## What has been changed

This version of Mcluster has been thought to work with MOCCA code. For this reason, some output functions and GPU and SSE/BSE version has been deleted in the make file. Moreover, the number of objects for each populatin as defined as  $n = n_{singles} + n_{binaries}$ , so that  $n_{star} = n_{singles} + 2 \cdot n_{binaries}$ . Finally, for the same reason, the optimal sampling IMF (Kroupa et al. (2011)) has been deleted.

**Constants** The solar radius in km value, the parsec in km value and the gravitational constant has been modified to correct values.

**Input variables** The input values for variables has been modified (see Readme.pdf document for better explanation).

Structure of main.c The main structure of Mcluster has been changed as:

- 1 Generation of stellar masses for all the populations;
- 2 Generation of positions and velocities for all populations;
- 3 Solve the Jeans equations in case of multiple stellar populations;
- 4 Scaling the system;
- 5 Generating binaries properties;
- 6 Output

Star\_array will be constructed as: Nbinaries\_1\_pop, Nsingle\_1\_pop, Nbinaries\_2\_pop, Nsingle\_2\_pop, ...

What is new In this version of Mcluster it has been added:

- Multiple stellar population;
- Energy calculation supposing spherical symmetry;
- Possibility to apply semi-major axis uniform distribution in log(a) for low mass stars and Sana et al., (2012); Oh, S., Kroupa, P., & Pflamm-Altenburg, J. (2015) period distribution for high mass stars;
- new eigenevolution and feeding algorithm Kroupa (2013)

**Bugs** In the case of presence of binaries, the maximum mass allowed for single is egual to msort value. This will lead to a wrong single IMF distribution (in particular for small binary fraction).