Toy Model For Hyperfine Measurement

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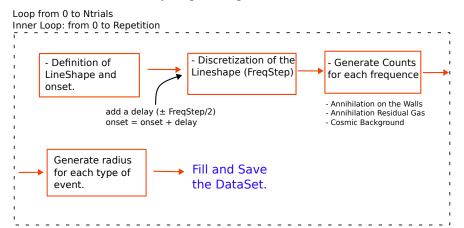






Scheme of the Simulation

Scheme of the Monte Carlo Toy for generating the events



In this simulation, the data are created and analyzed using RDataFrame framework.



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A brief introduction about the Monte Carlo

We have developed a Monte Carlo Toy that produces two .root dataset files. The variables are columns of values that are shown in the figure below:

```
* Row * runNumber * random.ra * delay.del * frequence * type.type * radius.ra *

* 0 * 0 * 0.4849736 * 2.4987087 * -25 * 2 * 2.8792768 *

1 * 1 * 0.2899349 * -1.685450 * 15 * 0 * 2.0739069 *

2 * 1 * 0.0197818 * -1.685450 * 15 * 0 * 1.8959179 *

3 * 1 * 0.2412778 * -1.685450 * 15 * 0 * 2.8919173 *

4 * 1 * 0.3846191 * -1.685450 * 15 * 0 * 3.3842529 *

5 * 1 * 0.4549068 * -1.685450 * 15 * 0 * 1.9130180 *

6 * 1 * 0.3739825 * -1.685450 * 15 * 0 * 1.047382 *
```

Figure: Structure of the dataset.

- runNumber: identifies which run the event belong to (from 0 to Repetition 1)
- random: values uniform distributed from 0 to 1, can be used to randomize the selection or for sub-sampling in the data
- delay: store the onset delay
- frequence: the frequency of the event
- type: type of the event: 0 annihilation on the walls, 1 residual gas annihilation, 2 cosmic event
- radius: radius of the annihilation vertex.

A brief introduction about the Monte Carlo

The Annihilation on the walls are generated as function of the frequency, using the two line-shapes of the transitions (c \rightarrow b) and (d \rightarrow a). The Annihilation on the residual gas and the cosmic background are generated uniformly on the frequency spectrum. All the parameters of the simulation are loaded from the ToyConfiguration.txt file. The parameters are chosen to reproduce the runs 4b.

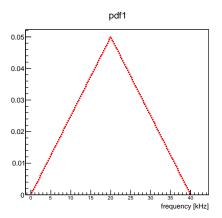
```
Nstack = 20
NHbar = 14
Repetition = 5
# cosmic rate is expressed in event/second
CosmicRate = 0.051028571
Efficiency = 1
pwall cb = 1
pwall ad = 1
C = 0.5
FrequencyStep = 5
TimeStep = 8
SweepStep = 24
# The following are in kHz units
x cb start = 0
x cb end = 40
x cb peak = 20
x da start = 1420000
x da peak = 1420020
x da end = 1420040
delav = 2.5
```

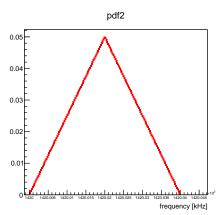
Figure: Parameters of the Toy.



Triangular Line-shape Pdfs

For this first use of the toy, we have chosen simple line-shapes, triangular with a symmetric rise and fall.

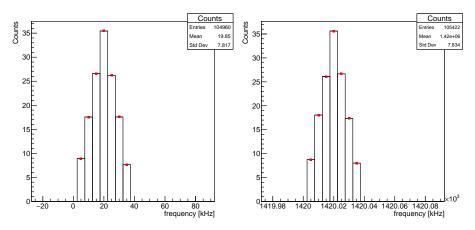






Triangular Line-shapes Simulation

We sample at the given frequency step of $5\,\text{kHz}$ the Pdfs, to simulate the experimental line-shapes. We applied the onset finding algorithm to this distribution.



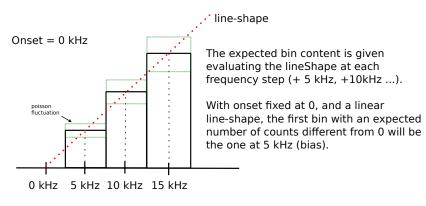
The onset is fixed ad $f=0\,\mathrm{kHz}$. In this sample the cosmic background is set to zero.



A simple Onset finding Algorithm

The first algorithm that is tested is quite simple: the onset is identified by the first bin with a content over a given threshold $(>N\mu_{cosmic})^{1}$

Before showing the plot with the simulated data, it is useful to remind how this algorithm deals with the frequency step: $\frac{1}{2}$





 1 Where the μ_{cosmic} is computed from the Poisson distribution of the cosmic counts expected per bin.

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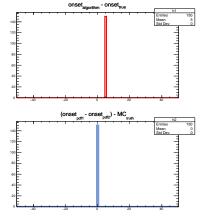
Alpha 2

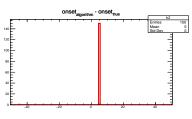
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Consistency Check 1

We have tested the algorithm with a dataset without cosmic background and delay fixed to zero. The algorithm identifies the onset at frequency $5\,\mathrm{kHz}$.

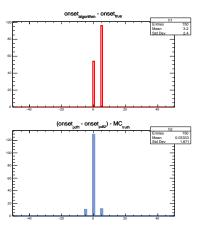


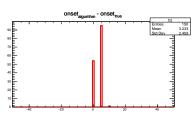




Consistency Check 2

We have tested the algorithm with a dataset without cosmic background. The delay is uniform distributed in $-2.5\,\mathrm{kHz}$ and $2.5\,\mathrm{kHz}$.

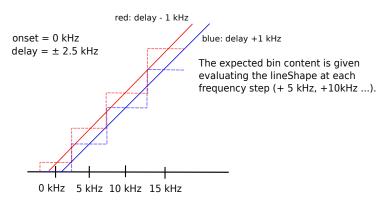






Consistency Check 2

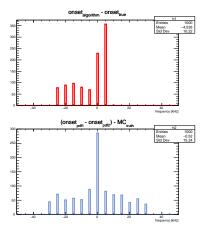
With the delay, two bins (frequency = 0 kHz and frequency = 5 kHz) are populated.

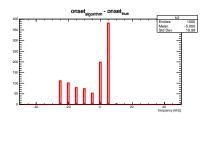




algorithm test: threshold $> 3\mu_{cosmic}$

In this case, we have applied the algorithm to simulated data with cosmic background (using a rate of $0.051\frac{event}{s}$, from passcut1). Each bin has an expected cosmic background of $dwelltime \cdot rate = 0.408$. The delay is uniform distributed in $-2.5\,\mathrm{kHz}$ and $2.5\,\mathrm{kHz}$.

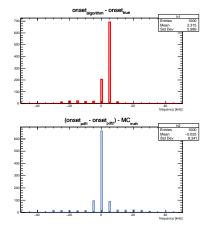


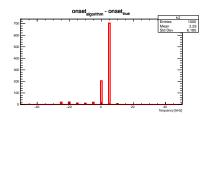




algorithm test: threshold $> 5\mu_{cosmic}$

In this case, we have applied the algorithm to simulated data with cosmic background (using a rate of $0.051\frac{event}{s}$, from passcut1). Each bin has an expected cosmic content of dwelltime \cdot rate = 0.408. The delay is uniform distributed in $-2.5\,\mathrm{kHz}$ and $2.5\,\mathrm{kHz}$.

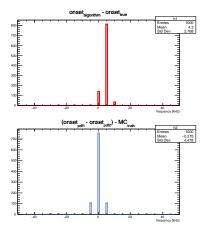


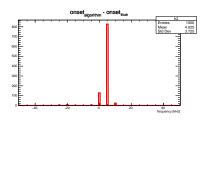




algorithm test: threshold $> 8\mu_{cosmic}$

In this case, we have applied the algorithm to simulated data with cosmic background (using a rate of $0.051\frac{event}{s}$, from passcut1). Each bin has an expected cosmic content of $dwelltime \cdot rate = 0.408$. The delay is uniform distributed in $-2.5\,\mathrm{kHz}$ and $2.5\,\mathrm{kHz}$.







Next step

- different line-shapes (e.g. quadratic rise, etc.)
- different onset-finding algorithm
- simulation of repetition/runs with Bfield drift.



Prob algorithm is triggered by the cosmic for threshold $> 3\mu_{cosmic}$ baseline over threshold is P=6.37% (Poisson distribution, 1 - P(k=0) - P(k = 1)), calculation bring to 6.37%

