Nuclear Engineering 101: Midterm 2 Study Guide

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Nuclear Electromagnetic Moments

Ref: [1, pp. 71-75],[2, Lec 10]

A distribution of electric charge and/or current will create an electric potential. This can be described by the "multipole moment".

Note the values of L are the *order of the moment*, not angular momentum.

Electric multipoles

- Monopole moment: L=0, this is just the spherical nucleus, the electric potential looks like $V=\frac{Q}{R}$.
- Dipole moment: L=1. The electric dipole is just linear translation of the nucleus. V=0, see below.
- Quadrupole moment: L=2, the electric potential is from the nucleus squishing in one direction and then the other. The amount of quadrupole moment can be calculated:

$$eQ = e \int \psi^* (3z^2 - r^2) \psi \ d^3 \vec{r}$$

If the nucleus is spherical, $\langle z \rangle^2 = \langle x \rangle^2 = \langle y \rangle^2$, so eQ = 0. The higher the value of eQ the more deformed the nucleus is.

All electric moments have parity determined by $(-1)^L$. If you want to do some kind of electromagnetic operation on a state (ψ) with some operator O, then you have to evaluate $\int \psi^* O \psi \ dv$. It doesn't matter what the parity of ψ is because they will always multiply (even times even or odd times odd are both always even). If O has negative parity, the whole thing is an odd function and the integral goes to 0. Therefore, to keep things based in reality, all moments with odd angular momentum must not be a thing because they have negative parity (L=1,3,5 etc). This is why the dipole doesn't happen.

Magnetic multipoles

- Monopole moment: L=0 doesn't happen, as far as we know.
- Dipole moment: L=1, nucleons and nuclei have dipole magnetic moments due to their angular momentum (they are charges moving in a loop, therefore magnetic field).

$$\mu = \frac{e\hbar}{2m}l = g_l l\mu_n$$

Here l is the angular momentum, g_l is a constant (1 for protons, 0 for neutrons) and μ_n is the "nuclear magneton" (just a constant).

• Quadrupole moment: L = 3 doesn't happen.

Just like before, all the moments with odd parity have to not exist. For magnetic moments, the parity is determined by $(-1)^{L+1}$, so all the *even* order moments can't exist.

Shell Model

Evidence

- Ionization energies "jump" at "magic numbers" like 2, 10, 18, 36, 54, 86. Something about these make the nucleus more tightly bound, this wouldn't be seen in a liquid drop [2, Lec. 12].
- α -decay: see a big jump in the α -decay of Radon after N=128. This is because the daughter has N=126 (magic number) [2, Lec. 12].

Shells

- You can use the 3D square well to get a pretty good approximation of what we see, or a parabolic $(1/r^2)$ one to get better. But the simple harmonic oscillator gives the best approximation but it's still wrong. [2, Lec. 12].
- Spin-orbit interaction: due to the interaction of the spin and the angular momentum $(\vec{l} \bullet \vec{s})$, you can get two different values of j, $l \pm \frac{1}{2}$. Now two different nucleons will see a *different* potential, so it splits all the states in two. [2, Lec. 13-16], [1, pp. 123-125].
- The size of the split gets bigger with increasing l. The potential is negative so the $J = l + \frac{1}{2}$ states will occur at *lower* energies. High spin states that enter a different lower shell are called *the intruder*. [2, Lec. 13-16]
- Evidence: Magnetic dipole $\mu = \mu_N (g_l l_z + g_s s_z)/\hbar$ should be different for the different spins in the same l level. Observations show this is true. [2, Lec. 13-16].
- Works well for spherical nuclei and ones that are close to magic numbers.

Independent Particle Model

- All shells are completely full or empty except for a single particle in the lowest energy of an otherwise empty shell. The nucleus J^{π} depends only on that last nucleon. [2, Lec. 13-16].
- The total angular momentum of the nuclei J is equal to the J value of the shell with the single particle in it. The parity is equal to the parity of the shell with the single particle in it, $(-1)^l$. For example, if there is a single particle in the $1p_{3/2}$ state, then $J^{\pi} = (\frac{3}{2})^-$ because $j = \frac{3}{2}$ and l = 1. [2, Lec. 13-16].
- The value J^{π} is referred to as the spin-parity or the total angular momentum of the nucleus. Both of those things mean the same thing when you're talking about the nucleus. For a nucleon, spin is the intrinsic angular momentum; the nucleus doesn't have any intrinsic angular momentum, only angular momentum due to it's components. We call it spin anyway when referring to the nucleus just to be confusing.
- The independent nucleon always gives the total angular momentum of the shell to the nucleus when in the **ground state**.
- When you are in an **excited state**, you can line up the nucleons in more interesting way. If you have three nucleons in the $f_{7/2}$ shell and we're in an excited state, we use the notation $(f_{7/2})^3$. You can add up their angular momentum any way you want, remember each can have angular momentum from -j to +j. So these can have $m=\pm\frac{1}{2},\pm\frac{3}{2},\pm\frac{5}{2},\pm\frac{7}{2}$. They can't have the **same** value of m, but they can all add to get, for example $\frac{7}{2}+\frac{5}{2}+\frac{3}{2}=\frac{15}{2}$. [1, pp. 149-151]
- In excited states, the single nucleons can also jump around to other states, changing the J^{π} of the nucleus. [2, Lec. 13-16]

α -decay

- The α -decay half-life can be estimated using a barrier tunneling model. You can calculate the frequency of collision with the potential well and use the transmission probability to calculate a half-life. This isn't the real model and only represents an electromagnetic baseline to give us insight into what is happening. It isn't *nuclear*. [2, Lec. 18]
- Any angular momentum carried away by the α -particle is **purely orbital**. The spins all couple pairwise to 0. The angular momentum change from α -decay can range from:

$$|I_i - I_f| \le l_\alpha \le I_i + I_f$$

With parity change:

$$\Delta \pi = (-1)^{l_{\alpha}}$$

The simplest decay is any $0 \to X$ state, because the only value for l_{α} is X. If that doesn't give the right parity change, the decay is *forbidden*. In this case it really means *forbidden*, unlike β -decay, which plays fast and loose with the term. [1, pp. 257-258].

- α -decay can populate many different states. Each one has a different Q-value (given by the Q value for decay to the ground state, minus the excitation energy). [1, pp. 257-258].
- The intensity depends on how similar the initial and final wave functions are, and the angular momentum (l_{α}) . [1, pp. 257-258].
- The hinderence factor (H) tells us about the similarity between the parent and daughter states.:
 - -1 < H < 4 Favored: α -particle is from two low-lying pairs of nucleons.
 - -4 < H < 10: favorable overlap between initial and final state.
 - -10 < H < 100: spin projections are parallel. Unfavorable overlap between initial and final states.
 - -100 < H < 1000: spin projections are parallel, change in parity.
 - -H > 1000: spin-flip, change in parity.

Any spin flip \rightarrow over 1000. Any parity change \rightarrow over 100. Non-favorable overlap \rightarrow over 10. See Fig. 1. [2, Lec. 18]

| H | Nuclear State Overlap | Parity Change | Spins |
|------------|-----------------------|---------------|-----------|
| >10 | Favorable | None | Parallel |
| 10 - 100 | Unfavorable | None | Parallel |
| 100 - 1000 | Unfavorable | Yes | Parallel |
| >1000 | Unfavorable | Yes | Spin-flip |

Table 1: Summary of α -decay hinderance factors.

- α -decay spectrum has sharp peaks, each represents a decay to different excited states. Higher E_{α} , lower excitation energy. The excited states then usually release γ -ray photons. [1, pp. 262-264]
- In an even-even nucleus, the alpha decay to the ground state is usually *very strong*. The 0⁺ → 0⁺ decay is not inhibited by differences between the wave functions (they are very similar. By the same logic, the decay of odd-A nuclei may be *very weak* because the two states may be significantly different. [1, pp.264]

β -decay

- Energy release distribution is continuous because there are two particles emitted so they can share any proportion of the energy. [1, 2, pp. 273, Lec 19]
- Three kinds:

$$\beta^{-}: n \to p + e^{-} + \bar{\nu}_{e}$$
$$\beta^{+}: p \to n + e^{+} + \nu_{e}$$
$$EC: p + e^{-} \to n + \nu_{e}$$

Each has a different energy release:

$$Q_{\beta^{-}} = [m(^{A}X) - m(^{A}X')]c^{2}$$

$$Q_{\beta^{+}} = [m(^{A}X) - m(^{A}X') - 2m_{e}]c^{2}$$

$$Q_{\epsilon} = [m(^{A}X) - m(^{A}X')]c^{2} - B_{n}$$

For β^- , the electron masses all cancel when using atomic masses because $-Zm_e + (Z+1)m_e - m_e = 0$, that is the electrons accounted for in the atomic masses for the original atom, final atom and leftover electron all cancel. They don't cancel for β^+ because $-Zm_e + (Z-1)m_e - m_e = -2m_e$. [1, pp. 274-276]

• Fermi theory of beta decay: There are three things that go into determining the spectrum of beta decay energy values:

 $N(p) \propto \text{Statistical factor} * \text{Coulomb Interaction (electrodynamic)} * \text{Nuclear Matrix Element}$ $N(p) \propto p^2 (Q_\beta - T_e)^2 * F(Z_d, T_e) * |M_{fi}|^2 * S(p, q)$

These terms are:

- Statistical factor: related to the number of final states accessible. Decays like to happen when there are a lot of final states that are open.
- Fermi function (F): the influence of the nuclear coulomb field. This is the **electrodynamic** part.
- Matrix element $(M_{fi}^2 * S(p,q))$: This is the **Nuclear** part. It is kind of like the hinderance factors for α -decay and represents the effect of the initial and final states. The S function only comes into play in **forbidden** decays, when the electron and neutrino have angular momentum $(l \neq 0, s)$ and q are terms for the electron and neutrino angular momenta.

[1, pp.281-282].

• You can plot:

$$(Q-T_e) \propto \sqrt{\frac{N(p)}{p^2 F(Z',p)}}$$

Which should give you a straight line that intersects the x-axis at the Q value. For forbidden decays, the line is **not** straight, you have to put that S(p,q) factor in on the bottom to make it straight. This is the shape factor, which then gives you a straight line. [1, pp. 282]

• Total decay rate: you can integrate some messy function to get a messy function for total decay rate:

$$\lambda = \text{Matrix Element Term} * \int F$$
-function and statistical factor

The full equation is in Krane, page 282. What is important is that the term outside the \int is the **nuclear** part, and the integral is the **electrodynamic** part. Those are the two things (and some constants) that determine the decay rate. You can just crunch the integral and look up values, which is f, or the Fermi Integral. Then, if you do $\frac{f}{\lambda}$ or just $ft_{1/2}$, then the integrals cancel and all you have is the Matrix Element part, the **nuclear** part. This lets us compare $ft_{1/2}$ values, which are only dependent on nuclear properties. It might be useful to think about the ft value as just a corrected half-life. Corrected to remove electrodynamic stuff. [1, 2, pp. 282-283, Lecs. 19-21].

• ft values have a huge range (10³ to 10²⁰) so we use $\log ft$ instead.

References

- [1] Kenneth S. Krane. Introductory Nuclear Physics. John Wiley & Sons, Inc., 3rd edition, 1988.
- [2] Lee Bernstein. Nuclear engineering class lectures. Fall 2015.