





Operational research for urban solar development

"PV failure detection based on operational time series"



07/12/2023 Alexandre Mathieu



Agenda



Curriculum

PV performance model steps



Curriculum Plan

Today —

Project: groups of 2 for the project

Day	Time	Duration	Content
Monday	9h45-11h15	1h30 + 1h30	50% Lecture / 50 %
27/11/2023	12h30-14h		Hands-on
Tuesday	8h-9h30	1h30 + 1h30	50% Lecture / 50 %
05/12/2023	9h45-11h15		Hands-on
Thursday	8h-11h	6h	25% Lecture / 75 %
07/12/2023	12h45-15h45		Project
Monday	8h-11h	6h	10% Lecture / 90 %
11/11/2023	12h30-15h30		Project
Friday 22/12/2023	8h-9h30	1h30	100 % Project



Agenda

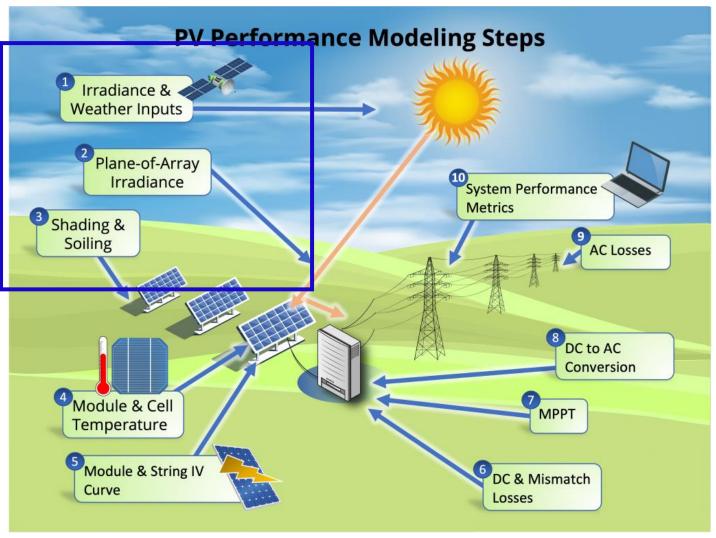


Curriculum

PV performance model steps



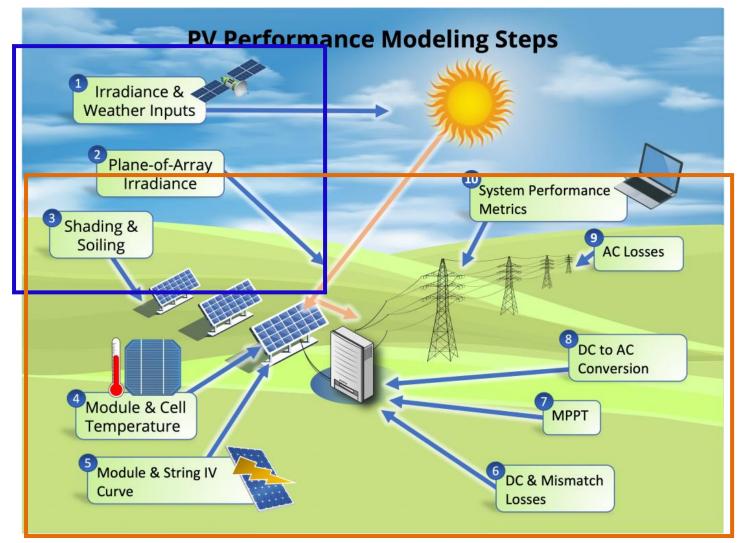
27/11/2023 & 05/12/2023





27/11/2023 & 05/12/2023

Today





Notebook recap 05/12/2023

The notebook is now corrected and can be read online:

https://github.com/AlexandreHugoMathieu/pvfault_detection_solar_academy/blob/master/notebooks/python_intro2_horizon_mask.ipynb



Notebook recap 05/12/2023

Python commands

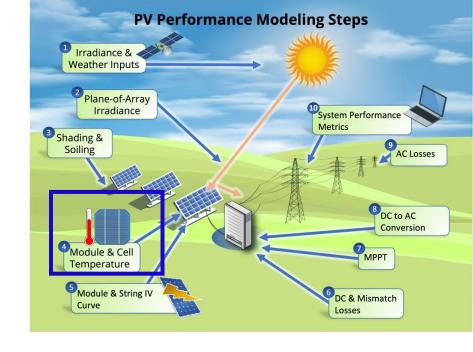
```
# Apply filters
filter_1 = weather_data["ghi"] > 800
filter_2 = (weather_data.index > pd.to_datetime("20220701").tz_localize("CET"))
filter = filter_1 & filter_2 # Combine with a "and" condition
weather data.loc[filter, "dhi"]
# If statement
a=1
if a<0: # assertion: is "a" under 0? Do not forget the ":" at the end of the line
              print("a is lower than 0") # Line non-executed since the assertion above is wrong, do not forget the "tab"
indendation after a "if"
# Loop over all elements of a list or pd. Series which allow to perform task on each of the element
For element in ["a","b"]:
              print(element)
for index, row in df.iterrows(): # Loop over all rows and index of the dataframe one by one
              print(row["column1"] + row["column2"])
# Plot with matplotlib
plt.plot(x, y, linewidth=0, marker="o") # scatter plot with no line in that case
# Function, useful to store few lines of code you want to reuse and apply with different inputs
def my func name(argument1, argument2): # Define the function "" with the "def" command and a small increment
tab to the right
              y = argument1 + argument2
              return y
my_func_name(1,2) # Apply the function with two arguments and return 3
my func name(1,3) # Return 4
```

import matplotlib.pyplot as plt



4. Module and Cell temperature

The hotter a module is, the less efficient it is!





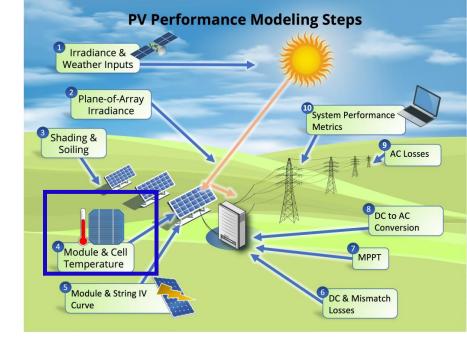
4. Cell temperature

Ross model:

Model to estimate the cell temperature T_c [°C] as function of ambient temperature and irradiance G_{POA} [W/m²].

$$T_c = T_a + G_{POA} \cdot k_{Ross}$$

 k_{ROSS} , typically in the range 0.02-0.05 K/m²/W.





 $k_{\it Ross}$ can be fitted from datasheet values. NOCT conditions:

 $G_{POA} = 800 \text{ W/m}^2$

 $T_a = 20^{\circ} \text{C}$

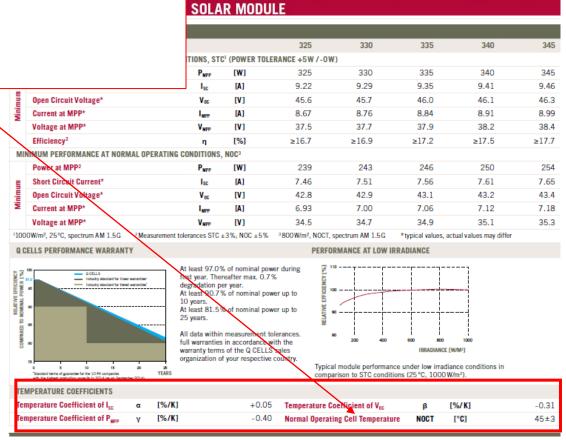
4. Cell temperature

Ross model:

Model to estimate the cell temperature T_c [°C] as function of ambient temperature and irradiance G_{POA} [W/m²].

$$T_c = T_a + G_{POA} \cdot k_{Ross}$$

 k_{ROSS} , typically in the range 0.02-0.05 K/m²/W.



L-G5 325-345



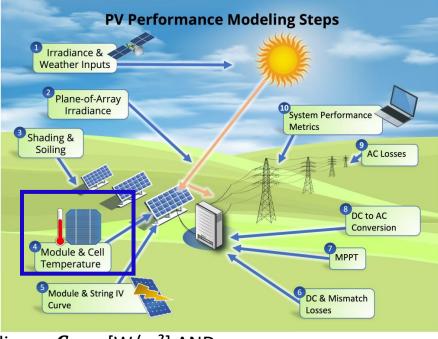
4. Cell temperature

Faiman model:

Model to estimate the cell temperature T_c [°C] as function of ambient temperature and irradiance G_{POA} [W/m²] AND wind WS $\left[\frac{m}{s}\right]$.

$$T_m = T_a + \frac{G_{POA}}{U_0 + U_1 \cdot WS}$$

 U_0 is the constant heat transfer component $[\frac{W}{Km^2}]$ U_1 is the convective heat transfer component $[\frac{W}{Km^2}]$





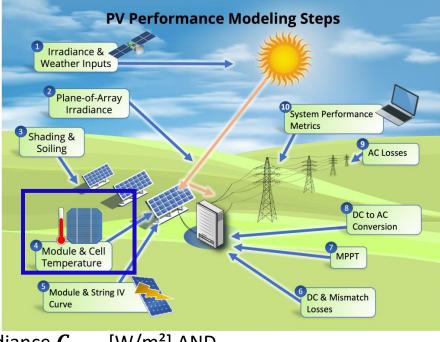
4. Cell temperature

Faiman model:

Model to estimate the cell temperature T_c [°C] as function of ambient temperature and irradiance G_{POA} [W/m²] AND wind WS $\left[\frac{m}{s}\right]$.

$$T_m = T_a + \frac{G_{POA}}{U_0 + U_1 \cdot WS}$$

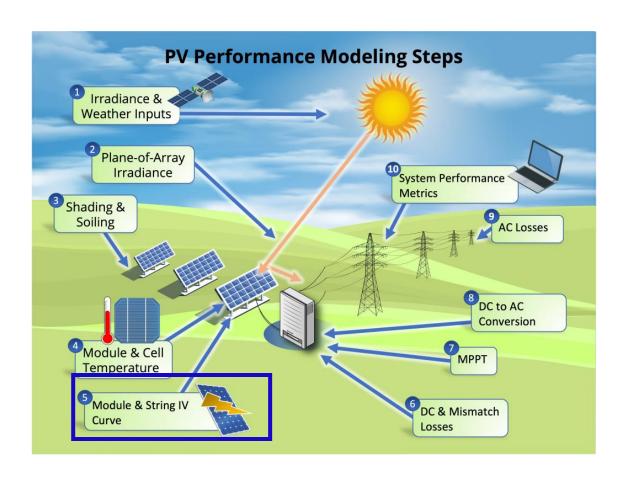
 U_0 is the constant heat transfer component $[\frac{W}{Km^2}]$ U_1 is the convective heat transfer component $[\frac{W}{Km^2}]$



In some cases, $T_c \simeq T_m$ can be assumed Between T_c and T_m , only few degrees of difference



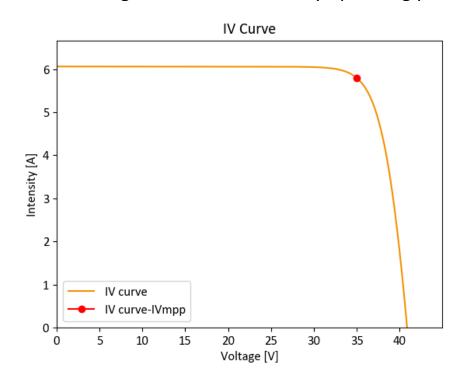
5. Module and String IV Curve

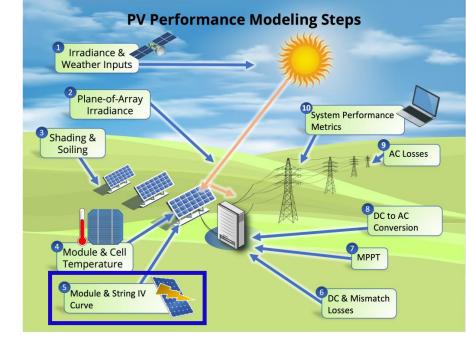




5. Module and String IV Curve

For a fixed irradiance and module temperature, the PV module has its I, current which depends on V, voltage and it can take many operating points.

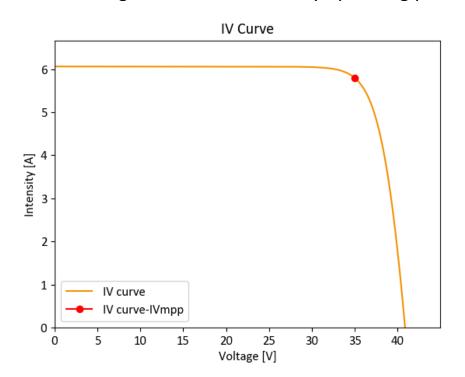




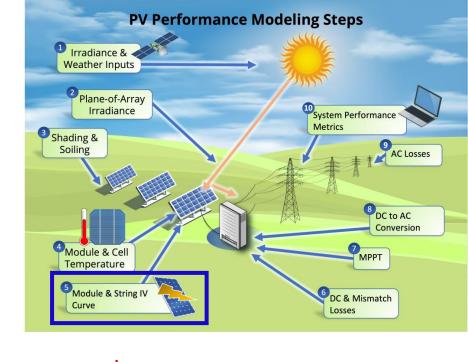


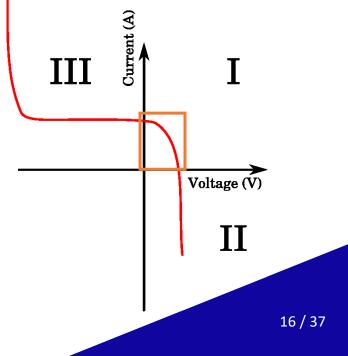
5. Module and String IV Curve

For a fixed irradiance and module temperature, the PV module has its I, current which depends on V, voltage and it can take many operating points.



In reality, the IV characteristics go out of the 1st quadrant and the module can potentially consume power.

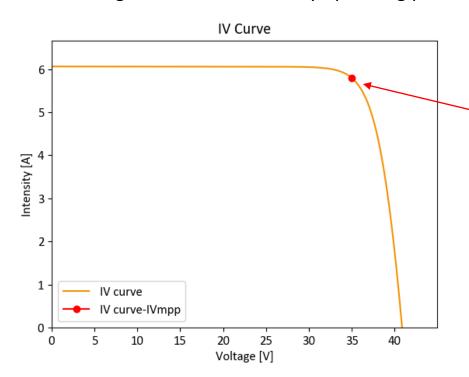


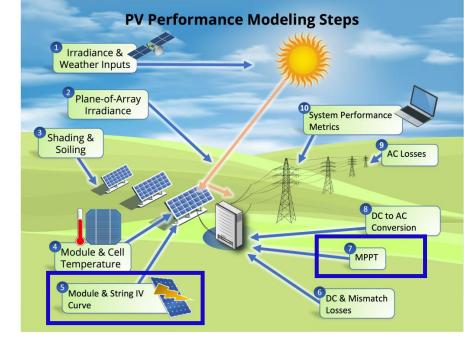




5. Module and String IV Curve

For a fixed irradiance and module temperature, the PV module has its I, current which depends on V, voltage and it can take many operating points.



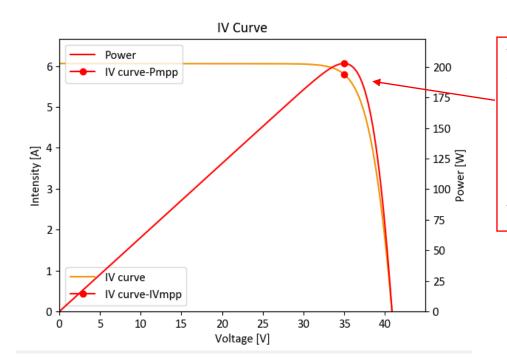


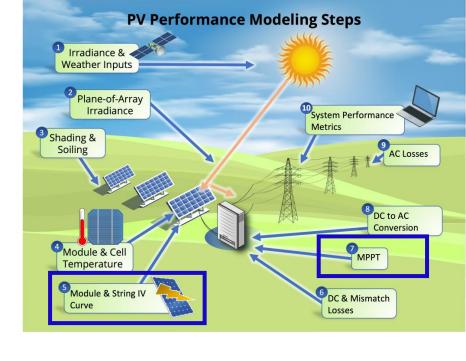
Then, the inverter is constantly searching for the operating point which maximizes the power MPP: Maximum Power Point.



5. Module and String IV Curve

For a fixed irradiance and module temperature, the PV module has its I, current which depends on V, voltage and it can take many operating points.





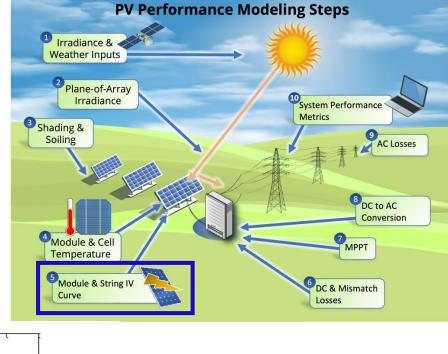
Then, the inverter is constantly searching for the operating point which maximizes the power MPP: Maximum Power Point.

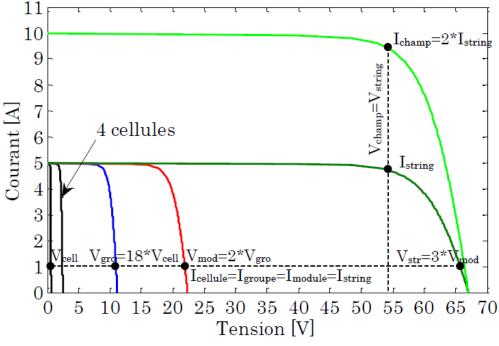
Especially, it changes the voltage with the MPP-Tracker (MPPT) to maximize power.



5. Module and String IV Curve

By the way... the IV curves can be summed up when the modules are connected in series or parallel! The inverter, then, maximizes the power of the PV array IV curve.





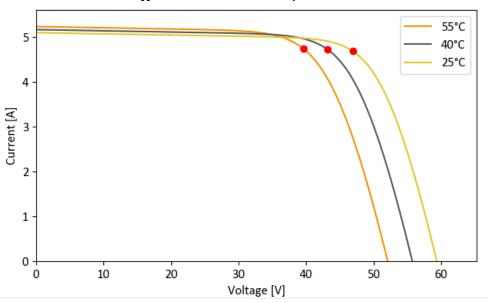


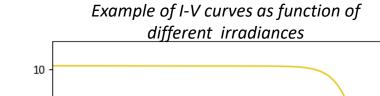
5. Module and String IV Curve

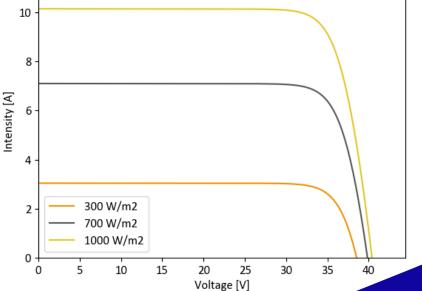
The IV curves' dependencies:

- Higher cell temperatures mostly decrease the voltage
- Higher irradiance level mostly increase the current

Example of I-V curves as function of different module temperature





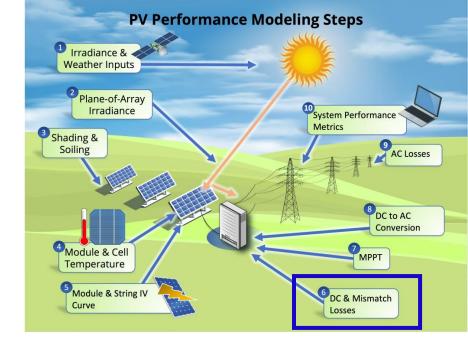




6. DC & Mismatch Losses

Not the focus of this class. However, keep in mind that:

- **DC wiring losses** are around 0.5%-2%.
- **Mismatch losses** refers to the fact that PV modules have different IV curves and this can entail significant losses.

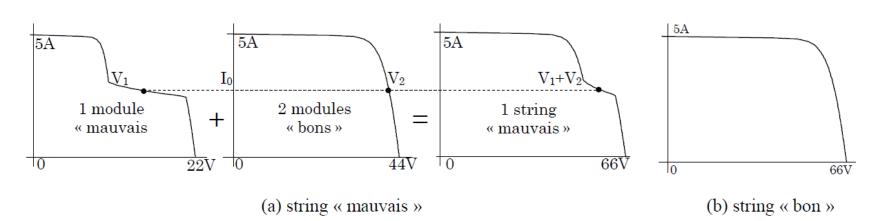


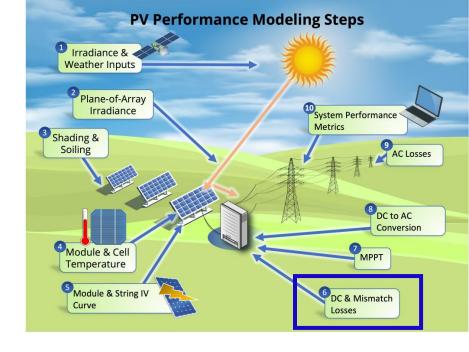


6. DC & Mismatch Losses

Not the focus of this class. However, keep in mind that:

- **DC wiring losses** are around 0.5%-2%.
- Mismatch losses refers to the fact that PV modules have different IV curves and this can entail significant losses.
 For instance, if one of them has a very degraded IV curve (shading or other), it can significantly degrade the IV curve at the array level.





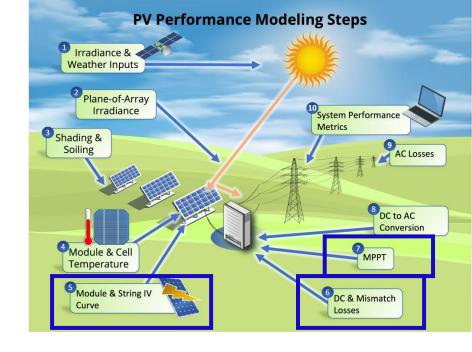


5./6./7. Power model

Constant efficiency model:

$$P_{dc} = \eta \cdot G_{POA} \cdot A$$

- P_{dc} , DC power in [W]
- η efficiency around 20% (from datasheet)
- G_{POA} the irradiance in the plane of array [W/m²]
- A, the PV installation area [m2]





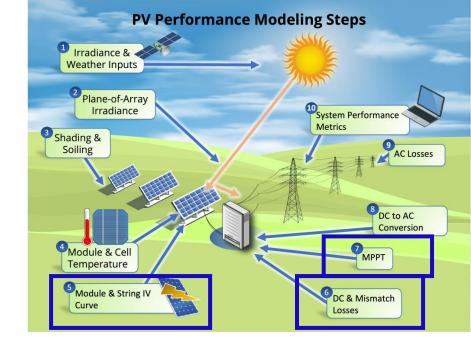
5./6./7. Power model

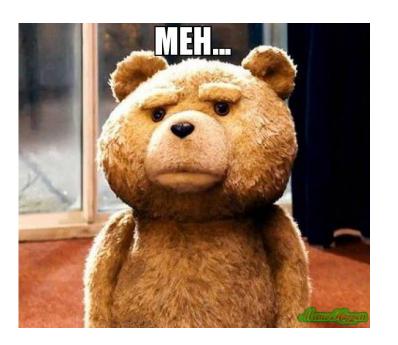
Constant efficiency model:

$$P_{dc} = \eta \cdot G_{POA} \cdot A$$

With:

- P_{dc} , DC power in [W]
- η efficiency around 20% (from datasheet)
- G_{POA} the irradiance in the plane of array [W/m²]
- A, the PV installation area [m2]





Not really precise for instantaneous values

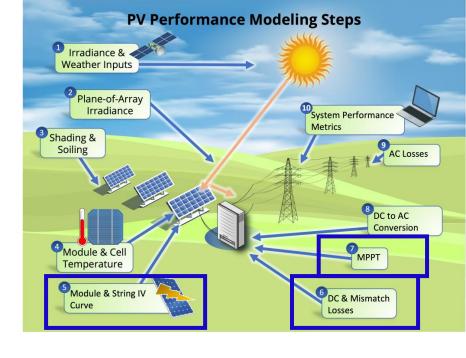


5./6./7. Power model

The <u>PVWatts power model</u> enables to take into account the effect of the cell temperature

$$P_{dc} = P_{dc0} \cdot \frac{G_{POA}}{1000 \, W/m^2} \cdot \left(1 + \gamma_{pdc} \cdot (T_{cell} - 25^{\circ}C)\right)$$

- P_{dc0} Nominal DC power [Wp] (installed capacity)
- G_{POA} the irradiance in the plane of array [W/m²]
- γ_{pdc} , the temperature coefficient (negative, usually between -0.2 -0.5 % W/m²/°C)
- T_{cell} , the cell temperature [°C]





5./6./7. Power model

The Huld model (used in PVGIS) enables to take into account the module temperature and non-lineary with irradiance.

$$P_{dc} = \eta_{Huld}(G, T_m) \cdot G' \cdot P_{dc0}$$

$$\eta_{Huld}(G) = 1 + k_1 \cdot \ln(G') + k_2 \cdot \ln(G')^2 + k_3 \cdot T_m' + k_4 \cdot T_m' \cdot \ln(G') + k_5 \cdot T_m' \cdot \ln(G')^2 + k_6 \cdot T_m'^2$$

- P_{dc0} Nominal DC power [Wp] (installed capacity)
- G_{POA} the irradiance in the plane of array [W/m²]
- $G' = \frac{G_{POA}}{1000 W/m^2}$ the normalized irradiance
- $T_m' = T_m 25$ °C, the module temperature delta [°C]
- $k_1 \dots k_6$, the model coefficients



5./6./7. Power model

The Huld model (used in PVGIS) enables to take into account the module temperature and non-lineary with irradiance.

$$P_{dc} = \eta_{Huld}(G, T_m) \cdot G' \cdot P_{dc0}$$

$$\eta_{Huld}(G) = 1 + k_1 \cdot \ln(G') + k_2 \cdot \ln(G')^2 + k_3 \cdot T_m' + k_4 \cdot T_m' \cdot \ln(G') + k_5 \cdot T_m' \cdot \ln(G')^2 + k_6 \cdot T_m'^2$$

- P_{dc0} Nominal DC power [Wp] (installed capacity)
- G_{POA} the irradiance in the plane of array [W/m²]
- $G' = \frac{G_{POA}}{1000 W/m^2}$ the normalized irradiance
- $T_m' = T_m 25$ °C , the module temperature delta [°C]
- $k_1 \dots k_6$, the model coefficients

Coefficient	c-Si	CIS	CdTe
k ₁	-0.017237	-0.005554	-0.046689
k ₂	-0.040465	-0.038724	-0.072844
<i>k</i> ₃	-0.004702	-0.003723	-0.002262
K ₄	0.000149	-0.000905	0.000276
k ₅	0.000170	-0.001256	0.000159
<i>k</i> ₆	0.000005	0.000001	-0.000006

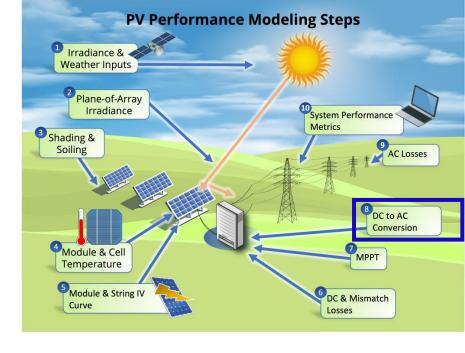


8. Inverter model

The PVWatts inverter model enables to calculate a generic AC/DC efficiency

$$\eta = \frac{\eta_{nom}}{\eta_{ref}} \cdot \left(-0.0162 \cdot \frac{P_{dc}}{P_{dc0}} - \frac{0.0059}{\frac{P_{dc}}{P_{dc0}}} + 0.9858 \right)$$

- η_{nom} The nominal inverter efficiency [-], by default 96%
- η_{ref} The reference inverter efficiency [-], by default 96.4%
- P_{dc} The DC power $\left[\frac{W}{m^2}\right]$
- P_{dc0} The DC input power limit [W/m2]





8. Inverter model

The Sandia inverter model enables to include the voltage and be more precise

$$P_{AC} = \left[\frac{P_{AC0}}{A - B} - C \cdot (A - B)\right] \cdot (P_{dc} - B) + C \cdot [P_{dc} - B]^2$$

Where:

- $A = P_{dc0} \cdot [1 + C1 \cdot (V_{dc} V_{dc0})]$
- $B = P_{s0} \cdot [1 + C2 \cdot (V_{dc} V_{dc0})]$
- $A = C_0 \cdot [1 + C3 \cdot (V_{dc} V_{dc0})]$

Parameters:

- V_{dc} : DC input voltage (V).
- V_{dc0} : DC voltage level (V) at which the AC power rating is achieved
- P_{AC} : AC output power (W)
- P_{AC0} : Maximum AC power rating for inverter at reference conditions (W).
- P_{dc0} : DC power level (W) at which the AC power rating is achieved.
- P_{s0} : DC power required to start the inversion process (W)
- C_0, C_1, C_2, C_3 : Empirical coefficients



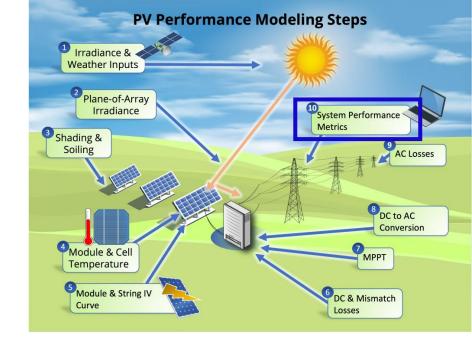
10. Performance Metrics

Performance Ratio PR (IEC 61724):

$$PR = \frac{E_{out}}{E_{POA}} / \frac{P_0}{G_{ref}}$$

Real efficiency normalized by efficiency in STC conditions (1000 W/m2, cell temperature of 25 °C, and AM1.5 spectrum)

- E_{out}: Energy produced [Wh] over the period of time T
- E_{POA} : Irradiation received [Wh/m²] over the period of time T
- P_0 : Installation DC rated power [Wp]
- $G_{ref} = 1000 \frac{W}{m^2}$, reference irradiation





10. Performance Metrics

Performance Ratio PR (IEC 61724):

$$PR = \frac{E_{out}}{E_{POA}} / \frac{P_0}{G_{ref}}$$

Real efficiency normalized by efficiency in STC conditions (1000 W/m2, cell temperature of 25 °C, and AM1.5 spectrum)

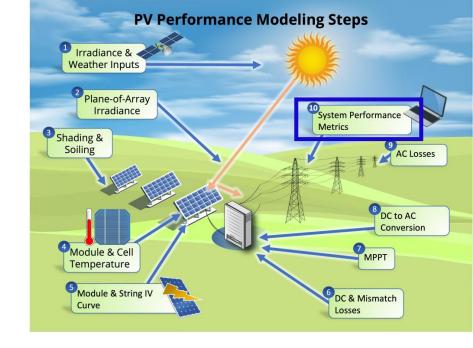
With:

- E_{out}: Energy produced [Wh] over the period of time T
- E_{POA} : Irradiation received [Wh/m²] over the period of time T
- P₀: Installation DC rated power [Wp]
- $G_{ref} = 1000 \frac{W}{m^2}$, reference irradiation

Energy Performance Index EPI:

$$EPI = \frac{PR}{PR_{expected}}$$

PR divided by expected (modelled) $PR_{expected}$





Time for some hands-on exercises!



Use the following notebook:

https://github.com/AlexandreHugoMathieu/pvfault_detection_solar_academy/blob/master/notebooks/dc_power_estimation.ipynb

Follow the python tutorial and estimate the AC power for one year.



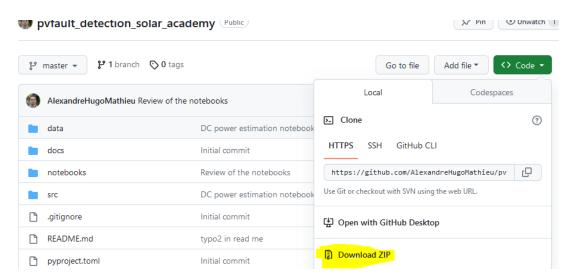
Resources

- Modeling guide PVPMC: https://pvpmc.sandia.gov/modeling-guide/
- Python / Pvlib tutorial: https://pvsc-python-tutorials.github.io/PVSC48-Python-Tutorial/
- To go further:
 - The Use of Advanced Algorithms in PV Failure Monitoring: https://iea-pvps.org/wp-content/uploads/2021/10/Final-Report-IEA-PVPS-T13-19 2021 PV-Failure-Monitoring.pdf



How to install Python and import the course repository to use the notebooks on your local PC.

- 1. Install python: www.python.org/downloads/, download and install the 3.9.13 "release" (Add python to your Path)
- 2. Go to https://github.com/AlexandreHugoMathieu/pvfault_detection_solar_academy, click on the green "Code" button and then download the folder as the zip

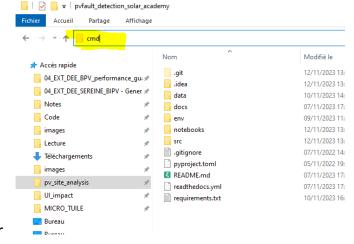




How to install Python and import the course repository to

use the notebooks on your local PC.

- 1. Install python: www.python.org/downloads/, download and install the 3.9.13 "release" (Add python to your Path)
- 2. Go to https://github.com/AlexandreHugoMathieu/pvfault_detection_solar_academy, click on the green "Code" button and then download the folder as the zip
- 3. Unzip it and put it in adequate location in your PC.
- 4. Let's create a virtual environment where you will find all the functions for this course:
 - 1. Go in the folder and open the command line from that same folder by writing "cmd" in the path bar (with Windows)





How to install Python and import the course repository to use the notebooks on your local PC.

- 1. Install python: www.python.org/downloads/, download and install the 3.9.13 "release" (Add python to your Path)
- 2. Go to https://github.com/AlexandreHugoMathieu/pvfault_detection_solar_academy, click on the green "Code" button and then download the folder as the zip
- 3. Unzip it and put it in adequate location in your PC.
- 4. Let's create a virtual environment where you will find all the functions for this course:
 - 1. Go in the folder and open the command line from that same folder by writing "cmd" in the path bar (with Windows)
 - 2. In the command bar: execute the following line to create the "solar_env" environnement that you will use in your notebooks
 - 1. "pip install virtualenv"
 - "python –m virtualenv solar_env"
 - 3. "call solar_env\Scripts\activate" (you should have a 'solar_env' on the left of the command at this point)
 - 4. "pip install –r requirements.txt" (load all the libraries, take a little time, be patient)
 - 5. "python -m ipykernel install --name=solarkernel" (create a kernel for the notebooks)



How to start a notebook

- 1. Go in the folder and open the command line from that same folder by writing "cmd" in the path bar (with Windows)
- 2. In the command bar, exexute:
 - "call solar_env\Scripts\activate" (go in the virtual env)
 - 2. "jupyter notebook" (open the notebooks browser)
- 3. Browse to the notebooks folder, choose one and pick the solarkernel when asked.

