

PHIOT - Quantum physics

Classical physics:- (traits)

- 1) All phenomenon - particle or wave
- 2) momentum & position exact. & are "objective" properties. (subjective in Q-physics)
- 3) wave exhibit features; not for particles; like interference, diffract! (slide)!
- 4) Doctrine of "determinism" → exact conditions gives exact results.

Quantum p. is bizarre:- (traits)

(slides)

Book: Modern Physics by Serway, Moses and Moyer 3rd edition.

(Seniors buy & sell).

↓ {more, on slides}

* ----- *

revise Faraday's law

Ampere's law

E

all E.M. laws.

→ degrees of freedom:-

for each trans.

$$\text{trans of rigid rot.} = \frac{1}{2} K_B T \rightarrow KE = \frac{1}{2} \frac{1}{2} m v_x^2$$

for each vib. = $K_B T$

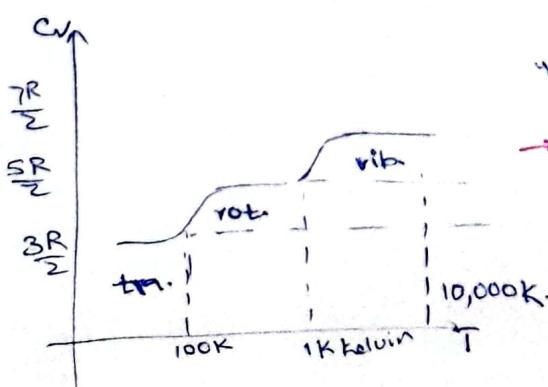
... 3 times in rotation

(Supporting material)

rigid diatomic $C_V = \frac{5R}{2}$.

non rigid diatomic $C_V = \frac{7}{2} R$.

* Specific heat of $H_2(g)$



"quantisation of energies!"

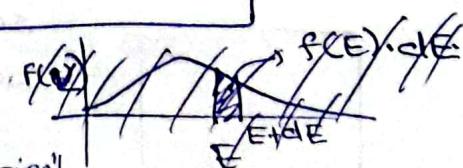
→ no 'classical' explanation.

bolzmann distribution.

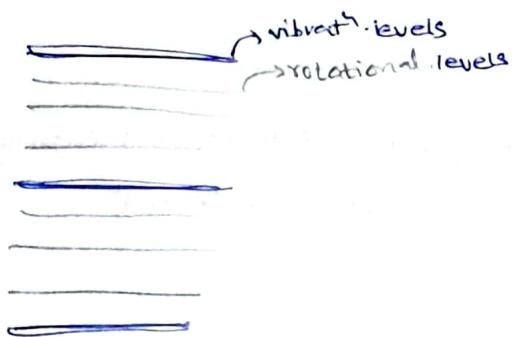
$$f(E) = A \cdot e^{-E/k_B T}$$



$$\frac{dE}{k_B T} = \frac{dE}{\frac{1}{2} k_B T} = \frac{2}{k_B T} dE$$



Quantisation of energy:- no quantisation for translational.



$$\text{rotation frequency, } \nu = \omega / 2\pi$$

$$\text{energy} = h\nu$$

* Thermal energy; $\frac{3}{2} k_B T \rightarrow \text{in } 10^{-21} \text{ J order}$

→ Triatomic linear:-

CO_2 (case)
3 trans; 2 rot; 4 vib...

so; expected



but no!... all vibration are not same energy level.

Ozone (another example)

→ Blackbody radiation:-

$$J(\nu, T) = \frac{c}{4} \cdot u(\nu, T)$$

power radiated
per unit area
per unit volume.
just, on outside
surface.

A particular vibration mode, will contribute; if $k_B T \approx h\nu_{\text{vib}}$.

$k_B T$ → used as a scale factor
for energy values in
molecular scales.
(E also as a unit of energy).

→ Wein's exponential law:-

$$u(\nu, T) = A \cdot \nu^3 \cdot e^{-K \nu / T} \Rightarrow \nu_{\text{max}} = (\alpha / T)$$

not valid; at longer λ .

Waves in a box:-



Some derivation there!

we get:

$$dN = \frac{8\pi\nu^2}{C^3} d\nu$$

$$U(v, T) = \frac{8\pi v^2}{C^3} k_B T$$

Rayleigh-Jeans law.

give good results at high wavelengths.
low

Didn't understand anything!.

2/8/19

CLASS-2:-

Kirchoff theorem

$$e(v) = J(v, \tau) \cdot A(v)$$

\downarrow absorbance.

for blackbody: $A(\nu) = 1$.

2) Stephan's law.

$$\int_{-\infty}^{\infty} f \cos \omega v dv \propto T^4$$

$$\therefore e(v) = \tau(v, T)$$

↳ universal; for all bodies

→ Energy density of black body :-

$J(\nu, T)$: power emitted per unit area; per ν interval; T - temperature

$U(v, T)$: energy density per unit frequency

$$\overline{U}(v, \tau) = \frac{c}{4} \cdot u(v, \tau)$$

some variation up more

→) Weinstaw:-

$$\frac{u(v,T)}{dv} = A v^3 e^{-Bv/T} dv$$

$$\rightarrow h \sigma \times v^2 \times \underline{e^{-\beta v_F d_S}}$$

classical! MB-

Should have been BE statistics.

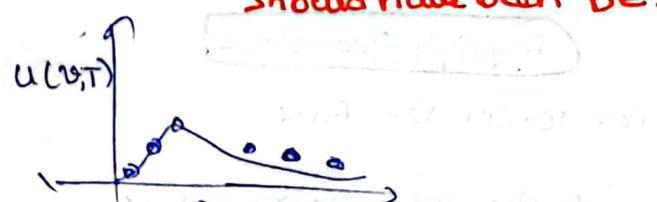
$$u(v, T) = f(\phi) \cdot g(v_H) \quad v = v_0 \dots$$

$$\frac{du(v,t)}{dv} = 0.$$

$$T_0 \dots d \langle v_i \cdot g_{K^0} \rangle e^{V_{T_0}}$$

$$e^{t_0 u_1} \cdot e^{-\frac{u_1}{T_0}} + f(u_1) - \frac{1}{T_0} \cdot e^{-\frac{u_1}{T_0}} = 0.$$

$$\frac{V}{T_0} = \text{constant}$$



Don't agree
for longer wavelength
(i.e. above 6000nm).

2nd idea:

2/8/19



Light standing in a box.



$$nx \left(\frac{\lambda_x}{2} \right) = L$$

$$\lambda_x = \frac{2L}{nx}$$

$$\therefore k_x = \frac{2\pi}{2L} \cdot nx = \boxed{\left[nx \cdot \frac{\pi}{L} \right]}.$$

$$A_x = A \sin \left(\frac{nx\pi}{L} \cdot x \right)$$

now! we need to know the no. of modes; per unit volume; in the frequency interval ν and $\nu + d\nu$.

Lamda

$$\text{no. of modes} = \frac{8\pi\nu^2}{C^3} d\nu \quad \text{per unit volume.} = g(\nu) d\nu$$

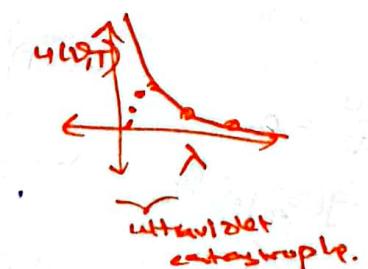
\rightarrow now; $U(\nu, T) = \text{no. of modes} \times \text{Energy of each mode.}$

$$U(\nu, T) = \frac{8\pi\nu^2}{C^3} \cdot K_T$$

(Rayleigh-Jeans law)

For lower $\nu \rightarrow$ fine

higher frequency $\rightarrow \propto$.



6/8/19

Special session:-

$$E_y = (\cos k_z z \cdot \cos k_x x) e^{i\omega t}$$

$$k_z \cdot z = \pm \pi$$

$$k_x \cdot x = \pm n\pi$$

$$\frac{10}{k} = c$$

$$\omega^2 = c^2(k_x^2 + k_z^2)$$

$$\omega^2 = \pi^2 c^2 (l^2 + m^2)$$

→ Planck hypothesis:-

- oscillators of walls of black body, emit or absorb radiation.
- oscillators of all frequencies are present.
- energy of these oscillators is quantised. (θ is not!)

Planck's equation:-

$$u(\nu, T) = \frac{8\pi h}{c^3} \frac{\nu^3 d\nu}{e^{\frac{h\nu}{kT}} - 1}$$

$$\lambda_{\text{max}}, T = 2.898 \times 10^{-3} \text{ m.k.}$$

$$\frac{dU(\lambda, T)}{d\lambda} = 0 \text{ for the result}$$

Total power:-

$$J(\lambda, T) = \int u(\lambda, T) d\lambda$$

$$\text{Integrate } \Sigma, \text{ we get } = C \cdot T^4$$

LASER:

(Fluke formula)
Weins:

$$h\nu \times \nu^2 \times \frac{e^{-h\nu/kT}}{\text{mistake!}}$$

Raileigh:

$$\frac{1}{K} \times \nu^2 \times \frac{1}{K} \rightarrow \text{mistake!}$$

Planck:

$$h\nu \times \nu^2 \times \left[\frac{1}{e^{h\nu/kT} - 1} \right] \rightarrow \text{correct!}$$

9/8/19

Planck int-energy = no. of modes \times avg. of each mode.

$$\frac{8\pi\nu^2}{c^3} d\nu \quad \neq kT.$$

but! energy is quantised.

$$E = nh\nu.$$

↓
will learn later
what is average.

as n is discrete.

this is the difference!

Quantum average :-

$$\langle E \rangle = \frac{\sum_{n=0}^{\infty} n h\nu e^{-\frac{(nh\nu)/kT}{1}}}{\sum_{n=0}^{\infty} e^{-\frac{(nh\nu)/kT}{1}}} \\ = \frac{h\nu}{e^{\frac{h\nu}{kT}} - 1}$$

Classical average :-

$$\langle E \rangle = \frac{\int_0^{\infty} E \cdot e^{-E/kT} dE}{\int_0^{\infty} e^{-E/kT} dE} \rightarrow \text{taking continuous variation.}$$

$$= kT$$

$$\therefore \text{Energy} = \frac{8\pi\nu^2}{c^3} \cdot \frac{h\nu}{e^{\frac{h\nu}{kT}} - 1} = \frac{8\pi\nu^2}{c^3} \cdot \frac{1}{\left(e^{\frac{h\nu}{kT}} - 1\right)}$$

if $h\nu \gg kT$
wien's law

planck's formula.
↓
if $h\nu \ll kT$
rayleigh-jean's law.

Now; trying to get wein's displacement law:-

$$U(\nu, T) = \frac{8}{\lambda^5} \cdot \frac{1}{(e^{\alpha/\lambda T} - 1)} \quad (\text{use } \alpha = \frac{C}{\lambda})$$

$$\frac{dU(\nu, T)}{d\lambda} = 0.$$

* we get $(\lambda \cdot T = 2.898 \times 10^{-3} \text{ Km})$
(after solving) Year!

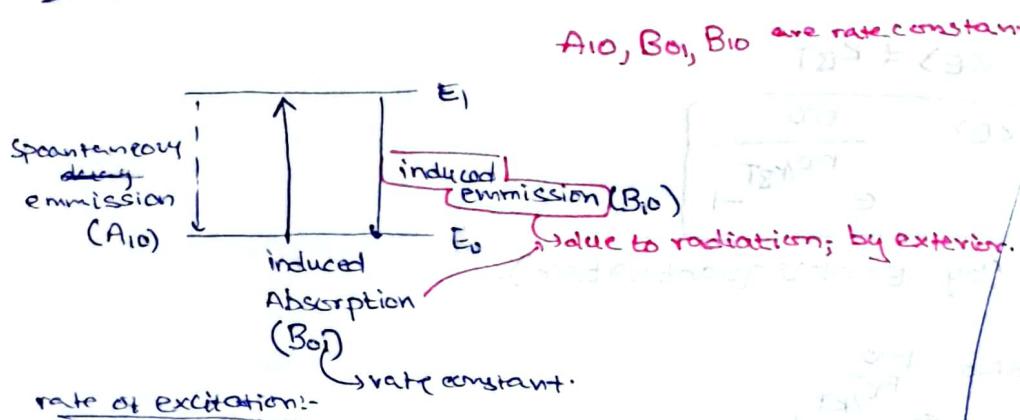
try to get stephen's law:-

$$J(\nu, T) = \frac{f}{4} \cdot U(\nu, T)$$

$$\text{Net radiation} = \int_0^{\infty} J(\nu, T) \cdot d\nu = \underbrace{\left(2\pi^3 \dots\right)}_{\text{constant}} T^4$$

Einstein's approach to Planck's law:-

in atoms:-



Rate of emission:-

$$R_{10} = N_1 (B_{10} \cdot U(0, T) + A_{10})$$

∴ equilibrium:-

$$R_{10} = R_{01}$$

$$\therefore \frac{N_0}{N_1} = \frac{A_{10} + B_{10} \cdot U(0, T)}{B_{01} \cdot U(0, T)} = \left(\frac{g_0}{g_1} \right) e^{\left(\frac{E_1 - E_0}{kT} \right)}$$

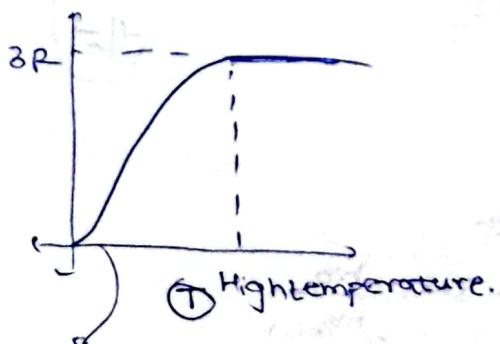
degeneracies

ii → Specific heat of solids:-

1) Dulong-Petit's law:-

$$TC_V = 3R$$

but experimentally-



$$C_V = \alpha \cdot T^3$$

$$C_V \rightarrow 0 \text{ as } T \rightarrow 0.$$

BLATTI,
BLATTI,
BLATTI

Einstein's model:-

At low temperatures, this formulae falls off rapidly as compared to expt. data.

$$U = 3Na \times \langle E \rangle.$$

now! $\langle E \rangle \neq k_B T$.

$$\langle E \rangle = \frac{h\nu}{e^{\frac{h\nu}{k_B T}} - 1}$$

(by $E = nh\nu$ quantisation).

$$\therefore U = 3Na \cdot \frac{h\nu}{e^{\frac{h\nu}{k_B T}} - 1}$$

$$\therefore C_V = \frac{\partial U}{\partial T} = 3$$



explained!
using quantisation.

finally!

$$C_V = 3R \cdot \frac{x^2 e^x}{(e^x - 1)^2}; x = \frac{h\nu}{k_B T} = \frac{T_E}{T}$$

$$T_E = \frac{h\nu}{E_F} = \text{einstein temperature}$$

for every metal!....

enough; if we know T_E ...

$$* T_E = \frac{h\nu}{k_B}$$

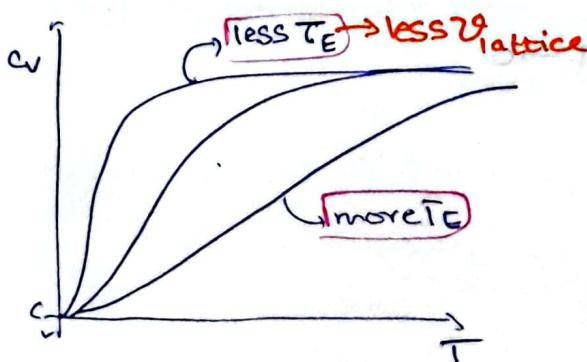
const. for a given metal.

obtained experimentally.

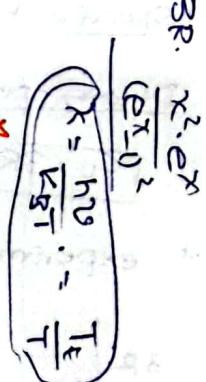
lattice frequencies... einstein; took same \rightarrow for all molecules

Quanta of lattice vibrations are called phonons.

* more T_E ; less C_V at αT



1836



• frank-herz experiment

$$U = \text{Energy} = 3Na \cdot \frac{h\nu}{e^{\frac{h\nu}{k_B T}} - 1}$$

$$C_V = \left(\frac{\partial U}{\partial T}\right) = 3Na \cdot h\nu \cdot \frac{1}{e^{\frac{h\nu}{k_B T}}} \cdot \frac{h\nu}{k_B T} \cdot \frac{1}{T^2}$$

$$\left(3Na \cdot K \cdot \left(\frac{h\nu}{k_B T}\right)^2 \cdot \frac{e^{\frac{h\nu}{k_B T}}}{(e^{\frac{h\nu}{k_B T}} - 1)^2} \right)$$

X ————— X

photoelectric effect:-

Compton:-

particle nature of light!

so; does photon has momentum.

Bit of special Relativity:-

Intrinsic rest mass = m_0 .

-rest mass energy = $m_0 c^2$.

$$m = \frac{m_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} = \gamma m_0.$$

* momentum = $\gamma m_0 v$

$$\therefore \text{Net energy} = \sqrt{p^2 c^2 + m_0^2 c^4} = \underline{\underline{\gamma m_0 c^2}}$$

Compton scattering:-

→ Demonstrates momentum conservation in scattered photons.

Classical: (Thompson scattering),

same λ ; initially & finally. (assume $v \ll c$)

→ Classical solution

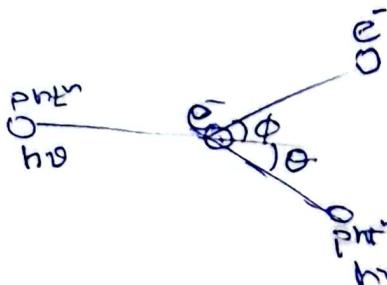
$$\frac{p_f}{p_i} = \frac{\sin \theta}{\sin (\theta - \phi)}$$

Compton:-

$$\text{for light } E^2 = p^2 c^2 + m_0^2 c^4$$

or photon

$$\therefore E = pc = h\nu$$



PCons. along y:-

$$p_e \sin \phi = \frac{h\nu'}{c} \cdot \sin \phi$$

PCons. along z:-

$$\frac{h\nu'}{c} = p_e \cos \phi + \frac{h\nu}{c} \cdot \cos \phi$$

For e^- : $E^2 = p^2 c^2 + m_0^2 c^4$

Energy eqn:

$$\frac{h\nu}{c} = E_e + h\nu$$

finally

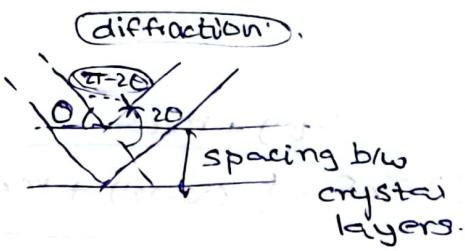
$$\lambda - \lambda_c = \lambda_c \cdot (1 - \cos \phi)$$

$\lambda_c = \frac{h}{m_0 c}$

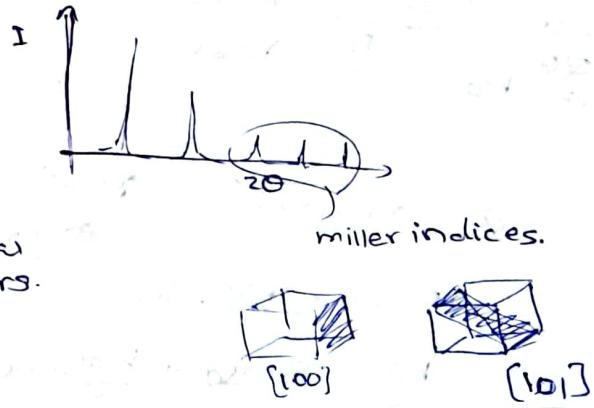
$$2.43 \cdot 10^{-12} \text{ m}$$

so; to notice change; λ must be small.
∴ high energy x-rays $\sim 10^{-12} \text{ m}$.

* X-ray diffraction:-



RANDOM:-



* Photon duality:-

refraction:- can be explained; using only wave nature.

* photon can behave like a particle & a wave; depending upon experiment.

* Single slit electron diffraction:-

$\Delta P_y = Psin\theta$

for first minimum;

$sin\theta = \frac{\lambda}{d}$

$\Delta y = n\lambda / sin\theta$

$\therefore \Delta y \cdot \Delta P_y = \lambda P = \frac{h}{d} \cdot P = h$

$\therefore [\Delta y \cdot \Delta P_y = h]$

Microscope-

Resolution \propto wavelength of light used.

So we can use the e-beam; as we can change its wavelength.

Read the text book.

$\frac{d\theta}{d\lambda} = \text{resolution}$

electron gun

(horizontal) θ

(transverse) \rightarrow (vertical)

Group velocity & phase velocity:-

~~(from - to) wave (from - to) particle = wave~~

Electron gun \rightarrow Britton's source

If a wave is given; how to give particle?

wave packet-

- spatial beats by superposition of sinusoidal waves of nearby wavelengths

see slide images

sinusoidal waves

2 waves:



3 nearby wavelengths:



5:



many!

nearby

wavelengths.



Wave packet.
(refers to particle position)

group velocity \rightarrow $v_g = \frac{\lambda}{T}$

$$\Delta\lambda = 0.5$$

$$\frac{\Delta\lambda}{\lambda} = 0.5$$

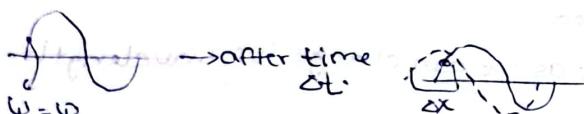
$$v_g = \frac{\lambda}{T}$$

$$v_g = \frac{\lambda}{T}$$

group velocity \rightarrow $v_g = \frac{\lambda}{T}$

group velocity

Phase velocity, group velocity:-


 $\psi = \psi_0$ → after time Δt .
 phase position moves.

$$\therefore \text{phase velocity} = \frac{\Delta \omega}{\Delta k}.$$

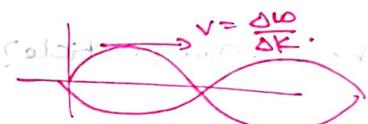
Group velocity:-

is the velocity with which the envelope moves.

$$\left. \begin{array}{l} \text{Eq: } A \sin(kx - \omega t) \\ A \sin((k + \Delta k)x - (\omega + \Delta \omega)t) \end{array} \right\}$$

$$\text{Sum.} = 2A \sin\left(\frac{kx - \omega t}{2}\right) \cos\left(\frac{\Delta k x - \Delta \omega t}{2}\right)$$

fast oscillating; slow oscillating



$$\therefore \text{Group velocity} = \frac{\Delta \omega}{\Delta k} = \frac{dk \omega}{dk} = v_{\text{phase}} \text{ (here).} = v_g$$

↳ velocity of wave packet.

$$\begin{aligned} \frac{\omega}{k} &= c \\ \therefore \omega &= kc. \\ \frac{d\omega}{dk} &= c. \end{aligned}$$

$$\text{Eq: } v \propto \sqrt{\lambda} \text{ for a water ripple.}$$

$$\lambda \omega = c \sqrt{\lambda}$$

$$v = \frac{c}{\sqrt{\lambda}}.$$

$$\omega = \frac{2\pi c}{\sqrt{\lambda}}.$$

$$k = \frac{2\pi}{\lambda},$$

$$\therefore \frac{d\omega}{dk} = \frac{1}{\lambda \sqrt{\lambda}} \cdot \lambda^2 = \frac{1}{\lambda} \cdot k \cdot \omega = v_{\text{group}}.$$

$$\boxed{v_g = \frac{1}{2} v_p}$$

relation b/w ω , k is dispersion relationship.

e.g: (i) $\frac{\omega}{k} = c$ for light in vacuum.

(ii) $\omega = A \sin(\frac{ka}{2})$; for 1D-atom chain.

Particle to wave:-

$$\lambda = \frac{h}{P} = \frac{h}{mv}$$

$$\text{General true... } \boxed{\nu = \frac{E}{h}} = \frac{mc^2}{h}$$

$$\text{speed } (v_p) = \lambda\nu = \frac{c^2}{\nu} \\ \approx v_p \neq v.$$

particle $\rightarrow v$

$$\omega = 2\pi\nu = 2\pi \frac{mc^2}{h} = \frac{2\pi mc^2}{h \cdot \sqrt{1 - \frac{v^2}{c^2}}} \\ K = \frac{2\pi}{\lambda} = \frac{2\pi m v}{h \sqrt{1 - \frac{v^2}{c^2}}} = \frac{2\pi m v}{h \cdot \sqrt{1 - \frac{v^2}{c^2}}}.$$

physics of dispersive wave.

$$V_g = \frac{d\nu}{dk} = \frac{d\omega/dv}{dr/dk} = V,$$

(after differentiation)

$$V_g \neq v$$

→ Relationship b/w v_g & v_p :

$$v_p = \frac{\omega}{k}.$$

$$V_g = \frac{d\omega}{dk} = \frac{d(c v_p \cdot k)}{dk} = \left[v_p + k \cdot \frac{dv_p}{dk} \right] = \left(v_p - \lambda \cdot \frac{dv_p}{d\lambda} \right)$$

$$\boxed{V_g = v_p + k \cdot \frac{dv_p}{dk}} = v_p - \lambda \cdot \frac{dv_p}{d\lambda}$$

Non-dispersive medium:-

for which $\frac{dv_p}{dk} = 0$.

$$\therefore \boxed{V_g = v_p}$$

→ Dispersive medium:-

$$V_g \neq v_p.$$

Normal dispersion: $V_g < v_p$ \Rightarrow $\frac{1}{V_g} > \frac{1}{v_p}$

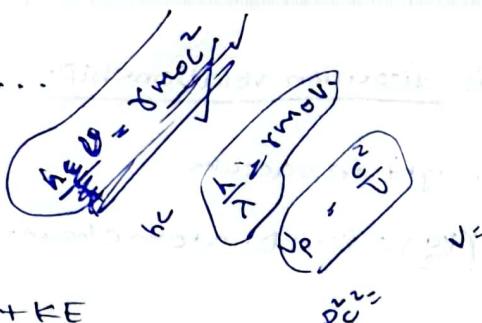
Anomalous dispersion: $V_g > v_p$ \Rightarrow $\frac{1}{V_g} < \frac{1}{v_p}$

$$\frac{1}{V_g} - \frac{1}{v_p} = q$$

Dispersion relations \Rightarrow $\omega = 2\pi f$ \Rightarrow $\omega = 2\pi c / \lambda$

* electron has $\lambda_{\text{deB.}} = 2 \times 10^{-12} \text{ m.}$

Find KE, phase vel., $v_g.$



$$\text{So, } E_0 = mc^2$$

$$E = \sqrt{(pc)^2 + (mc^2)^2} = E_0 + KE_{\text{rest}}$$

$$\therefore KE = \sqrt{E_0^2 + p^2 c^2 - E_0}$$

find $v_g:$

$$1 - \frac{v^2}{c^2} = \frac{E_0^2}{E^2}$$

$$\therefore v = c \sqrt{1 - \left(\frac{E_0}{E}\right)^2}$$

$$v = v_g \sqrt{\frac{E}{E_0}}$$

$$v_g = \frac{c}{\sqrt{1 - \left(\frac{E_0}{E}\right)^2}}$$

$$v_p = 1.3c$$

Dispersion relation for de Broglie waves:-

Phase velocity:

$$v_p = \frac{c\omega}{E} = \frac{p}{E}$$

$$h\omega = E$$

$$\therefore v_p = \frac{p}{E}$$

→ Heisenberg's uncertainty relation:-

uncertainty in wavelength of component waves.

$$\Delta x \cdot \Delta p_x \geq \frac{\hbar}{2} \rightarrow \text{hence; we can measure } x, p_y, p_z \text{ precisely, simultaneously.}$$

$$\Delta E \cdot \Delta t \geq \frac{\hbar}{2}$$

means:-

if a system is known to exist in energy E ; over a limited period Δt , then this energy is uncertain by at least an amount $\frac{\hbar}{4\pi\Delta t}$.

→ to know energy of an excited state with exact precision;

it has to stay in that state as $\Delta t \rightarrow 0$

* can e^- be found in nucleus?... cause; we have β -decay $n \rightarrow p$... physicists doubted.

* estimating atomic size; using uncertainty.

$$\Delta x = \frac{\text{radius}}{2}$$

$$P = \Delta P = \frac{\hbar}{2\Delta x}$$

$$KE = \sqrt{(pc)^2 + (mc^2)^2} - mc^2 \approx 19.75 \text{ MeV.}$$

experimental results show that no e^- particle in atom passes energy greater than a meV.

β -decay $e^- \rightarrow 1 \text{ meV}$

$\xrightarrow{\text{Potential Energy}}$

classical world

quantum world.

- 1) The observer is objective & passive \wedge the observer is not objective & passive.
- 2) events happen; independent of whether observer is present or not. 2) the act of observation changes the physical system irreversibly.
- 3) this is known as Objective reality. 3) this is called subjective reality.

Fourier transformations:-

any $f(x) = \sum a_n \sin nx + b_n \cos nx.$

whenever ' h ' is comparable to $\{($ momentum scale \times length dimension in experiment $)\}$ then wave nature is prominent.

if $f(x) = \sum a_n \cos nx.$ (take some n')

$$\int_{-\pi}^{\pi} f(x) \cdot \cos n' x dx = \sum a_n \int_{-\pi}^{\pi} \cos nx \cos n' x dx.$$

$$= \sum a_n \frac{1}{2} \int_{-\pi}^{\pi} [\cos(n-n')x + \cos(n+n')x] dx$$

$\cancel{\int_{-\pi}^{\pi} \cos(n-n')x dx}$ its $\int_{-\pi}^{\pi} \cos(n+n')x dx$ is always zero

$$= \frac{1}{2} \cdot a_{n'} \cdot 2\pi$$

$$= \pi a_{n'}$$

$$n = n'$$

$$\therefore a_{n'} = \frac{1}{\pi} \cdot \int_{-\pi}^{\pi} f(x) \cdot \cos n' x dx$$

in general: $f(t) = \sum a_n \sin nt + \sum b_n \cos nt.$

$$f(t) = \int_{-\infty}^{\infty} g(\omega) e^{i\omega t} d\omega$$

Here $g(\omega) = \int_{-\infty}^{\infty} f(t) e^{-i\omega t} dt.$

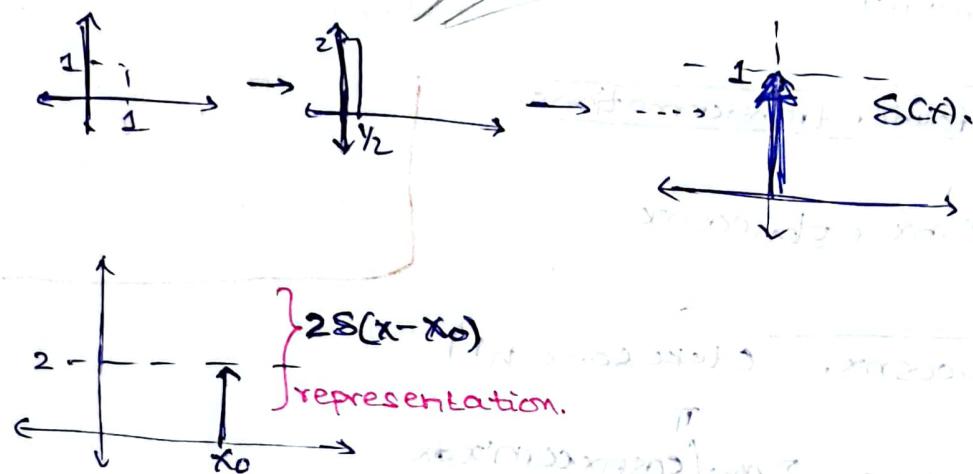
S function:-

$$\delta(x) = 0 \text{ for } x \neq 0 \\ = \infty \text{ for } x = 0.$$

$$\delta(x-x_0) = \int_{-\infty}^{\infty} e^{i(t-x_0)} dt$$

$$\int_{-\infty}^{\infty} \delta(x) dx = \int_{-\infty}^{\infty} \delta(x) dx = 1.$$

such that



* δ function makes sense; when it occurs as an integral.

$$\int_{-\infty}^{\infty} f(x) \cdot \delta(x-x_0) dx = \int_{x_0-\epsilon}^{x_0+\epsilon} f(x) \delta(x-x_0) dx \\ = \underline{f(x_0)}$$

examples

-2

1) $f(t) = \alpha$.

$$g(\omega) = \int_{-\infty}^{\infty} f(t) e^{-i\omega t} dt$$

$$= \alpha \cdot \int_{-\infty}^{\infty} e^{-i\omega t} dt$$

$$= \alpha \cdot \delta(\omega)$$

2) $f(t) = \cos \omega_0 t$.

$$f(t) = \frac{1}{2} (e^{i\omega_0 t} + e^{-i\omega_0 t})$$

$$g(\omega) = \int_{-\infty}^{\infty} \frac{1}{2} (e^{i\omega_0 t} + e^{-i\omega_0 t}) e^{-i\omega t} dt$$

$$g(\omega) = \int_{-\infty}^{\infty} f(t) e^{-i\omega t} dt$$

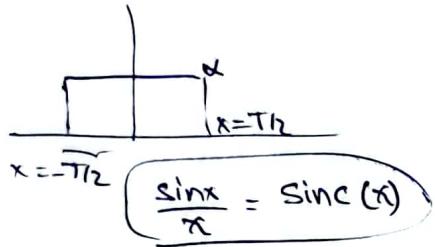
$$\frac{1}{2} (e^{i\omega_0 t} + e^{-i\omega_0 t}) e^{-i\omega t}$$

$$(e^{it(\omega_0 - \omega)}) + \frac{1}{2} + e^{-it(\omega_0 + \omega)}$$

$$g(\omega) = \frac{1}{2} \delta(\omega - \omega_0) + \frac{1}{2} \delta(\omega + \omega_0)$$

$$f(t) = \frac{1}{2} \cos \omega_0 t$$

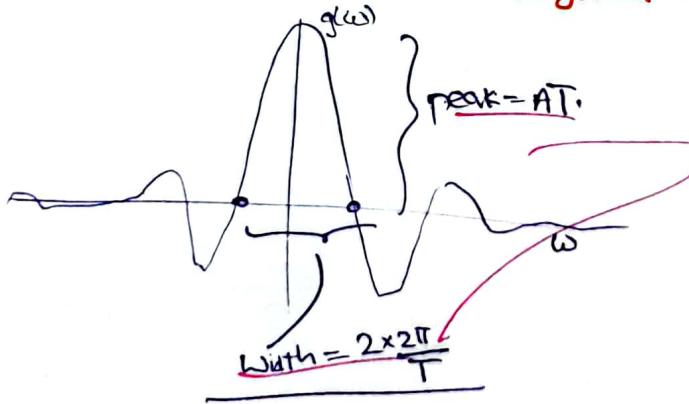
3)



$$g(\omega) = AT \cdot \text{sinc}\left(\frac{\omega T}{2}\right)$$

here:

Used widely in digital communications.



both are inversely related.
hint at $\omega \propto 2\pi / T$.

4

$$(\cos \frac{\omega}{2}) \cos \omega$$

Tut 5:-

- electron interference.

(YDSE)

- Davisson Germer.

Q2

$$\frac{1}{c^2} \cdot \frac{\partial^2 u}{\partial t^2} = \frac{\partial^2 u}{\partial x^2}$$

$e^{i(kx-\omega t)}$ is sorn.

$$\frac{1}{c^2} \cdot \frac{\partial^2 u}{\partial t^2} = \Sigma \frac{\partial^2 u}{\partial x^2}$$

$\hookrightarrow e^{i(\vec{k} \cdot \vec{r} - \omega t)} = 0$ is sorn.

$\sin(kx - \omega t)$

$\sin(k(x+y) + \omega t)$

$$\boxed{\cos^2 \left(\frac{k(x+y)}{2} \right)}$$

$$\therefore \vec{k}_1 = k_0(\hat{x} + \hat{y} + \hat{z}), k^2 = k_0(\Sigma)$$

$$\begin{aligned} y_1 &= A \cdot e^{i(\vec{k}_1 \cdot \vec{r} - \omega t)} \\ y_2 &= A \cdot e^{i(\vec{k}_2 \cdot \vec{r} - \omega t)} \end{aligned}$$

$$y = A \cdot e^{i(k_0 z - \omega t)} \cdot [1 + e^{ik_0(x+y)}]$$

$$\text{Intensity} = y y^* = y_1 y_2^* \quad \text{is max at } x=y=0$$

$$\propto \cos^2 \left(\frac{k}{2}(x+y) \right).$$

$$I = |y_1|^2 + |y_2|^2 = |y_1 + y_2|^2$$



Gaussian function:-

$$* \boxed{f(x) = A \cdot e^{-\frac{(x-\mu)^2}{2\sigma^2}}}$$

Normalization,

(imp. in Q.phy.)

i.e.

$$\int_{-\infty}^{\infty} f(x) dx = 1$$

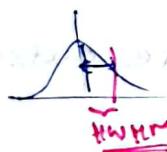
$$\int_{-\infty}^{\infty} e^{-\alpha x^2} = \left(\frac{\pi}{\alpha}\right)^{1/2}$$

$$\therefore \text{now } A \cdot \int_{-\infty}^{\infty} e^{-\frac{(x-\mu)^2}{2\sigma^2}} = 1$$

$$A = \frac{1}{\sqrt{2\pi}\sigma}$$

$$f(x) = \frac{1}{\sqrt{2\pi}\sigma} \cdot e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$

HWHM half width half maxima.



$$\text{HWHM} = \sigma \cdot \sqrt{2 \ln 2}$$

\therefore width $\propto \sigma$

$$* f(x) = A \cdot e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$

Simpler

$$\therefore \text{standard deviation} = \frac{\sigma}{\sqrt{2}}$$

F.T (Fourier transform) of Gaussian

$$a(k) = \int_{-\infty}^{\infty} e^{-\frac{(x-\mu)^2}{2\sigma^2}} \cdot e^{-ikx} dx$$

(I doubt in proof of this)

$$* \boxed{a(k) = \sqrt{2\pi} \sigma_x \cdot e^{-\frac{(k\sigma_x)^2}{2}}}$$

$$\therefore \sigma_k = \frac{1}{\sigma_x} \quad \therefore \boxed{\sigma_k \cdot \sigma_x = 1}$$

the origin of uncertainty relation.

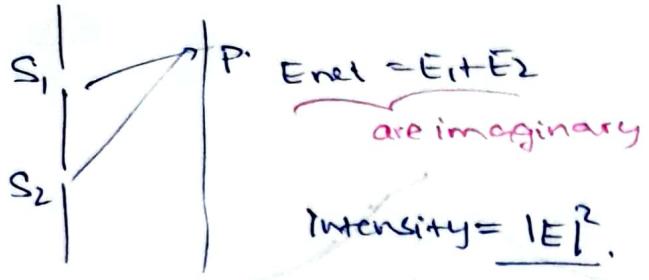
* for Gaussian func:

$$\langle x \rangle = \mu$$

$$\langle x^2 \rangle = \mu^2 + \sigma^2$$

$$\therefore \text{variance } (\sigma^2) = \langle x^2 \rangle - \langle x \rangle^2 = \sigma^2$$

satisfied!



$$\text{Intensity} = |E|^2 = |E_1|^2 + |E_2|^2$$

$$= |E_1|^2 + |E_2|^2 + 2\text{Re}(E_1^* E_2)$$

Similarly:- for ψ waves.

$$\Psi = \Psi_1 + \Psi_2 \quad (\text{for matter waves}).$$

$$\therefore P = \Psi \cdot \Psi^* = |\Psi|^2 = \text{probability}$$

* wave fun: $\Psi(x, t)$

$$\text{Average: } \langle O \rangle = \int_{-\infty}^{\infty} \Psi^*(x, t) \cdot O \cdot \Psi(x, t) dx.$$

Superposition:

$$\Psi = C_1 \Psi_1 + C_2 \Psi_2$$

→ Wave packets:-

to form a true wave packet that is zero everywhere outside Δx ;

we need to add infinite harmonic waves;

with continuous variation
wavelengths (k).
amplitudes, $a(k)$.

* now; if $\Psi(t)$ is given; we can find $a(k)$ position momentum.

$$\hbar k = p$$

Gaussian wave packet

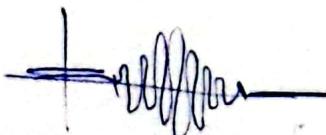
$$\Psi(x) = \int_{-\infty}^{\infty} a(k) e^{-ikx} dk.$$

$$a(k) = A e^{-\frac{(k-k_0)^2}{2\sigma_k^2}} \rightarrow \text{this } k_0 \text{ will decide } x_0??....$$

We know; if $a(k)$ is gaussian; $\Psi(x)$ is gaussian.

$\Psi(x)$ will be

$$\sigma_x = \frac{1}{\sigma_k}$$



$$\Psi(x) = \underbrace{\text{const.}}_{\text{oscillatory}} \underbrace{e^{ik_0 x}}_{\text{envelope}} e^{-\frac{\sigma_k^2 x^2}{2}}$$

$$\therefore \Delta x \cdot \Delta k =$$

$$\Delta x \cdot \Delta p = \hbar.$$

the uncertainty is minimum for Gaussian wave packet.

$$\therefore \text{in general; } \boxed{\Delta x \cdot \Delta p \geq \hbar}$$

this is the uncertainty relation for wave packet.

* normalize $\psi(x) = C e^{ikx} e^{-\frac{\sigma^2 k^2}{2}}$.

means

$$\int_{-\infty}^{\infty} |\psi|^2 dx = 1. \text{ not } \int_{-\infty}^{\infty} \psi^* \psi dx = 1$$

we get.

$$C = \frac{\sqrt{\sigma k}}{\sqrt{4\pi}}$$

imaginary coming.

for contributions in Q.mechanics; we take $|\psi|^2$;

not $\psi^* \psi$.

so; we have a ' $\sqrt{2}$ ' factor.

* In quantum physics; average is

$$\langle x \rangle = \frac{\int_{-\infty}^{\infty} \psi^*(x) \cdot x \cdot \psi(x) dx}{\int_{-\infty}^{\infty} \psi^*(x) \cdot \psi(x) dx}$$

hence; we get $\frac{1}{\sqrt{2}}$ factors in

standard deviations.

Basics:-

- wave function
- normalisation
- Operators (\hat{O}): turns functions into functions.

Eg: derivative operator: $\hat{O} = \frac{d}{dx}$

* $\hat{A}\psi = \alpha\psi$ α is scalar.

α is called eigen value.

$\psi(x)$ is eigen function of A .

Eg. in slides

doubt:

if $\hat{A} = x$.

$\hat{A}(e^x) = e^x$ or $x e^x$? ✓

$$[A, B] = AB - BA, \rightarrow \text{commutator.}$$

if $[A, B] = 0$; operators are commute.

Expectation value:-

$$\langle \hat{O} \rangle = \int_{-\infty}^{\infty} \psi^*(x, t) \hat{O} \psi(x, t) dx \quad \text{quantum mechanical contribution}$$

$\int_{-\infty}^{\infty} \psi^* \psi dx = 1$ here.

$$\text{de Broglie wave. } \psi(x,t) = A \cdot e^{-i(Kx - \omega t)}$$

we will write; for any general; as we have Fourier with us

$$1) \frac{\partial}{\partial t} \psi(x,t) = A \cdot i\omega \cdot e^{i(Kx - \omega t)}$$

$$\omega = \frac{E}{\hbar}$$

$$\frac{\partial}{\partial t} \psi(x,t) = -\frac{A \cdot E}{\hbar} \cdot \frac{\partial}{\partial x} \psi(x,t)$$

$$2) \frac{\partial}{\partial x} \psi(x,t) = +iK \cdot \psi(x,t).$$

$$K = \frac{P_x}{\hbar}$$

$$i\hbar \frac{\partial}{\partial x} \psi(x,t) = P_x \psi(x,t).$$

$$i\hbar \frac{d}{dt} \psi(x,t) = E \cdot \psi(x,t)$$

$$\hat{E} = +i\hbar \frac{d}{dt}$$

$$\therefore \left[\hat{P}_x = -i\hbar \frac{\partial}{\partial x} \right]$$

only component!

I) Kinetic energy :- (consider a non-relativistic particle)

$K = \frac{P^2}{2m}$ } how much true? \rightarrow true in non-relativistic domain.

$$\text{Ansatz: } K \approx \frac{1}{2m} \left(\frac{\partial \psi}{\partial x} \right)^2$$

$$\hat{K} = -\frac{\hbar^2}{2m} \cdot \frac{\partial^2}{\partial x^2}$$

funny!

Facts:-

- 1) for every physical quantity (**observable**); there is an operator.
- 2) Heisenberg uncertainty involves non-commuting operators; these cannot be simultaneously measured precisely.
 \hat{x}, \hat{p} are non commuting.

$$x \frac{\partial}{\partial x} f(x) \neq \frac{\partial}{\partial x} (xf(x))$$

→ checking commutability:-

$$1) \quad \hat{x}\hat{p}_x - \hat{p}_x\hat{x}:$$

$$= x \cdot -i\hbar \frac{d}{dx} (\psi(x)) + i\hbar \frac{d}{dx} (x \cdot \psi(x))$$

$$= i\hbar \cdot \psi(x) - [x, p_x] = i\hbar \neq 0$$

∴ noncommutable.

2) if $[\hat{x}, \hat{p}]$ is done;

$$[\hat{x}, \hat{p}] = 0.$$

* let A, B, C be operators.

$$[A, B] = C.$$

then commutation

general uncertainty relation

$$\Delta A \cdot \Delta B \geq \frac{1}{2} | \langle C \rangle |$$

Expectation.

$$\Delta x \cdot \Delta p_x \geq \frac{1}{2} | \langle i\hbar \rangle | = \frac{\hbar}{2}$$

tut-6: tough guy!!

Uncertainty & Fourier:-

P8) $\Delta x = 10^{-14} \text{ m}$

$\Delta p_x \approx 10^{-20} \text{ Ns}$

$$E = \sqrt{(pc)^2 + (mc^2)^2}$$

$= 19.6 \text{ MeV}$. High!

for a particle in nucleus; $E_{\max} \approx 4 \text{ MeV}$.

for proton; we also have pot. energy.

b)

ground state energy:-

$$\left(\Delta x \approx \frac{p}{2m} \right) + \omega \frac{1}{2} k x^2 \quad \text{(probabilistic, minimum of } H_0 = \frac{p^2}{2m} + \frac{1}{2} kx^2 \text{)}$$

both ≈ 0 .

$$\Delta p_x \cdot \Delta x = \frac{\hbar}{2}$$

$$\text{put } \Delta p_x = \frac{\hbar}{2(\Delta x)}.$$

i.

$$\left(\Delta x \approx \frac{p}{2m} \right) + \omega \frac{1}{2} k x^2 = \text{const.}$$

and it is a parabola. $\omega = \sqrt{\frac{k}{m}}$ (constant)

so it is simple harmonic motion. $\omega = \sqrt{\frac{k}{m}} = \text{const.}$

so it is simple harmonic motion. $\omega = \sqrt{\frac{k}{m}} = \text{const.}$

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Non commuting Operators & uncertainty

$$[A, B] = C.$$

then $\Delta A \cdot \Delta B \geq \frac{1}{2} |C|$

$x P_x$

we get

$$[x P_x] = \frac{i\hbar}{2}$$

another operator.

$\therefore \Delta x \cdot \Delta P_x \geq \frac{\hbar}{2}$

K.P:

commuting.

\therefore no uncertainty restriction.

- can measure both accurately.

* $\psi(x) = e^{ix}$ is continuous, (how to say?) draw $\text{Re}(\psi(x))$.

Finding $\psi(x)$:-

$$\hat{E}\psi = \hat{P}\psi + V(x)\psi$$

we will have $\psi(x, t)$;

but we showed $\psi(x)$ also same eqn.

$$\psi(x, t) = \psi(x) \cdot \phi(t).$$

$$\boxed{\frac{d}{dt}(\psi(x)) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi(x) + V(x) \psi(x).}$$

$$(E)\psi(x) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} (\psi(x)) + V(x) \psi(x)$$

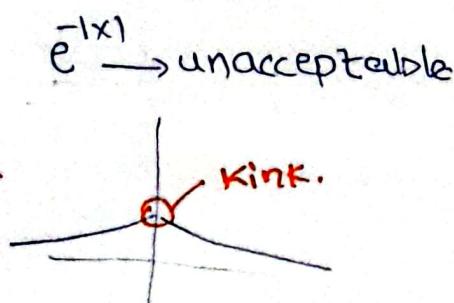
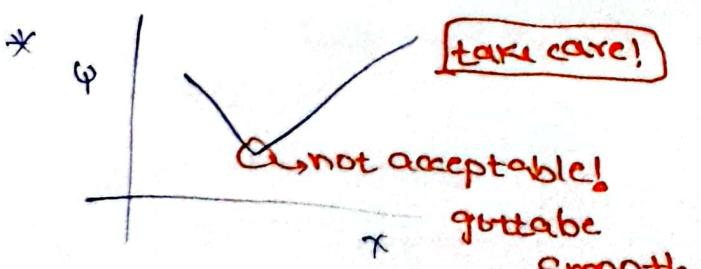
} solve this to get $\psi(x) \rightarrow \infty$ or 0
 put boundary cond. to
 eigen value. get limited solution.

now; Hamiltonian (H) operator:

$$\hat{H} = \hat{P}^2 + V(x) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x)$$

Boundary conditions :-

- 1) finite, single valued, continuous.
- 2) differentiable (so as to apply P_x).



* $e^x \rightarrow$ not acceptable.

* $e^{-ix} \rightarrow$ acceptable...

* Hermitian operator:

an operator \hat{O} is hermitian if

$$\int_{-\infty}^{\infty} \psi^* \hat{O} \psi dx = \int_{-\infty}^{\infty} (\hat{O} \psi)^* \psi dx$$

Properties:

1) \hat{O} expectation value is real.

$$\langle \hat{O} \rangle = \int_{-\infty}^{\infty} \psi^* \hat{O} \psi dx. \quad (\text{proof: show } \langle \hat{O} \rangle = \langle \hat{O} \rangle^*)$$

2) eigenvalues of \hat{O} is real.

(proof: $\hat{O}\psi = \alpha \cdot \psi$)

$$\int_{-\infty}^{\infty} \psi^* \hat{O} \psi dx = \int_{-\infty}^{\infty} \alpha \cdot \psi \psi^* dx.$$

= α (eigenvalue is a

number, not a function)

$$\int_{-\infty}^{\infty} (\hat{O} \psi)^* \psi dx = \int_{-\infty}^{\infty} \psi^* \hat{O} \psi dx.$$

$$= \alpha^*$$

(eigen value is a number; not a function)

* $\hat{x}, \hat{p}_x, \hat{k}, \hat{E}$ are hermitian operators.

Hence, their eigen values (observables)

are real quantities.

(What is eigen value? $\rightarrow S(x-x_0)$)

Solving ψ :

1) free particle:-

no force is there

$V = V_0$ (constant).

$$\therefore -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V_0 \psi = E \psi.$$

$$\therefore \ddot{\psi} = -\frac{2m(E-V_0)}{\hbar^2} \psi$$

$$\therefore \text{let } \psi(x) = A e^{ikx}$$

$$\therefore e^{ikx} \cdot k^2 = -\frac{2m(E-V_0)}{\hbar^2} e^{ikx}$$

$$k^2 = \frac{2m(E-V_0)}{\hbar^2}$$

we will get

$$p_x = K \hbar = \sqrt{\frac{2m(E-V_0)}{\hbar^2}}$$

obvi!

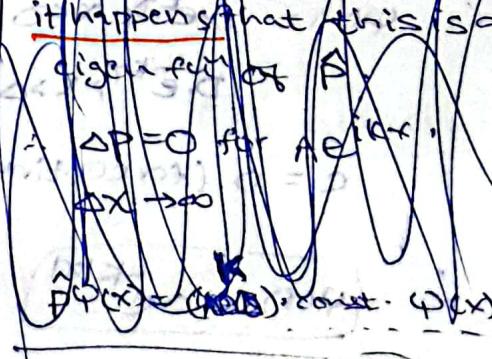
$E > V_0$:

$$\lambda = \pm ik$$

$$\therefore \psi(x) = A e^{ikx} + B e^{-ikx} \quad \text{most general.}$$

$$\text{as } \pm k = \sqrt{\frac{2m(E-V_0)}{\hbar^2}}$$

as $\pm k = \sqrt{\frac{2m(E-V_0)}{\hbar^2}}$



here $\Delta R_x = 0$

$$\therefore \Delta X = \infty$$

- the $\psi(x) = A \cdot e^{ikx}$ cannot be normalized.

time dependent!

$$\psi(x,t) = \psi(x) \cdot \phi(t)$$
$$\downarrow e^{-i\omega t} = e^{-\frac{iE}{\hbar}t}$$

$$\therefore \psi(x,t) = A \cdot e^{i(kx - \omega t)} + B \cdot e^{-i(kx + \omega t)}$$

wave moves \rightarrow wave \leftarrow

$$\langle \hat{P}_x \rangle = \hbar k \cdot \frac{|A|^2 - |B|^2}{|A|^2 + |B|^2}$$

case-ii: $E < V_0$ (tunneling effect). (we read; α -particles don't have needed energy; but still! they escape).

$$\lambda = \pm \sqrt{\frac{2m(V_0 - E)}{\hbar^2}} = \pm K$$

real.

$$\psi(x) = C \cdot e^{Kx} + D \cdot e^{-Kx} \quad \text{general.}$$

$$\text{as } x \rightarrow \infty \quad e^{Kx} \rightarrow \infty$$

$$x \rightarrow -\infty \quad e^{-Kx} \rightarrow \infty$$

$$\therefore \psi(x) = C \cdot e^{Kx}; x < 0$$

$$D \cdot e^{-Kx}; x > 0.$$

$$C = D \quad (\text{for continuity}).$$

$$\psi(x) = C \cdot e^{-K|x|}$$

not diff. at $x=0$ if $C \neq 0$.

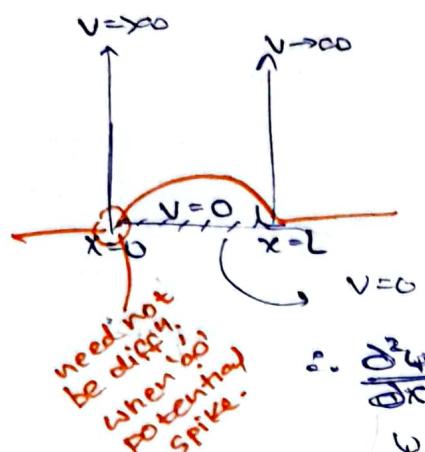
$$\therefore C = 0$$

- no physical situation; for $E < V_0$ everywhere

But!; for some regions; $E < V_0$ possible!
elsewhere; $E > V_0$

26. Sept.

2) particle in a box



$$\psi(x) = 0 \text{ for } x \in R - [0, L]. \quad \text{obvio}$$

$$\psi(x) = A \sin(kx) + B \cos(kx).$$

equivalent.

$$\psi(0) = 0.$$

$$\Rightarrow B = 0$$

$$\psi(L) = 0$$

$$\Rightarrow A \cdot \sin(kL) = 0$$

$\neq 0$

$$kL = n\pi$$

$$L = \frac{n\pi}{k}$$

$$k = \sqrt{\frac{2mE}{\hbar^2}}$$

$$\therefore k^2 = \frac{n^2\pi^2}{L^2}$$

$$E_n = \frac{n^2\pi^2\hbar^2}{2mL^2}$$

quantised!

for one stationary state $\psi_n(x)$

normalize:

$$|A|^2 = \frac{2}{L}.$$

$$A = \frac{e^{i\theta}}{\sqrt{L}} \quad \left. \begin{array}{l} \text{let us take real.} \\ \end{array} \right.$$

$$*\hat{H}\psi_n = E_n \psi_n$$

is a number.

but generally:

$$\hat{H}\psi \neq E\psi$$

if $\psi = c_1\psi_1 + c_2\psi_2 + \dots$

27. Sept.

* Now least energy $\neq E_0$
 $= E_1$

\therefore Quantum systems posses "zero-point energy".

(in classical phy.; $\epsilon = 0$)

* one more approach:

$$\frac{n\lambda_d}{2} = L$$

$$\lambda_d = \frac{2L}{n} = \frac{\hbar}{P} \Rightarrow P = \frac{n\hbar}{2L}$$

$$KE = E = \frac{P^2}{2m} = \frac{n^2\pi^2\hbar^2}{2mL^2}$$

≈

25th Sept. 2019.

Tut. 7:

Operators.
wavefn.

Sch. eqn for finding $\psi(x)$.

Particle in a box.

doubts:-

$$\nabla \langle \hat{P} \rangle = \int_{-\infty}^{\infty} \psi^* \hat{P} \psi dx$$

What if $\psi(x)$ is not eigen function.

$$\hat{P}\psi = P\psi$$

Wont be same for all.

✓ got it.

(a) $\psi(x) = A \left[\sin\left(\frac{\pi x}{L}\right) + \sin\left(\frac{2\pi x}{L}\right) \right]$

$$= \sqrt{\frac{L}{2}} A \cdot (\psi_1(x) + \psi_2(x))$$

→ particle in box.

$$\psi(x) = A \cdot \left(\sin \frac{\pi}{L} x + \sin \frac{2\pi}{L} x \right)$$

What about \hat{E}, \hat{E} , \hat{P}, \hat{P} . } need to collapse!

$$\int_{-\infty}^{\infty} \psi(x) \cdot \psi^*(x) dx$$

$$= \frac{1}{2} A^2 \left(\int_{-\infty}^{\infty} \psi_1^2(x) dx + \int_{-\infty}^{\infty} \psi_2^2(x) dx \right)$$

$$+ \boxed{\int_{-\infty}^{\infty} \psi_1(x) \cdot \psi_2(x) dx}$$

both are orthogonal.

$$\langle x \rangle = \int_{-\infty}^{\infty} \psi(x) \psi^* = 0.$$

$$\langle x^2 \rangle = \int_{-\infty}^{\infty} \psi(x)^2 dx = ?$$

$$\langle p \rangle = \int_{-\infty}^{\infty} \psi^* \left(-i\hbar \frac{d}{dx} \psi \right) dx$$

$$\langle p^2 \rangle = \int_{-\infty}^{\infty} \psi^* \left(-\hbar^2 \frac{d^2}{dx^2} \psi \right) dx$$

energy measurement is made $\rightarrow \neq E$ $\hat{H}(E) + (\text{Energy}).C_E$

$$= \phi$$

where $\hat{H}\phi = E\phi$ as $E \neq \text{eigen.}$

constant.

$$\psi = \frac{1}{\sqrt{2}} (\psi_1(x) + \psi_2(x))$$

$$\left| \int \psi^* \psi dx \right|^2 = 1 \quad \dots$$

$$\text{7) } \hat{H}\psi = -i\hbar \frac{\partial}{\partial t} \psi.$$

$$\psi_h(x, t) = \psi_0(x) \cdot e^{-iE_h t / \hbar}$$

only for ψ_0 , not for any ψ .

$\psi_{in} = \frac{1}{\sqrt{2}} (\psi_1(x) + i\psi_2(x)) \rightarrow$ not a stationary state.
not a possible situation for one dim. box.

$$\psi(x, t) = \frac{1}{\sqrt{2}} \left(\psi_1(x) e^{-iE_1 t / \hbar} + \psi_2(x) e^{-iE_2 t / \hbar} \right) \quad (\text{Bohr knows}).$$

(ψ_1, ψ_2) are independent

diff. with respect to each other \Rightarrow

one independent part ψ_1 and one independent part ψ_2

one independent part ψ_1 and one independent part ψ_2

one independent part ψ_1 and one independent part ψ_2

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one independent part ψ_1 and one independent part ψ_2

one independent part ψ_1 and one independent part ψ_2

Friday
27-sept-19

* zero point energy; by uncertainty.

$$\Delta x_{\max} = L$$

$$\Delta p_{\min} = \frac{\hbar}{L}$$

$$\therefore K_{\min} = \frac{p^2}{2m} = \frac{\hbar^2}{2mL^2}$$

Quantum particle can't have zero energy.
Where will the Ψ go?

→ Orthogonality:

Ψ_1, Ψ_2 are orthogonal if

$$\int_{-\infty}^{\infty} \Psi_1^* \cdot \Psi_2 dx = 0$$

$$\Psi_n = \sqrt{\frac{2}{L}} \cdot \sin\left(\frac{n\pi}{L} \cdot x\right)$$

* Ψ_l, Ψ_m are orthogonal for $l \neq m$.

for a Hermitian operator's domain;

eigen fun's of different eigen values are orthogonal.

→ time dependence:

(write),

$$\Psi(x,t) = W(x) \cdot \underline{\phi(t)}.$$

now; don't think $\Psi(x) \cdot \underline{\phi(t)} = \underline{\Psi(kx - \omega t)}$

as $\underline{\Psi(x)} \neq e^{ikx}$ (ain't true all the time).

∴ $\Psi(x,t)$; if $= \Psi(x) \cdot e^{-i\omega t}$.

then $(|\Psi(x)|)^2$ doesn't change; because of $\underline{\phi(t)}$...

take care.

* $\Psi(x)$ → used to find P, E, KE hence "stationary state!"

!! $\Psi(x)$ has \xrightarrow{x} used to find probability in B, B^*, b, b^* .
all data in it!

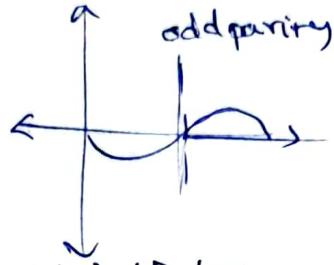
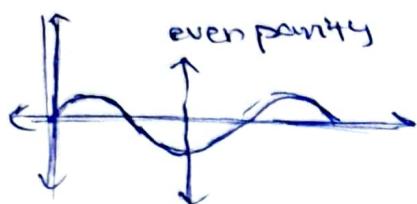
Uncertainty:

$\langle p \rangle = 0$. (as how much right; that much left).

$\langle p^2 \rangle \neq 0$.

$\langle x \rangle = \frac{L}{2}$? check!

Symmetry:



* discussion: particle in 1D box.

* Colours from Quantum dots:

nanoscale semiconductor arrangement is called quantum dot.

"size dependent colour"

$$\text{as } E_2 - E_1 = \frac{3\hbar^2\pi^2}{2mL^2}$$

this increases; colour becomes red.

Superposition:-

orthonormal eigen fns.

* $\Psi(x,t) = \sum C_n \Psi_n(x) \cdot e^{-iE_n t/\hbar}$

NOW; $|\Psi(x,t)|^2 = \Psi_{x,t} \cdot \Psi_{x,t}^*$

= time dependent!

\therefore this $\Psi(x,t)$ is not a stationary state.

* But $\int_{-\infty}^{\infty} |\Psi(x,t)|^2 dx = |C_1|^2 + |C_2|^2 + \dots$ (time independent!)

$\therefore \sum |C_n|^2 = 1$.

* see <slides>.

Copenhagen \rightarrow A place where
Quantum mechanics
postulates were
formulated.

* if a particle is in state Ψ and an ideal measurement of variable A will yield one of the eigenvalues ' α ' ($\alpha \hat{A}$) with probability $P(\alpha)$.

The state of system will change from Ψ to Ξ ($\hat{A}\Xi = \alpha\Xi$).

$P(\alpha) = \left| \int_{-\infty}^{\infty} \xi_{\alpha}^* \cdot \Psi dx \right|^2$ \rightarrow Copenhagen interpretation.

yields the famous
 $|\Psi(x)| \rightarrow$ to find particle.

\rightarrow measurement:

$$\Psi(x,t) = \frac{1}{\sqrt{2}} \cdot \Psi_1(x) e^{-iE_1 t/\hbar} + \frac{1}{\sqrt{2}} \cdot \Psi_2(x) e^{-iE_2 t/\hbar}$$

only corresponds to phase; no change in

$$\frac{|\Psi_1(x,t)|^2}{\Xi} + \frac{|\Psi_2(x,t)|^2}{\Xi} = 1$$

if system is not disturbed;

$\Psi(x,t)$ will "evolve" as such.

i.e. individual phase keep changing.

Ehrenfest theorem: "differentiating expectation values!"

$$\frac{d}{dt} \langle A \rangle = \langle [A, H] \rangle + \langle \frac{\partial A}{\partial t} \rangle$$

Mathematical ground:

$$[A, BC] = B[A, C] + [A, B]C \quad (\text{rel. order of } B, C \text{ won't change}).$$

$$[\bar{A}, BC] = A[B, C] + [A, C]B$$

$$[\hat{x}, p] = i\hbar \cdot -i\hbar \frac{\partial \psi}{\partial x} - (-i\hbar \frac{\partial (x^n \psi)}{\partial x})$$

$$[\hat{x}, p] = i\hbar \cdot n x^{n-1} \psi$$

$$[\hat{x}, \hat{p}] = i\hbar$$

$$\text{to do } [\hat{x}, p^2] = [\hat{x}, pp] = p[\hat{x}p] + [\hat{x}p]p \\ = 2i\hbar \cdot \hat{p}$$

→ equation of motion of "operator average":

$$\boxed{\frac{d \langle A \rangle}{dt} = \langle \frac{\partial A}{\partial t} \rangle + \frac{i}{\hbar} \langle [H, A] \rangle} \quad (\text{proof involved}; \hat{H}\psi = i\hbar \frac{\partial \psi}{\partial t})$$

if A is not explicitly time dependent;

$$\boxed{\frac{\partial \langle A \rangle}{\partial t} = \frac{i}{\hbar} \langle [H, A] \rangle} \quad (\hat{x}, \hat{p})$$

$$\int \psi^* H A \psi = \int (H \psi)^* A \psi$$

Hermitian.

$$\boxed{\frac{\partial \langle p \rangle}{\partial t} = -\langle V(x) \rangle}$$

if $[H, A] = 0$;

A is constant of motion.

Tut-8:

postulate is

$$\hat{H}\psi = i\hbar \frac{\partial \psi}{\partial t}$$

(lignes).

this will give.

$$\hat{H}\psi_{(x)} = E\psi_{(x)}$$

$$\text{if } \psi(x, t) = \psi(x) - \phi(t)$$

then $\psi(x, t)$ is stationary state.

$$\psi(x, t) = \sum c_i \psi_i(x, t)$$

satisfies

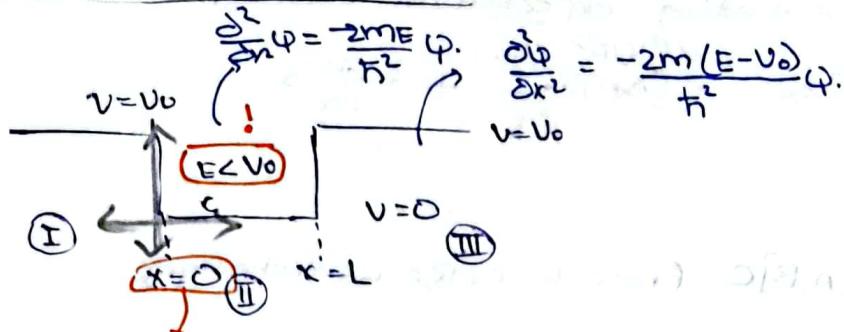
$$\hat{H}\psi = i\hbar \frac{\partial \psi}{\partial t}$$

$$\psi(x) = \sum c_i \psi_i(x)$$

doesn't satisfy.

$$\hat{H}\psi = E\psi.$$

* Body in a finite square well:-



here; ψ gotta be differentiable...

$\rightarrow \psi$'s differentiability is overlooked; only in case of $V \rightarrow \infty$.

Step 1: divide into region 1, 2, 3.

$$\text{R I: } \frac{\partial^2 \psi}{\partial x^2} = \frac{2m(V_0 - E)}{\hbar^2} \psi = \alpha^2 \psi. \quad (\text{R II: } \frac{\partial^2 \psi}{\partial x^2} = -\frac{2mE}{\hbar^2} \psi = -K^2 \psi. \quad \text{R III: } \frac{\partial^2 \psi}{\partial x^2} = -\frac{2mE}{\hbar^2} \psi = -K^2 \psi)$$

$$\therefore \psi = A \cdot e^{\alpha x} + B \cdot e^{-\alpha x} \quad \text{at } x=0 \quad B=0; \text{as boundary conditions.}$$

$$\psi_1 = A \cdot e^{\alpha x} \quad \left[\langle [A, H] \rangle \frac{\partial}{\partial x} + \langle [B, H] \rangle = \cancel{A} \alpha e^{\alpha x} \right]$$

$$\text{R II: } \frac{\partial^2 \psi}{\partial x^2} = -\frac{2mE}{\hbar^2} \psi = -K^2 \psi. \quad \text{Now; we have } \alpha, K. \dots$$

$$\text{R III: } \psi_3 = B \cdot e^{-\alpha x} \quad \left[\cancel{B} \cdot \cancel{e^{-\alpha x}} + D \cdot \cancel{e^{-\alpha x}} \right] \quad \text{(interdependent).}$$

on total; A, B, C, D, E are unknowns.

eqns: 1) cont. at $x=0$

2) diff. at $x=0$

3) cont. at $x=L$

4) diff. at $x=L$

5) Normalisation!

Yippe!

(S-S) var. equ?

$$A \cdot \alpha = C \cdot K.$$

$$B \cdot e^{\alpha L} = C \cdot \cancel{e^{KL}} + D \cdot \cancel{e^{-KL}}$$

$\Rightarrow C \cancel{e^{KL}}$ as initial part is zero.

$$C = N \left(\frac{\alpha}{K} \right)$$

$$D = \frac{\alpha}{\alpha - K}.$$

$$C \sin KL + D \cos KL \quad \cancel{D \sin KL} = 0 \quad \cancel{C \cos KL}$$

$$= B \cdot e^{-\alpha L}$$

$$C \cdot \cancel{e^{KL}} - D \cdot \cancel{e^{-KL}} = -\alpha B e^{-\alpha L}$$

$$\frac{\alpha}{\alpha - K} \cdot \frac{K}{\alpha} \cdot D \sin KL - C \cos KL = C \sin KL + D \cos KL.$$

at 10 Oct 2019.

Tut 8:

Q7) Hermitian if $H = H^\dagger$ dagger. (we know, that integration definition.)

$$\hat{C} = \hat{A}\hat{B}; A = A^\dagger \text{ as } A \text{ is Hermitian.}$$

$$\boxed{\hat{C} = \hat{C}^+} \Rightarrow \hat{A}\hat{B} = (\hat{A}\hat{B})^+ = (\hat{B}^+\hat{A}^+) = \hat{B}\hat{A}$$

$$\therefore \underline{\hat{A}\hat{B} = \hat{B}\hat{A}}.$$

SYNTAX: product of operators \hat{A}, \hat{B} is $\hat{A}\hat{B}$ or $\hat{B}\hat{A}$.

$$\downarrow \quad \downarrow$$

$$\hat{A}(\hat{B}\psi) \quad \hat{B}(\hat{A}\psi)$$

$$* [A, BC] = [A, B] + [A, C]$$

$$[A, BC] = B[A, C] + [A, B]C$$

B, C order won't change.

rodan.

$$\frac{\tan kL - \frac{\alpha}{k}}{\frac{\alpha}{k} + \tan kL + 1} = \frac{\alpha}{k}.$$

Satisfying value is E.

$$\alpha \tan kL - \tan kL + 2\alpha$$

$$\alpha^2 + \tan^2 kL + 2\alpha \tan kL = 0$$

$$\tan kL = \frac{2\alpha}{1-\alpha^2}$$

$$\frac{2\tan kL}{1-\tan^2 kL} = \frac{2\left(\frac{\alpha}{k}\right)}{1-\left(\frac{\alpha}{k}\right)^2}$$

α, k ; have only one variable $\Rightarrow E$ (energy);
solve numerically, or graphically to
get

"QUANTISED" values of E.

$$\boxed{\tan kL = \frac{\alpha}{k}}$$

$$\cot \frac{kL}{2} = -\left(\frac{\alpha}{k}\right)$$

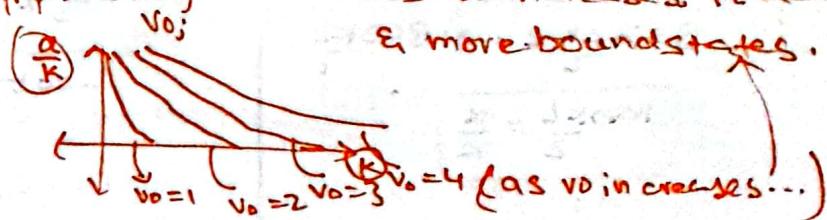


No. of ext points gives us

no. of bound states.

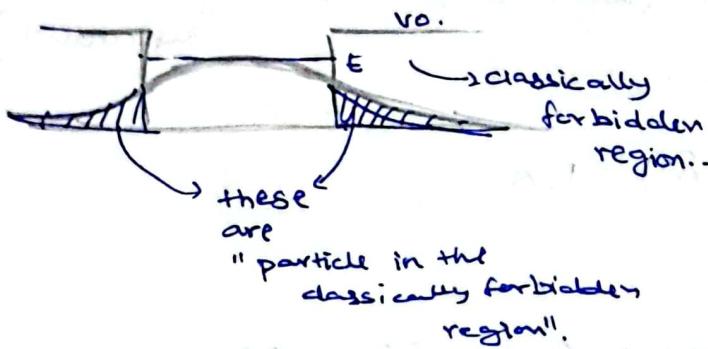
* (not infinite here...).

\rightarrow as V_0 increases; it admits more & more bound states.



$V_0=1 \quad V_0=2 \quad V_0=3 \quad V_0=4$ (as V_0 increases...)

diagrams:-



* Particle is not "contained" by a finite well.

(think; what happens with \bar{E} in a metal volume.)

→ Approximate energy expression:-



$$\text{penetration depth } (S) = \frac{1}{\alpha} = \frac{\hbar}{\sqrt{2m(V_0 - E)}}$$

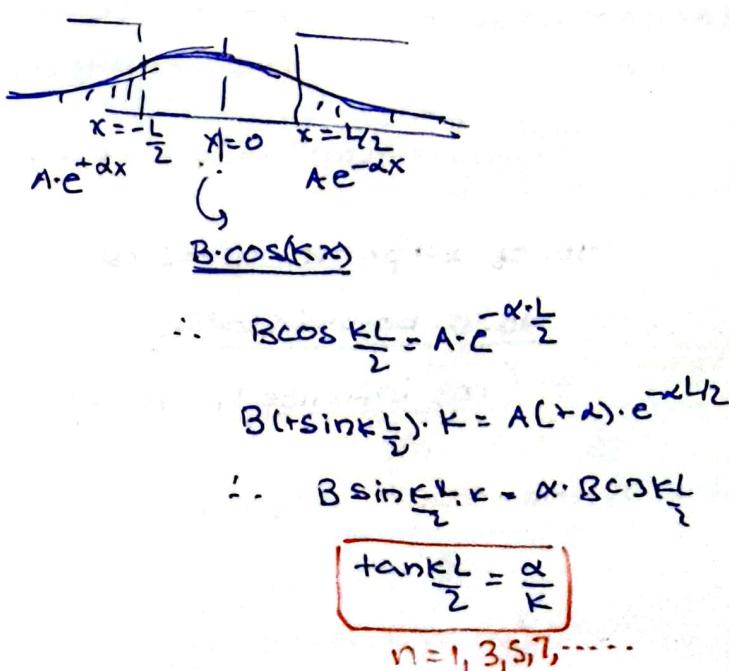
∴ effective dimension of well = $L + 2S$.

$$\therefore \text{Average energies} = \frac{n^2 \hbar^2 \pi^2}{2m(L+2S)^2}$$

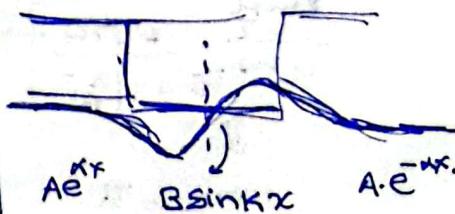
we won't use,
i.e.

* comparison of \bar{E} ; infinite/infinite well :- see lec 15, slide 18..

* even parity states:-



odd parity state:-



$$\therefore B \sin \frac{kL}{2} = A \cdot e^{-\frac{\alpha L}{2}}$$

$$B \cos \frac{kL}{2} \cdot k = A(-\lambda) \cdot e^{-\lambda L/2}$$

$$B \cos \frac{kL}{2} \cdot k = -\alpha \cdot B \sin \frac{kL}{2}$$

$$\boxed{-\cot \frac{kL}{2} = \frac{\alpha}{k}}$$

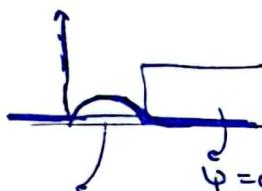
$$n = 2, 4, 6, \dots$$

→ How many states for a given value V_0, L :- (semi infinite well case).

* if $E = V_0$;

$$K = \frac{\sqrt{2mV_0}}{\hbar}$$

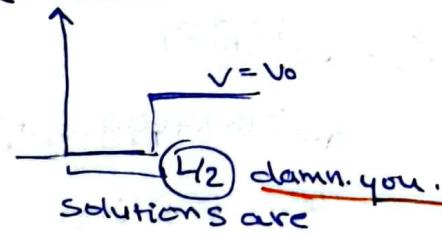
ζ



$\zeta = 0$; as $E - V = 0$.

$$\psi = A \cos(kx) + A \sin(kx)$$

$\therefore A \cos(kL) = 0$. $\therefore A \sin(kL) = 0$ too!; continuity of $\psi/x = v$.
 Σ slope = 0



* for no. of solutions; $\frac{\alpha}{K} \rightarrow 0$; $\alpha \rightarrow 0$
 $E \rightarrow V_0$

$$\text{i.e. } K \rightarrow (K_0 = \frac{\sqrt{2mV_0}}{\hbar})$$

\therefore if $K_0 \in [n\pi, -(n+1)\pi]$

then n , bound states.

$$V_0 = \frac{\hbar^2 k_0^2}{2m}$$

See that length of well
nicely, j_l

We finally get

$$n_{\max} = \{ \text{something} \}$$

$$n_{\max} = \frac{1}{2} + \sqrt{\frac{V_0}{\left(\frac{\pi^2 \hbar^2}{2mL^2}\right)}}$$

L = length of semi infinite well.

$$\frac{KL}{2} = \frac{\pi}{2}$$

$$\frac{\pi k}{2} = \frac{\pi}{2}$$

* Some example, - if we get

$$n_{\max} = 3.76$$

we have 4 bound states.

* plane waves. $\psi(x) = A e^{ikx}$.

(cannot normalise.)

Quantum mechanics in 3D:-

$$\star \psi(\vec{r}) = A \cdot e^{i(\vec{k} \cdot \vec{r})}$$

$$= A \cdot e^{i(k_x x + k_y y + k_z z)}$$

$$= (\psi(x) \psi(y) \psi(z)).$$

* for a cuboid :- (standard, $V=0$).

$$\hat{H}\psi = E\psi.$$

$$\frac{-\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \psi = E\psi.$$

$$\therefore \Sigma (\hat{H}_x \psi) = E\psi.$$

$$\therefore \text{since } H_x \psi = E_x \psi.$$

$$\text{we have } \psi_x = \sqrt{\frac{2}{L_x}} \cdot \sin(n_x \pi x)$$

$$\therefore \psi(x, y, z) = \sqrt{\frac{8}{V_{\text{cuboid}}}} \cdot \sin \frac{n_x \pi}{L_x} x \cdot \sin \frac{n_y \pi}{L_y} y \cdot \sin \frac{n_z \pi}{L_z} z.$$

$$\begin{aligned} \sum E &= E_x + E_y + E_z \\ &= \frac{\hbar^2 \pi^2}{2m L^2} \left(\sum \frac{n_x^2}{L_x^2} \right) \end{aligned}$$

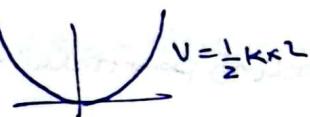
Degeneracy :-

if $g=n$; n -fold degeneracy.

degeneracy level.

Quantum Harmonic Oscillator:-

$$* F = -KX \Rightarrow V(X) = \frac{1}{2} KX^2.$$



in quantum mechanics; we rarely bring up "force".

Sch. equn:

$$\left(\frac{-\hbar^2}{2m} \frac{\partial^2}{\partial X^2} + \frac{1}{2} KX^2 \right) \Psi = E \Psi.$$

$$\frac{\partial^2 \Psi}{\partial X^2} + \frac{2m}{\hbar^2} \left(E - \frac{1}{2} KX^2 \right) \Psi = 0.$$

Hermite polynomials.

What will be the solutions? tough! $\Psi(X) = A e^{-\alpha X^2} (\sum C_n X^n)$

put back in the
to solve...

Ground state!:-

$$\Psi(X) = A \cdot e^{-\alpha X^2}$$

Put in Sch. equn;

$$\left(4\alpha^2 - \frac{mk}{\hbar^2} \right) X^2 + \left(\frac{2mE_0}{\hbar^2} - 2\alpha \right) = 0.$$

$$\therefore 4\alpha^2 = \frac{mk}{\hbar^2}; \quad \frac{2mE_0}{\hbar^2} = 2\alpha.$$

$$\text{we know; } \frac{k}{m} = \omega^2.$$

$$\therefore 4\alpha^2 = \frac{m\omega^2}{\hbar^2}$$

$$\alpha = \frac{m\omega}{2\hbar}$$

$$\Rightarrow \frac{2mE_0}{\hbar^2} = \frac{2m\omega}{2\hbar}$$

$$\boxed{E_0 = \frac{1}{2}\hbar\omega} !!$$

$\omega = \sqrt{\frac{k}{m}}$ is a constant
at any situation.

Ground state energy.

1st e.s!:-

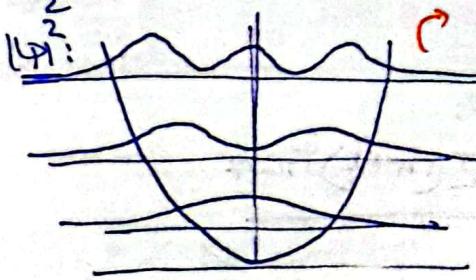
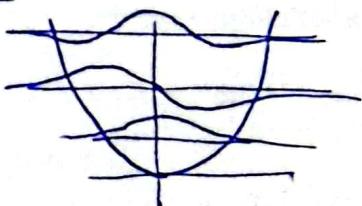
$$\Psi = Ax \cdot e^{-\alpha x^2}.$$

$$\text{we again get } \alpha = \frac{m\omega}{2\hbar} \quad \& \quad E_1 = \frac{1}{2} \cdot (3) \hbar\omega.$$

Energies are:- $\frac{1}{2}\hbar\omega, \frac{3}{2}\hbar\omega, \frac{5}{2}\hbar\omega, \frac{7}{2}\hbar\omega, \dots$

zero point energy $= E_0 = \frac{1}{2}\hbar\omega$.

$\Psi:$

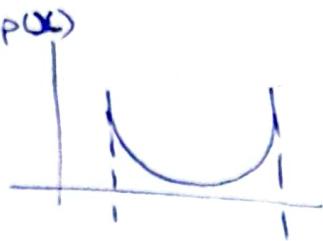


$|\Psi(x)|$ extends beyond
the (classical)
turning
points.

* Classical probability of finding particle:-

$$P(x) \propto \frac{1}{V(x)}$$

$$P(x) = K \cdot \frac{1}{\sqrt{A^2 - x^2}}$$



* Quantum probability vs classical p.



* Uncertainty in G.S:-

$$\langle x \rangle = 0$$

$$\langle x^2 \rangle = \frac{\hbar}{2m\omega}$$

$$\Delta x = \sqrt{\langle x^2 \rangle - \langle x \rangle^2} = \sqrt{\frac{\hbar}{2m\omega}}$$

$$\therefore \Delta p \cdot \Delta x = \frac{\hbar}{2}$$

$$\langle p \rangle = 0$$

$$\langle p^2 \rangle = \left(\sqrt{\frac{\hbar m \omega}{2}} \right)^2$$

$$\therefore \Delta p = \sqrt{\frac{\hbar m \omega}{2}}$$

(i.e. G.S. is a minimum uncertainty state).

→ in 3D:-

$$\Psi = \Psi(x) \Psi(y) \Psi(z)$$

$$\therefore E_{\text{net}} = \underbrace{(n_x + n_y + n_z) + \frac{3}{2}}_{\text{i.e. only when}} \hbar \omega \text{ only when isotropic!}$$

i.e. only when

$$\omega_x = \omega_y = \omega_z$$

$$\text{i.e. } k_x = k_y = k_z \dots$$

don't write blindly.

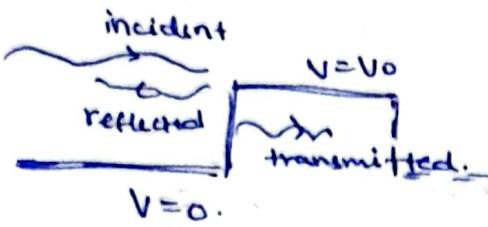
otherwise;

$$\underline{E = \sum (n + \frac{1}{2}) \hbar \omega_x}$$

Free state problems: Now; the particle is not bounded!

- Scattering,
tunneling.

* eg: A particle, coming from left; onto a step potential.



* If a particle is bounded-

() probability current is 0:

wild thoughts:

1) Far free particle;
maybe

$\psi(x,t) \neq \psi(x) \cdot \psi(t)$
as it gotta be a
traveling wave.

(but even then;

$\hat{H}\psi = i\hbar \frac{\partial}{\partial t} \psi$ holds no)

$$i\hbar \frac{\partial}{\partial t} \psi = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V(x) \cdot \psi \quad (1)$$

$$-i\hbar \frac{\partial}{\partial t} \psi^* = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi^*}{\partial x^2} + V(x) \cdot \psi^* \quad (2)$$

$$\psi^* \times (1) - \psi \times (2)$$

$$\therefore i\hbar \left(\frac{\partial}{\partial t} (\psi^* \psi) \right) = -\frac{\hbar^2}{2m} \left[\psi^* \frac{\partial^2 \psi}{\partial x^2} - \psi \frac{\partial^2 \psi^*}{\partial x^2} \right]$$

$$\text{identity: } \frac{\partial}{\partial x} \left(\psi^* \frac{\partial \psi}{\partial x} - \psi \frac{\partial \psi^*}{\partial x} \right) = \psi^* \frac{\partial^2 \psi}{\partial x^2} - \psi \frac{\partial^2 \psi^*}{\partial x^2}$$

$$\therefore i\hbar \cdot \frac{\partial}{\partial t} (\psi^* \psi) = -\frac{\hbar^2}{2m} \cdot \frac{\partial}{\partial x} \left(\psi^* \frac{\partial \psi}{\partial x} - \psi \frac{\partial \psi^*}{\partial x} \right)$$

$$\therefore \frac{\partial}{\partial t} (\psi^* \psi) + \frac{\hbar}{2mi} \cdot \frac{\partial}{\partial x} \left(\psi^* \frac{\partial \psi}{\partial x} - \psi \frac{\partial \psi^*}{\partial x} \right) = 0$$

if a particle is

"moving",

the
prob.current
is finite.

$$\boxed{\frac{\partial P}{\partial t} + \frac{\partial J}{\partial x} = 0} \rightarrow \text{think like } \boxed{J dt = P dx} ?? \dots$$

$J(x,t)$ is the "current" associated with
probability density P .

$J =$

$$\boxed{J = \frac{\hbar}{2im} \left(\psi^* \frac{\partial \psi}{\partial x} - \psi \frac{\partial \psi^*}{\partial x} \right)} \quad \boxed{J = \frac{\hbar}{m} \cdot \text{Im}(\psi^* \frac{\partial \psi}{\partial x})}$$

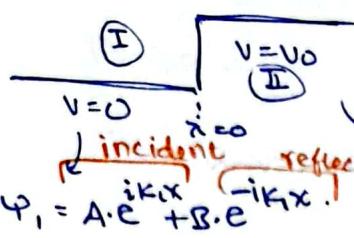
units: \vec{s} , [Time] $^{-1}$

* if $\psi = A \cdot e^{ikx}$

$$J = \frac{\hbar}{m} \cdot \text{Im}(A^2 e^{-ikx} \cdot ik \cdot e^{ikx}) = \frac{P}{m} |A|^2 = V \cdot |A|^2$$

(case-1)

$$E > V_0.$$



$$J = \frac{\hbar}{m} \cdot \text{Im} \left(\Psi \frac{\partial \Psi^*}{\partial x} \right)$$

$$k_1 = \sqrt{\frac{2mE}{\hbar}}$$

Determine \hbar using Ψ .

so if $\psi_1 = A \cdot e^{ik_1 x} + B \cdot e^{-ik_1 x}$; $\psi_2 = C \cdot e^{ik_2 x}$
 then $P = \hbar k_1$ $P = -\hbar k_1$ $P = \hbar K_2$.

continuity at 0.

$$A + B = C$$

diff.

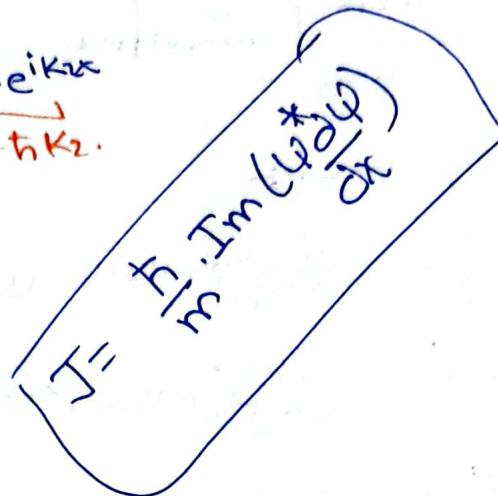
$$(A - B) k_1 = C k_2$$

$$A - B = C \cdot \frac{k_2}{k_1}$$

$$2A = C \frac{(k_1 + k_2)}{k_1}$$

$$C = \left(\frac{2k_1}{k_1 + k_2} \right) A \Rightarrow B = \left(\frac{k_1 - k_2}{k_1 + k_2} \right) A.$$

we cannot normalize the Ψ_1 or Ψ_2 .



lets see probability current.

Here $J = \nabla \cdot |\Psi(x)|^2$ (as $A \cdot e^{ikx}$ form).
 $= P/m$.

$$J_i = J_{\text{incident}} + J_{\text{reflected}}$$

$$J_i = +\frac{\hbar k_1}{m} (|A|^2 - |B|^2)$$

$$J_2 = J_{\text{trans}} = \frac{\hbar k_2 |C|^2}{m}$$

$$J_{\text{incident}} = |J_{\text{reflected}}| + J_{\text{transmitted}}$$

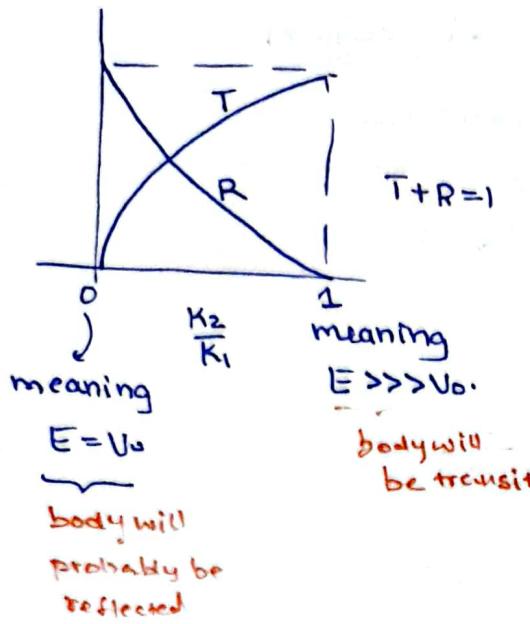
$$\text{Reflection coefficient (R)} = \frac{J_R}{J_i} = \left(\frac{k_1 - k_2}{k_1 + k_2} \right)^2$$

$$\text{Transmission coefficient (T)} = \frac{J_T}{J_i} = \frac{2k_1 k_2}{(k_1 + k_2)^2}$$

$$\therefore R + T = 1$$

* $K_1 = K \cdot \sqrt{E}$ (our case is $E > V_0$).
 $K_2 = K \cdot \sqrt{E - V_0}$

$$\therefore \frac{K_2}{K_1} = \sqrt{1 - \frac{V_0}{E}} \quad \text{&} \quad R = f\left(\frac{K_2}{K_1}\right); T = f\left(\frac{K_2}{K_1}\right)$$



Conclusion:-

classically;

if $E > V_0$; the body "goes" be transmitted;

but

Quantumly;

the body has finite probability of reflecting,

case-II) $E < V_0$:

$$E < V_0 \quad V = V_0.$$

$$\textcircled{I} \quad \frac{\partial^2 \psi}{\partial x^2} = K_2^2 \cdot \psi.$$

$$\frac{\partial^2 \psi}{\partial x^2} = -K_1^2 \psi.$$

$$\psi = A \cdot e^{+ikx} \text{ right.} \quad \textcircled{II} \quad \psi = C \cdot e^{K_2 x} + D \cdot e^{-K_2 x} \quad \psi_{\infty}(\infty) = 0 \quad \therefore C = 0.$$

$$\psi = A \cdot e^{+ikx} + B \cdot e^{-ikx} \quad P = \hbar K_1 \quad \text{But!}$$

left.

$$P = -\hbar K_1.$$

e^{ikx} (imaginary) \rightarrow not normalizable function

$e^{K_2 x}$ \rightarrow normalizable in $-\infty$ to 0.

$$\therefore \psi_{\infty} = C \cdot e^{-K_2 x}$$

real function; probability current

$= 0$!

Solving:

continuity: $A + B = C$

slope: $A K_1 - B K_1 = C (-K_2)$

$$A - B = -\frac{K_2 \cdot C}{K_1}$$

$$\therefore B = \left(\frac{i k + \alpha}{i k - \alpha} \right) A; C = \frac{2 i k}{i k - \alpha} \cdot A.$$

note here: $J_{\text{reflected}} = J_{\text{incident}}$.

"try it out; $\Im(\frac{d^2 \psi}{dx^2}) = 0$

$J_{\text{transmission}} = 0$

if $E < V_0$:

reflection coefficient = 1.
transmission coefficient = 0.

But note: $\psi \neq 0$ in the transmission region.

2. The particle can penetrate the barrier....

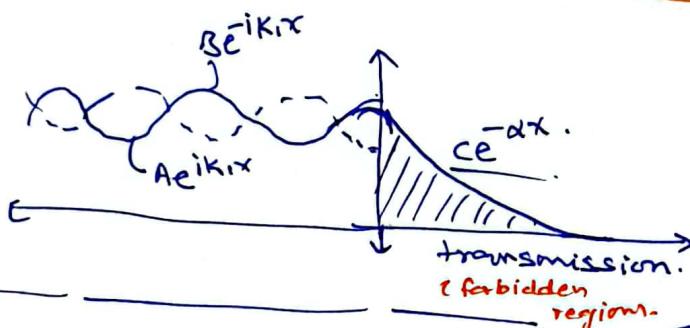
(existence is there; current isn't there).

* on right:

$$|\psi(x)|^2 \propto \left(\frac{4k^2}{k^2 + \alpha^2} \right) e^{-2\alpha x}.$$

like water in a pit;
of a flowing river.

* Penetration depth $S = \frac{1}{\alpha} = \frac{h}{\sqrt{2m(V_0 - E)}}$



* STEP POTENTIAL:

$E < V_0$.

$$\boxed{V = V_0 \dots}$$

Solve to obtain
reflection &
transmission
coefficients.

(I) $x=0$ (II) $x=L$ (III)

$$Ae^{ikx} + Be^{-ikx}$$

$$Ce^{\alpha x} + De^{-\alpha x}$$

$$Fe^{ikx}$$

no e^{-ikx} term

Note

$$x \rightarrow \infty$$

$$x \rightarrow -\infty$$

as only transmitted
wave is there.

hence both
terms in
general equation.

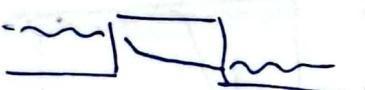
Transmission
coefficient; =

$$= \frac{(F)^2}{(A)}.$$

$$\boxed{\frac{V_0 \psi_{\text{left}}}{V_0 \psi_{\text{right}}}}$$

$$\sinh(x) = \frac{e^x - e^{-x}}{2}$$

$$\frac{1}{T} = 1 + \frac{1}{4} \left(\frac{d}{K} + \frac{K}{\alpha} \right)^2 \sinh^2(\alpha L)$$

in  case.

$$K = \frac{\sqrt{2mE}}{\hbar}; \quad \alpha = \frac{\sqrt{2m(V_0 - E)}}{\hbar}$$

obvious.

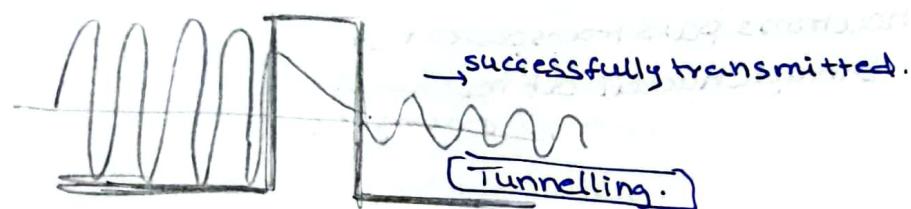
$$T = \left[1 + \frac{1}{4} \left(\frac{V_0^2}{E(V_0 - E)} \right) \sinh^2(\alpha L) \right]^{-1}$$

$$\left(1 + \frac{1}{4} \cdot \frac{V_0^2}{E(V_0 - E)} \sinh^2(\alpha L) \right)^{-1}$$

*

two factors: V_0 (i.e. α) height
 L width.

(increase in any;
decrease T). 



$$\text{Barrier penetration depth} = \frac{1}{\alpha} = \frac{\hbar}{\sqrt{2m(V_0 - E)}}$$

Limiting cases:-

1) $\alpha L \gg 1$

2) $\alpha L \ll 1$

3) if $E > V_0$:

Simple modification; $\alpha = ik$; $K = \sqrt{\frac{2m(E-V_0)}{\hbar^2}}$

$$T = \left[1 + \frac{1}{4} \cdot \frac{V_0^2}{E(E - V_0)} \sin^2(kL) \right]^{-1}$$

→ Transmission

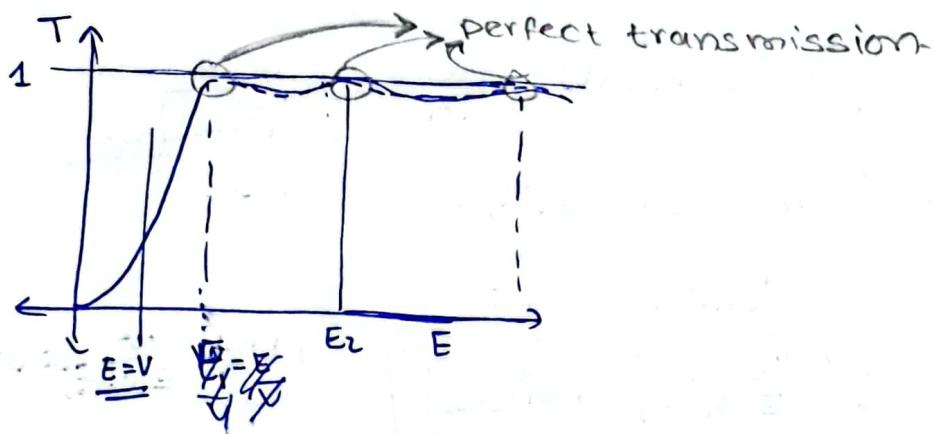
Resonance:-

∴ for $kL = n\pi$; $T = 1$

transmission resonance. as $T=1$.

B=0 in this case?.

Transmission resonance:
finally; for  thing;



- * Ramsauer effect: noble gases; transparent to $E \propto$ certain energy.

Size resonance: neutrons pass transparently through nuclei at resonant energies

$$KL = n\pi$$

$$\text{i) } \sqrt{2m(E-V)} = \frac{n\pi h}{L}$$

$$E = V_0 + \frac{n^2 \pi^2 h^2}{2m L^2}$$

- scanning tunneling microscope

for transmission resonance.

$$\text{ii) or; } L = \frac{n\pi}{K} = \frac{n\pi}{2\pi} \cdot \lambda = n\left(\frac{\lambda}{2}\right)$$

λ = wavelength of particle in barrier region.

i. for transmission resonance;

$$\text{Length barrier} = n \times \frac{\lambda}{2}.$$

→ α -decay:

$$\lambda = \text{decay rate} = f \cdot T(E)$$

$$(f = \frac{1}{2R} \cdot \frac{V^2}{\lambda})$$

frequency

Barrier length = L .

$$1 + \frac{V^2}{4E(V-E)} \cdot \sinh^2(\alpha L)$$

V is appropriately taken.

$$\underline{\underline{t_{1/2} = \frac{\ln 2}{\lambda}}}.$$

tut - last 2nd:

Statistics:

$$MB: \frac{N_i}{g_i} = e^{\alpha} \cdot e^{-\beta E_i}$$

$$EB: \frac{N_i}{g_i} = \frac{1}{e^{\alpha + \beta E_i} - 1}$$

$$FD: \frac{N_i}{g_i} = \frac{1}{e^{\alpha + \beta E_i} + 1}$$

Q) even if e is kept in a box;
we will have spin degeneracy;

$$\delta = Y_2 (\text{intrinsic to } e)$$

not to atom.)

Intro to quantum statistics:-

Founders:

of
statistical
physics:-

L. Boltzmann

J. W. Gibbs

J. C. Maxwell

- ignore particle-particle
interaction.

- * How many ways can the particles be distributed in energy levels; with constraints of fixed no. of particles & fixed total energy E .

Equilibrium distribution is the most probable way of distribution.

distinguishable particles

indistinguishable particles

microstate can be more;
for each macrostate.

macrostate = 1; for each macrostate.

called multiplicity.

1) Maxwell Boltzmann (MB):

→ Assumes that each microstate is equiprobable.

- * if single degeneracy; there is observed exponential decay in occurrence of

(hence; not
the real gases
case)

higher energy
states.

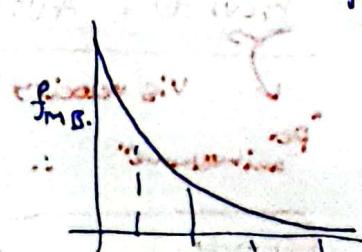
$$\therefore f_{MB} = A \cdot e^{-\frac{E_i}{K_B T}}$$

for discrete
energy levels.

proof is O.K.

$$\frac{N_i}{g_i} = A \cdot e^{-\frac{E_i}{K_B T}}$$

proof ✓



is when; density
of states
comes into
picture).

* if energy levels are large; & are closely spaced;

$$f_{MB} = A e^{-E_i/k_B T} \quad (\text{for discrete})$$

i.e.

$$n_i = g_i \cdot A \cdot e^{-E_i/k_B T}$$

now:

$$\frac{n(E) dE}{dE} = g(E) dE \cdot A \cdot e^{-E/k_B T}$$

$n(E) \rightarrow$ no. of particles per unit volume.
 dE with E to $E+dE$ energy.

$g(E) dE \rightarrow$ density of states (no. of states in E to $E+dE$ energy band). per unit volume

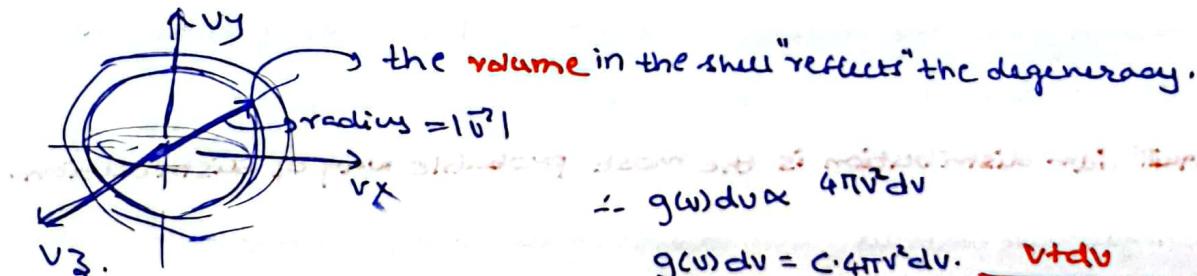
→ calculating density of states:-

$$E = \frac{1}{2}mv^2 \quad (\text{for gas molecules}).$$

So; now try to see possible values of $|\vec{v}|$.

∴ we see the different "directions" to reflect upon the
& this is how:- degeneracy of $|\vec{v}|$.

$$\vec{v} = v_x \hat{i} + v_y \hat{j} + v_z \hat{k}$$



$$\therefore g(v) dv \propto 4\pi v^2 dv$$

$$g(v) dv = C \cdot 4\pi v^2 dv. \quad \frac{v dv}{E+dE} \dots E+dE$$

$$\text{now: } v = k \cdot \sqrt{E}.$$

$$\frac{g(E) dE}{E} \dots \frac{v dv}{E}$$

$$\boxed{g(v) dv = g(E) dE \text{ feel karo.}}$$

$$\therefore g(E) dE = A v^2 dv$$

$$= A E \cdot \frac{1}{\sqrt{E}}$$

$$\boxed{g(E) dE = A E^{1/2}}$$

∴ now

$$\boxed{n(v) dv = A v^2 e^{-\frac{1}{2} \frac{mv^2}{k_B T}} dv}$$

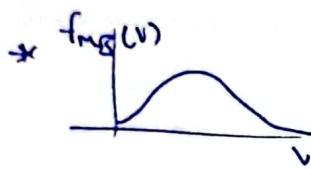
v is velocity

"no minima" ∴ $\int_{v=0}^{\infty} n(v) dv = \frac{N}{V}$ (! not N !)

$$\boxed{\frac{n(v) dv}{N/V} = 4\pi \left(\frac{m}{2\pi k_B T} \right)^{3/2} v^2 \exp(-mv^2/2k_B T) = f_{MB}(v)}$$

our chemistry formula.

fraction of molecules.



$$\begin{aligned} \text{Standard deviation.} &= \sqrt{\langle f_{\text{rel}}^2 \rangle - \langle f_{\text{rel}} \rangle^2} \\ &= \sqrt{\frac{3kT}{m} - \frac{8}{\pi} \frac{RT}{m}} \\ &= \sqrt{3 - \frac{8}{\pi}} \cdot \sqrt{\frac{RT}{m}}. \end{aligned}$$

$$\langle \text{Kinetic energy} \rangle = \frac{3}{2} k_B T$$

Statement of equipartition theorem:

γ -function:-

$$\gamma(a) = \int_0^\infty e^{-x} x^{a-1} dx$$

$$\gamma(a) = (a-1) \gamma(a-1)$$

$$\gamma(\frac{1}{2}) = \sqrt{\pi}$$

$$\gamma(\frac{1}{2}) = \int_0^\infty e^{-x} \cdot \frac{1}{2} x^{-\frac{1}{2}} dx$$

$$\gamma(\frac{1}{2}) = \int_0^\infty e^{-u} \sqrt{u} du = \int_0^\infty e^{-t^2} \frac{1}{t} \cdot 2tdt$$

$$\gamma(\frac{1}{2}) \cdot \gamma(\frac{1}{2}) = \int_0^\infty \int_0^\infty e^{-(x+y)} \frac{1}{xy} dx dy$$

Maxwell-Boltzmann assumptions:-

- 1) distinguishable.
- 2) particles must be far apart.
- 3) no restriction on no. of particles in a vessel.
- 4) classical particles; no wave nature.

* * Validity: (How to decide distinguishable or indistinguishable)
by 4th postulate:

(Adiabatic \ll average distance b/w particles)

i.e.
$$\left[\frac{\hbar}{\sqrt{2m \cdot \frac{3}{2} k_B T}} \ll \left(\frac{V}{N} \right)^{1/3} \right] !$$
 don't forget $\left(\frac{1}{3} \right).$

\therefore MB valid when: high T;
low particle density,
high mass.

P.T.O.

He at 273K ✓✓

Helium at 4.2K XX

when $\lambda_{db} \ll d$:



particles are distinguishable.

MB statistics applies.

if $d < \lambda_{db}$:



particles are indistinguishable.

Bose-Einstein & FD statistic -

fermi
dirac

* indistinguishable *.

$$\Psi(1,2) = \pm \Psi(2,1)$$

∴ two-particle wavefunction is either symmetric or anti-symmetric.

wrt interchange of particles.

+ } Bosons → can be at the same place at same time.
photon, → integral spins. $Hg (=1)$, $H (=1)$, He

- } fermions → electrons
protons.
can't be at the
same place
at the same
time.

$$e^-(=\frac{1}{2}), p^+(\frac{1}{2}), n^+(\frac{1}{2}) \dots$$

unit of spin = $\hbar = 1.2 \times 10^{-31}$ Jouls second.

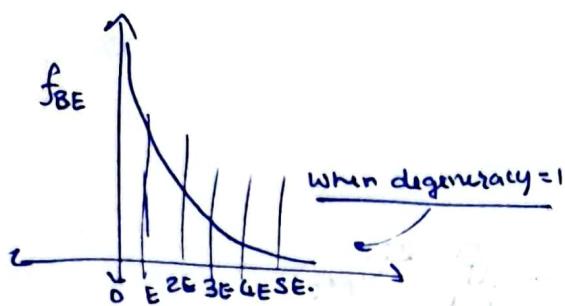
* Boson: more particles in same state is possible.

fermions: if $\Psi_a = \Psi_b$; $\Psi = \frac{1}{\sqrt{2}} (\Psi_a(1) \Psi_b(2) - \Psi_b(1) \Psi_a(2)) \underline{\underline{= 0}}$.

Fermions avoid each other.

→ Boson distribution:-

each macrostate; gives only one microstate.



In bosonic systems;

Lower energy levels; one more occupied.
than MB.

→ Fermion distributions

No more than two particles can be in a single level.

(1,1)

$$E = 4E_0$$

✓

1

$\rightarrow 4E$; not at all occupied.

Fermi distribution

These both are
equi probable....
(as indistinguishable)

probability of occupation of energy levels:

sizable at lower energies.

Sudden dip at higher energy levels

→ Counting microstates: with degeneracy

1) MIS distribution:-

$$n_i = \frac{N! \cdot \prod_{j=1}^k (g_j)^{n_j}}{n_1! n_2! \cdots n_k!}$$

$$3E - - - - - = 9 = 4$$

$$2E - \text{---} - g=2$$

$$E \leftarrow -g =$$

0 - 9-1

2) Bosonic distribution:-

$$- - - \quad g_i$$

$$\text{no. of microstates: } N_i t g_{i-1} c_{g_{i-1}}$$

$$\therefore \text{count} = \sum_{i=1}^K n_i + g_{i-1} C_{g_{i-1}}$$

Fermionic:

--- Fermi degeneracy
 Many states appear equal
 $g_i \geq N_i$ ways = $g_i C_{N_i}$

$$\therefore \text{no. of microstates} = \prod_{i=1}^n g_i C_{N_i}$$

* In case of large no. of particles under the constraint of

- (1) Fixed no. of particles.
- (2) Fixed energy.

it can be (rigorously) shown that:

B.E -

$$f(E) = \frac{1}{B e^{E/kT} - 1}$$

$f(E)$ = probability of finding a particle in an energy state of Energy E .

FD:

$$f_{FD}(E) = \frac{1}{H e^{E/kT} + 1}$$

against classical particle:-

$$f_{MC}(E) = A \cdot e^{-E/kT}$$

B.H
2
by normalisation.

$$\int f dE \neq 1$$

if $g_i = 2$
 then each state has $f(E)$
 $\therefore \text{net} = g_i \cdot f(E)$
 only E states.

$$f(E) = \frac{N_e dE}{g(E) dE}$$

essentially.

\rightarrow To get that gas equation;

we still need to take degeneracy

Density of states into account.

$$\int g_i(E) f dE = \frac{N}{V}$$

after normalisation, we get:-

$$- B = 1, \therefore f_{BE}(E) = \frac{1}{e^{E/kT} - 1}$$

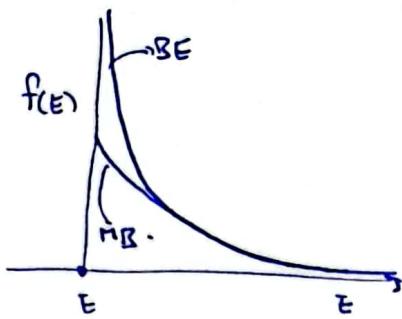
$$- H = e^{-EF/kT}$$

$$f_{FD}(E) = \frac{1}{e^{(E-E_F)/kT} + 1}$$

E_F : Fermi energy.

Shouldn't be -.

Bose-E.statistics:-



as $E \rightarrow \infty$ $f_{BE} = f_{MB}$. (i.e. all classical particles behave as Bosons at low T.)

as $E \rightarrow 0$, $f_{MB} = \text{finite}$

$$f_{BE} \rightarrow \infty$$

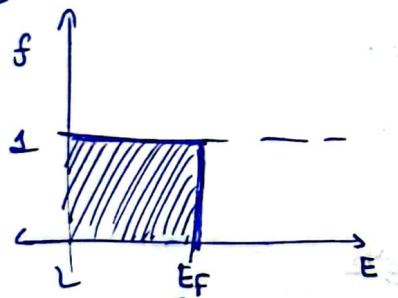
∴ at low temperatures, all particles drop to ground state

Hence:

Bose-einstein condensate

* Fermi-dirac statistics:-

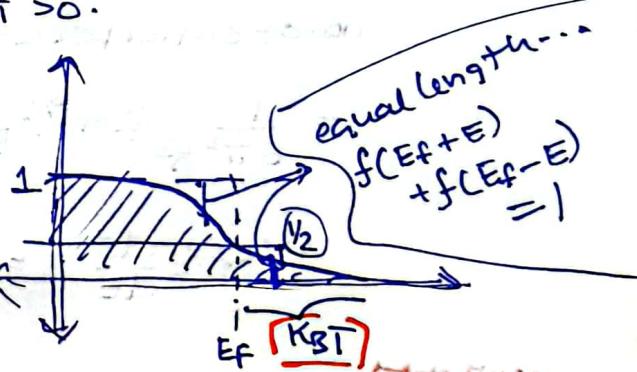
1) $T=0^+$



defines highest energy at $T=0K$.

→ $T>0$; the FD graph smears out.

2) $T>0$.



* Fermi temperature:

$$E_F = K_B \cdot T_F$$

$$T_F = \frac{E_F}{K_B}$$

for $T > T_F$;

$$f_{FD} \rightarrow f_{MB}$$

at high temperatures:-

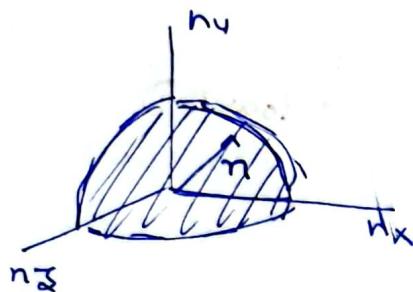
$$\frac{f_{BE}}{f_{FD}} \rightarrow f_{MB}$$

Density of states; quantum picture:-

3D:

$$\frac{E}{\hbar} = \frac{n_x^2 + n_y^2 + n_z^2}{2mL^2}$$

$$E = \frac{\pi^2 \hbar^2}{2mL^2} (n_x^2 + n_y^2 + n_z^2)$$



No. of states \propto volume of $\frac{1}{8}$ th sphere

In that n_x, n_y, n_z it was proportional,
here equal.

$$S = \frac{1}{8!} \frac{4}{3} \pi n^3$$

$$n = \sqrt{2mE} \frac{L}{\pi \hbar}$$

$$S = \frac{1}{6} \pi (2mE)^{3/2} \left(\frac{L}{\hbar \pi}\right)^3$$

$$V = L^3$$

$$\text{No. of states per unit volume} = \frac{S}{V}$$

$$= \frac{1}{6\pi^2 \hbar^3} (2m)^{3/2} E^{1/2}$$

$$g(E) = \frac{ds}{dE} = \frac{1}{4\pi^2 \hbar^3} (2m)^{3/2} E^{1/2}$$

No. of states
per unit energy
per unit volume

$$\boxed{g(E) = \frac{2\pi}{\hbar^3} (2m)^{3/2} E^{1/2}}$$

2D:

$$E = \frac{\hbar^2 \pi^2}{2mL^2} (n_x^2 + n_y^2)$$



$$S = \frac{1}{4} \pi n^2$$

$$\boxed{g(E) = \frac{2\pi m}{\hbar^2}}$$

1D:

$$S = n = \sqrt{2mE} \frac{L}{\pi \hbar}$$

$$\boxed{g(E) = \frac{(2m)^{1/2} \cdot E^{-1/2}}{\hbar}}$$

$$\begin{aligned}
 3D \quad g(E) &= \frac{2\pi}{h^3} (2m)^{3/2} E^{1/2} \\
 2D \quad g(E) &= \frac{2\pi m}{h^2} \\
 1D \quad g(E) &= \frac{(2m)^{1/2}}{h} E^{-1/2}
 \end{aligned}$$

without considering
particle's spin.

if E are filled up:

$$g(E) \times 2 = \text{density of states}.$$

* if they just ask: "Number of states per unit energy"; $[g(E) \times \text{Volume}]$ is our answer.

* $g(p)dp = g(v)dv = g(E)dE = g(k)dk = \dots$

Blah!

Phase space -

System of n particles; defined by q_i, p_i

$$q_i = (x_i, y_i, z_i)$$

$$p_i = p(x_i, p_x, p_y, p_z)$$

instantaneous
state;

defined by a point in phase space.

Red box.

$$g(p)dp = \frac{4\pi p^2 dp}{h^3}$$

? B-E statistics
application.

→ Planck law for blackbody:

$$g(p)dp = \frac{C_1 \pi p^2}{h^3} dp.$$

to account for the two polarisations:

$$g(p)dp = 2 \times \frac{4\pi p^2}{h^3} dp.$$

ν = frequency.

$$g(\nu)d\nu = \frac{8\pi \nu^2}{c^3} d\nu !$$

No. of photons occupied $E \pm dE$ to $h(\nu + d\nu)$

$$h\nu_{d\nu} = \frac{8\pi \nu^2}{c^3} d\nu \frac{1}{e^{\frac{h\nu}{k_B T}} - 1}$$

↑ B-E statistics.

$$\therefore U(\nu)d\nu = n(\nu)d\nu(h\nu) = \frac{8\pi \nu^2}{c^3} \cdot \frac{h\nu}{e^{h\nu/k_B T} - 1} d\nu$$

F.D. Statistics appln:-

Free e⁻ gas theory of metals:-

{pending}.

* Density of e⁻:

$$n_e = Z_c \cdot N_A \cdot \frac{dm}{A}$$

contribution from each atom



now; even for e⁻; FD statistics applied:-

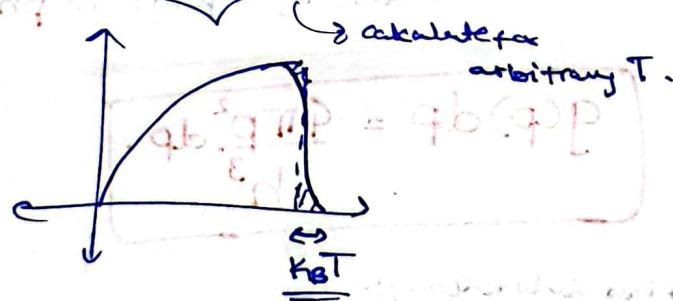
3D:

$$g(E) = \frac{8\pi}{h^3} \cdot \sqrt{2E \cdot m} e^{3/2} dE$$

Spin has 2 orientations.

$$\therefore g(E) \text{ is } \frac{8\pi}{h^3} \cdot \sqrt{2E \cdot m} e^{3/2} dE$$

$$\therefore n_e = \frac{8\sqrt{2}\pi m e^{3/2}}{h^3} \int_0^\infty E^{1/2} \frac{1}{e^{(E-E_F)/k_B T + 1}} dE.$$



said; at T=0

$$\text{now } n_e = \left(\dots \right) \cdot \int_0^\infty \sqrt{E \cdot 1} dE$$

we have ..

$$E_F = \frac{\hbar^2}{2me} \left(\frac{3n_e}{8\pi} \right)^{2/3}$$

E_F increase with n_e

at e⁻ fill more states.

$$\frac{1}{2} m_e v_F^2 = E_F$$

↳ fermi velocity

$$T_F = \frac{E_F}{k_B}$$

↳ fermi temperature.

Tut - 12:

1)

f.d....

$$n(E) = \frac{\int_{-E_F}^E q(E) \cdot f_{FD}(E) dE}{\int_{-E_F}^E q(E) f_{FD}(E) dE}$$

$q(E)$ not mentioned!

use particle in 3D-box!

$$q(E) \approx (\text{const.}) \cdot E^{1/2}$$

$$\text{Sp. heat of Cu} = \frac{dQ}{dT} = \left(\frac{3}{2} k_B \times \frac{3}{2} \frac{kT}{E_F} \right) N_A$$

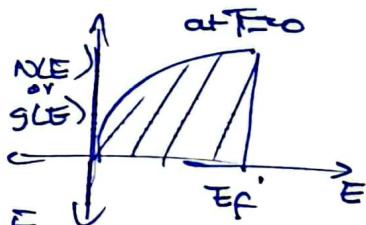
contd.

No. let us calculate E_{total} .

$$E_{\text{total}} = \int_0^{E_F} E n(E) dE \quad \text{at } T=0$$

↓

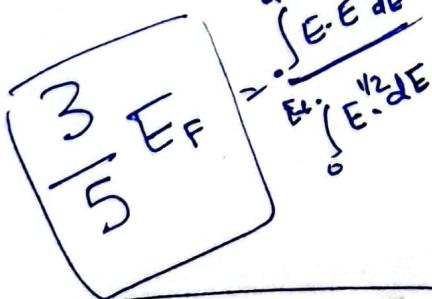
$$n(E) dE = g(E) dE$$



- finally

$$E_{\text{total}} = N_E \cdot \left(\frac{3}{5} E_F \right)$$

Average \bar{E} energy (in metal) at $T=0 = \frac{3}{5} E_F$



(This is \bar{E} according to classical theory;

but quantum effects make this impossible;
all \bar{E} could be in $E = \text{constant}$).

Heat capacity due to \bar{E} :

$$\begin{aligned} \text{Say excited no. of electrons} &= n(G^+) dE \\ &= g \end{aligned}$$

