## Radhyd Outline

## Initialization or Re-Initialization radhyd

# Program radhyd

## .....Initialization

#### • Call initialize

#### 1. Call model\_initialize

- (a) Assign unit 'nread' to Data3/reset.d
- (b) Assign unit 'nprint' to Data3/superdump.d
- (c) Assign unit 'nrstd1' to Data3/rstdmp1.d
- (d) Assign unit 'nrstd2' to Data3/rstdmp2.d
- (e) Call **genrd** 
  - i. Assign values to array dimensions nz, nez, nnu, and nnc
  - ii. Call **mgfld\_array\_initialize** Initialize the dimensions of the mgfld arrays
    - A. Call set\_abem\_arrays
    - B. Call set\_brem\_arrays
    - C. Call set\_e\_advct\_arrays
    - D. Call set\_eos\_snc\_arrays
    - E. Call set\_incrmnt\_arrays
    - F. Call set\_mdl\_cnfg\_arrays
    - G. Call set\_nu\_dist\_arrays
    - H. Call set\_pair\_arrays
    - I. Call set\_scat\_a\_arrays
    - J. Call set\_scat\_e\_arrays
    - K. Call set\_scat\_i\_arrays
    - L. Call set\_scat\_n\_arrays
    - M. Call set\_scat\_nn\_arrays
    - N. Call set\_t\_cntrl\_arrays
  - iii. Initialize the dimensions of the non-mgfld arrays
  - iv. Call set\_boundary\_arrays
  - v. Call set\_convect\_arrays
  - vi. Call set\_eos\_bck\_arrays
  - vii. Call set\_eos\_ls\_arrays
  - viii. Call set\_hydro\_arrays
  - ix. Call set\_mgfld\_remap\_arrays
  - x. Call set\_psi0p\_arrays
  - xi. Call set\_nucbrn\_arrays
  - xii. Call set\_shock\_arrays

```
xiii. Initialize q0
      xiv. Call read_init—Reads 'head' (problem description) and 'nrst' (cycle number)
      xy. If nrst = 0
          A. Call mgfld_var_initialize Initializes some of the mglfd variables
          B. Call init_var Initializes some of the non-mglfd variables
          C. Call mgfld_read—Reads in mglfd keys
             ....Open 'Data3/transport_keys.d'
             .....Call read_transport_keys
             .....Close 'Data3/transport_keys.d'
             .....Open 'Data3/edit_keys.d'
             .....Call read_edit_keys
             .....Close 'Data3/edit_keys.d'
         D. Call model_read—Reads in non-mglfd keys
             .....Open 'Data3/hydro_keys.d'
             .....Call read_hydro_keys
             .....Close 'Data3/hydro_keys.d'
             .....Open 'Data3/initial_model.d'
             .....Call read_initial_model
             .....Close 'Data3/initial_model.d'
      xvi. If nrst \neq 0
          A. Call reset_var—Initializes use
          B. Call readst—Reads in restart file from rstdmp1 or rstdmp2
          C. Call readst—Reads in changes from reset.d
   (f) Close unit 'nrstd1'
   (g) Close unit 'nrstd2'
   (h) Close unit 'nread'
   (i) Call genst_hy
        i. Sets quantities at inner edge of configuration
        ii. Compute rest masses of zones assuming Newtonian description
       iii. Loads equation of state tables
2. If \log r /= 0
   .....lagrangian = .true. if jj = 1 and kk = 1
  ELSE
  .....lagrangian = .false. END IF lagr \neq 0
3. Check that 1D arrays are large enough to accommodate all 3 sweeps
4. Set number of zones per PE
5. Call evh1_load
   (a) Set EVH1 globals
   (b) Set the geometry and boundary conditions
   (c) Get radial limits from MGFLD
   (d) IF rezn = ye
       .....IF lagrangian
       ......Call lagregrid
```

......Load radial grid into sweep arrays, offseting for ghosts

- ........Call coord\_bc—Set coordinates in ghost zones of sweep arrays
  .......Call volume—Calculate zone volumes
  ......Call paraset—Set up coefficients for parabola subroutine
  ......Set jm and jnumax to imax+1
  .....IF not lagrangian
  ......Call eulregrid
  ......Set jnumax & jm to maximum zone number
  .....END IF lagrangian
  .....Put state variables into 1D arrays, padding with 6 ghost zones
  .....Load EVH1 boundary conditions from MGFLD
- .....Call e\_compose—Compute the total energy e(n)
- .....Reload values changed by rezoning into MGFLD variables
- .....Build a j grid
- .....Build a k grid
- .....Set transverse velocities initially to zero

ELSE iF rezn = no

- .....Set imax to maximum zone number
- .....Call mgfld\_to\_evh1\_restart—Load evh1 arrays
- i. Load quantities from MGFLD to EVH1 arrays
- ii. Put radial grid and sweep arrays, offsetting for ghosts
- iii. Call **coord\_bc**—Set coordinates in ghost zones of sweep arrays
- iv. Call volume—Calculate zone volumes
- v. Call paraset—Set up coefficients for parabola subroutine
- vi. Load EVH1 boundary conditions from MGFLD
- vii. Put state variables into 1D arrays, offsetting for ghosts
- viii. Compute the total energy e(n)

END IF rezn

#### 6. Call mgfld\_setup

- (a) Call **genst** if nrst = 0, otherwise Call **genrst**
- (b) Call time\_step\_select
- (c) Call **mgfld\_reset**
- (d) Call **mgfld\_edit**

## 7. Call mgfld\_to\_evh1

- (a) Load arrays zte, zei, and zye
- 8. Call svel\_init

# ......Problem Cycling—Program radhyd

- Set nmin = 7, nmax = imax + 6
- Call cycle
  - 1. Update cycle number
  - 2. Print cycle number to 'Data3/cycle.d'

- Zero increment variables
- Call evh1\_evolve\_xy—Perform hydro step with x-sweeps preceding y-sweeps
  - 1. Compute and print the total energy
  - 2. Save pre-Lagrange step variables in \_i arrays for mgfld (xxx make these multi-D arrays)
  - 3. Call sweepx
    - (a) Set nmin = 7, nmax = imax + 6, ntot = imax + 12
    - (b) Loop over j, k (i.e., different (radial) rays)
    - (c) Put state variables for a given radial ray into 1D arrays, padding with 6 ghost zones
    - (d) Load grid coordinates in xa0, dxa0, xa, and dx
    - (e) Call **eos\_result** (Computes T from ei, and then p, s, and gamma)
    - (f) Call sweepbc
      - i. Call **coord\_bc** (Loads ghost coordinates with boundary coordinates)
      - ii. Load ghost zones with state variables
    - (g) Call **volume** (Computes volume elements)
    - (h) Call **paraset** (Updates parabolic coefficients with initial grid coordinates for later use in obtaining parabolic interpolants of flow variables inside each grid zone)
    - (i) Call **e\_compose** (Computes e(n) from ei(n), ekin(n), and egrav(n), which is needed for subroutine **evolve** to advance the total energy)
    - (j) Call **ppm** 
      - i. Call **flatten** (Calculate flattening coefficients for smoothing near shocks)
      - ii. Call **parabola** (Computes parabolic interpolants for flow variables in each grid zone)
      - iii. Call **states** (Integrate parabolae over causal domain to get input states for Riemann problem)
      - iv. Call **riemann** (Obtain the zone face averages, *umid* and *pmid*)
      - v. Call **evolve** (Lagrangian update is performed—mass conservation ( $\rho$  update), momentum conservation (u update; energy conservation (e update))
        - A. Grid positions are updated from umid and dt
        - B. Call **zone\_center** (Calculates volume averaged zone centers, which are used to calculate external forces)
        - C. Call **forces** (Calculate forces using zone-centered coordinates at t (0) and at t+dt (1))
        - D. Calculate dvolume and average area based on geometry of sweep
        - E. Update the density from the new zone positions
        - F. Update the velocity due to pressure gradients and forces
        - G. Update the energy due to net work performed on the zone surfaces by pressure and by the zone-centered forces times the zone-centered displacement
        - H. Call sweepbc (Grid change requires updated boundary conditions
          .....Call coord\_bc (Loads ghost coordinates with boundary coordinates)
          .....Load ghost zones with state variables
        - I. Call paraset (Grid change requires updated parabolic coefficients
        - J. Call **e\_decompose** (Extract the internal energy from the updated total energy)

- K. Call **etotal** (Total fluid energy check)
- L. Compute and store dei, the internal energy change due to the hydro step
- (k) sweepx  $\leftarrow$
- (1) Save Lagrangian updated variables (except T) in \_l arrays for mgfld
- (m) For a Lagrangian run, updated the coordinates to zxa, xdz, zxc
- 4.  $evh1_evolve_xy \leftarrow$
- 5. Call **sweepy** (This has not yet been interfaced with MGFLD)
- radhyd  $\leftarrow$
- Call **evh1\_to\_mgfld\_hydro** (Loads results of EVH1 Lagrangian hydro (roi, ye0i, t0i, rol, dei, u0l, x0l) to advance temperature and compute pseudoviscosities)
  - 1. Call hydro\_t\_change (Compute temperature change from dei and store in dtmpmn(j, 1))
  - 2. Call mgfld\_hydro
    - (a) Call **pseudo** Computes pseudoviscosities for editting purposes
    - (b) Call **snuc** Updates composition from nuclear reactions
    - (c) Call **nsetest** Flashes or deflashed zones, as appropriate
- Ramp up an explosion, if criteria are satisfied
- Call evh1\_to\_mgfld\_transport (Loads arrays \_\_i into arrays for transfer to MGFLD transport)

```
1. roi \rightarrow rhop

t0i \rightarrow tp

ye0i \rightarrow yep

u0i \rightarrow up

xai \rightarrow rp

psi0p \text{ from } \mathbf{psi0p_module}

dtnphn\_aetr \rightarrow dtime
```

2. Call mgfld\_transport\_in

```
(a) rhop \rightarrow rho

tp \rightarrow t

yep \rightarrow ye

up \rightarrow u

psi0p \rightarrow psi0

up \rightarrow u

compute dr

compute dmrst, rstmss\ rho \rightarrow rhoa

t \rightarrow ta

dr \rightarrow dra

r \rightarrow ra

u \rightarrow ua

dtime \rightarrow dtnphn\_aetr
```

- (b) Call **mgfld\_transport** (Do the neutrino transport)
  - i. Call agr\_nu\_cal (Transfer the lapse functions to transport variables (GR runs))

- ii. Call **enu\_cal** (Updates the neutrino bin energies on the basis of the new lapse functions (GR runs))
- iii. Call **gamgr\_nu\_call** (Transfers the GR gammas for transport modules (GR runs))
- iv. Call **extrap** (At the moment, a dummy subroutine)

```
v. r \rightarrow r\_nu

rho \rightarrow rho\_nu

t \rightarrow t\_nu

ye \rightarrow ye\_nu

ra \rightarrow ra\_nu

rhoa \rightarrow rhoa\_nu

ta \rightarrow ta\_nu

yea \rightarrow yea\_nu
```

- vi. Call **eddington** (Compute Eddington factors)
- vii. Call **comvcf\_cal** (Compute the stress-energy coupling (for editing purposes)
- viii. Call **w\_cal** (Update the relativistic enthalpy using *rhoa\_nu*, *ta\_nu*, and *ye\_nu* (GR runs))
- ix. Call **agr\_cal** (Update the lapse functions *rhoa\_nu*, *ta\_nu*, *ye\_nu*, and *ra\_nu* (GR runs))
- x. Call agr\_cal\_cal (Transfer the lapse functions to transport variables (GR runs))
- xi. Call **gamgra\_nu\_call** (Transfers the GR gammas for transport modules (GR runs))
- xii. Call **enua\_cal** (Updates the neutrino bin energies on the basis of the new lapse functions (GR runs))
- xiii. Transfer updated neutrino energies, lapse functions, and GR gammas into original arrays

```
unuea 
ightharpoonup unue
dunuea 
ightharpoonup dunue
unubea 
ightharpoonup unu
dunua 
ightharpoonup dunu
unuba 
ightharpoonup unub
ncoefaa 
ightharpoonup ncoefa
ecoefaa 
ightharpoonup ecoefa
ecoefaea 
ightharpoonup ecoefae
agra_n u 
ightharpoonup agr_n u
agrajmh_n u 
ightharpoonup agrg_n u
```

- xiv. Call eqstz (Update thermodynamic quantities for transport)
- xv. Call **nu\_adv** Perform the source and transport step
  - A. Transfer variables from calling statement to transport module

```
r\_in \rightarrow ra\_nu

rho\_in \rightarrow rhoa\_nu

t\_in \rightarrow ta\_nu

ye\_in \rightarrow yea\_nu
```

- B. Call **pre\_trans** (Compute area, vol, dr, drjmh, c\_r, c\_e)
- C. Call **nu\_abemtr** (Advance neutrino occupation probabilities due to emission, absorption and transport)

```
.....Initialize increment variables dye_emabtr, dt_emabtr, dt_emabtrk, dye_emabtrk,
dpsi_e mabtr
....ta_nu \rightarrow t_nu_0 \quad yea_nu \rightarrow ye_nu_0
.....Call dtau_aetr (Get time step for transport)
....psi0 \rightarrow psi0_0
.....psi0\_0 \rightarrow psi0\_i
....psi1 \rightarrow psi1\_i
....t_nu_0 \rightarrow ta_nu
....t_nu_0 \rightarrow t_nu_i
\dots ye_nu_0 \rightarrow yea_nu
\dots ye_nu_0 \rightarrow ye_nu_i
.....Iterate
\dots t_n u_i \to ta_n u
\dots ye_nu_i \rightarrow yea_nu
..........Call c_psi_set (Update thermodynamic quantities for transport)
......Update absorption and emission scattering rates
.....Update isoenergetic scattering rates
......Compute inverse mean free paths
......Compute diffusion coefficients
.....Compute psi1_i
......Compute d_ye, d_ye_t, d_ye_ye, d_ye_psi, d_t, d_t, d_t, d_t, d_t
.......Call pre_c_psi_set (Implement switches iyenu, itnu)
.........Call a_psi_set (Compute recursion coefficients)
.........Call psi_bd (Implement boundary conditions)
.........Call d_sub (Compute dpsi0\_iph(j), this iterations increment of pis0)
.....psi0_{-i} + dpsi0_{-i}ph \rightarrow psi0_{-i}p1
.....dpsi\_emabtr + dpsi0\_iph \rightarrow dpsi_emabtr
.....t_nu_i + cf_t + cf_t - t_psi(j,k) \times dpsi0_iph \rightarrow t_nu_i
.....ye\_nu\_i + cf\_ye + cf\_ye\_psi \times dpsi0\_iph \rightarrow ye\_nu\_i
.....dt\_emabtrk + cf\_t + cf\_t\_psi \times dpsi0\_iph \rightarrow dt\_emabtrk
......dye\_emabtrk + cf\_ye + cf\_ye\_psi \times dpsi0\_iph \rightarrow dye\_emabtrk
.....Test for convergence
.....End iteration
....t_nu_0 \rightarrow ta_nu (Restore initial value to ta_nu
....ye_nu_0 \rightarrow yea_nu (Restore initial value to yea_nu
.....dt_emabtr + dt_emabtrk \rightarrow dt_emabtr (Add changes to t due to all energy
zones)
....dye\_emabtr + dye\_emabtrk \rightarrow dye\_emabtr (Add changes to ye due to all
energy zones)
....psi0 + dpsi\_emabtr \rightarrow psi0 (Update psi0)
.....Compute psi1
....dc \rightarrow dcr
```

- D.  $dye\_emabtr \rightarrow dye$
- E.  $dt_-emabtr \rightarrow dtmpnn$

F. Call nu\_scat (Advance neutrino occupation probabilities due to inelastic scattering and pair production)

```
xvi. \mathbf{mgfld\_transport} \leftarrow
xvii. Call t_adv (Update the temperatures)
xviii. Call ye_adv (Update the electron fractions)
 xix. Switch arrays
       r \rightarrow rr
       ra \rightarrow r
       ra\_nu \rightarrow r\_nu
       etc.
 xx. Call eqstz (Recompute thermodynamic functions)
```

- xxi. Call gammaz (Recompute thermodynamic gammas)
- xxii. Call **nu\_stress** (Compute neutrino stresses)
- (c)  $mgfld_transport_in \leftarrow$
- (d) Call **mgfld\_reset** (Reset thermodynamic and rate tables)
  - i. Call **eqstt** (Compute the internal energy prior to resetting tables)
  - ii. Call esrgnz (Reset EOS tables)
  - iii. Call eqstz (Recompute thermodynamic quantities)
  - iv. Call **gammaz** (Recompute thermodynamic gammas)
  - v. Call abemset, bremset, scataset, scatiset, scatinset, scatnset, pairset (Recompute rate tables)
  - vi. Call **nucset** (Recompute nuclear reaction rate tables)
- (e)  $mgfld_transport_in \leftarrow$
- $3. \ evh1\_to\_mgfld\_transport \leftarrow$
- radhyd  $\leftarrow$
- Call mgfld\_transport\_to\_evh1
  - 1. Call mgfld\_transport\_out (Load results of MGFLD transport into arrays for export)

```
(a) ta \rightarrow tp
     yea \rightarrow yep
     aesv \rightarrow ep
     psi0 \rightarrow psi0p
     stress \rightarrow nu\_stress (Combine stresses of different neutrino flavors)
```

2.  $mgfld_transport_to_evh1 \leftarrow$ 

```
3. tp \rightarrow zte
    ep \rightarrow zei
    yep \rightarrow zye
    nu\_stress \rightarrow znu\_str
```

- radhyd  $\leftarrow$
- Call evh1\_to\_mgfld\_e\_advct (Load variables for export to energy advection)

```
1. r0i \rightarrow rhop

r0l \rightarrow rhoap

t0i \rightarrow tp

zte \rightarrow tap

ye0i \rightarrow yep

zye \rightarrow yeap

u0i \rightarrow up

xai \rightarrow rp

xal \rightarrow rap

dtnphn\_aetr \rightarrow dtime
```

2. Call **mgfld\_nu\_energy\_advct\_in** (Export variables to neutrino energy advection modules)

```
(a) rhop \rightarrow rho

rhoap \rightarrow rhoa

tp \rightarrow t

tap \rightarrow ta

yep \rightarrow ye

yeap \rightarrow yea

up \rightarrow u

psi0p \rightarrow psi0

rp \rightarrow r

rap \rightarrow ra

Compute dmrst and rstmss

ua \rightarrow u

dtime \rightarrow dtj
```

(b) Call **nu\_energy\_advct** (Perform the energy advection step)

```
i. r\_in \rightarrow r\_nu

ra\_in \rightarrow ra\_nu

rho\_in \rightarrow rho\_nu

rhoa\_in \rightarrow rhoa\_nu

t\_in \rightarrow t\_nu

ta\_in \rightarrow ta\_nu

ye\_in \rightarrow ye\_nu

yea\_in \rightarrow yea\_nu
```

- ii. Call eddington (Compute Eddington factors)
- iii. Call **comvcf\_cal** (Compute the stress-energy coupling (for editing purposes)
- iv. Call  $\mathbf{w}$ \_cal (Update the relativistic enthalpy using rho\_nu, t\_nu, and ye\_nu (GR runs))
- v. Call  $\operatorname{agr\_cal}$  (Update the lapse functions  $\operatorname{rho\_nu}$ ,  $\operatorname{t\_nu}$ ,  $\operatorname{ye\_nu}$ , and  $\operatorname{r\_nu}$  (GR runs))
- vi. Call agr\_cal\_cal (Transfer the lapse functions to transport variables (GR runs))
- vii. Call enu\_cal (Compute the neutrino bin energies))
- viii. Call **nu\_U** (Compute the initial neutrino energy per unit mass)
- ix. Call  $\mathbf{w}$ \_cal (Update the relativistic enthalpy using rhoa\_nu, ta\_nu, and ye\_nu (GR runs))

- x. Call **agr\_cal** (Update the lapse functions  $rhoa\_nu$ ,  $ta\_nu$ ,  $ye\_nu$ , and  $ra\_nu$  (GR runs))
- xi. Call agra\_cal\_cal (Transfer the lapse functions to transport variables (GR runs))
- xii. Call enua\_cal (Compute the neutrino bin energies)
- xiii. Call e\_advct (Perform the neutrino energy advection)
- xiv.  $psi0\_a \rightarrow \psi 0$  (Restore updated neutrino distribution to psi0 array)
- xv. Call **rebal** (Prevent overfilling of neutrinos states)
- xvi. Call **nu\_Ua** (Compute the final neutrino energy per unit mass)
- xvii. Call **nu\_stress** (Update the neutrino stresses)
- 3.  $mgfld_nu_energy_advct_in \leftarrow$
- $evh1_to_mgfld_e_advct \leftarrow$
- Call mgfld\_e\_advct\_to\_evh1
  - 1. Call mgfld\_nu\_energy\_advct\_out
    - (a)  $nu\_stress \rightarrow znu\_str$
    - (b)  $psi0p \rightarrow psi0p\_module$
- $\bullet \ evh1\_to\_mgfld\_e\_advct \leftarrow$
- Call mgfld\_to\_evh1 (Load undated T, ye, and ei for EVH1)
- Loop over y and z and put x variables in 1D arrays, padding with 6 ghost zones
- Call **pre\_remap\_psi** (Prepare for *psi*0 remap)
  - 1. Divide psi0's by Larangian updated rho and store in array  $psi0_re$ , padding with 6 ghost zones)
  - 2. Put initial (Lagrangian updated) and final (Eulerian) grids, rho's and psi0's into arrays, padding with 6 ghost zones)
  - 3. Call **coordbc\_psi** (Find grid coordinates of ghost zones)
  - 4. Call **CALL sweepbc\_r** (Fill *rho* ghost zones with boundary values)
- Call remap\_psi
  - 1. Call **paraset** (Updates parabolic coefficients for later use with Lagrangian grid coordinates in obtaining parabolic interpolants of flow variables inside each grid zone)\*\*\*\*\*
  - 2. Call **parabola** (Computes parabolic interpolants for *rho* and *psi0* in each grid zone)
  - 3. Calculate the volume of the overlapping subshells
  - 4. Integrate over the parabolic profiles to calculate the total mass and neutrino number in the overlapping subshells
  - 5. Calculate the volumes before and after remap
  - 6. Advect psi0 by moving the subshell quantities into the appropriate Eulerian zone
  - 7. Restore psi0 by multiplying remapped  $psi0_re$  by remapped rho

- Call **pre\_remap\_comp** (Prepare for composition remap)
  - 1. Put initial (Lagrangian updated) and final (Eulerian) grids, *rho*'s into arrays, padding with 6 ghost zones)
  - 2. Find use nonnse boundary
  - 3. Put initial (Lagrangian updated) xn's into array comp, padding with 6 ghost zones)
  - 4. Call **coordbc\_psi** Find grid coordinates of ghost zones
  - 5. Call **sweepbc\_r** (Fill *rho* ghost zones with boundary values) \*\*\*\*\*
  - 6. Fill left and right ghosts with boundary rho's and comp's

## • Call **remap\_comp** (Remap composition)

- 1. Call **paraset** (Updates parabolic coefficients for later use with Lagrangian grid coordinates in obtaining parabolic interpolants of flow variables inside each grid zone)\*\*\*\*\*
- 2. Call **parabola** (Computes parabolic interpolants for *rho* and *comp* in each grid zone)
- 3. Calculate the volume of the overlapping subshells
- 4. Integrate over the parabolic profiles to calculate the total mass and mass of each specie in the overlapping subshells
- 5. Calculate the volumes before and after remap
- 6. Advect comp by moving the subshell quantities into the appropriate Eulerian zone
- 7. Restore comp to xn
- 1. Add six ghost zones to psi0 array, and divide by the updated density
- 2. Find ghost coordinates

#### • Call **remap** (Remap flow variables)

- 1. Call **paraset** (Updates parabolic coefficients for later use in obtaining parabolic interpolants of flow variables inside each grid zone)\*\*\*\*\*
- 2. Call parabola (Computes parabolic interpolants for flow variables in each grid zone)
- 3. Call **e\_compose** (Computes e(n) from ei(n), ekin(n), and egrav(n), which is needed remap the total energy)
- 4. Call **parabola** (Computes parabolic interpolants for *e* variables in each grid zone)
- 5. Calculate the volume of the overlapping subshells
- 6. Calculate the total mass (fluxr), etc. in overlap
- 7. Advect mass, etc. by moving the subshell quantities into the appropriate Eulerian zone
- 8. Reload Eulerian grid coordinates
- 9. Call **paraset** (Updates parabolic coefficients for later use with Eulerian grid coordinates in obtaining parabolic interpolants of flow variables inside each grid zone)
- 10. Call **e\_decompose** (Extract the internal energy from the total energy)
- Call eos\_result (Computes T from ei, and then p, s, and gamma)
- $\bullet$  Fill in multi-D arrays, zro, etc.

- Put updated values into zone arrays (\_\_e) for **mgfld\_edit**
- Call  $evh1\_to\_mgfldseted$  (Transfer remapped variables to mgfld variables (rho, t, ye, etc.) for edit)
- Call **mgfld\_edit** (Edit)
- Call **mgfld\_terminate** (Examine termination criteria)
- Call time\_step\_select (Compute timestep for next cycle)
- Repeat, but call sweepy before sweepx