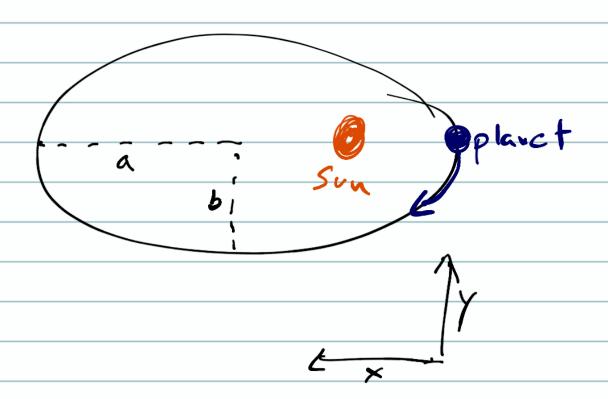
Root-finding examples

-> Kepler discovered that planets have elliptical orbits.



He found that the orbit could be described by these equations $x(t) = a(\cos(E) - e)$ $y(t) = a(1-e^{2} \sin E)$

where E=wt+esin(E)

-> Great! Except that the equation for E is implicit and not landytically solvable. So, in practice a root-finding method is used to solve this equation. 0=wt+esin(E)-E=f(E) $\frac{dF}{dE} = e\cos(E) - 1$ Pseudocode for Newton-Raphson Pick E guess las long as ell, should be stafe that F' +0) En = Eguess En=Egress error=Inf while errors 10-3

Calc f(En), dF En

En+ = En - ((En)) dF En

error=|En+1-En| Two other common uses for root-finding methods. 1. Backword Evler solves. 1 = f(xi) It we can't rearrange Mis x. ag(xi), Then we need a method to solve $x_{i+1} - f(x_{i+1}) + x_i = 0$

7. Finding The steady-state of systems of expled ODES

An example! Lotka-Volterra model

(aka Predator-Prey model)

Consider an ocosystem will two species

-> Species X always has enough food and

grows exponentially due to reproduction.

dx = ax

-> Species y ents species x and grows according to how much it ents and its current size: dx = 5 x y

-> Species x only dies by being speed caten by y: Bxy

-> Species y dies naturally due to its
own population size (i.e. how much the
ecosystem can support). By

The model for population sizes of species x and y:

$$\frac{dx}{dt} = \alpha x - \beta xy$$

$$\frac{dy}{dt} = \delta xy - \delta y$$

Strady states of population

$$\frac{dx}{dt} = 0 = \alpha \times - \beta \times \gamma$$

$$\frac{dx}{dt} = 0 = \delta \times \gamma - \delta \gamma$$

(We can solve)

$$0 = \alpha - \beta \gamma - 3 | \gamma = \alpha/\beta$$

$$0 = \delta \chi - \delta - 3 | \chi = \gamma/\delta$$

$$0 = \delta \chi - \delta - 3 | \chi = \gamma/\delta$$
Also,
$$|\chi = \gamma = 0$$

But, if this model was stightly more complicated dx = dx-Bxy-by 立と=Dxy-xy+2x Suddenly not so easy to solve f(x,y) = xx-Bxy-by fz(x,y)=5xy-8y+7x $J = \begin{bmatrix} \frac{\partial f_1}{\partial x} & \frac{\partial f_2}{\partial x} \\ \frac{\partial f_1}{\partial y} & \frac{\partial f_2}{\partial y} \end{bmatrix} = \begin{bmatrix} \frac{\partial f_1}{\partial x} & \frac{\partial f_2}{\partial y} \\ -\frac{\partial f_3}{\partial y} & \frac{\partial f_2}{\partial y} \end{bmatrix}$ occan from

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