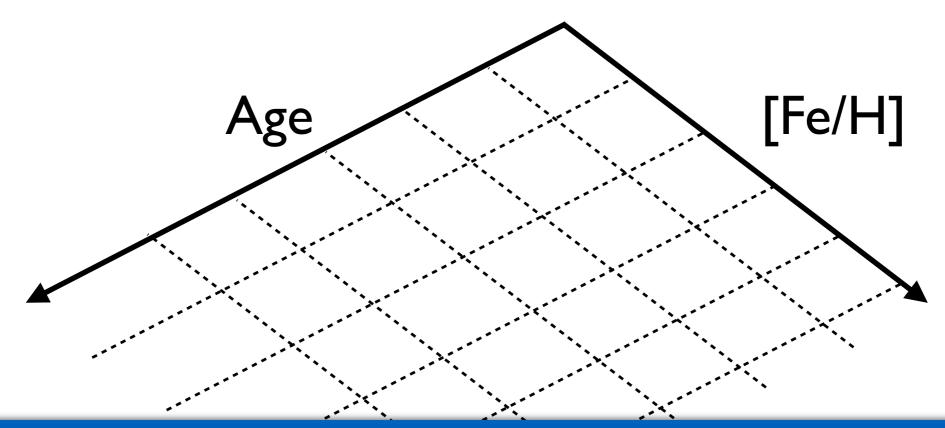
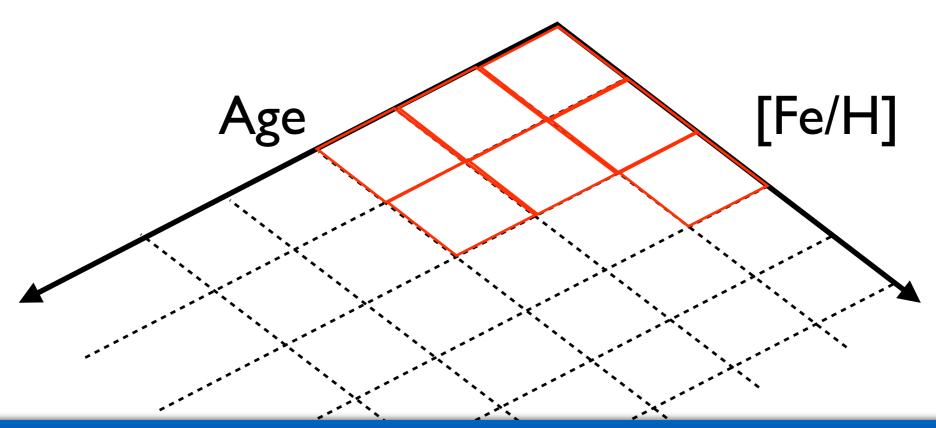
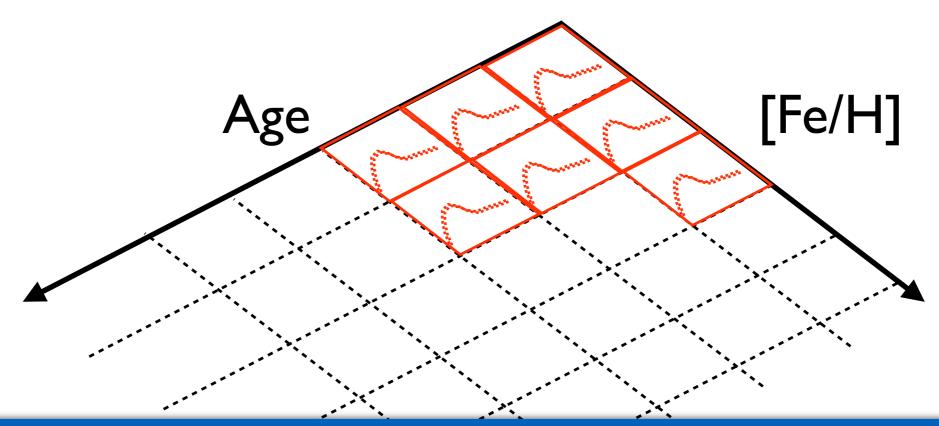
Step1: Create a grid of age and [Fe/H]



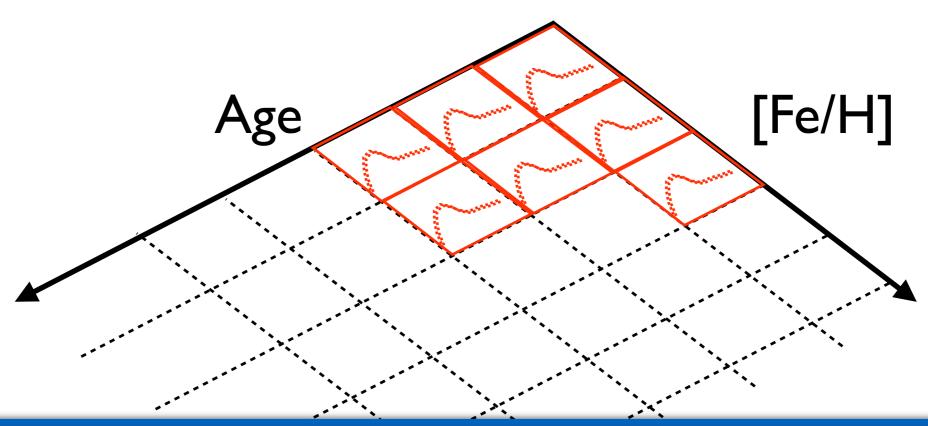
Step2: at each point assign an isochrone of that age and metallicity

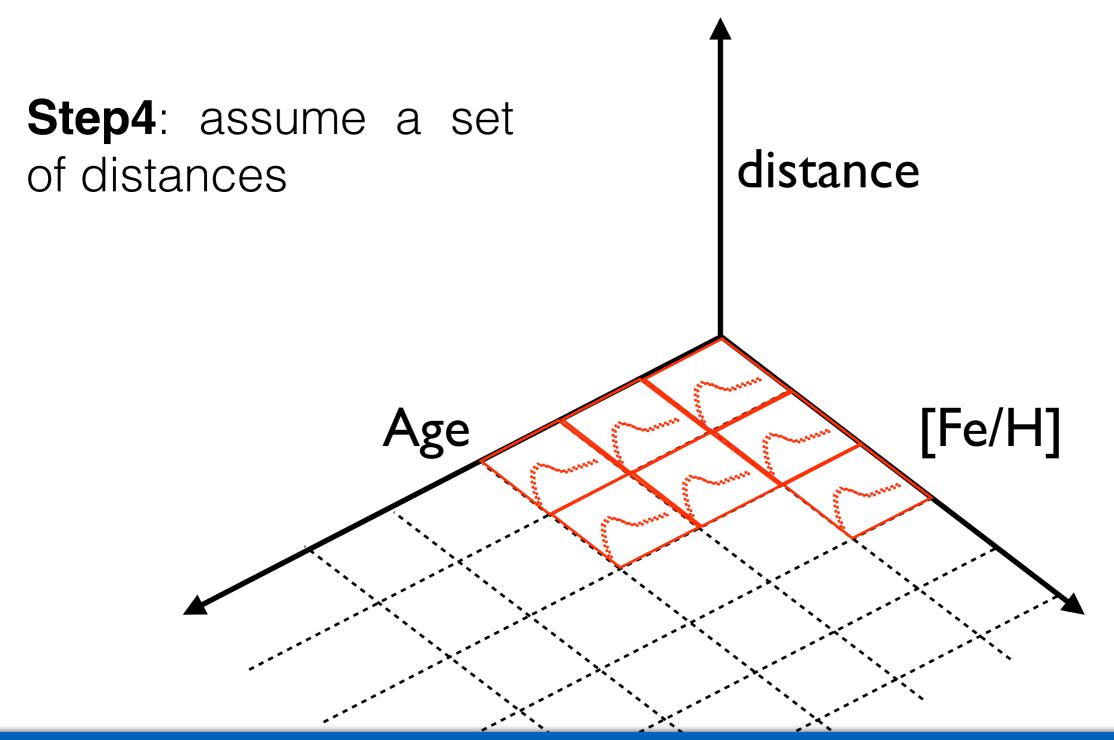


Step3: populate each isochrone with an IMF



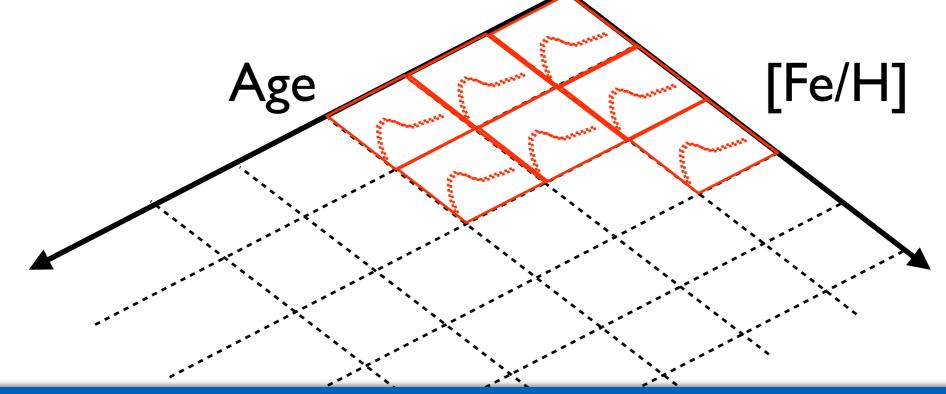
at the end of Step 1-3, at each point on the grid we will have a synthetic population of stars described by absolute magnitudes (M_K) and colours (J-K)





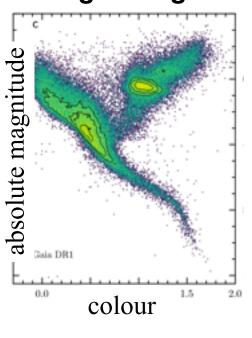
at a given (age, [Fe/H], distance) we will be able to compute apparent magnitudes m_K

distance

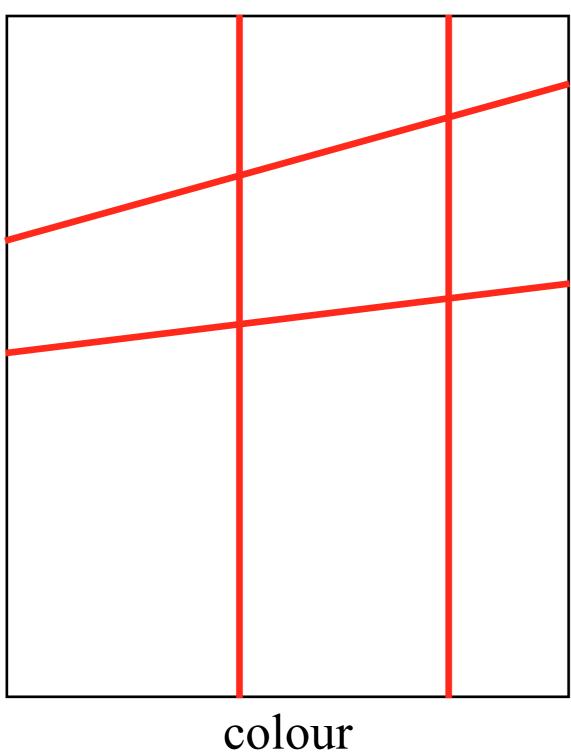


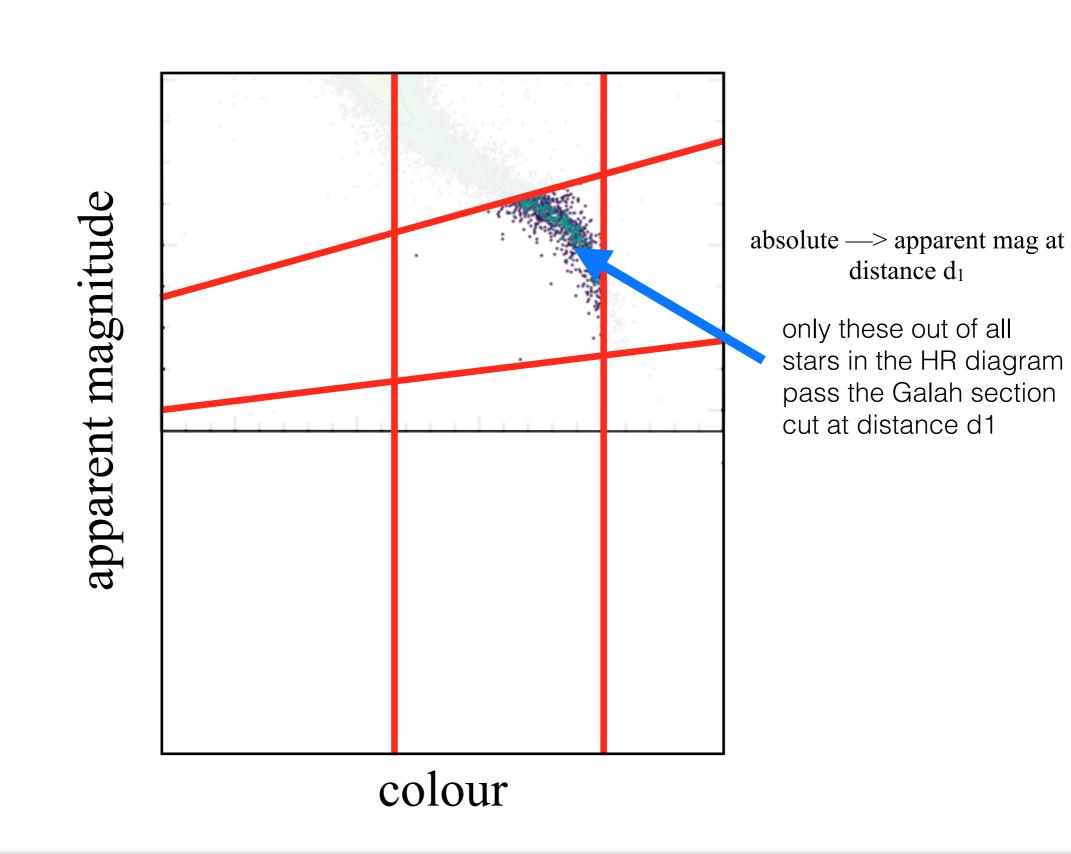
1) This is our synthetic HR diagram at a given age and [Fe/H]

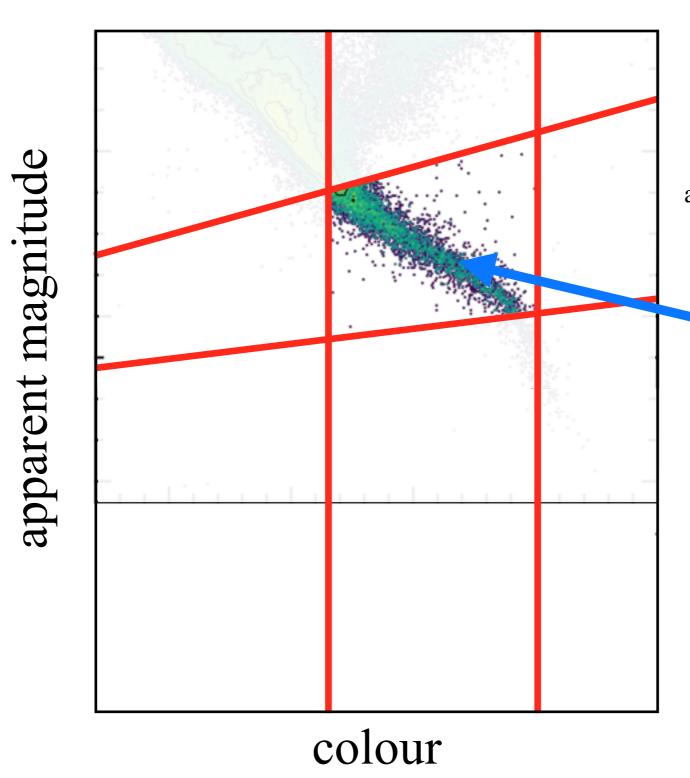
apparent magnitude



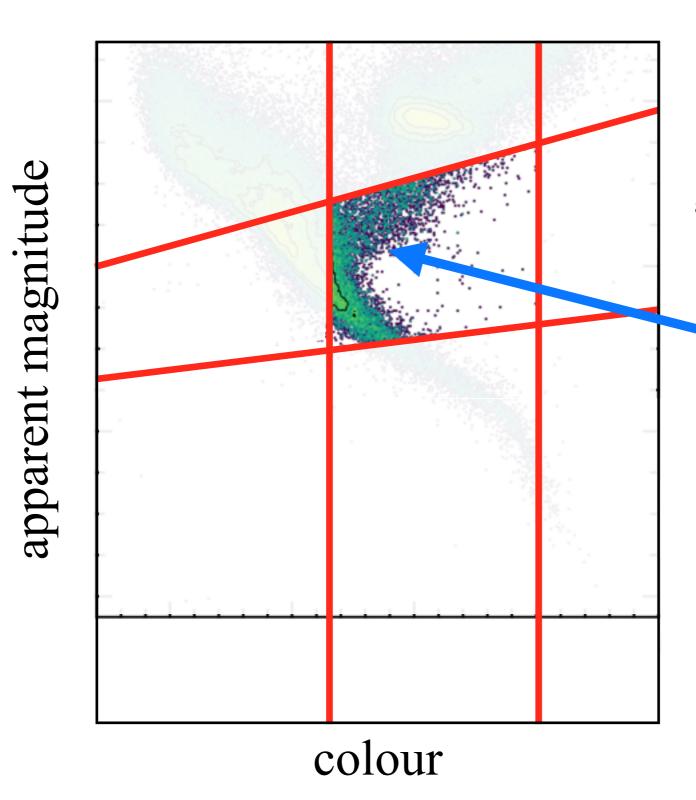
2) Now we apply a set of distances and at each distance use the same m_K and J_K selection of Galah



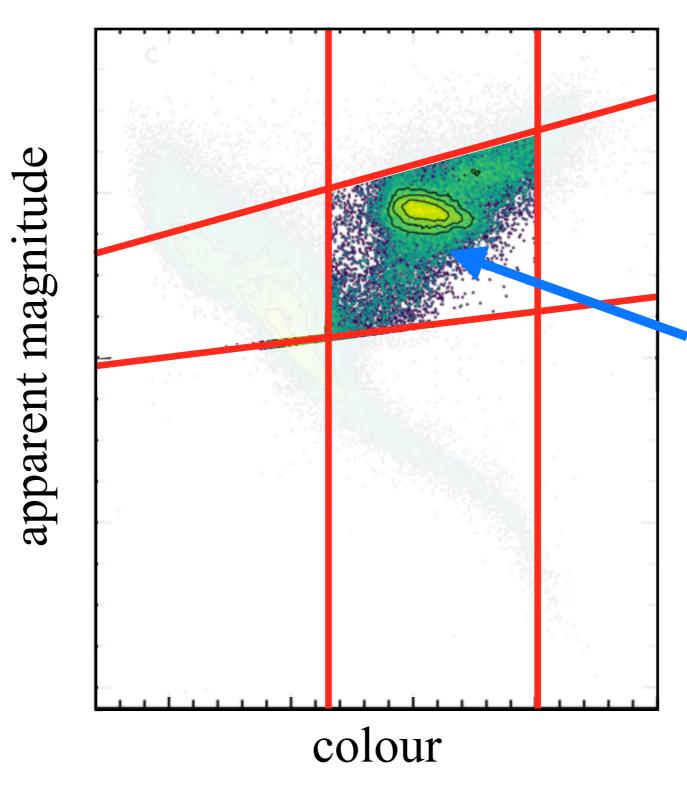




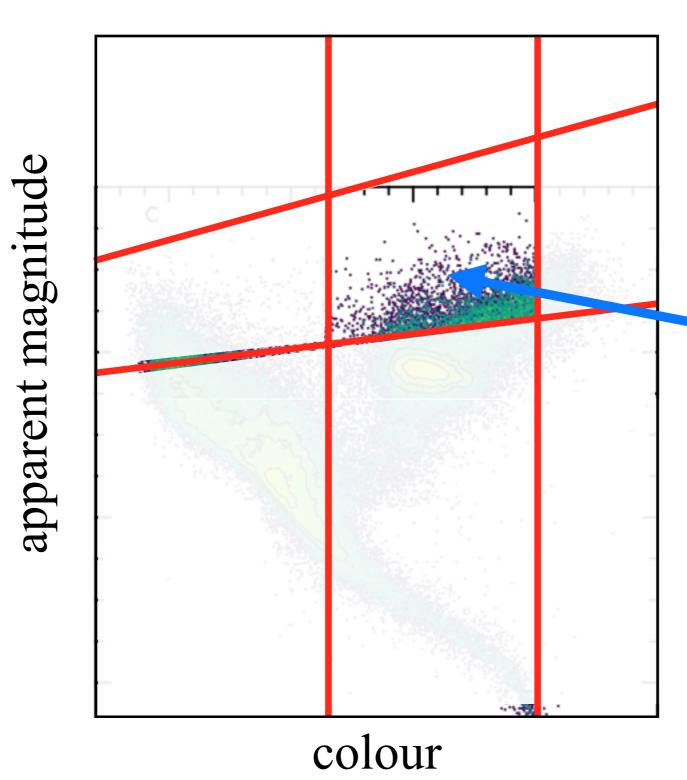
absolute \longrightarrow apparent mag at distance $d_2 > d_1$



absolute \longrightarrow apparent mag at distance $d_3 > d_2$



absolute \longrightarrow apparent mag at distance $d_4 > d_3$



absolute —> apparent mag at distance $d_5 > d_4$