```
In [1]:
         import array
         import numpy as np
         import matplotlib.pyplot as plt
         binfpath="/fs/lustre/cita/hqchen/data/z-ifrit-a=0.1401.bin"
         with open(binfpath, 'rb') as f:
             binarydata=f.read()
         lllen=8; flen=4; ilen=4
         N=1024
         firstfield=lllen*2+3*ilen
         nextfield=firstfield
         xHI=array.array('f',binarydata[nextfield+lllen:nextfield+lllen+N**3*flen])
         xHI=np.reshape(xHI,(N,N,N),order='F')
         nextfield+=lllen+N**3*flen
         delta=array.array('f',binarydata[nextfield+lllen:nextfield+lllen+N**3*flen]
         delta=np.reshape(delta,(N,N,N),order='F')
         nextfield+=lllen+N**3*flen
         T=array.array('f',binarydata[nextfield+lllen:nextfield+lllen+N**3*flen])
         T=np.reshape(T,(N,N,N),order='F')
```

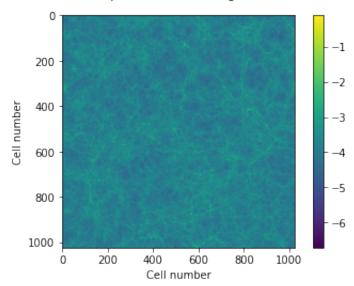
```
In [2]:
         fig1, ax1 = plt.subplots()
         im1 = ax1.imshow(np.log10(xHI[0]))
         ax1.set(xlabel = "Cell number", ylabel = "Cell number")
         fig1.colorbar(im1)
         fig1.suptitle("A plot of a slice of $\log(xH I)$")
         fig2, ax2 = plt.subplots()
         im2 = ax2.imshow(np.log10(delta[0]))
         fig2.colorbar(im2)
         ax2.set(xlabel = "Cell number", ylabel = "Cell number")
         fig2.suptitle("A plot of a slice of $\log(\Delta)$")
         fig3, ax3 = plt.subplots()
         im3 = ax3.imshow(np.log10(T[0]))
         fig3.colorbar(im3)
         ax3.set(xlabel = "Cell number", ylabel = "Cell number")
         fig3.suptitle("A plot of a slice of $\log(T)$")
```

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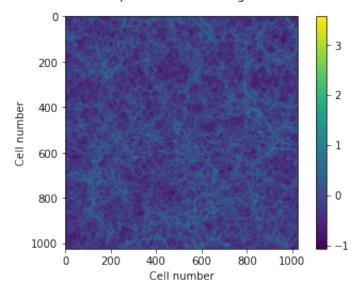
```
/tmp/ipykernel_2004116/591149048.py:8: RuntimeWarning: divide by zero encoun
tered in log10
  im2 = ax2.imshow(np.log10(delta[0]))
/tmp/ipykernel_2004116/591149048.py:14: RuntimeWarning: divide by zero encou
ntered in log10
  im3 = ax3.imshow(np.log10(T[0]))
Text(0.5, 0.98, 'A plot of a slice of \(T)')
```

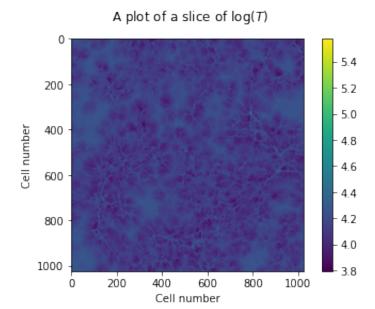
Out[2]:

A plot of a slice of $log(xH_I)$



A plot of a slice of $log(\Delta)$





Task 2

```
In [3]:
    import astropy
In [4]:
    from astropy.cosmology import WMAP9 as cosmo
    from astropy import units as u
    crt_dns = cosmo.critical_density(6)
    mean_gas_density = crt_dns*cosmo.Ob(6)
    print(mean_gas_density)
```

1.432778841086186e-28 g / cm3

```
In [6]:
    sigma = 4.48e-18
    nH = (0.76*mean_gas_density/p)
    nHI = xHI*nH
    nHI = nHI.astype('float128')
    cml = (40*0.68)/1024*u.Mpc
    ds = cml*(1/(1+6))
    ds = ds.to(u.cm)
    dT = sigma * nHI * ds
    print(dT.max())

2.910085329003365 1 / cm2
In [7]:
    print(nHI.max())
```

5.547614273382351e-05 1 / cm3

```
In [8]:
         from numpy import float128
         print(nHI)
        [[[6.74571377e-09 6.38827302e-09 5.75865577e-09 ... 1.07600648e-08
           8.21360135e-09 7.04414260e-09]
          [7.67662289e-09 6.92509960e-09 6.13529849e-09 ... 1.13985843e-08
           9.96054794e-09 8.55018278e-09]
          [9.01550923e-09 7.51253726e-09 6.50632348e-09 ... 1.14072840e-08
           1.12177583e-08 1.05493712e-081
          [5.40711653e-09 5.72221870e-09 6.19928953e-09 ... 6.03759931e-09
           5.55259971e-09 5.45984546e-091
          [5.71314507e-09 5.72291681e-09 5.85148463e-09 ... 7.31123873e-09
           6.16871665e-09 5.86614579e-09]
          [6.06233908e-09 5.96980865e-09 5.72571146e-09 ... 9.11899800e-09
           6.97984026e-09 6.31654506e-09]]
         [[5.48221868e-09 5.28250510e-09 5.17780352e-09 ... 8.64833005e-09
           7.15157933e-09 6.03492856e-09]
          [6.61523991e-09 5.69851810e-09 5.04244824e-09 ... 9.86796511e-09
           9.04403130e-09 7.89581023e-09]
          [8.48688053e-09 \ 6.67445255e-09 \ 5.43424594e-09 \ \dots \ 9.99739669e-09]
           1.04928599e-08 1.01232711e-081
          [5.35914690e-09 5.90746518e-09 6.73172940e-09 ... 5.94980287e-09
           5.42529799e-09 5.36209166e-09]
          [5.29865085e-09 5.49629009e-09 6.00415806e-09 ... 6.46318110e-09
           5.62216762e-09 5.38439249e-091
          [5.24459631e-09 5.30382804e-09 5.45351719e-09 ... 7.39922790e-09
           6.04202244e-09 5.43723200e-09]]
```

```
[[5.65969582e-09 5.40447687e-09 5.32718669e-09 ... 8.78586093e-09
  7.66012498e-09 6.51277920e-09]
 [6.74983269e-09 5.43955547e-09 4.94069496e-09 ... 1.01115383e-08
 9.74585923e-09 8.50905479e-091
 [8.27511659e-09 5.94726091e-09 4.65099470e-09 ... 1.02272972e-08
 1.10958682e-08 1.05951843e-081
 [5.74321213e-09 6.75065648e-09 8.07553402e-09 ... 6.12818774e-09
 5.51611823e-09 5.61805180e-09]
 [5.43424594e-09 5.91756999e-09 6.90652691e-09 ... 6.49798881e-09
 5.70652325e-09 5.57161073e-091
[5.37224709e-09 5.50367529e-09 5.88012572e-09 ... 7.37938377e-09
 6.22773211e-09 5.55938140e-09]]
[[1.19268533e-08 1.15134302e-08 1.07251497e-08 ... 1.40352689e-08
  1.33990881e-08 1.25596511e-08]
 [1.25206165e-08 1.14769465e-08 1.03723554e-08 ... 1.24436701e-08
 1.33240592e-08 1.31913547e-08]
 [1.21977752e-08 1.05859739e-08 8.71948558e-09 ... 1.04326725e-08
 1.20314079e-08 1.29868418e-08]
 [6.00654060e-09\ 6.70937395e-09\ 7.05845338e-09\ \dots\ 7.74368214e-09
 5.96088867e-09 5.85237769e-09]
 [7.55000151e-09 7.81013476e-09 7.81824294e-09 ... 1.09410241e-08
 8.33111624e-09 7.68858133e-09]
 [9.84165016e-09 9.77997416e-09 9.55050616e-09 ... 1.34060398e-08
  1.13394298e-08 1.01863371e-08]]
[[1.24550663e-08 1.11333485e-08 9.92762139e-09 ... 1.59916471e-08
  1.53827706e-08 1.38836107e-081
 [1.25435635e-08 1.07303881e-08 9.55896251e-09 ... 1.55223567e-08
 1.58943401e-08 1.49971093e-08]
 [1.19662289e-08 \ 9.15804232e-09 \ 7.71655273e-09 \ \dots \ 1.23242456e-08
 1.45179655e-08 1.47393902e-08]
 [5.99372241e-09 6.77149625e-09 7.40306660e-09 ... 6.24009866e-09
 5.38948841e-09 5.57586333e-091
 [7.96806532e-09 8.31029379e-09 8.14707146e-09 ... 9.95279947e-09
 7.95883093e-09 7.80632181e-091
 [1.04770210e-08 1.01515791e-08 9.23015087e-09 ... 1.40485517e-08
  1.17774750e-08 1.08903908e-08]]
[[9.89857885e-09 9.12902198e-09 7.99851918e-09 ... 1.50374380e-08
  1.27527651e-08 1.10960370e-08]
 [1.03382263e-08 8.87411211e-09 7.83998555e-09 ... 1.50668384e-08
 1.42406380e-08 1.26288322e-081
 [1.03281339e-08 8.21122015e-09 7.27207450e-09 ... 1.27465389e-08
 1.41608973e-08 1.31094859e-08]
 [5.77009107e-09 6.13660900e-09 6.66895561e-09 ... 6.04331385e-09
  5.52572210e-09 5.72868286e-09]
```

```
[7.17705095e-09 7.16960757e-09 6.77356216e-09 ... 8.78170692e-09 7.27096516e-09 7.18669479e-09]
[8.72241124e-09 8.30928037e-09 7.34680095e-09 ... 1.21840085e-08 9.77877868e-09 8.99050789e-09]]] 1 / cm3

In [9]: from numpy import float128

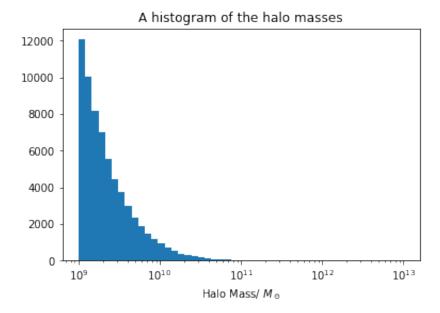
new_tau = dT.max().value print(new_tau) percent = float128(10) percent = float128(np.exp(-new_tau)*100) print(percent)

2.9100853290033651996 5.4471081707766847687
```

```
In [10]: halofpath="/fs/lustre/cita/hqchen/data/hlist_0.14005.list"
halodata = np.loadtxt(halofpath, skiprows = 63)
halomass = halodata[:, 10]

#mass = []
#for i in range(len(halodata)):
# mass = np.append(mass, float128(halodata[i][83:94]))
```

```
In [11]:
    plt.hist(halomass, bins = np.logspace(start =np.log10(10**9), stop = np.log
    plt.gca().set_xscale("log")
    plt.title("A histogram of the halo masses")
    plt.xlabel("Halo Mass/ $M_\odot$")
    plt.show()
```

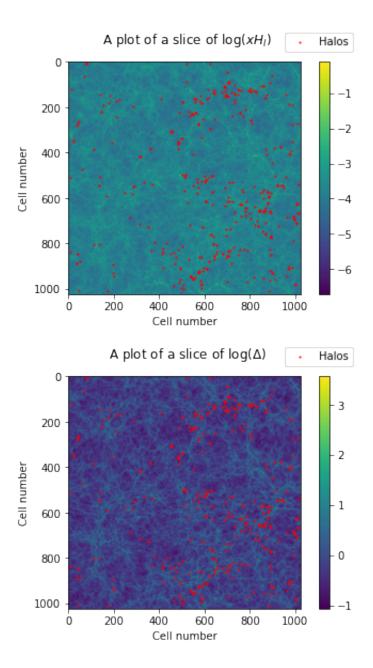


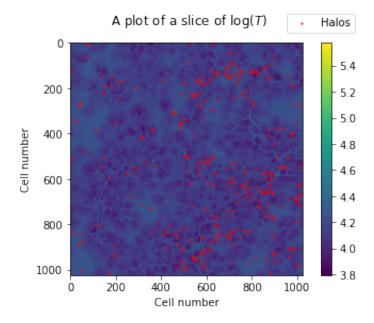
Task 6

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```
In [52]:
          x, y, z = halodata[:, 17], halodata[:, 18], halodata[:, 19]
          m = 1024*0.989391/40
          x, y, z = m*x, m*y, m*z
          x_{int} = x.astype(int)
          indices = np.where(x int ==0)
          y_int = y[indices]
          z int = z[indices]
          fig1, ax1 = plt.subplots()
          im1 = ax1.imshow(np.log10(xHI[0]))
          ax1.set(xlabel = "Cell number", ylabel = "Cell number")
          ax1.scatter(y_int, z_int, color='red', s=1, label= 'Halos')
          ax1.legend(loc = 'upper right', bbox to anchor = (1.25, 1.15))
          fig1.colorbar(im1)
          fig1.suptitle("A plot of a slice of $\log(xH I)$")
          fig2, ax2 = plt.subplots()
          im2 = ax2.imshow(np.log10(delta[0]))
          ax2.scatter(y int, z int, color='red', s=1, label= 'Halos')
          ax2.legend(loc = 'upper right', bbox_to_anchor = (1.25, 1.15))
          fig2.colorbar(im2)
          ax2.set(xlabel = "Cell number", ylabel = "Cell number")
          fig2.suptitle("A plot of a slice of $\log(\Delta)$")
          fig3, ax3 = plt.subplots()
          im3 = ax3.imshow(np.log10(T[0]))
          ax3.scatter(y int, z int, color='red', s=1, label= 'Halos')
          ax3.legend(loc = 'upper right', bbox_to_anchor = (1.25, 1.15))
          fig3.colorbar(im3)
          ax3.set(xlabel = "Cell number", ylabel = "Cell number")
          fig3.suptitle("A plot of a slice of $\log(T)$")
         /tmp/ipykernel 2004116/3369886743.py:24: RuntimeWarning: divide by zero enco
         untered in log10
           im2 = ax2.imshow(np.log10(delta[0]))
         /tmp/ipykernel 2004116/3369886743.py:32: RuntimeWarning: divide by zero enco
         untered in log10
           im3 = ax3.imshow(np.log10(T[0]))
         Text(0.5, 0.98, 'A plot of a slice of \ \log(T)$')
```

Out[52]:





Task 7

In []: