

Newton's Law of Cooling

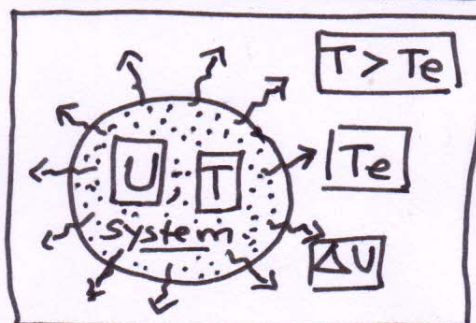
Statement: Rate of loss of heat is proportional to the excess temperature over the surroundings.

- i) The law is true for all excess temperatures in conditions of forced convection of air.
- ii) The law is approximately true in still air for a temperature excess of 20K-30K.
- iii) For natural convection the proportional relation is (excess temperature)^{5/4} (Dulong and Petit)

1. $\Delta U \rightarrow$ Energy lost as heat

2. $(T - T_e) \rightarrow$ Excess temperature

3. $\Delta U / \Delta t \propto (T - T_e)$ Newton's Law of Cooling ($T \rightarrow$ Kelvin unit)



4. $\Delta U \rightarrow du$ (Infinitesimal loss)

5. $du = mc dT$ $m \rightarrow$ mass, $c \rightarrow$ specific heat capacity

6. $\frac{du}{dt} = mc \frac{dT}{dt} \propto (T - T_e) \Rightarrow \frac{du}{dt} = KS(T_e - T)$

$K \rightarrow$ proportional constant, $S \rightarrow$ surface area

7. $\frac{dT}{dt} = \frac{KS}{mc} (T_e - T)$ $\frac{du}{dt}, \frac{dT}{dt} < 0$ Since $T_e < T$.
(loss) (P.T.O.)

8/. $m = \rho V$ $\rho \rightarrow$ Density, $V \rightarrow$ Volume.

9/. $\frac{dT}{dt} = \left(\frac{K}{\rho c}\right) \frac{S}{V} (T_e - T) \Rightarrow \frac{dT}{dt} \propto \frac{S}{V}$

For bodies of similar shape, small bodies ~~cool~~ cool faster than large ones.

Example: A sphere of radius r . $S = 4\pi r^2$

$V = \frac{4}{3}\pi r^3 \Rightarrow \frac{S}{V} \propto \frac{1}{r} \Rightarrow \frac{dT}{dt} \propto \frac{1}{r}$

10/. Integral Solution: $\frac{dT}{dt} = \frac{KS}{mc} (T_e - T)$ $\begin{matrix} T > T_e \\ \frac{dT}{dt} < 0 \end{matrix}$

$\Rightarrow \int \frac{dT}{T - T_e} = -\frac{KS}{mc} \int dt$ Write $r = \frac{KS}{mc}$

$\Rightarrow \ln(T - T_e) = -rt + A$ When $t = 0, T = T_0$

$\Rightarrow A = \ln(T_0 - T_e) \therefore \ln(T - T_e) = \ln(T_0 - T_e) - rt$

$\Rightarrow T - T_e = (T_0 - T_e) e^{-rt}$ i) When $t \rightarrow \infty$
 $T \rightarrow T_e$

$\Rightarrow T = T_e + (T_0 - T_e) e^{-rt}$ ii) When $t = 0$,
 $T = T_0$.

Stefan - Boltzmann Law (in Radiation)

Radiation wattage: $W = \epsilon \sigma S (T^4 - T_e^4)$

$W \rightarrow$ Energy radiation/time (Joule/sec).

$\sigma \rightarrow$ Stefan's Constant, $\epsilon = 1$ (for a black body)
(P.T.O.) (Joule m⁻² s⁻¹ K⁻⁴) Kelvin

$$\therefore W = \frac{dU}{dt} = \epsilon \sigma S (T^4 - T_e^4) \quad \left[\begin{array}{l} \epsilon \leq 1 \\ (\epsilon = 1 \text{ for a black body}) \end{array} \right]$$

$T \rightarrow$ Temperature of the body.

$T_e \rightarrow$ Environment temperature.

$$T = T_e + \Delta T$$

$$\Rightarrow T^4 - T_e^4 = (T_e + \Delta T)^4 - T_e^4 = T_e^4 \left(1 + \frac{\Delta T}{T_e} \right)^4 - T_e^4$$

$$\text{If } \Delta T \ll T_e, \left(1 + \frac{\Delta T}{T_e} \right)^4 \approx 1 + \frac{4\Delta T}{T_e} \rightarrow \text{Binomial Expansion}$$

$$\Rightarrow T^4 - T_e^4 \approx T_e^4 \left(1 + \frac{4\Delta T}{T_e} \right) - T_e^4 \approx 4 T_e^3 \Delta T$$

$$\text{Now } \frac{dU}{dt} = mc \frac{dT}{dt} = \epsilon \sigma S (T^4 - T_e^4) \quad \left(\text{linear approximation} \right)$$

$$\Rightarrow \frac{dT}{dt} = \frac{\epsilon \sigma S}{mc} (T^4 - T_e^4) \approx \frac{\epsilon \sigma S}{mc} \cdot 4 T_e^3 (T - T_e)$$

$$\Rightarrow \frac{dT}{dt} \approx \left(\frac{4\epsilon \sigma T_e^3}{mc} \right) S (T - T_e) \quad \left[\begin{array}{l} \text{Now } T > T_e \\ \text{but } dT/dt < 0 \end{array} \right]$$

$$\therefore \frac{dT}{dt} \approx \left(\frac{4\epsilon \sigma T_e^3}{mc} \right) S \times \left(\overset{\uparrow}{T_e - T} \right) \quad \left(\text{Forcing the negative sign} \right)$$

$$\text{Comparing with } \frac{dT}{dt} = \frac{kS}{mc} (T_e - T) \quad \left(\text{ignoring sign} \right)$$

$$k \approx 4\epsilon \sigma T_e^3 \quad (\epsilon \text{ is dimensionless}) \quad (\epsilon \leq 1)$$

$$\Rightarrow \text{Unit of } k \text{ is } \frac{\text{joule} \times \text{Kelvin}^3}{\text{m}^2 \text{sec} \times \text{Kelvin}^4} \quad \left[\begin{array}{l} \text{In the S.I. System} \end{array} \right]$$

$$\Rightarrow \text{Unit of } k \rightarrow \frac{\text{joule}}{\text{m}^2 \times \text{sec} \times \text{Kelvin}} \equiv \boxed{\text{Watt m}^{-2} \text{K}^{-1}} \quad (K \rightarrow \text{Kelvin})$$

Newton's Law of Cooling: An Application

Problem: A dead body has been found in a room whose temperature is $\theta_e = 25^\circ\text{C}$. When discovered, the body temperature was $\theta_1 = 28^\circ\text{C}$. One hour later the body temperature ~~became~~ $\theta_2 = 27^\circ\text{C}$. Normal human body temperature is $\theta_0 = 37^\circ\text{C}$. How long ago did the person die? $t_2 - t_1 = \Delta t = 1 \text{ hour}$

Solution: $T(\text{Kelvin}) = \theta(\text{Celsius}) + 273$

$\Rightarrow \Delta T = \Delta \theta$. $\theta_1 = \theta_e + (\theta_0 - \theta_e) \exp(-\lambda t_1)$

and $\theta_2 = \theta_e + (\theta_0 - \theta_e) \exp(-\lambda t_2)$ $\lambda = \frac{kS}{mc}$

$\therefore \ln\left(\frac{\theta_1 - \theta_e}{\theta_0 - \theta_e}\right) = -\lambda t_1$ and $\ln\left(\frac{\theta_2 - \theta_e}{\theta_0 - \theta_e}\right) = -\lambda t_2$

$\Rightarrow \ln\left(\frac{\theta_1 - \theta_e}{\theta_0 - \theta_e}\right) - \ln\left(\frac{\theta_2 - \theta_e}{\theta_0 - \theta_e}\right) = -\lambda t_1 + \lambda t_1 + \lambda \Delta t$

$\Rightarrow \lambda \Delta t = \ln\left(\frac{28 - 25}{37 - 25}\right) - \ln\left(\frac{27 - 25}{37 - 25}\right) = \ln\left(\frac{3}{12}\right) - \ln\left(\frac{2}{12}\right)$

$\Rightarrow \lambda \Delta t = \ln\left(\frac{3}{12} \times \frac{12}{2}\right) = \ln\left(\frac{3}{2}\right) \Rightarrow \lambda = \frac{\ln(1.5)}{\Delta t}$

$\Rightarrow \lambda t_1 = \ln\left(\frac{\theta_0 - \theta_e}{\theta_1 - \theta_e}\right) \Rightarrow t_1 = \frac{1}{\lambda} \ln\left(\frac{12}{3}\right)$

$\Rightarrow t_1 = \frac{\Delta t \times 2 \ln 2}{\ln(3/2)} \Rightarrow t_1 = \frac{2 \ln 2}{\ln(1.5)} \times 1 \text{ hour}$

$\Rightarrow t_1 \approx 3.42 \text{ hours} = 3 \text{ hours } 25 \text{ minutes}$ Death took place 3.42 hours ago.

(Time of Death) \rightarrow