

## \*Defects in Crystal

case of vacancies – reason to occur and equilibrium concentration

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- ✱ We can imagine to create a vacancy by transferring an atom from a lattice site in the interior to a lattice site on the surface of the crystal. This kind of vacancy is called Schottky defect.
- ✱ Instead if an atom is taken away from its site and transferred inside the crystal this is called Frenkel defect.
- ✱ At thermal equilibrium,  $T \neq 0$ , a certain number of vacancies are always present. A surprising consequence is that a certain concentration of vacancy at equilibrium lowers the free energy ( $G$ ) of a crystal at constant temperature and pressure, giving a more stable crystal.

## Estimation of number of vacancies at equilibrium

- ✱ Is possible to estimate the density of vacancy at the equilibrium. To do that we first need to define the free energy for a crystal.
- ✱ According to thermodynamics the change in the free energy of a chemical reaction in a close system is:

$$\Delta G = \Delta H - T\Delta S \quad (1)$$

which  $\Delta H$  is the enthalpy variation and  $\Delta S$  the entropy variation.

- ✱ From a thermodynamic point of view, a solid containing vacancies constitutes a solid solution where the vacancies are dissolved in. In analogy with equilibrium problems with liquid solvent and solution, we can treat the solid like a solvent and the vacancies as a solute, and the problem may be treated in terms of the thermodynamics of chemical reactions and solutions.

- ✦ Let's consider a reaction that starts from a perfect crystal with  $N$  sites and 0 vacancies ( $G_0$ ), and ends up in a crystal with  $N$  sites and  $n$  vacancies ( $G_n$ ). Hence the change in free energy is:

$$\Delta G = G_n - G_0 \quad (2)$$

- ✦ To evaluate properly the free energy we have to evaluate the entropy  $S$  and the enthalpy  $H$ .
- ✦ In addition vacancy increases the enthalpy of the crystal due to energy required to break bonds

- \* in thermodynamics the entropy of a system is:

$$S = k_b \ln(\Omega) \quad (3)$$

Where  $\Omega$  represents all the possible microstates of the system, which describe all of the possible ways that a system could be found.

- \* Hence, in this case the problem reduces to figure out how many ways we can distribute  $n$  vacancies in  $N$  site:

$$\Omega = \frac{N!}{(N-n)!(n!)} \quad (4)$$

If  $n = 0$  we obtain  $\Omega = 1$  and as a consequence the entropy of a perfect crystal  $S_0$  is null.

- \* Using stirling approximation for  $N \rightarrow \infty$ :

$$S = k_b (N \ln(N) - (N-n) \ln(N-n) - n \ln(n)) \quad (5)$$



- ✱ The enthalpy is defined as follow:

$$H = U + \sigma V$$

where the stress  $\sigma$  is a generalisation of the pressure  $p$ .

- ✱ In our case, the  $H_0$  is a constant independent on the number of vacancies, nonetheless,  $H_n$  has a linear dependence on the number of vacancies  $n$ .
- ✱ This is because the enthalpy shows how much energy and work are required to create a vacancy. If we assume that the enthalpy required to create a vacancy is  $H_{vac} = U_{vac} + \sigma V_{vac}$ , hence the entalpy to create  $n$  vacancies the  $H_n$  term is proportional to  $n$ :

$$H_n = n(U_{vac} + \sigma V_{vac}) \quad (6)$$

## Minimum of the free energy

- ✱ Now we have all the ingredients to estimate the number of vacancies for which we have the minimum in  $\Delta G = \Delta H - T\Delta S$ .
- ✱ Hence imposing  $0 = \frac{d}{dn}\Delta G$  we yield with:

$$\frac{n}{N - n} = e^{-\frac{H_{vac}}{k_B T}}$$

considering also the approximation  $N \gg n$

$$c_{n_{vac}} = \frac{n}{N} = e^{-\frac{U_{vac} + \sigma V_{vac}}{k_B T}} \quad (7)$$

- ✱ Here we have found that only for a certain density of vacancies the minimum is reached.

## Minimum of the free energy

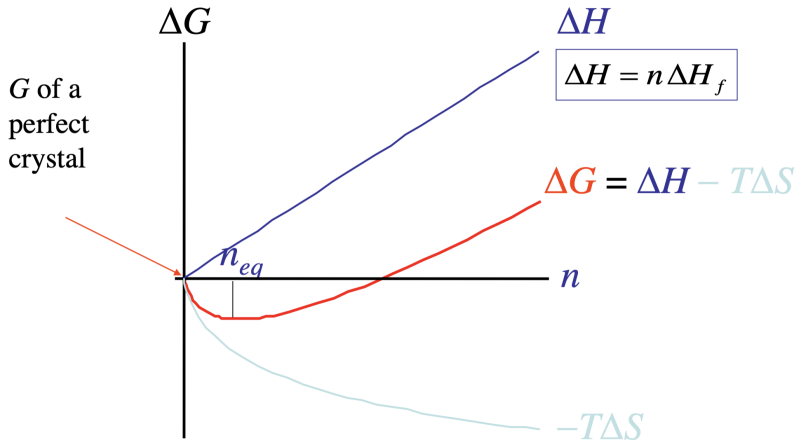


Figure 1: Caption

## Temperature dependence

- \* We can see that for  $T \rightarrow 0$ , theoretically we don't have any vacancies in the crystal. But when the temperature increases the vacancies also increase.

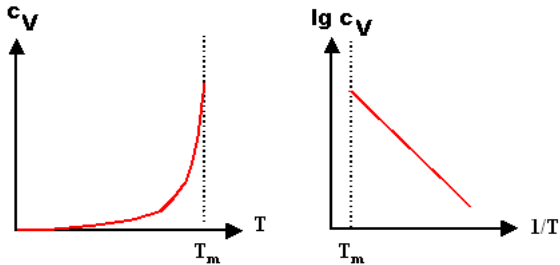
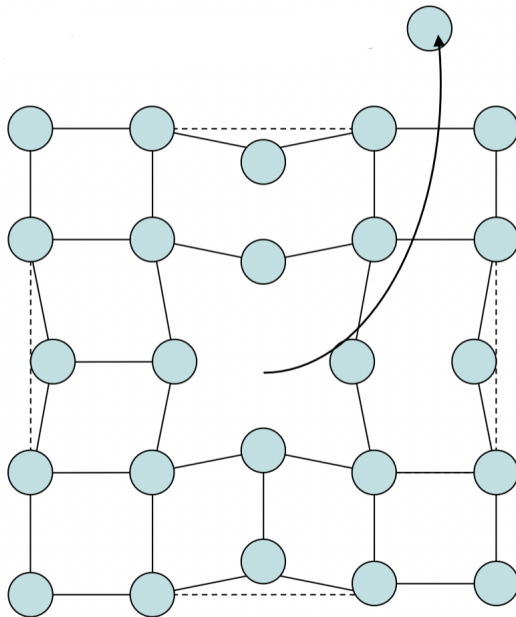


Figure 2: Caption

- \* Neighboring atoms tend to move into the vacancy, which creates a tensile stress field The stress/strain field is nearly spherical and short-range.
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## Pressure dependence



## How to find vacancies experimentally

- ✱ Vacancies are areas with low electron densities. Moreover, they are kind of attractive to a positron because they form a potential well for a positron - once it falls in there, it will be trapped for some time
- ✱ some positrons will be trapped inside vacancies and their percentage will depend on the vacancy concentration. The trapped positrons will enjoy a somewhat longer life span. The average life time of all positrons will thus go up with an increasing number of vacancies, i.e. with increasing temperature.

## How to find vacancies experimentally

