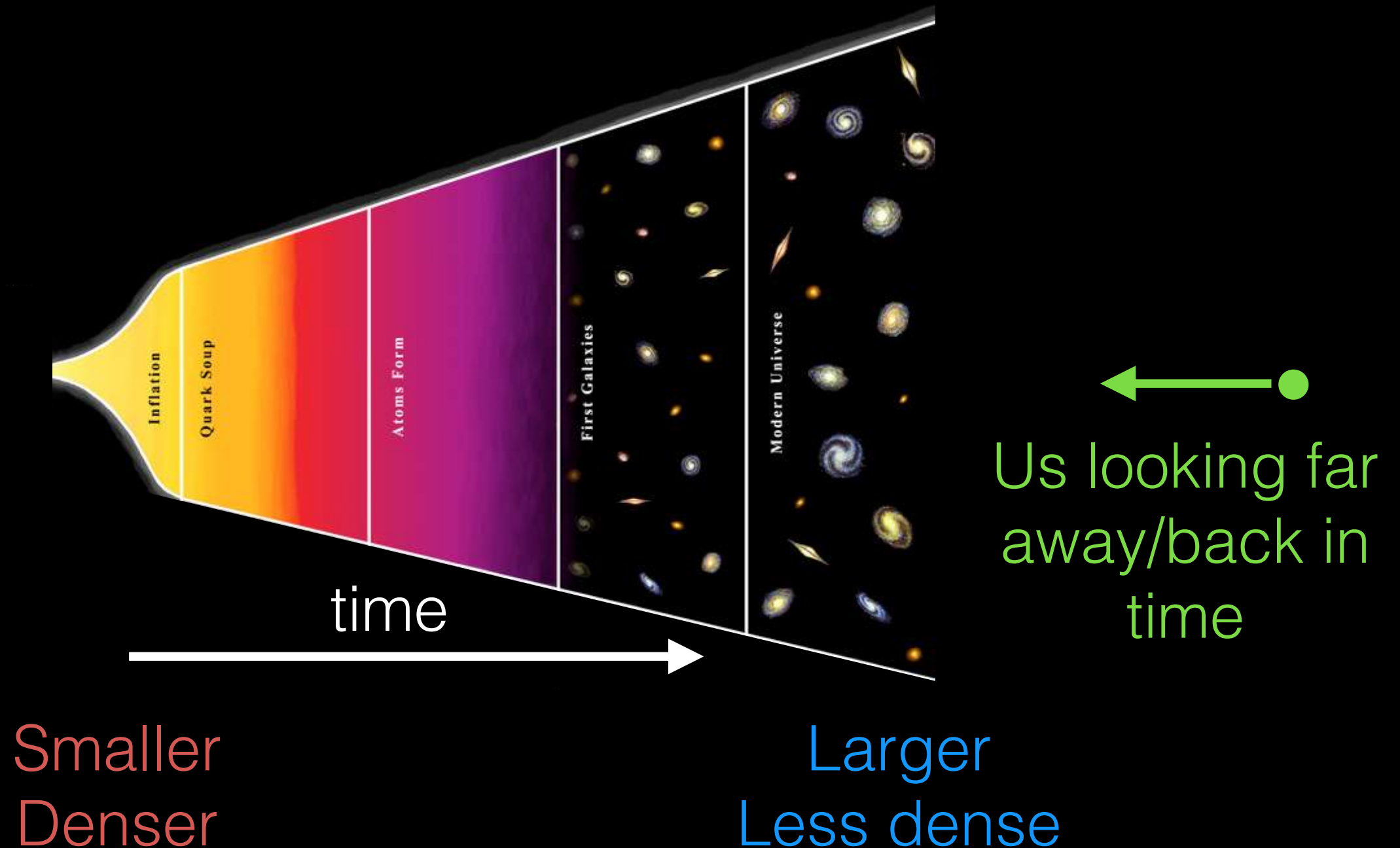


← ●
Us looking far
away/back in
time

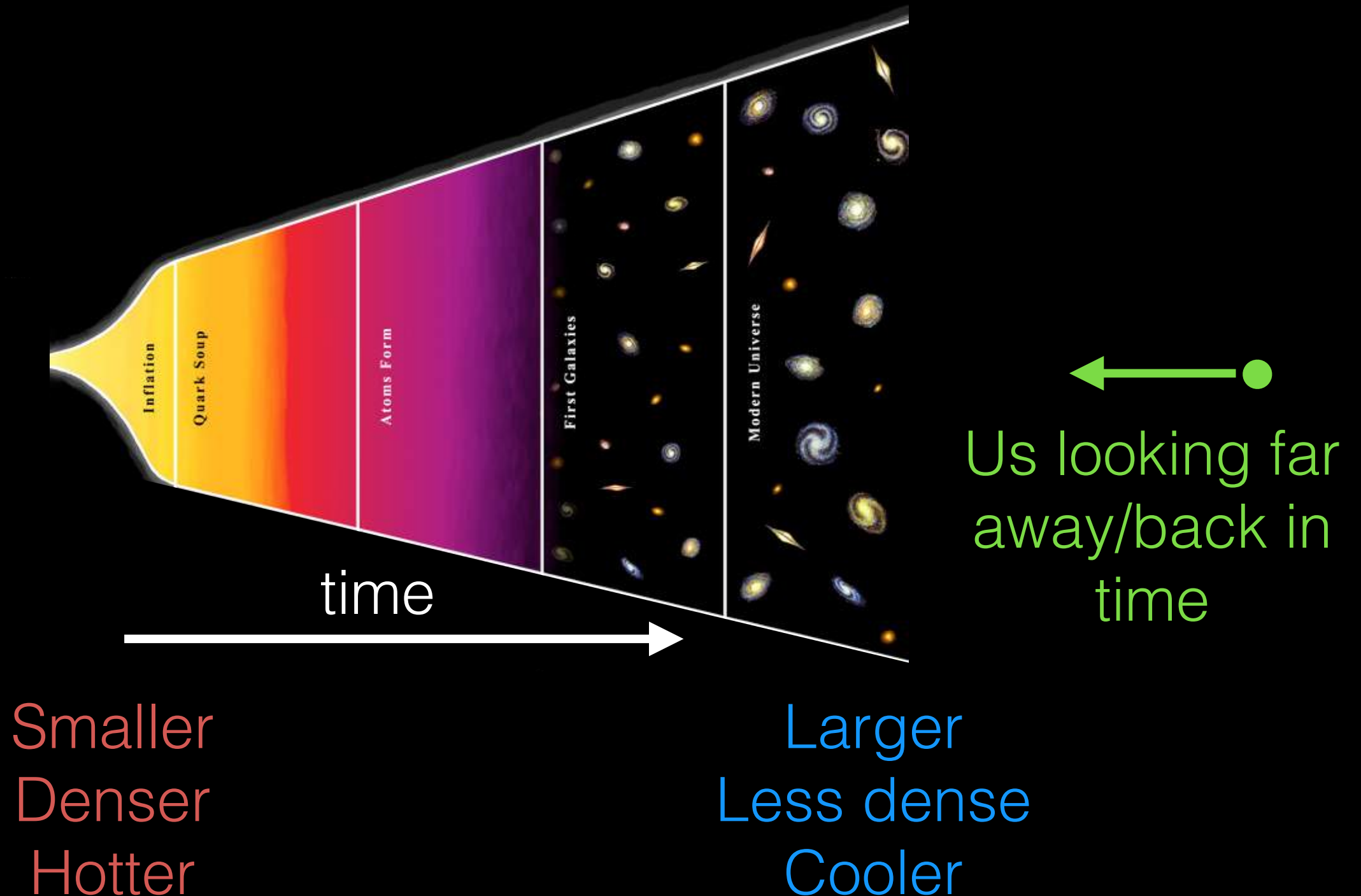
Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe

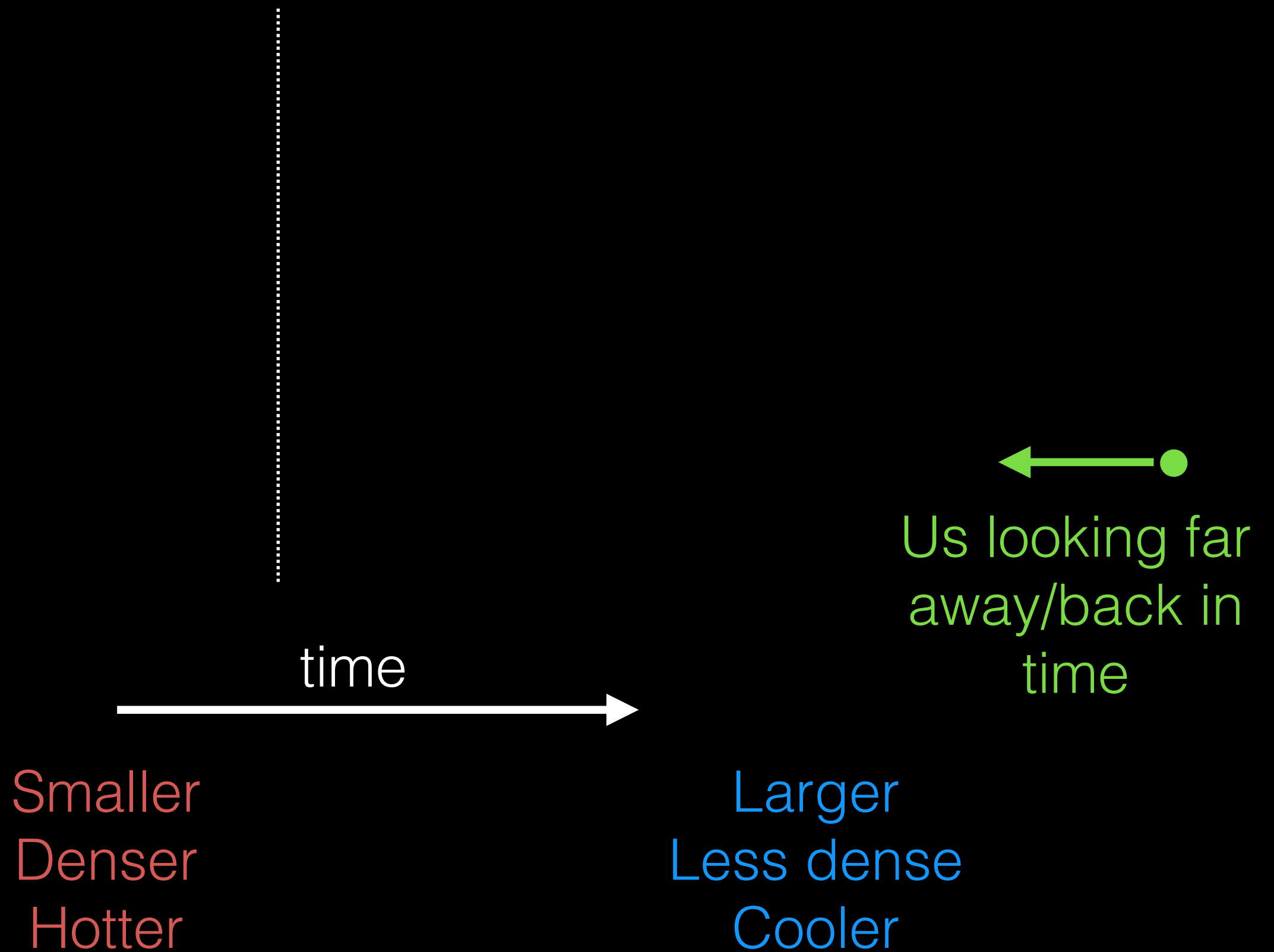


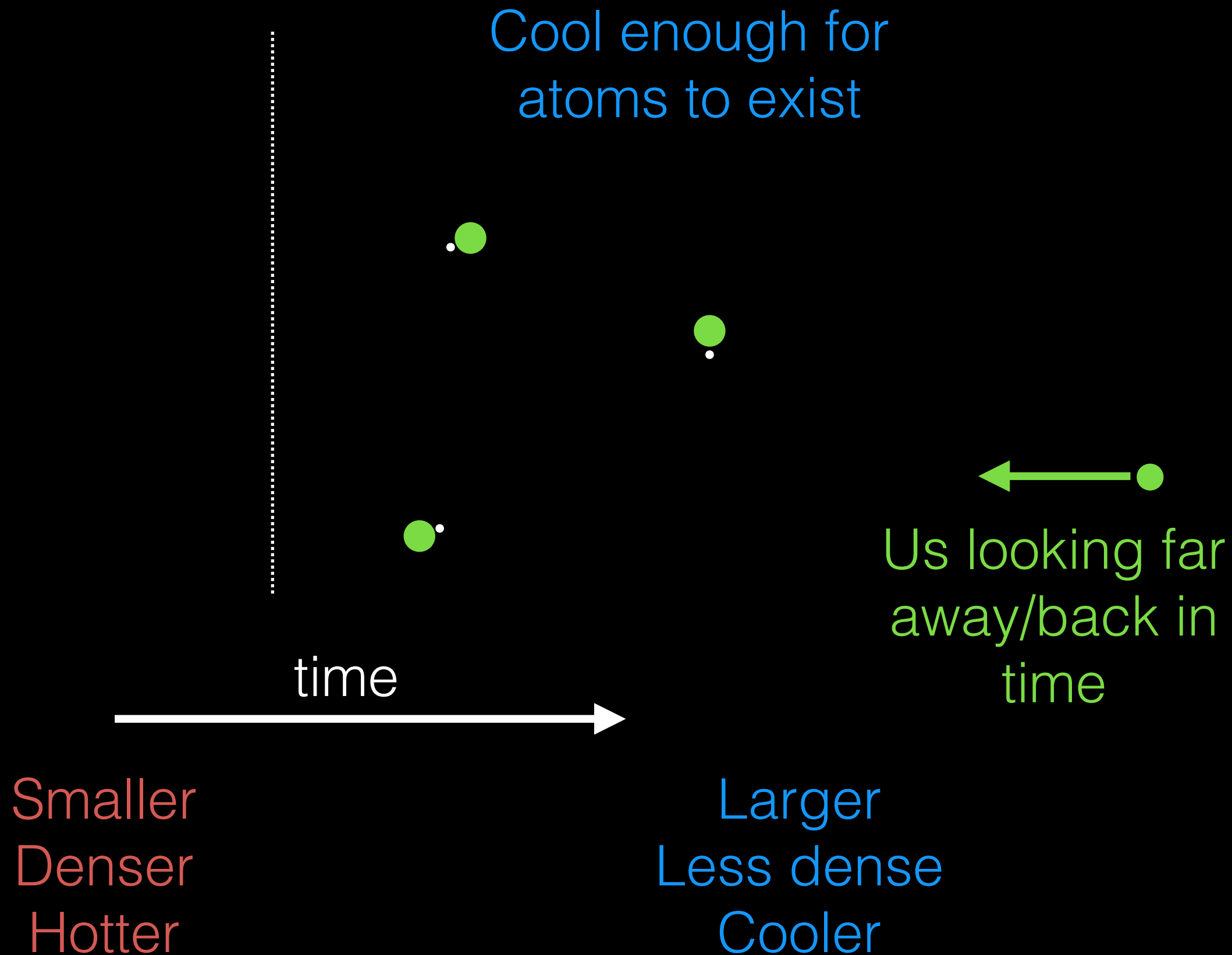
Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe



Our Universe is expanding, so as we look back in time, we see a smaller, denser, hotter Universe

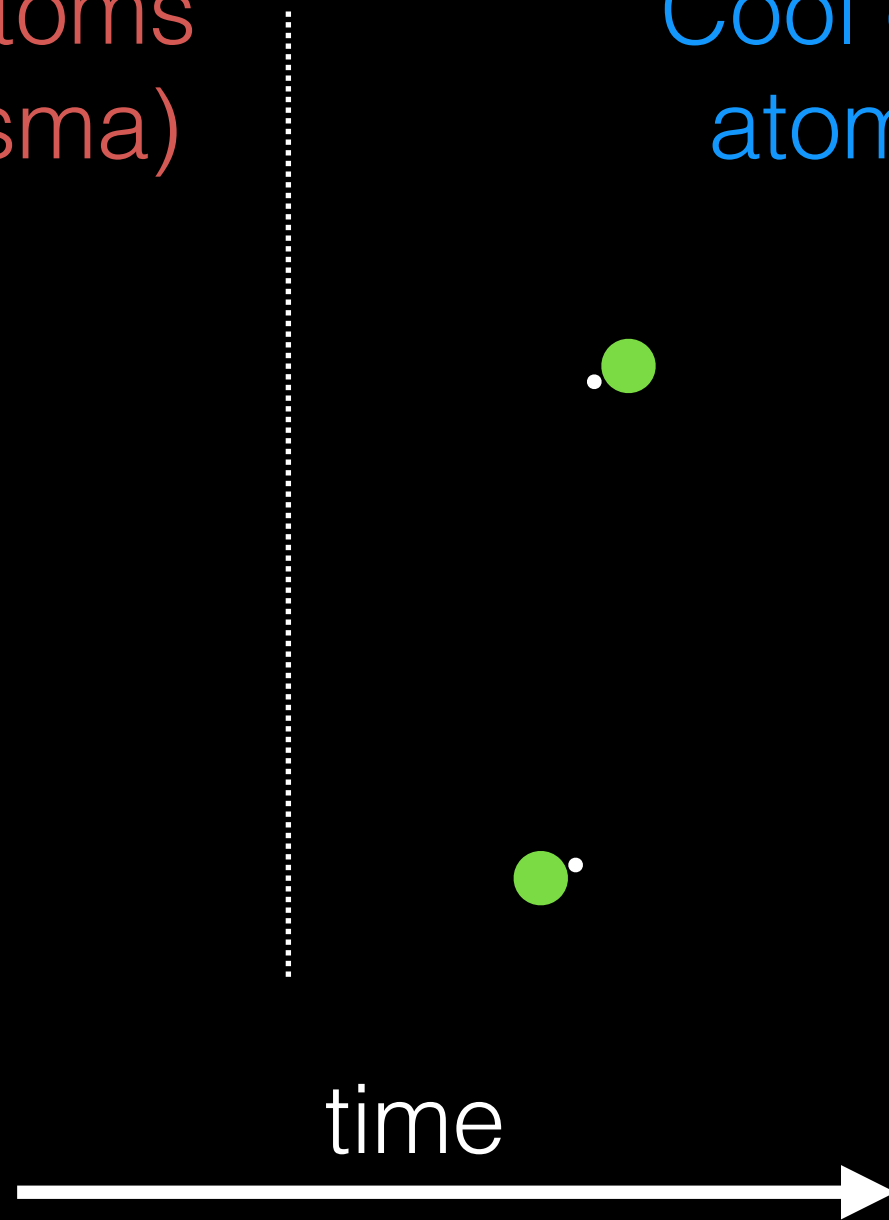






Too hot for atoms
to form (plasma)

Cool enough for
atoms to exist



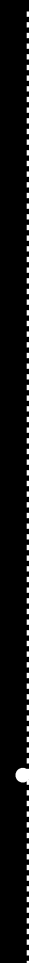
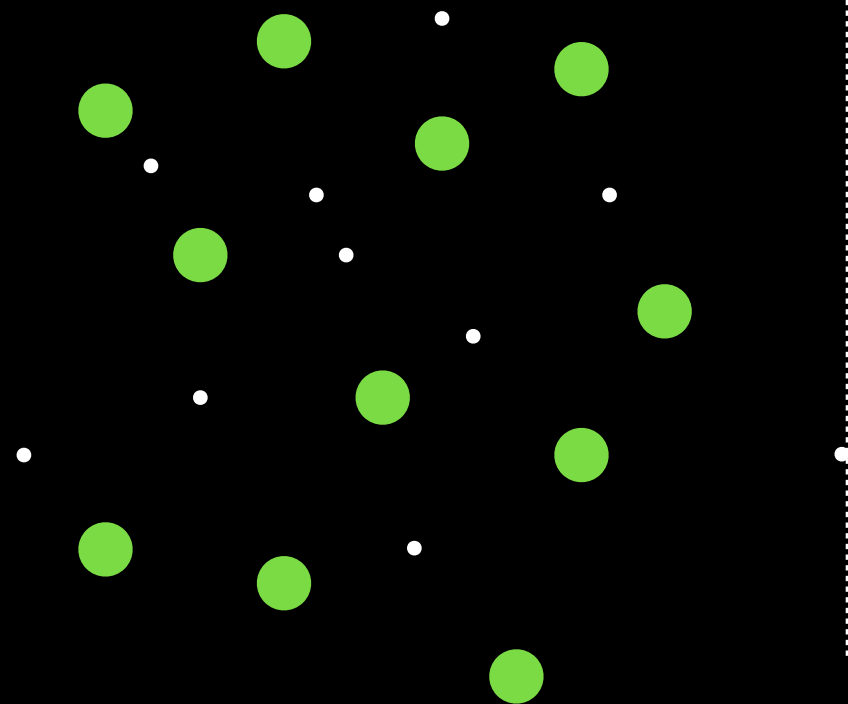
Us looking far
away/back in
time

Smaller
Denser
Hotter

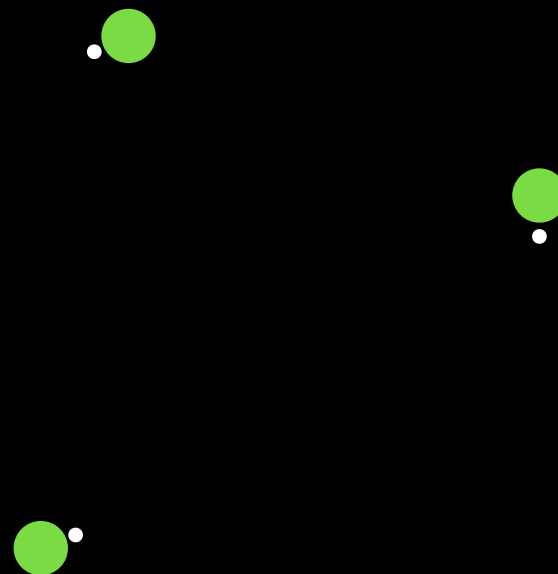
Larger
Less dense
Cooler

Too hot for atoms
to form (plasma)

Cool enough for
atoms to exist



time



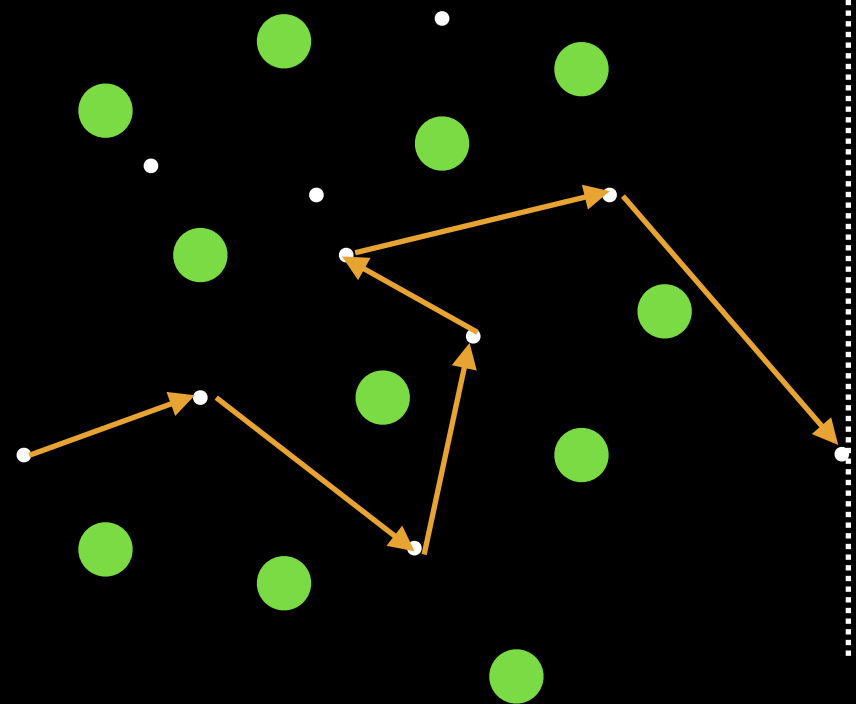
Us looking far
away/back in
time

Smaller
Denser
Hotter

Larger
Less dense
Cooler

Too hot for atoms
to form (plasma)
Not transparent

Cool enough for
atoms to exist



time



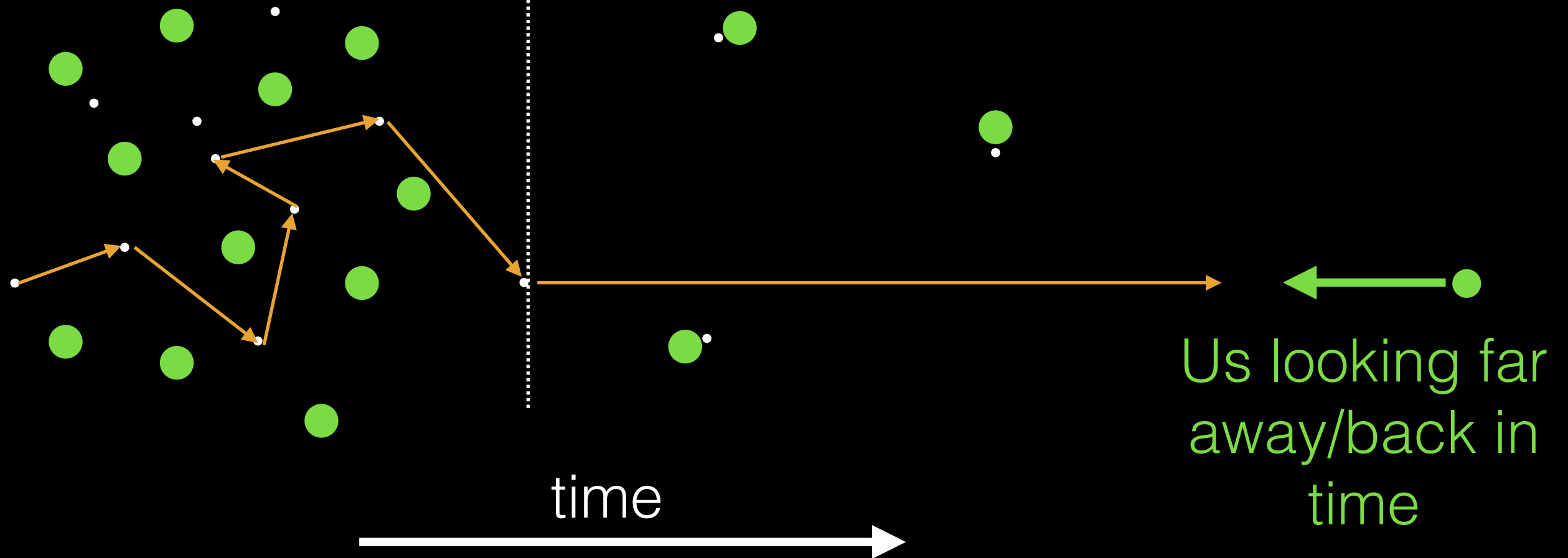
Us looking far
away/back in
time

Smaller
Denser
Hotter

Larger
Less dense
Cooler

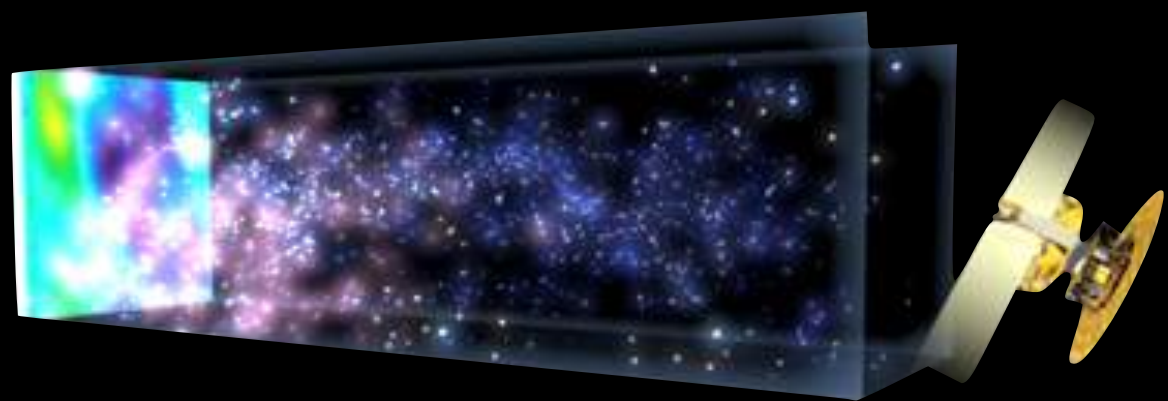
Too hot for atoms
to form (plasma)
Not transparent

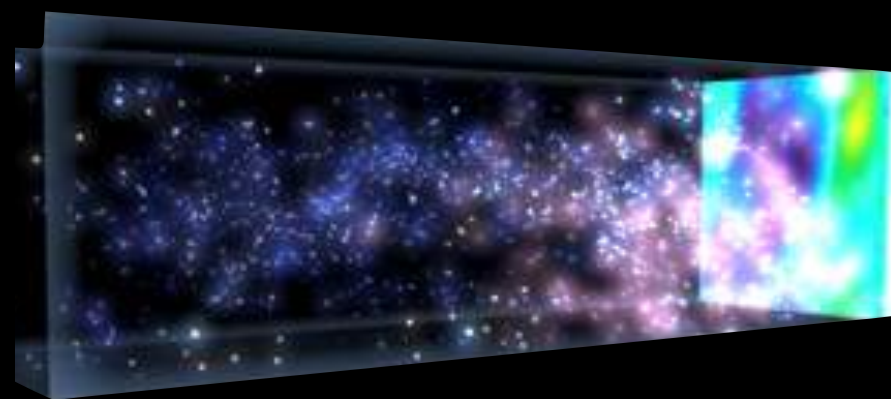
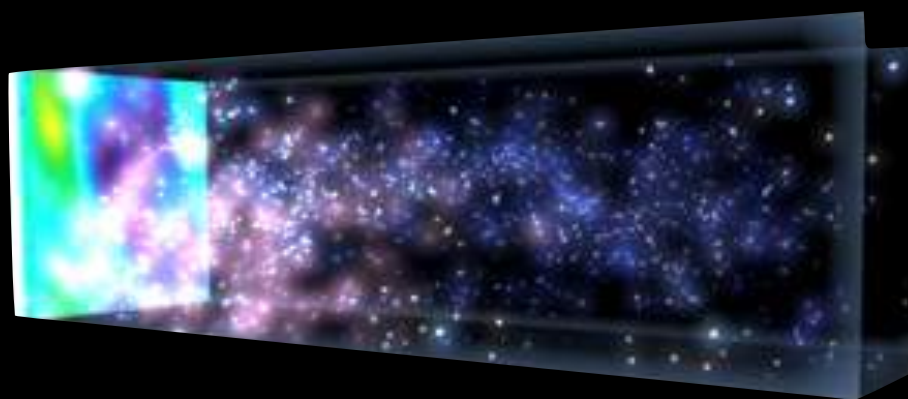
Cool enough for
atoms to exist
Transparent

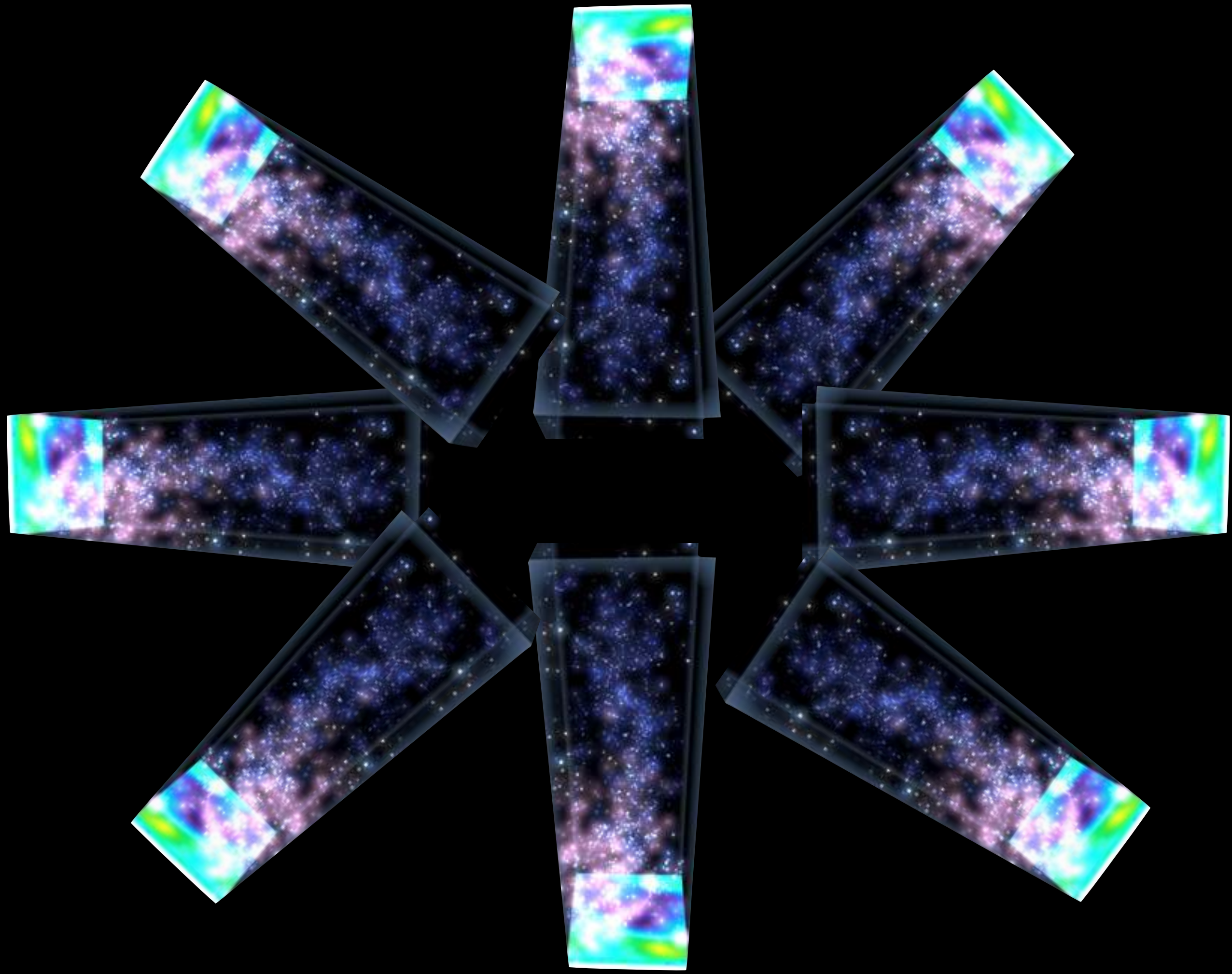


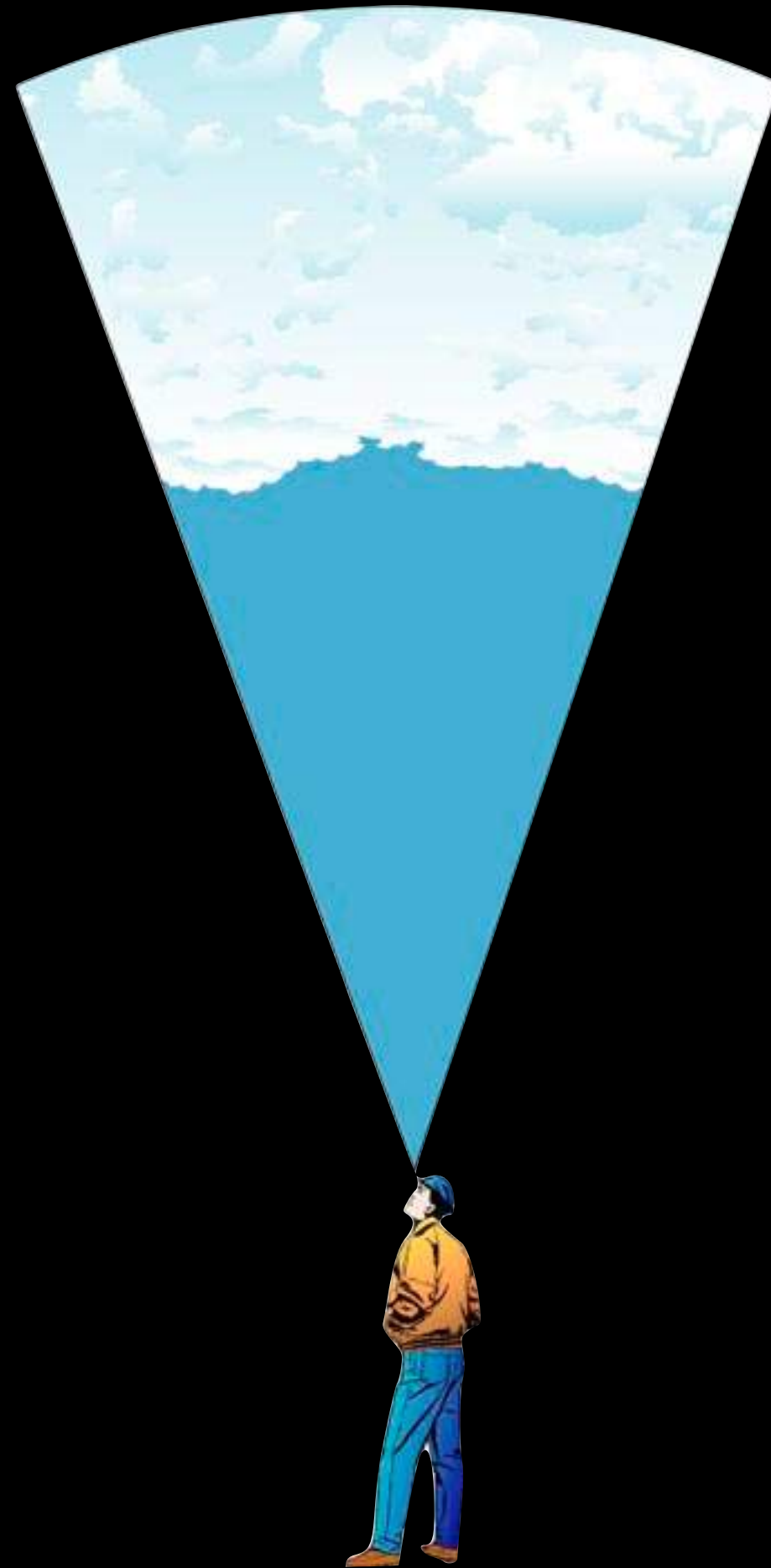
Smaller
Denser
Hotter

Larger
Less dense
Cooler

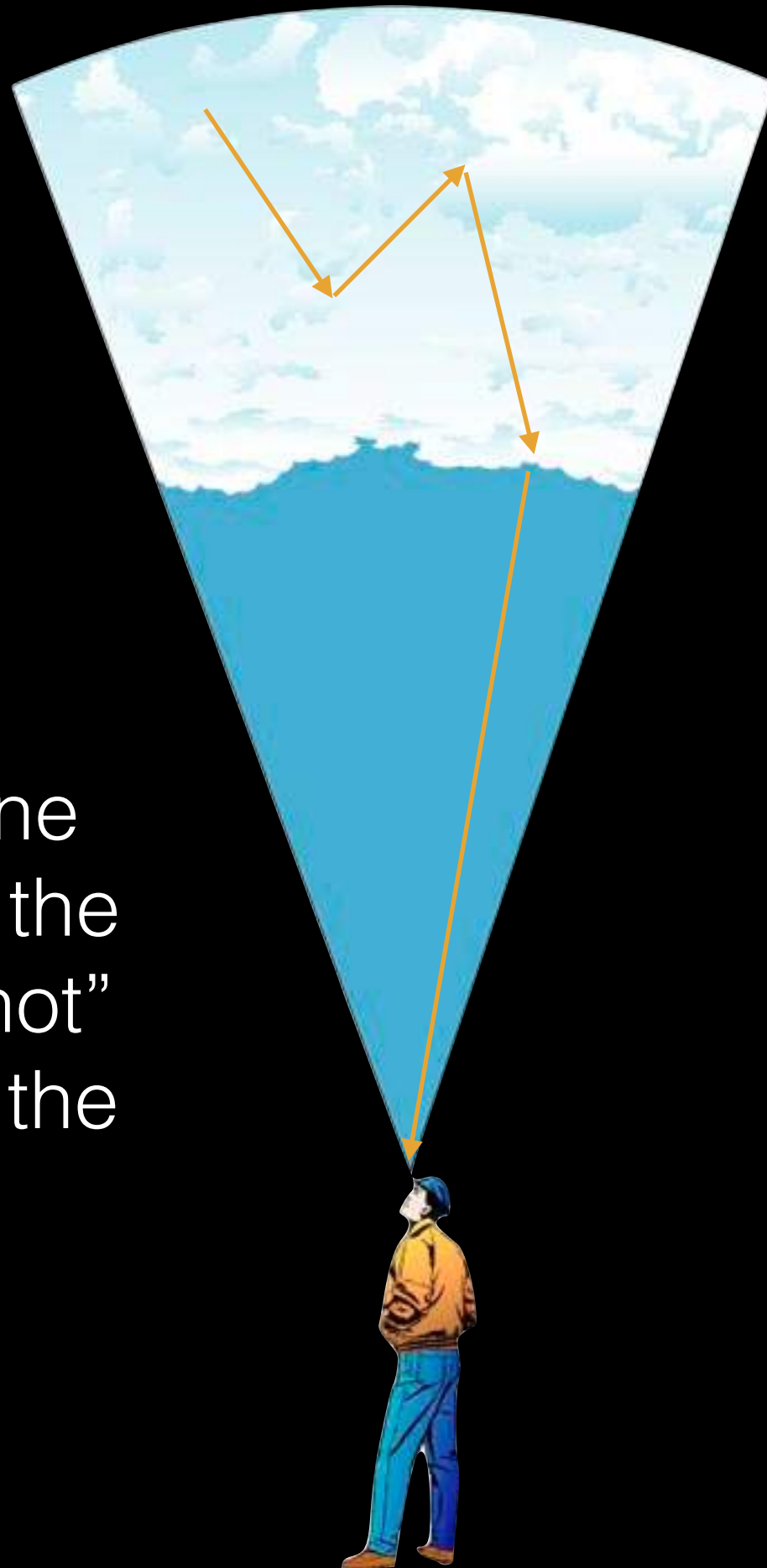




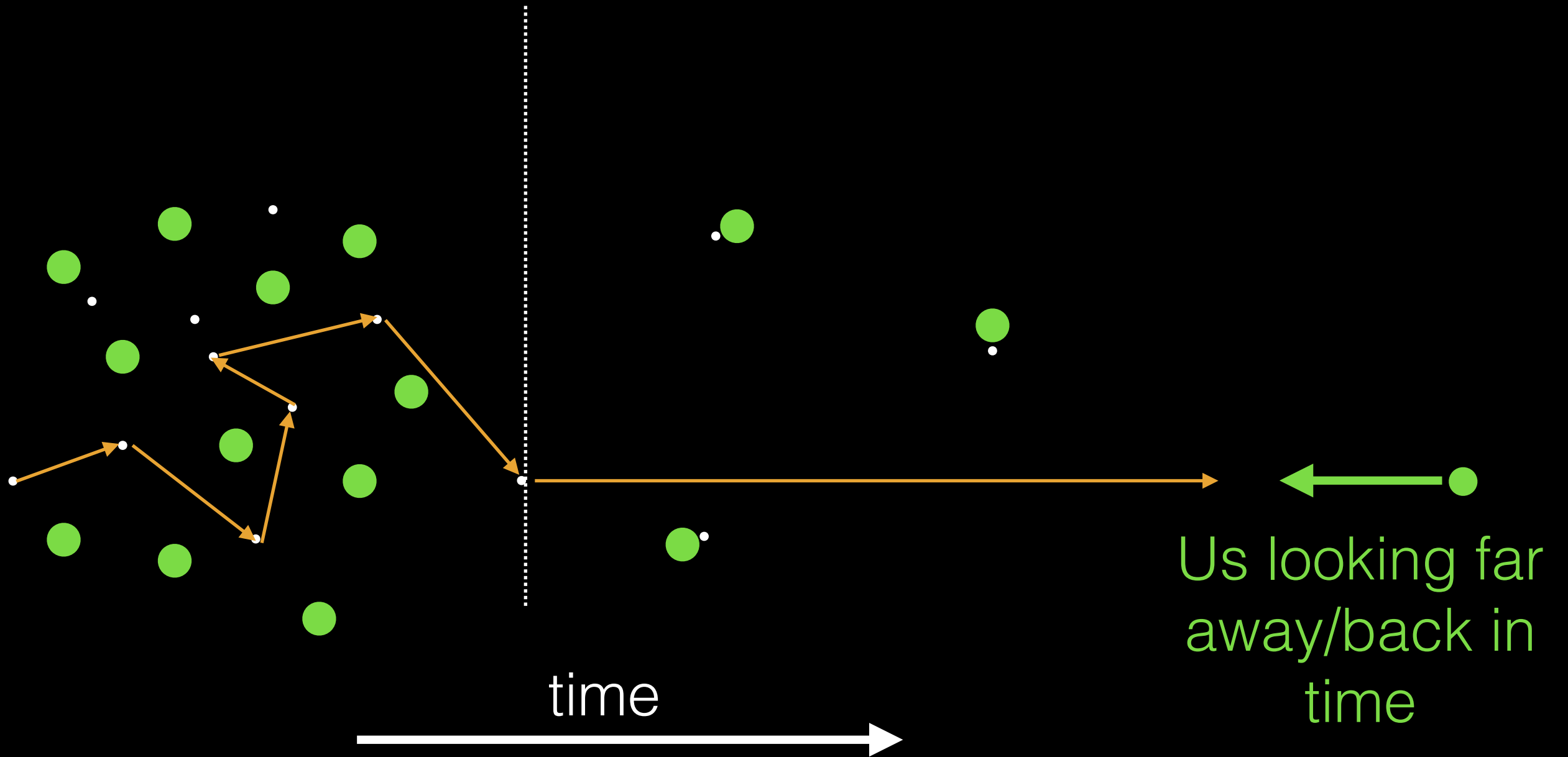




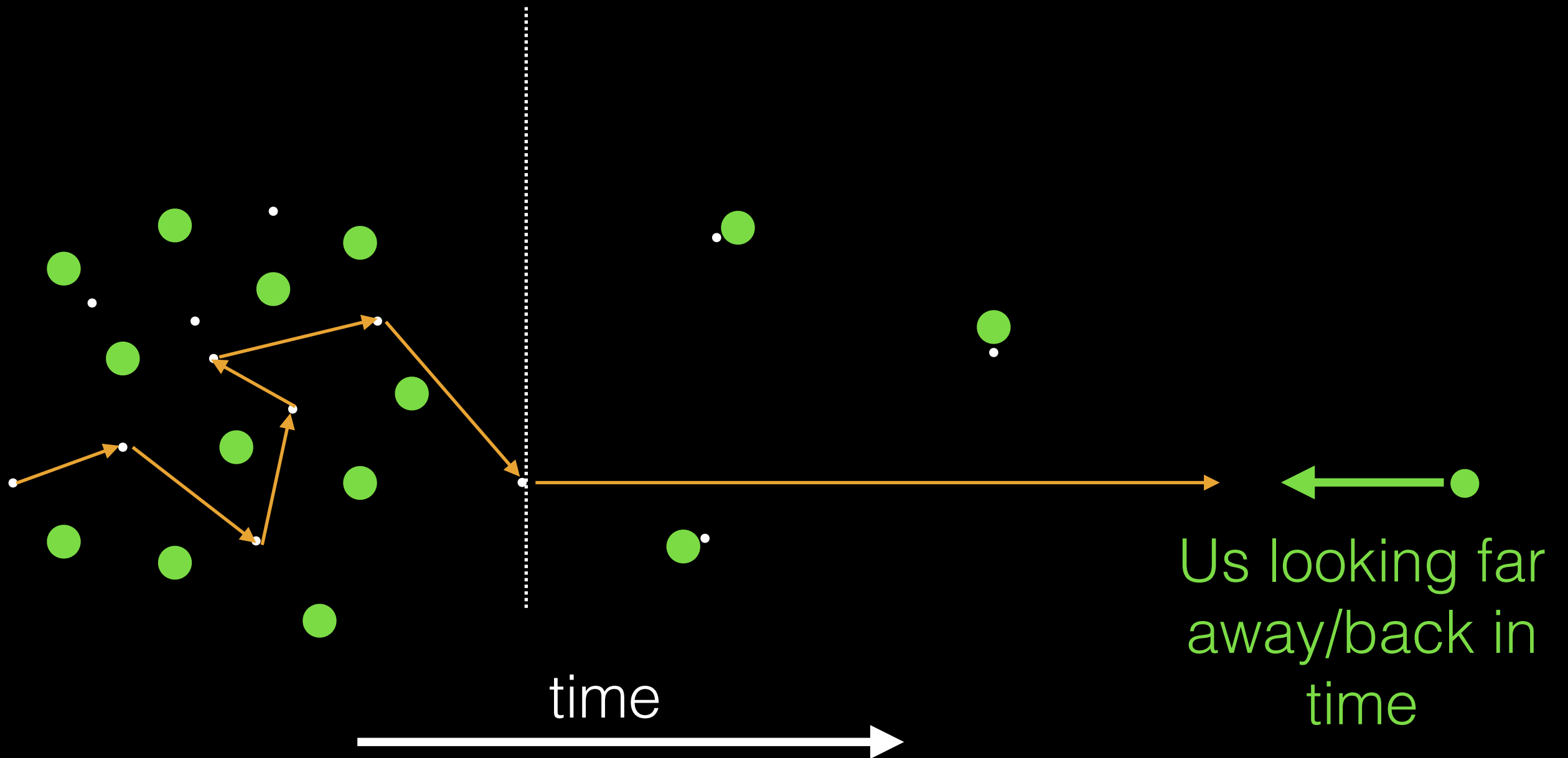
We see an outline
of the surface of the
cloud, a “snapshot”
of the exterior of the
cloud



400,000 years

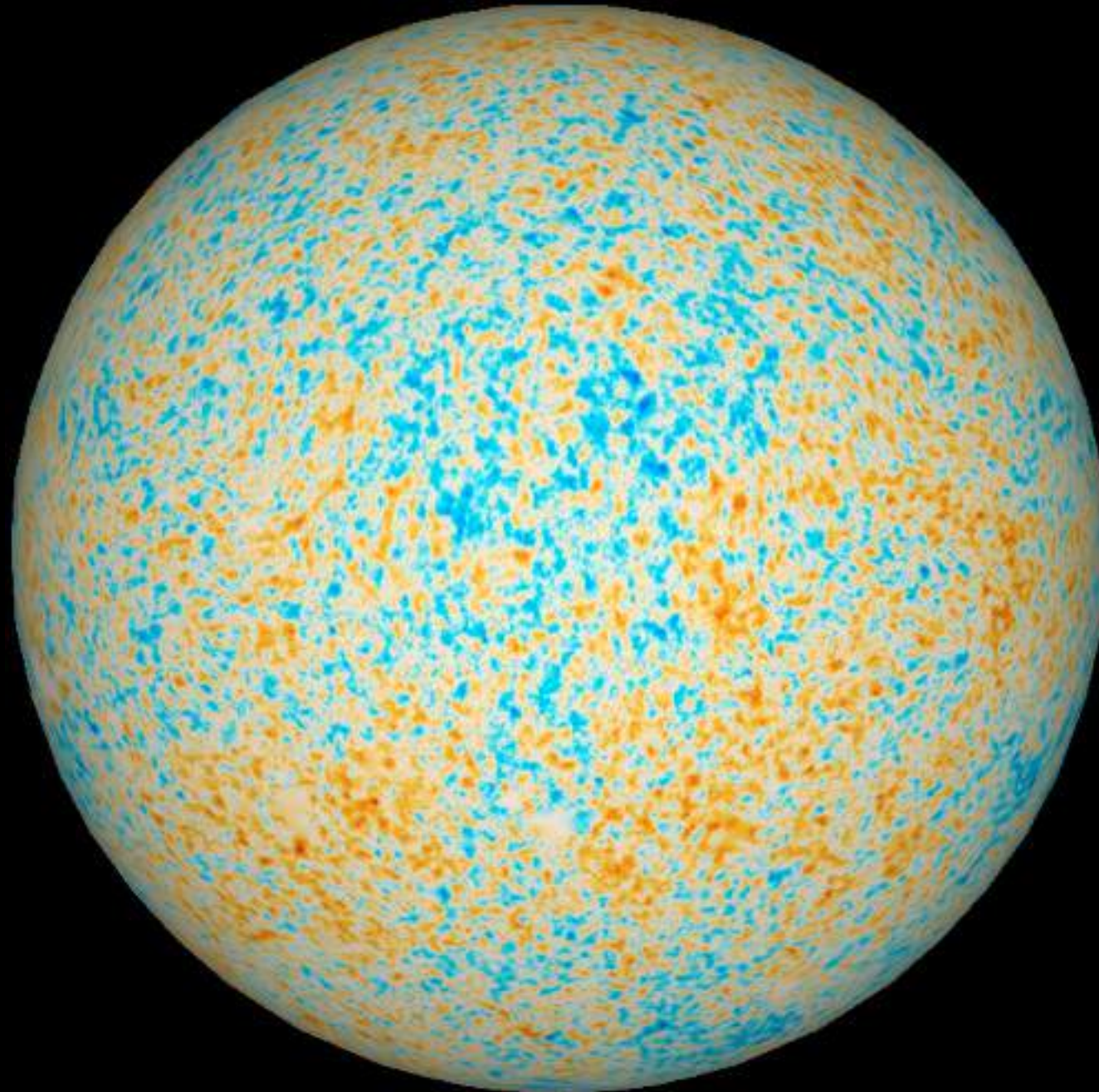


400,000 years



We see a snapshot of our Universe when it was only 400,000 years old. A mere baby!

The Cosmic Microwave Background
provides a snapshot of our baby Universe

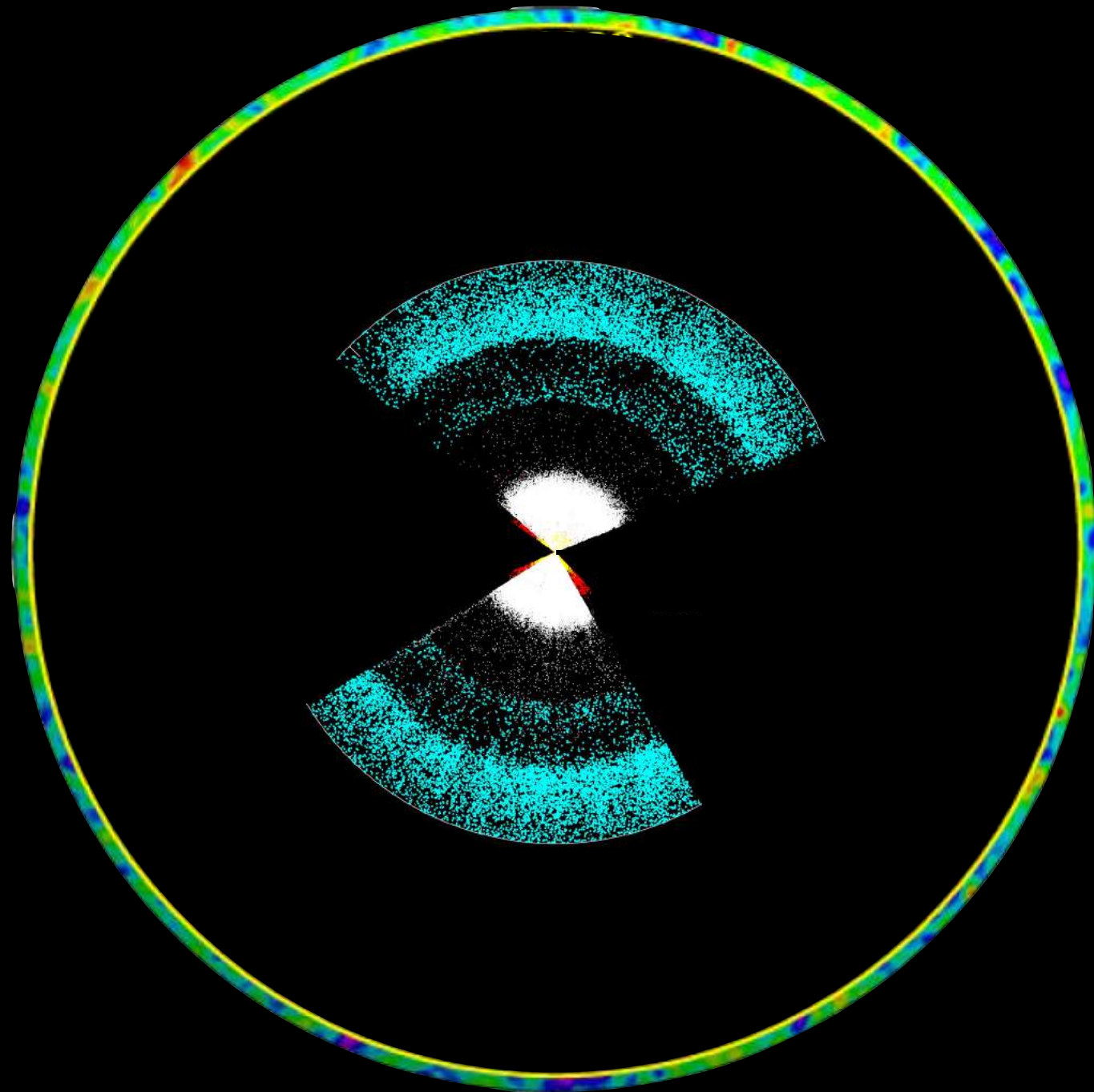




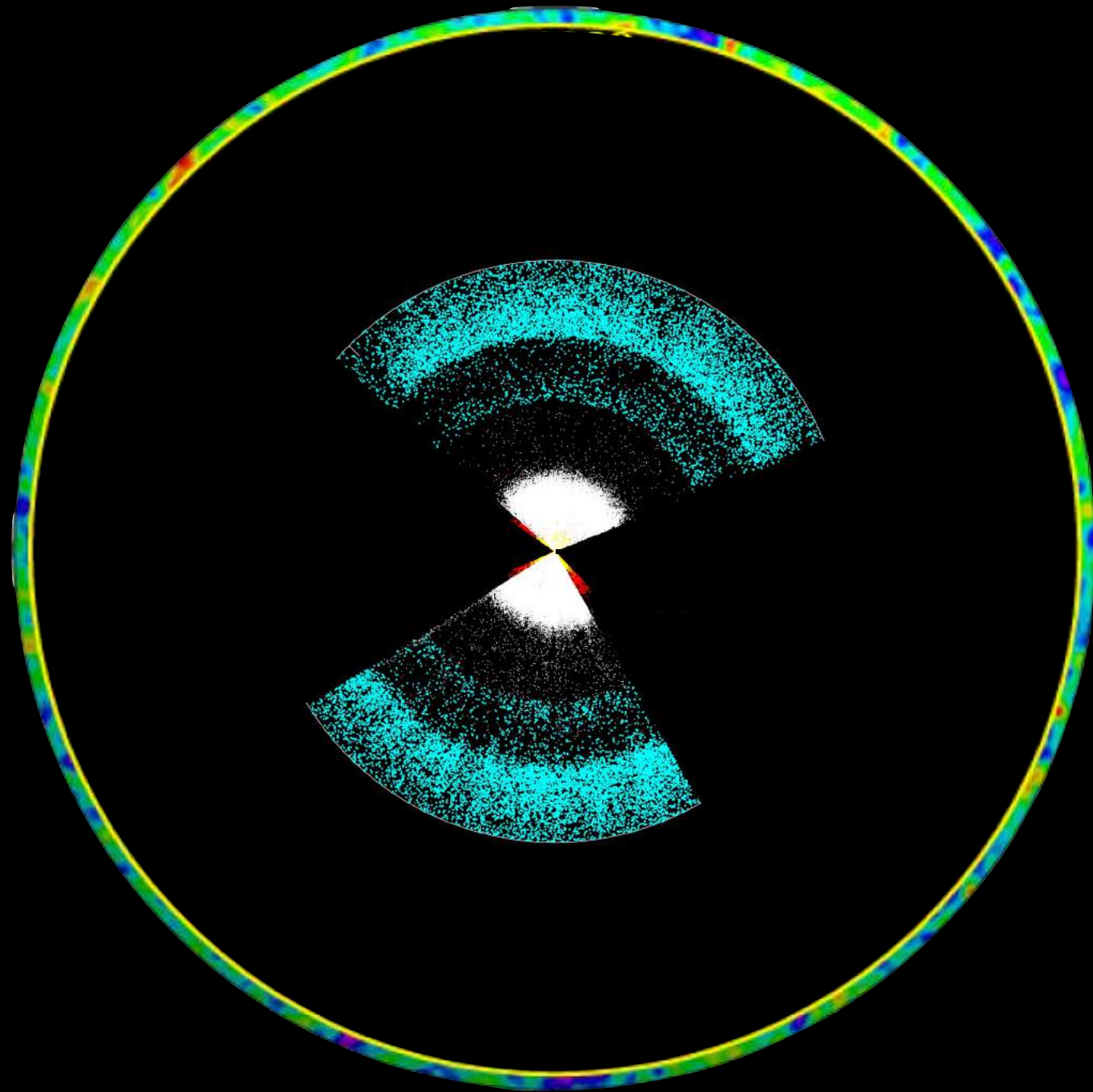
Alvarez et al. (2009)

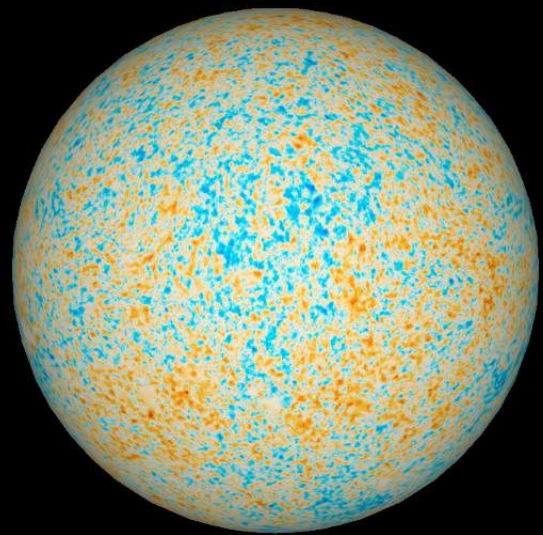
What about the real
thing?

We have yet to observe most
of the observable Universe

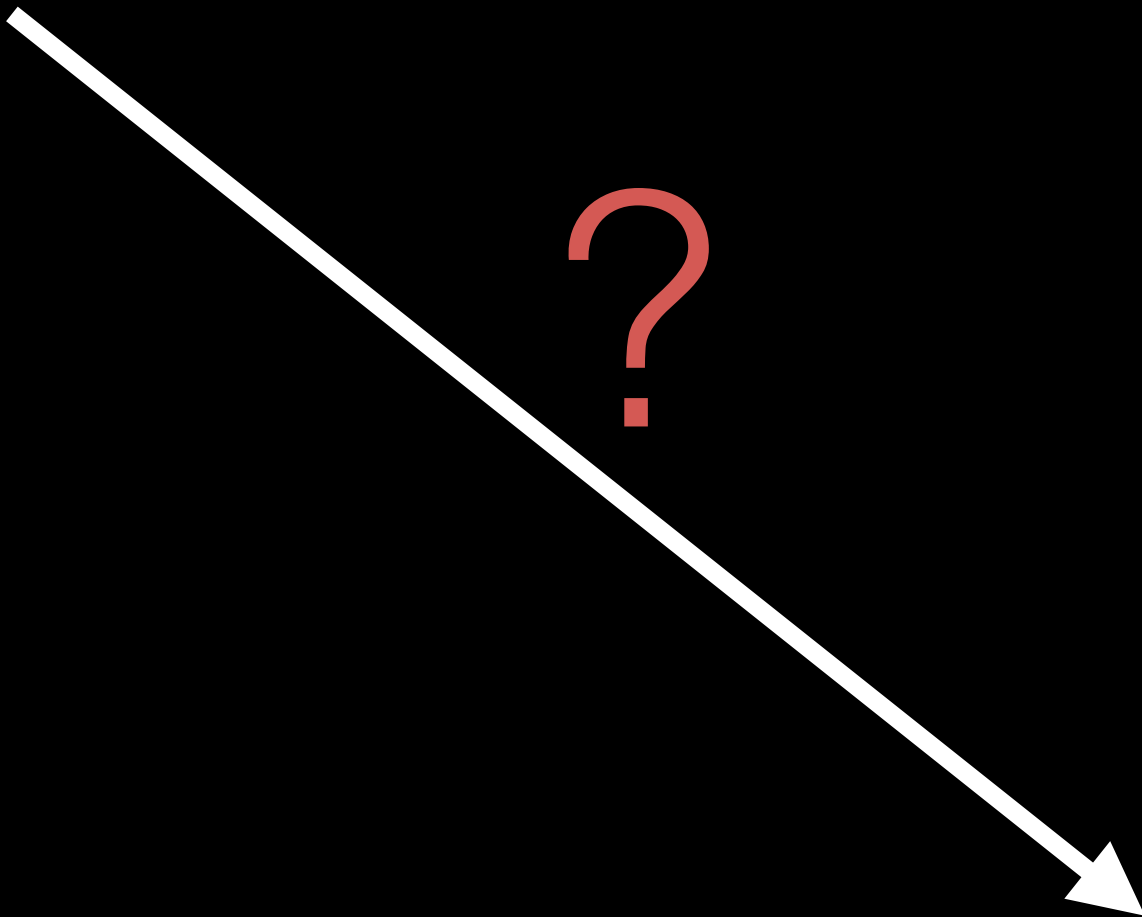


We have seen our grown-up Universe, our
baby Universe, but not our Universe as it
was growing up

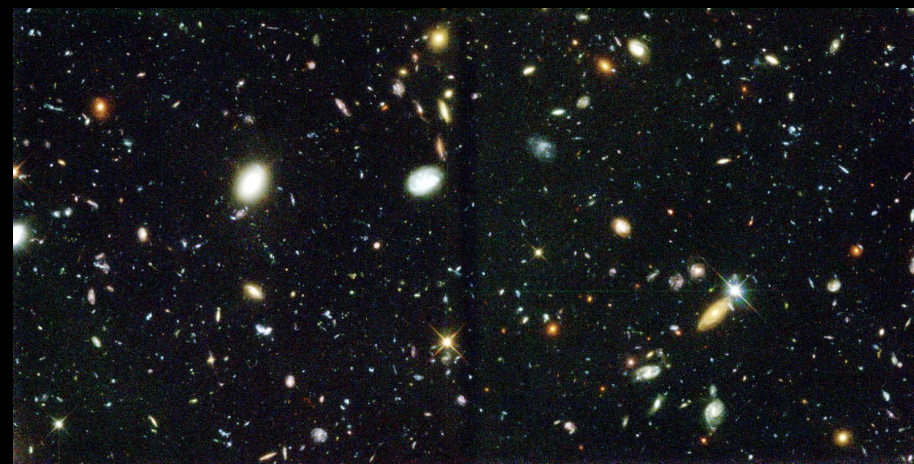




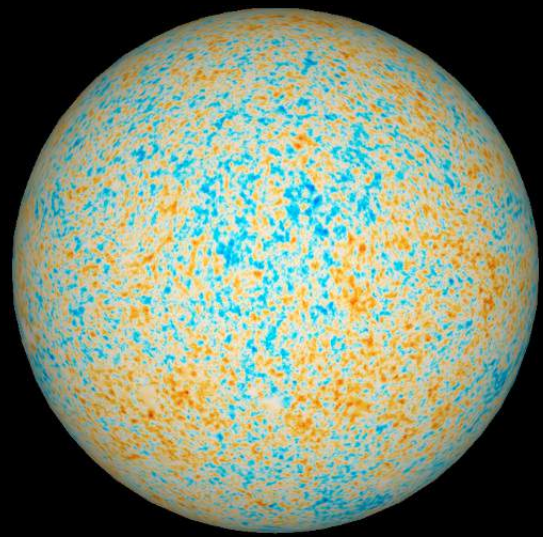
Baby Universe



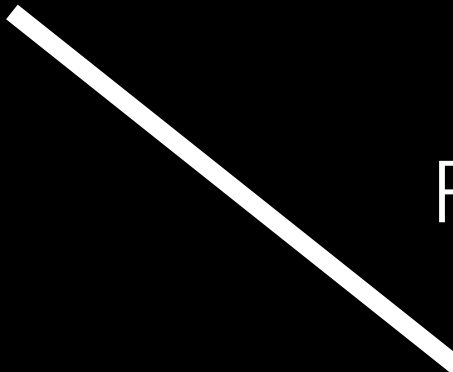
Time



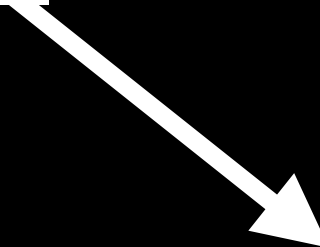
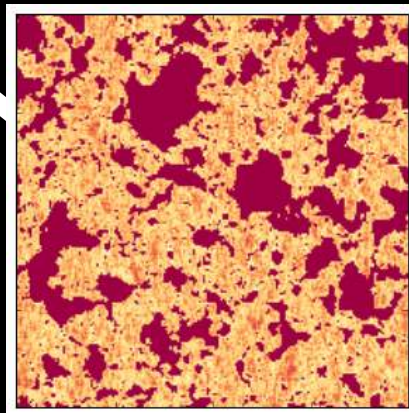
Mature Universe



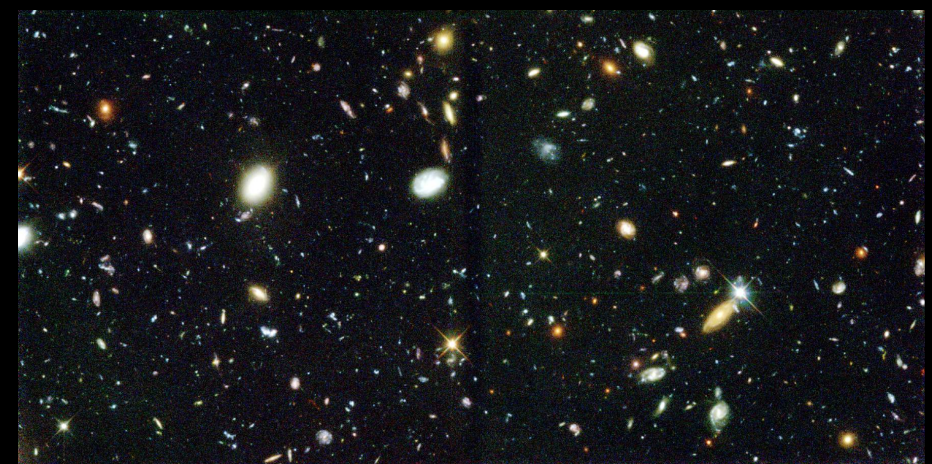
Baby Universe



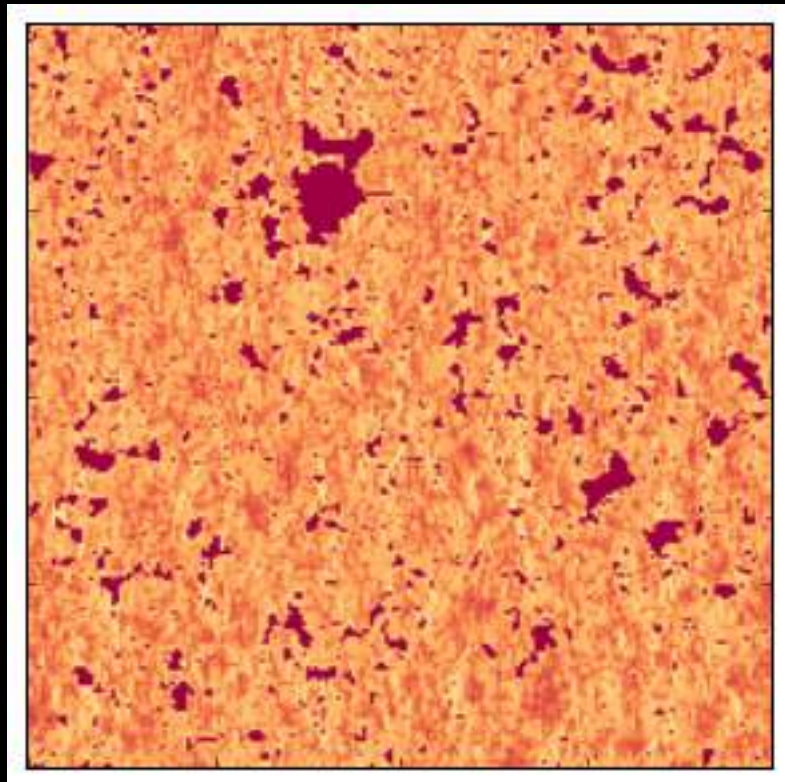
Reionization?



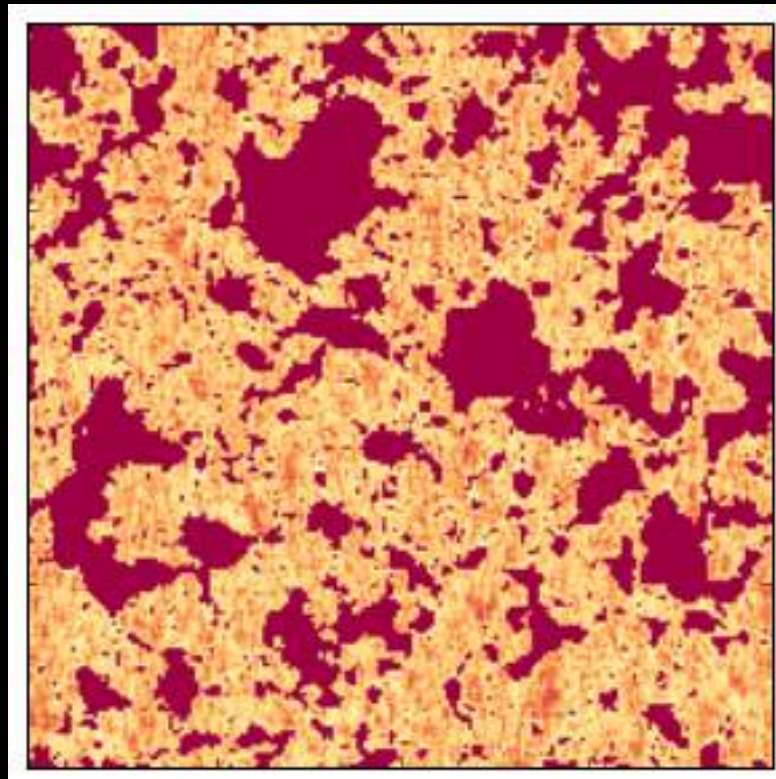
Time



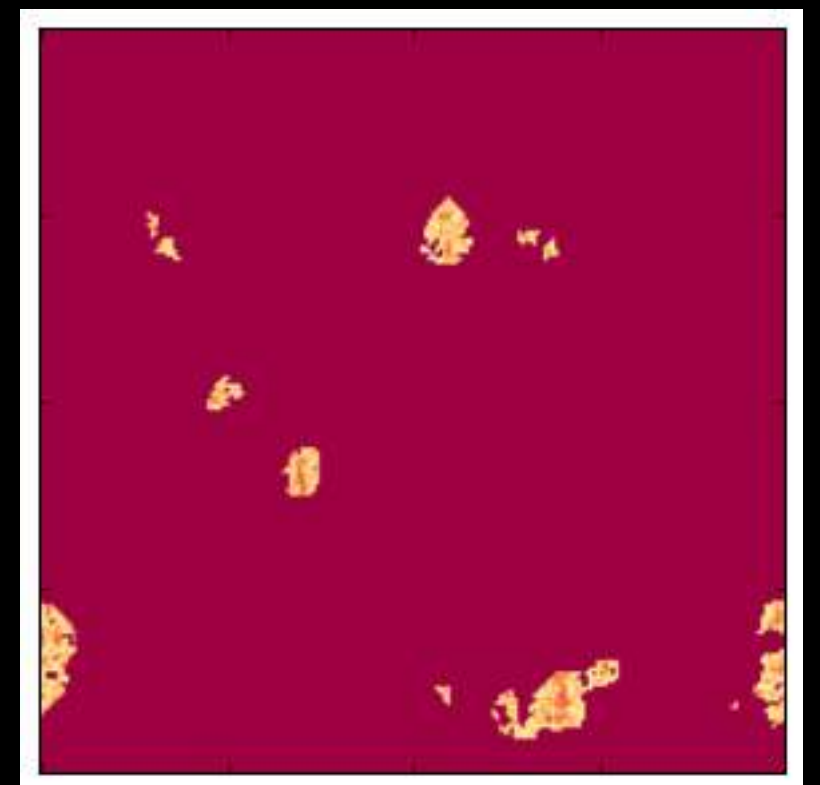
Mature Universe



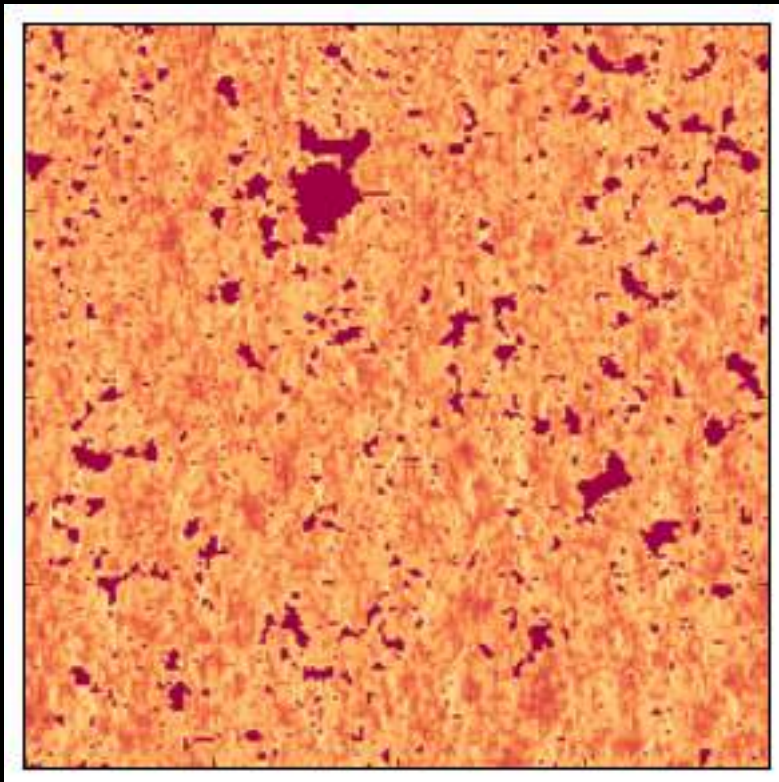
Beginning of
reionization



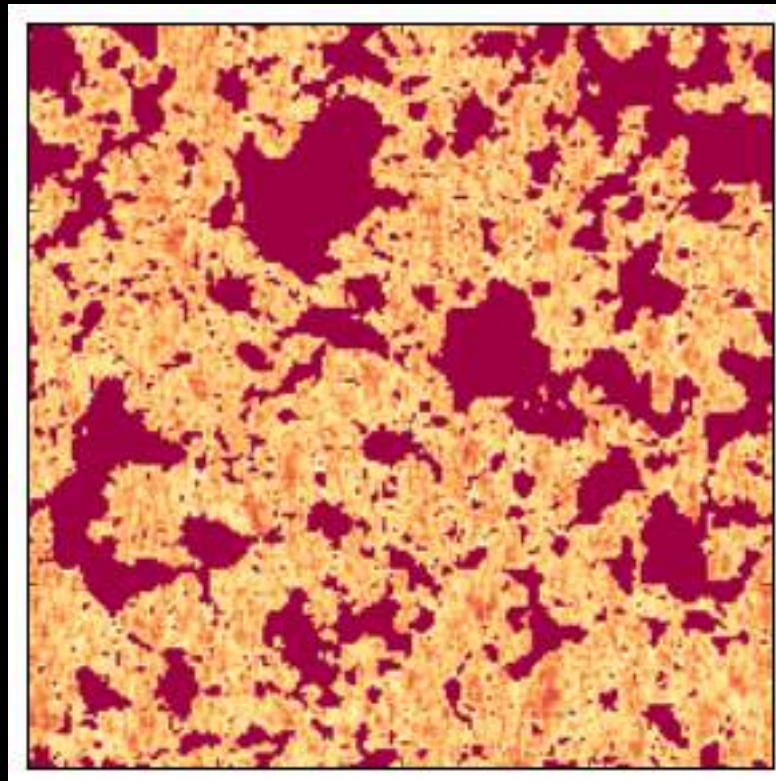
Middle of
reionization



End of reionization



Beginning of
reionization



Middle of
reionization

How did the first
galaxies affect their
environments?



End of reionization

What was the nature
of the first galaxies?