Studying Annihilation Distributions

Adriano Del Vincio, Germano Bonomi, Simone Stracka

University of Brescia, INFN Pisa

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Objective

A procedure to disantangle the contribution of annihilations of antihydrogen on the trap walls and on the residual gas has been developed and tested. It is based on a Maximum Likelihood fit of the radial distribution of the annihilations that has been tuned using an "ad-hoc" Monte Carlo simulation toy.

- The model/template for annihilations on the trap walls has been extracted from the 2-second mixing window.
- The model/template for annihilations of the residual gas has been obtained using the losses when antihydrogen is hold in the trap (microwaves - UW - losses).
- A model/template for the cosmic background is also used.

The PDFs used for the fit procedure and the data generation the data are listed below:

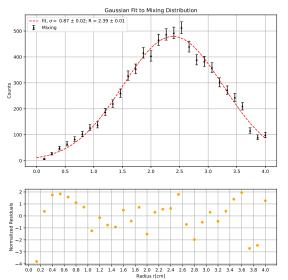
- PDF Mixing: the Normal distribution
- PDF Residual gas: Rayleigh distribution $\frac{r}{\sigma^2}e^{-\frac{r^2}{2\sigma^2}}$
- PDF Cosmic: $k \cdot x$

The factor k is the normalization constant. The Mixing, residual gas and cosmic data are fitted and the result are shown in the following slides.

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MIXING

Mixing dataset represents almost pure data of anti-hydrogen annihilation on the walls. The radius distribution is fitted with a Gaussian.

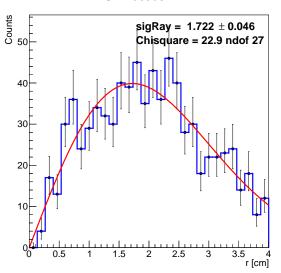




Residual Gas

The microwave losses represent an almost pure anti-hydrogen sample of annihilation events due to residual gas inside the trap.

UW losses PDF

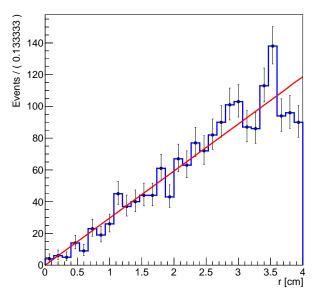




Cosmic

The cosmic distribution is obtained from dataset without anti-hydrogen.

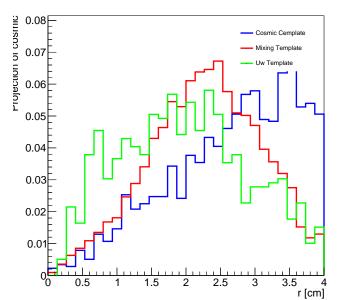






PDFs normalized plotted together







Radial Density

The histogram in r variable doesn't account for the different area of the bin which is $2\pi r \cdot dr$. So it is useful to divide per $2\pi r$ to obtain the radial density of the events.

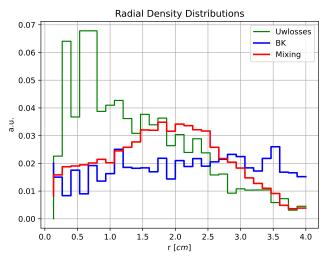


Figure: Radial density for r68465_uw_exp_freq4.vertex.csv dataset

Monte Carlo Simulation Toy

To study the accuracy of the algorithm to reconstruct the various parameter, we have developed a "toy" simulation tool. The model to generate the data is:

$$PDF_{total} = a \cdot Gauss_{mix} + b \cdot Rayleigh + c \cdot linearModel_{cosmic}$$
 (1)

where a, b, c represent the "weights" of the various contributions to the PDF used to generate the data. The number of annihilation is indicated as N_{sample} . Once the data are generated, they are fitted with the model:

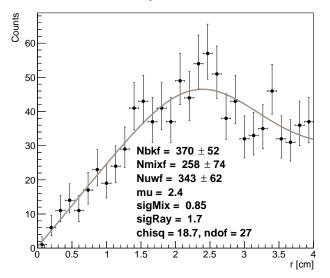
$$PDF_{fit} = Nfit_{mix} \cdot Gauss_{mix} + Nfit_{uw} \cdot Rayleigh + Nfit_{bk} \cdot linearModel_{cosmic}$$
 (2)

The parameters of the fit are $Nfit_{mix}$, $Nfit_{uw}$ and $Nfit_{bk}$. The "true value" are defined as:

- $Ngen_{mix} = a \cdot N_{sample}$
- $Ngen_{uw} = b \cdot N_{sample}$
- $Ngen_{bk} = c \cdot N_{sample}$



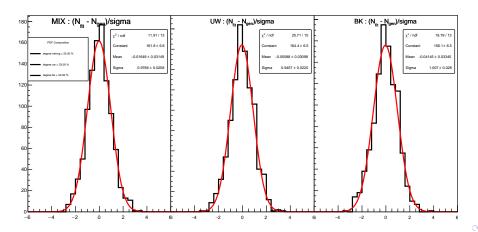
Toy Model Fit





Toy: $N_{sample} = 1000$, a = 33%, b = 33%, c = 33%

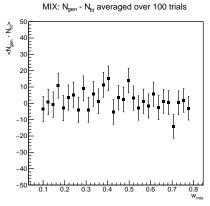
In this plot we have fixed the weight of each distribution to 33%, with $N_{sample}=1000$ and $N_{trials}=1000$, to ensure that the algorithm is able to reconstruct the parameters, and check the presence of a bias. The variable of the histograms are: $\frac{Nfit-Ngen}{\sigma_{fit}}$. The distributions are normal and the fit procedure is behaving as expected.

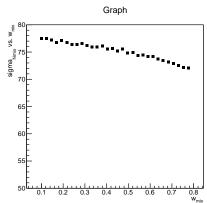


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weight variation a for mix.

Now we study how the coefficients of the fit $Nfit_{mix}$, $Nfit_{uw}$ and $Nfit_{bk}$ vary with the increment of the weight a. At fixed c=10%, a is raised from 10% to 80% and b is decreased accordingly.



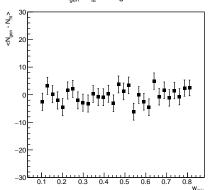


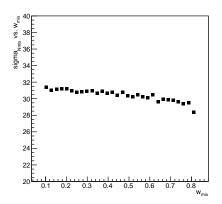
The number events is always $N_{sample} = 1000$. For each value of the weight a we iterate 100 times ($N_{trials} = 100$) to study the reconstructed coefficients with the AL variation of the weights.

$N_{sample} = 165.$

Now we have done the same plot with N=165, the same amount of data in r68465_uw_exp_freq4.vertex.csv after applying cut1. The value of c is fixed to reproduce the number of expected events from background (c=6%).

MIX: N_{gen} - N_{fit} averaged over 100 trials



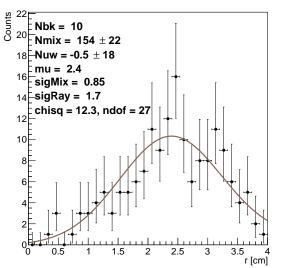


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Fit to Data

$$\label{eq:pdf} \begin{split} \mathsf{PDF} &= \mathsf{Gaussian} \; \big(\mathsf{Mixing}\big) + \mathsf{Rayleigh} \; \big(\mathsf{Residual} \; \mathsf{gas}\big) + \mathsf{linear} \; \mathsf{model} \; \big(\mathsf{cosmic} \; \mathsf{fixed}\big). \\ \mathsf{Data} \; \mathsf{taken} \; \mathsf{from:} \; \mathsf{r68465_uw_exp_freq4.vertex.csv} \end{split}$$

r68465_uw_exp_freq4.vertex.csv



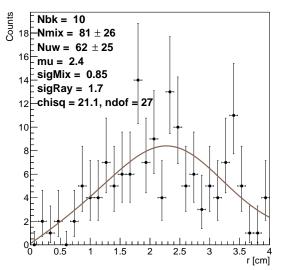


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Fit to Data

$$\label{eq:pdf} \begin{split} \mathsf{PDF} &= \mathsf{Gaussian} \; (\mathsf{Mixing}) + \mathsf{Rayleigh} \; (\mathsf{Residual} \; \mathsf{gas}) + \mathsf{linear} \; \mathsf{model} \; (\mathsf{cosmic} \; \mathsf{fixed}). \\ \mathsf{Data} \; \mathsf{taken} \; \mathsf{from:} \; \mathsf{r68465_uw_exp_freq5.vertex.csv} \end{split}$$

r68465_uw_exp_freq5.vertex.csv





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ADDITIONAL MATERIAL



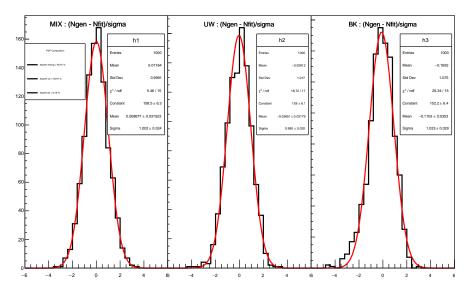
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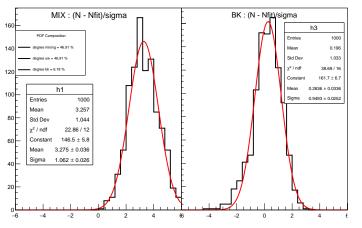
 $N_{mix} - N_{fit}$ for a = 46%, b = 46%, c = 6%.



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N_{uw} parameter of the fit model fixed

The Toy simulation is tested fixing the N_{uw} parameter of the fit model to 0. In the following plot the weight are a=46%, b=46%, c=6%, where c is fixed in such a way to reproduce the number of expected background events in r68465_uw_exp_freq4.vertex.csv, which correspond to 6% of the total events.



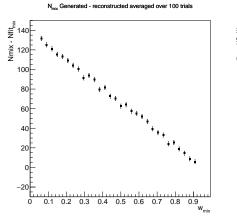


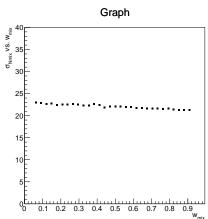
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N_{uw} parameter of the fit model fixed

We study the bias $N_{mix} - N_{reconstructed}$ with the parameter N_{uw} of the fit model fixed to 0. For small value of w_{mix} , corresponding to small contribution of Mixing (and, conversely, a significant contribution of $Residual\ gas\ PDF$) we observe a large bias.





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