Tutorial for the sdhdfProc software package

George Hobbs

This document can be edited at

https://docs.google.com/document/d/1jA5lOn5MlltnOrchrggwLSnBtKRPJVUL2F984NmVRKk/edit?usp=sharing

Introduction

Use case: Are there any bright spectral lines in my data?

Use case: Having a play with a scan across a source

Introduction

All the data files described here are available from:

/DATA/BRAHE_2/hob044/sdhdf_tutorial/

In each sub-directory the list of commands for the specific use-case is available in a text file called "commands".

Note that there are various ways to use the sdhdfProc software:

- Obtain the software from https://bitbucket.csiro.au/scm/cpda/sdhdf tools.git and use the makefile to install. In this case the software will be called as described in this tutorial e.g., to plot a spectrum use sdhdf plotSpectrum.
- Use the pre-installed software on the CSIRO machines. Type:
 - sdhdf to get usage details.
 - Run sdhdf plotSpectrum (note space instead of the underscore) to run the sdhdf plotSpectrum code.
- To use George's version of the code (which may be changed at any time), use e.g., /u/hob044/software/new_c/gitRepos/sdhdf_tools/c/sdhdfProc/src//sdhdf_plotSpectrum to plot the spectrum. Note that this code has been compiled under "buster"

In all cases the \$SDHDF environment variable needs to be set an pointing to the run time directory (e.g.,

```
setenv SDHDF
/u/hob044/software/new_c/gitRepos/sdhdf_tools/c/sdhdfProc/runtime/
)
```

An easy way to inspect an sdhdf file is to use "silx view"

```
$ silx view <filename.hdf>
```

The primary sdhdfProc documentation is available from https://docs.google.com/document/d/1PhVhzOmVZ7iqA97LxRHF0gP5OHUI4Mw0Hp3KVjU9yIA/edit?usp=sharing

Use case: Are there any bright spectral lines in my data?

We make use of observations of a region in the Orion nebula carried out as P595 by Macquarie University students (supervised by Jo Dawson). The files are in /DATA/BRAHE_2/hob044/sdhdf_tutorial/orion. We can first identify the ON and OFF source pointings using sdhdf describe:

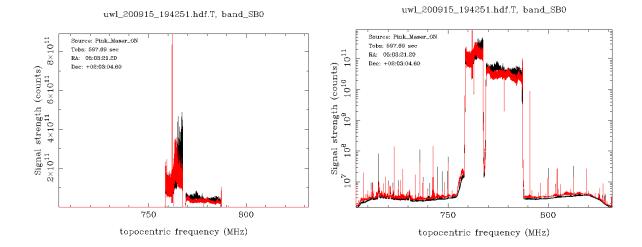
The first file is "on" source. The second file is "off" source. The files are large (>6GB in volume) and so let's start by average the data in time:

\$ sdhdf_modify -T -e T *.hdf

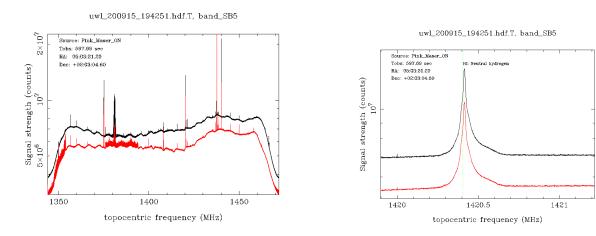
This produces uwl_200915_194251.hdf.T and uwl_200915_212041.hdf.T. We can now plot the "on" spectrum:

\$ sdhdf_plotSpectrum -f uwl_200915_194251.hdf.T

The output spectrum is as follows (left panel). Note that we're viewing the first sub-band (which has a lot of radio interference) on a linear scale. If we press 'l' then we can switch to a logarithmic y-scale (see right panel). The black and red traces represent AA and BB polarisation channels respectively.

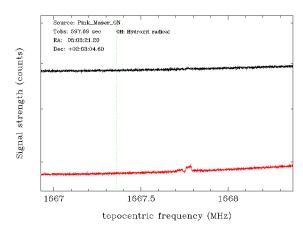


This sub-band is hard to work in, so let's use the '>' symbol to move through the sub-bands until we reach sub-band 5 (left panel below). We can zoom in around the HI line by moving the mouse cursor slightly to the left of 1420 MHz, pressing 'z', then moving the mouse to the right and clicking the mouse cursor (right panel). If we press 'm' we can plot the position of the rest frequency of the spectral line (as a vertical, green dashed line).



If we zoom-in around the OH line at 1667 MHz we see a signal that is of interest:

uwl_200915_194251.hdf.T, band_SB7

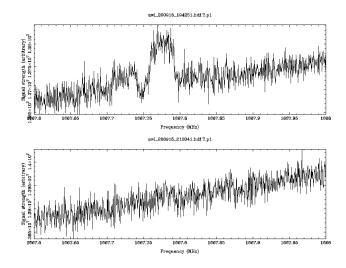


We can sum the polarisations using

\$ sdhdf_modify -p1 -e p1 *.T

And then determine whether the signal is, or is not, in the off source pointing:

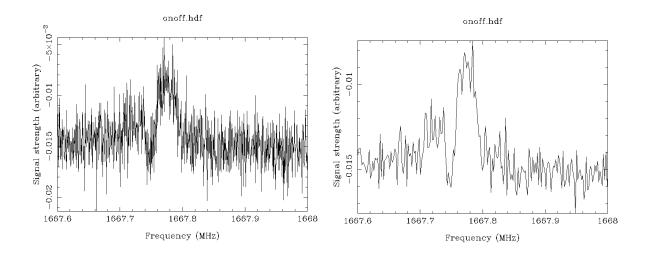
\$ sdhdf_plotMultiSpec -sb 7 -f0 1667.6 -f1 1668 -ny 2 uwl_200915_194251.hdf.T.p1 uwl_200915_212041.hdf.T.p1



This seems convincing. We can form the quotient [(ON-OFF)/OFF] using:

 $sdhdf_onoff-on\ uwl_200915_194251.hdf.T.p1-off\ uwl_200915_212041.hdf.T.p1-o\ onoff.hdf\\ sdhdf_plotMultiSpec-sb\ 7-f0\ 1667.6-f1\ 1668\ onoff.hdf$

The resulting profile is shown in the left-hand plot below. We can include the command line argument -av 4 when using sdhdf_plotMultiSpec to average 4 adjacent frequency channels together (right-hand panel):



Note that in this simple example we have not carried out many standard processing tasks including calibration, converting to the local standard of rest (and velocity), RFI flagging etc.

Use case: Having a play with a scan across a source

Request: The user wishes to explore the SDHDF format and software using a scan across a bright continuum source. We will use a P737 scan across PKS 0407-658.

Note: This scan was a great circle scan. Dhagu allows different types of scans.

Let us first check the basic properties of this file:

```
$ sdhdf_describe uwl_200720_024617.hdf
```

This gives basic information:

File Observer	RCVR	Bands	POL_TYPE	Beam CAL		SDHDF PID DEC_s	Sched_ID	Source	Tel		
uwl_200720	0_024617 UWL	.hdf 26	AABBCRCI	0 ON	2020-07-20-02:46:17 116.9 04:03:50.81		14710	0407scan_ROW5	Parkes		

We can see that the source name is 0407scan_ROW5, the project code (P737), that 4 polarisation products were recorded (AABBCRCI) and that the calibrator was ON. The integration time was 116.9 seconds.

In order to find out the sub-bands stored in the file we use the -band option:

```
$ sdhdf describe -band uwl 200720 024617.hdf
```

This gives:

	#	BandID	FO	F1	NCHAN	TDUMP	NPOL	NDUMP	TOBS	
[Pand]	000	band SB0	704.00	832.00	22760	0.999	4	117	116.933	uwl 200720 024617.hdf
		_		960.00		0.999				
		· · · · · — ·	832.00				4	117	116.933	uwl_200720_024617.hdf
		band_SB2	960.00	1088.00		0.999	4	117	116.933	uwl_200720_024617.hdf
		band_SB3	1088.00	1216.00		0.999	4	117	116.933	uwl_200720_024617.hdf
		band_SB4	1216.00	1344.00		0.999	4	117	116.933	uwl_200720_024617.hdf
		band_SB5	1344.00	1472.00		0.999	4	117	116.933	uwl_200720_024617.hdf
[Band]	006	band_SB6	1472.00	1600.00	32768	0.999	4	117	116.933	uwl_200720_024617.hdf
[Band]	007	band_SB7	1600.00	1728.00	32768	0.999	4	117	116.933	uwl_200720_024617.hdf
[Band]	008	band SB8	1728.00	1856.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	009	band SB9	1856.00	1984.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	010	band SB10	1984.00	2112.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	011	band SB11	2112.00	2240.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	012	band SB12	2240.00	2368.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	013	band SB13	2368.00	2496.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	014	band SB14	2496.00	2624.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	015	band SB15	2624.00	2752.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	016	band SB16	2752.00	2880.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	017	band SB17	2880.00	3008.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	018	band_SB18	3008.00	3136.00	32768	0.999	4	117	116.933	uwl_200720_024617.hdf
[Band]	019	band SB19	3136.00	3264.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	020	band SB20	3264.00	3392.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
[Band]	021	band SB21	3392.00	3520.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
		band SB22	3520.00	3648.00	32768	0.999	4	117	116.933	uwl 200720 024617.hdf
		band SB23	3648.00	3776.00		0.999	4	117	116.933	uwl 200720 024617.hdf
		_								

```
[Band] 024 band_SB24 3776.00 3904.00 32768 0.999 4 117 116.933 uwl_200720_024617.hdf
[Band] 025 band_SB25 3904.00 4032.00 32768 0.999 4 117 116.933 uwl_200720_024617.hdf
```

Showing that 26 sub-bands were recorded with the spectral dump time of 1 second.

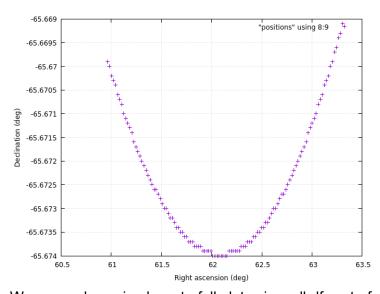
In particular for a scan we would like to know exactly where we were pointing at the centre of each spectral dump. We can obtain this with the -dump command line argument:

```
$ sdhdf describe -dump uwl 200720 024617.hdf
```

	ElapsedTime	MJD	UTC	AEST	RA	DEC	RADEG	DECDEG A	AZ	EL	GL	GB
[SDUMP]	0.500	59050.1154800	02:46:17.50	12:46:18.40	04:03:50.81	-65:40:11.5	60.962	-65.670 2	209.390	39.053	278.829	-41.327
[SDUMP]	1.499	59050.1154920	02:46:18.50	12:46:19.40	04:03:55.66	-65:40:12.0	60.982	-65.670 2	209.390	39.059	278.825	-41.319
[SDUMP]	2.499	59050.1155030	02:46:19.50	12:46:20.40	04:04:00.51	-65:40:12.5	61.002	-65.670 2	209.390	39.067	278.820	-41.312
[SDUMP]	3.498	59050.1155150	02:46:20.50	12:46:21.39	04:04:05.35	-65:40:13.0	61.022	-65.670 2	209.390	39.075	278.815	-41.304
[SDUMP]	4.497	59050.1155270	02:46:21.50	12:46:22.39	04:04:10.20	-65:40:13.5	61.042	-65.670 2	209.390	39.080	278.810	-41.297
[SDUMP]	5.497	59050.1155380	02:46:22.50	12:46:23.39	04:04:15.05	-65:40:14.0	61.063	-65.671	209.390	39.087	278.805	-41.289

We can store the positions in a text file using:

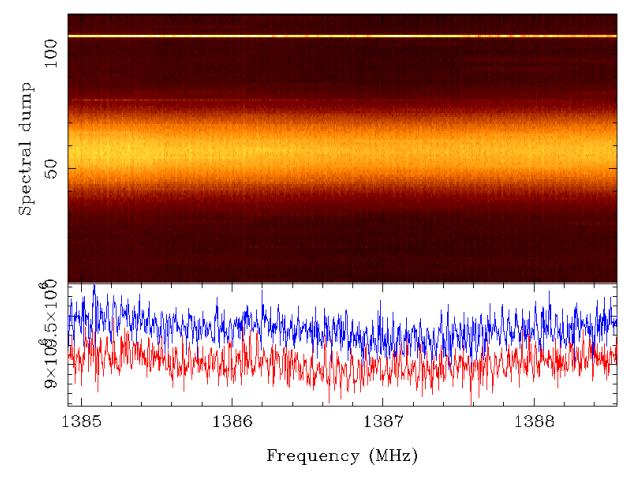
```
$ sdhdf describe -dump uwl 200720 024617.hdf | grep SDUMP > positions
```



We can make a simple waterfall plot using sdhdf_waterfall:

```
$ sdhdf_waterfall -band 5 uwl_200720_024617.hdf
```

Here we selected a relatively clean band (band 5) and then used 'x' to zoom-in on the x-axis to a clean part of that sub-band. The result is:

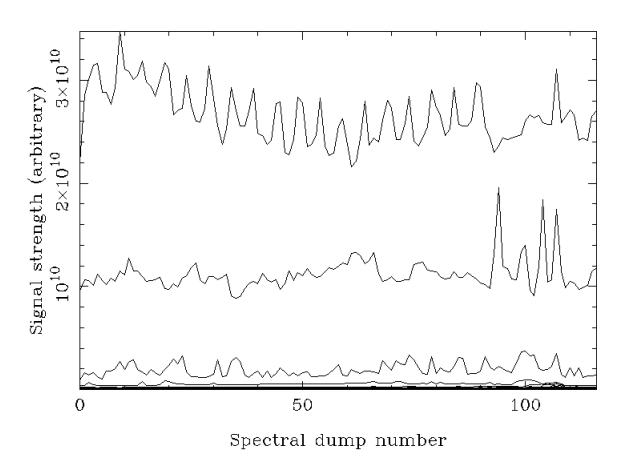


The top panel shows the uncalibrated signal strength as a function of frequency and time (colour scale). The bottom panel shows the min-hold and averaged spectrum averaged over these spectral dumps.

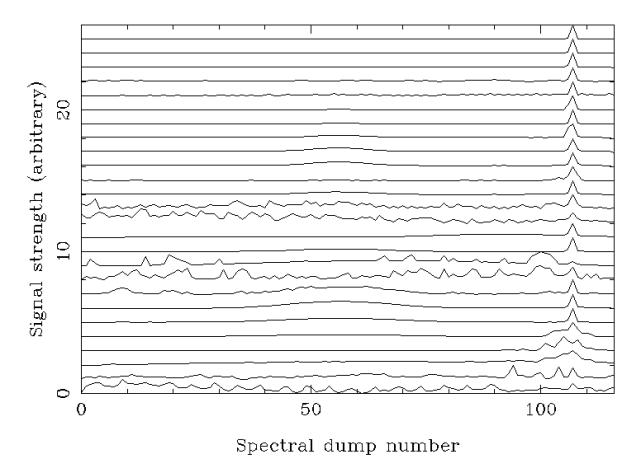
For a more detailed look at the scan we can use sdhdf_plotScan:

```
$ sdhdf plotScan -f uwl 200720 024617.hdf
```

The output is shown below. The x-axis is the spectral dump number and the y-axis is the signal strength. The data in each sub-band is averaged in frequency and then plotted. Clearly some sub-bands are at a much higher level than other bands.



Pressing 'n' allows us to normalise each plot and then pressing 'o' offsets each band from each other:



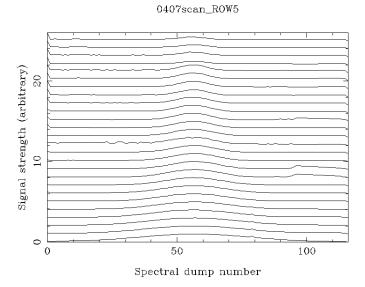
The "triangular" feature on the right hand side is broad-band RFI that lasted for one spectral dump. The slight "mound" in the middle is the continuum source. Pressing 'z' allows you to use the mouse to zoom in on a region and press 'u' to unzoom back to the original plot. Pressing 'x' toggles the x-axis from spectral dump number to various other parameters (e.g., right ascension, azimuth etc.)

Clearly we need to carry out some RFI flagging as we are averaging in frequency across wide bands that contain RFI and we have at least one spectral dump affected by strong broadband interference. Note that this observation was unfortunately affected by: mobile handsets, aircraft DME, satellites and WiFi/Bluetooth!

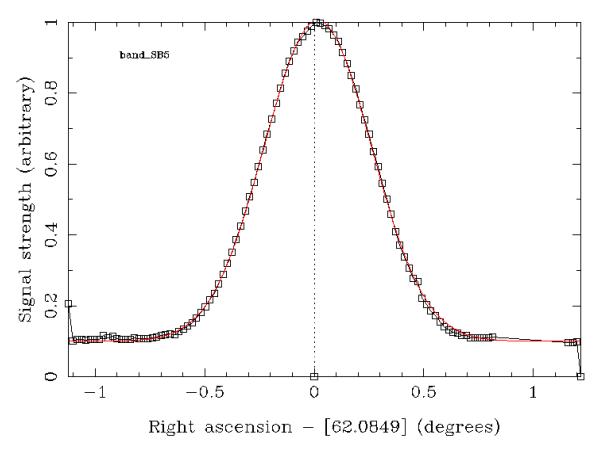
We can do this with sdhdf_flag (see later tutorial, or the documentation, on the use of this package)

After flagging we can plot the scan again (note that we're now using the -src option to specify the continuum source name):

\$ sdhdf_plotScan -src 0407-658 -f uwl_200720_024617.hdf.flag



We can use the '+' button to select a specific sub-band and then press 's' to select that sub-band:



Note that we've used 'x' to select a right ascension offset for the plot and 'f' to fit a Gaussian. The parameters for the fitted Gaussian (for each sub-band) are listed on the terminal.