

AE4872 Satellite Orbit Determination

Homework Assignment 3

Least squares estimation of an initial state vector and a dynamical parameter

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Time Spent
20 Hours



List of symbols

Table 1: List of symbols

Symbol/Acronym	Description	Unit
A	Information/Data Matrix	-
LSQ	Least Squares	-
WLSQ	Weighted Least Squares	-
MSL	Matlab Script Lines	-
r	Distance vector	m
pinv.	Pseudo Inverse	-
SVD	Singular Value Decomposition	-
t	Time	s
var.	Variance	Hz^2
x	State vector	-
\bar{y}	Observation vector	-
\hat{y}	Estimated observation vector	-
\hat{x}	Estimated state vector	-
ρ	Range measurement	m
$\tilde{\rho}$	Pseudo range	m
t_T	Transmitter time	s
t_R	Receiver time	s
x_{gps}	GPS x location	m
y_{gps}	GPS y location	m
z_{gps}	GPS z location	m
s	State vector	-
s_0	Initial state vector	-
c	speed of light, 299792458	m/s
P_{xx}	State covariance matrix	-
P_{yy}	Observation vector covariance matrix	-
\hat{S}	State vector	-
s	Estimated state vector	-
μ	Standard gravitational parameter	$3.986004419 \times 10^{14} m^3 s^{-2}$

Cover Image: 1895 Antique map of constellations:

<https://nl.pinterest.com/pin/361765782537705142/>.

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LSQ Estimate for initial state vector

Previously in assignment 2, the position of one of the SWARM satellites was analysed without taking into account the dynamics. Previous assignments along with the supporting Matlab files can be accessed at the git repository ¹. For this assignment same data is used, but a dynamical parameter is also estimated. Thus the model under consideration is dynamical which requires solution via numerical techniques. The purpose of this chapter is to outline dynamical model used followed by a discussion of the results obtained.

1.1. Problem definition

The observation vector and the state vector used in Assignment 2, is illustrated with the aid of Equations 1.1 and 1.2. Equation 1.1 illustrates the dependency of the observation vector on measured and estimated range measurements ρ along with the corrections for receiver (t_R) and transmitter clock times(t_T). This observation vector is further expanded in the linear matrix format in Equation 1.2.

$$\bar{y} = \bar{\rho} - (c \cdot t_R - c \cdot t_T) - \rho \quad (1.1)$$

$$\bar{y} = A \cdot \bar{x}$$

Where, the information matrix A is:

$$A = \begin{bmatrix} \frac{x-x_{GPS}}{\rho} & \frac{y-y_{GPS}}{\rho} & \frac{z-z_{GPS}}{\rho} \end{bmatrix}$$

Where, ρ represents:

$$\rho = \sqrt{(x - x_{GPS})^2 + (y - y_{GPS})^2 + (z - z_{GPS})^2} \quad (1.2)$$

And finally the state vector is represented by:

$$\bar{x} = \begin{bmatrix} \Delta x \\ \Delta y \\ \Delta z \end{bmatrix}$$

Least squares approach undertaken in this section involves using an initial state vector which consists of position and velocity. This is expressed with the aid of Equation 1.3. The state vector is significantly different than the one represented in Equation 1.2. However, The information presented in Equation 1.2 can be manipulated to accommodate the new state vector presented in Equation 1.3.

$$\bar{s} = \begin{bmatrix} x \\ y \\ z \\ \dot{x} \\ \dot{y} \\ \dot{z} \end{bmatrix} \quad (1.3)$$

¹Git Repository: <https://github.com/Alixir/Satellite-Orbit-Determination>

Differentiating the new state vector results in the expression presented in Equation 1.4.

$$\dot{\bar{S}} = F(\bar{S}, t) = \dot{\bar{S}}(t) = \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \\ -\frac{\mu \cdot x}{r^3} \\ -\frac{\mu \cdot y}{r^3} \\ -\frac{\mu \cdot z}{r^3} \end{bmatrix}$$

The latter follows from the Equations of motion:

Equations of motion can be expressed as:

$$\begin{aligned} \frac{\delta x}{\delta t} &= \dot{x} \\ \frac{\delta y}{\delta t} &= \dot{y} \\ \frac{\delta z}{\delta t} &= \dot{z} \\ \frac{\delta \dot{x}}{\delta t} &= \ddot{x} = \frac{-\mu x}{r^3} \\ \frac{\delta \dot{y}}{\delta t} &= \ddot{y} = \frac{-\mu y}{r^3} \\ \frac{\delta \dot{z}}{\delta t} &= \ddot{z} = \frac{-\mu z}{r^3} \end{aligned} \tag{1.4}$$

Where $r = \sqrt{x^2 + y^2 + z^2}$

\bar{S} can be re-written to facilitate iterative propagation with the aid of Euler integration. This is presented with the aid of Equation 1.5.

$$\dot{\bar{S}} = F(S^*, t) + \frac{\delta F(t)}{\delta S(t)|_{S^*}} [S(t) - S^*(t)]$$

$$\dot{\bar{S}} \approx S(t_0 + \Delta) - S(t_0)$$

Here, $S(t) - S^*(t)$ can be represented as the state vector used in the previous assignment.

However, the previous state vector will now include velocity terms as well.

The updates state vector, s can now be written as:

$$s = \begin{bmatrix} \Delta x \\ \Delta y \\ \Delta z \\ \Delta \dot{x} \\ \Delta \dot{y} \\ \Delta \dot{z} \end{bmatrix} \tag{1.5}$$

Thus the above equation for $\dot{\bar{S}}$ simplifies to:

$$s(t) = \phi(t, t_0)s(t_0)$$

$$\dot{s}(t) = \dot{\phi}(t, t_0)s(t_0)$$

Here, ϕ represents the state transition matrix.

ϕ can be propagated iteratively using:

$$\phi(t, t_0) = \frac{\delta F(t)}{\delta S(t)|_{S^*}} \phi(t, t_0)$$

Using the expressions outlined in Equation 1.5 the linear matrix expression for one epoch can now be written as shown in 1.6, which relates observations to the initial state.

$$\bar{y}_1 = H_1 \phi(t_1, t_0)s(t_0) + \epsilon_1$$

Where H is the updated information matrix, to facilitate the new state vector s

$$H(t_{i+1}) = \begin{bmatrix} \frac{x-x_{GPS}}{\rho} & \frac{y-y_{GPS}}{\rho} & \frac{z-z_{GPS}}{\rho} & 0 & 0 & 0 \end{bmatrix} \tag{1.6}$$

A key relation in the above expression is $\frac{\delta F}{\delta S}$. This is derived with the aid of equation of motion presented earlier in Equation 1.4. Equations 1.7 - 1.10 derive this key expression.

The gradient of $F(t)$ wrt $S(t)$ can be expressed as:

$$\frac{\delta F(t)}{\delta S(t)} = \begin{bmatrix} \frac{\delta F_1}{\delta S_1} & \cdots & \frac{\delta F_n}{\delta S_1} \\ \vdots & \ddots & \vdots \\ \frac{\delta F_1}{\delta S_k} & \cdots & \frac{\delta F_n}{\delta S_k} \end{bmatrix} \quad (1.7)$$

Where, k represents the vector size of states S .

While, n represents the vector size of time differentiation of S , F .

In this case $n = k = 6$.

$$\frac{dF_4}{dS_1} = \frac{d}{dx} \left(\frac{-\mu \cdot x}{r^3} \right) = \frac{\mu}{r^3} \left(\frac{3 \cdot x^2}{r^2} - 1 \right) \quad (1.8)$$

$$\frac{dF_4}{dS_2} = \frac{d}{dy} \left(\frac{-\mu \cdot x}{r^3} \right) = \frac{3 \cdot \mu \cdot x \cdot y}{r^5} \quad (1.9)$$

$$\frac{dF(S, t)}{dS} = \begin{bmatrix} 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 0 & 0 & 1 \\ \frac{\mu}{r^3} \left(\frac{3 \cdot x^2}{r^2} - 1 \right) & \frac{3 \cdot \mu \cdot x \cdot y}{r^5} & \frac{3 \cdot \mu \cdot x \cdot z}{r^5} & 0 & 0 & 0 \\ \frac{3 \cdot \mu \cdot x \cdot y}{r^5} & \frac{\mu}{r^3} \left(\frac{3 \cdot y^2}{r^2} - 1 \right) & \frac{3 \cdot \mu \cdot y \cdot z}{r^5} & 0 & 0 & 0 \\ \frac{3 \cdot \mu \cdot x \cdot z}{r^5} & \frac{3 \cdot \mu \cdot y \cdot z}{r^5} & \frac{\mu}{r^3} \left(\frac{3 \cdot z^2}{r^2} - 1 \right) & 0 & 0 & 0 \end{bmatrix} \quad (1.10)$$

Now that the fundamental expressions are outlined, Figure 1.1 outlines the algorithm used to conduct the Least Squares estimate for a given initial vector.

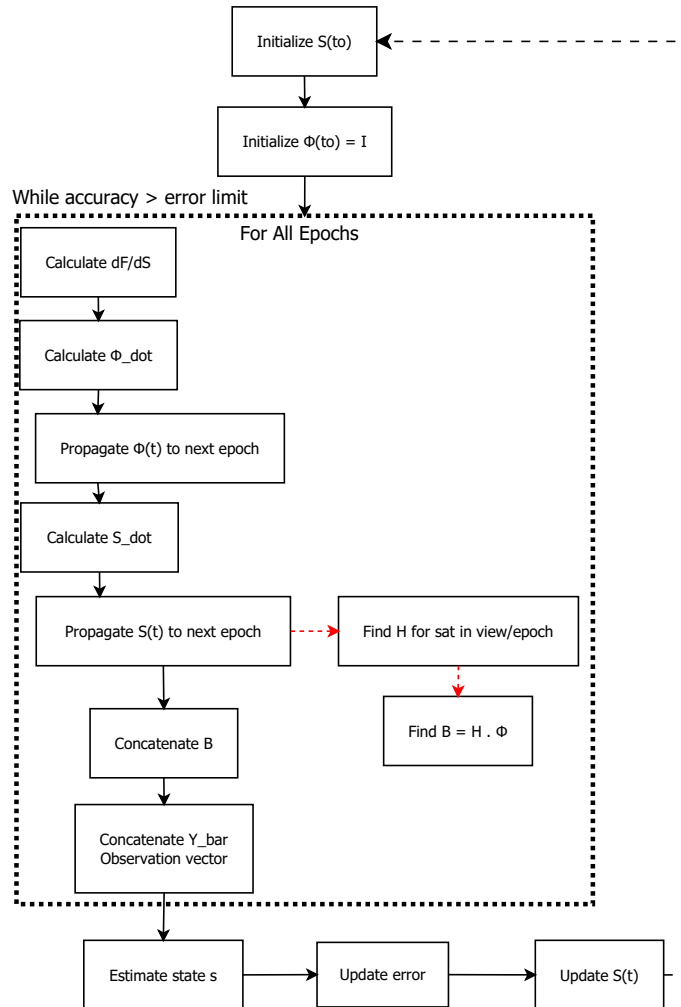


Figure 1.1: Flowchart illustrating steps undertaken to conduct LSQ for a given initial state vector.

As a first guess of initialisation of \bar{S} , the positions and velocities at the first epoch are taken from previous assignment. The velocity part is obtained via simple Euler scheme as indicated in the Matlab script in Appendix. Slight deviation is added to the initial vector to observe convergence. Equation 1.11, outlines the initial state vector \bar{S} along with modified \bar{S} to visualise convergence.

$$\bar{S}_0 = \begin{bmatrix} 1.94e6 \\ -5.84e6 \\ 2.93e6 \\ 0.001e6 \\ -0.0031e6 \\ -0.0069e6 \end{bmatrix}$$

Adding the vector [1000m 1000m 1000m 100 m/s 100 m/s 100m/s]' for visualising convergence. (1.11)

$$\bar{S}_0 = \begin{bmatrix} 1.93e6 \\ -5.81e6 \\ 3.0e6 \\ 0.001e6 \\ -0.0032e6 \\ -0.0068e6 \end{bmatrix}$$

1.2. Results

This section presents the results for the algorithm outlined in Figure 1.1. The algorithm initiates using the initialisation vector \bar{S} defined in Equation 1.11. ϕ is chosen to be identity of size 6x6. The error limit is defined by the norm of x,y,z components of \bar{S} . A limit of 10^{-3} is used for convergence of iteration. $\frac{dF}{ds}$ is calculated with the aid of Equation 1.10. This is used to calculate $\dot{\phi}$. With the aid of Runge-Kutta integration, this is used to propagate $\phi(t)$ to next epoch. Which is further used to estimate $\dot{\bar{S}}$, later used to propagate \bar{S} to next epoch. Updated information matrix H is calculated for all satellites in the view. The observation vector is noted for all epochs. The updated information matrix H is multiplied with the state transition matrix ϕ , and stored in B for all epochs. After looping through the epochs $H \cdot \phi$ for all epochs is appended in a combined information matrix of size 87x6. The observation vector is obtained with the aid of Equation 1.1 as indicated earlier. Now that the linear LSQ format indicated in Equation 1.1 is obtained, SVD method illustrated in Assignment 2² is used to estimate s. Interested reader is referred to the script presented in the Appendix. Once s is estimated, it is added to $\bar{S}(t_0)$ to update the initialisation vector for next iteration. Since, SVD guided LSQ was implemented no singularity warnings were presented by Matlab. The results of algorithm is presented in Figure 1.2. Figure 1.2 illustrates the evolution of the epoch wise residual norm for increasing iteration numbers. The residual norm obtained using $\bar{Y} - \hat{Y}$, where \bar{Y} represents the vector observations and \hat{Y} represents the vector of estimations. The error is defined as the norm of positional values of \bar{S} .

²Assignment 2: <https://github.com/Alixir/Satellite-Orbit-Determination>

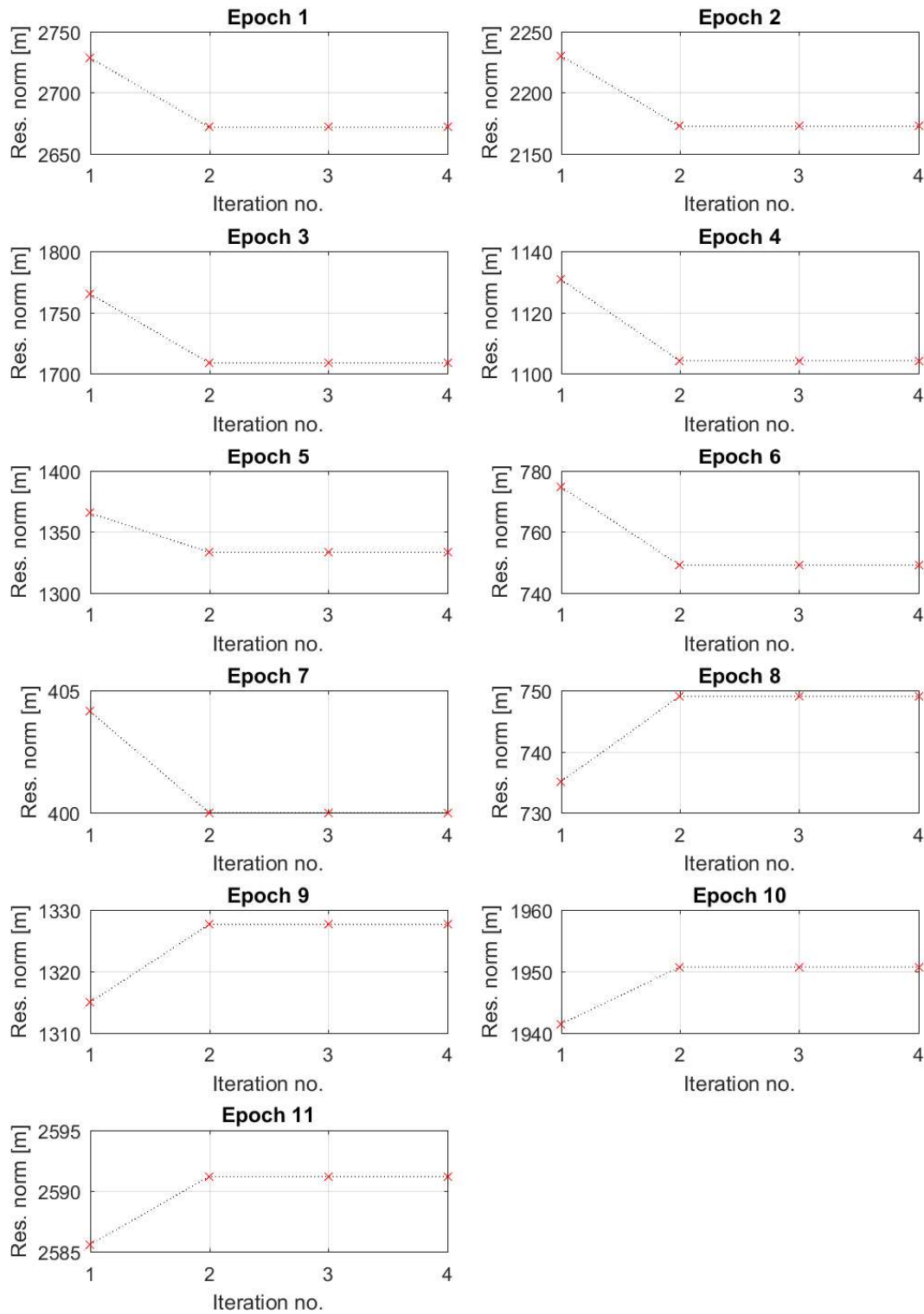


Figure 1.2: Evolution of epoch wise residual norm.

As seen in Figure 1.2, best results are obtained after 4 iterations. Note this depends on the definition of the error limit used. The error margin used here is 10^{-3} as indicated in the above discussion.

Estimating gravitational parameter

The aim of this chapter is to include a new dynamical parameter in the state vector \bar{S} defined earlier in Equation 1.3. The dynamical parameter is the gravitational constant μ . An estimate for μ is obtained. Results of the satellite position are compared with the results obtained in Chapter 1.

2.1. Problem Definition

The new state vector after inclusion of μ is represented in Equation 2.1. The dimension of the initial state transition matrix ϕ is now updated to an identity matrix of size 7x7. The gravitational parameter μ defined in Equation 2.2.

$$\bar{S} = \begin{bmatrix} x \\ y \\ z \\ \dot{x} \\ \dot{y} \\ \dot{z} \\ \mu \end{bmatrix} \quad (2.1)$$

$$V(r) = -\frac{\mu}{r} \quad (2.2)$$

Based on this definition of μ , the definition of information matrix H, $F(\bar{S}, t)$ and $\frac{dF}{dS}$ requires to be updated. Gravitational parameter μ , is no longer a constant for all calculations. It is rather updated every iteration. An initial value of $3 \cdot 10^{14}$ is chosen for μ . However, the choice of initial μ does not impact the solution, it always converges to the same value. This will be discussed later. For other elements of \bar{S} , same initialisation values from Equation 1.11 is used. The updated F is presented in Equation 2.3 and the corresponding update of $\frac{dF}{dS}$ due to the influence of $\frac{dF}{d\mu}$ is expressed in Equation 2.4. Finally, the updated matrix H is presented in Equation 2.5.

$$F(\bar{S}, t) = \dot{\bar{S}}(t) = \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \\ -\frac{\mu \cdot x}{r^3} \\ -\frac{\mu \cdot y}{r^3} \\ -\frac{\mu \cdot z}{r^3} \\ \dot{\mu} \end{bmatrix} \quad (2.3)$$

$$\frac{dF}{d\mu} = \begin{bmatrix} 0 \\ 0 \\ 0 \\ -\frac{x}{r^3} \\ \frac{y}{r^3} \\ -\frac{z}{r^3} \\ 0 \end{bmatrix} \quad (2.4)$$

$$\frac{dF(S, t)}{dS} = \begin{bmatrix} 0 & 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 1 & 0 \\ \frac{\mu}{r^3} \left(\frac{3 \cdot x^2}{r^2} - 1 \right) & \frac{3 \cdot \mu \cdot x \cdot y}{r^5} & \frac{3 \cdot \mu \cdot x \cdot z}{r^5} & 0 & 0 & 0 & -\frac{x}{r^3} \\ \frac{3 \cdot \mu \cdot x \cdot y}{r^5} & \frac{\mu}{r^3} \left(\frac{3 \cdot y^2}{r^2} - 1 \right) & \frac{3 \cdot \mu \cdot x \cdot y}{r^5} & 0 & 0 & 0 & -\frac{y}{r^3} \\ \frac{3 \cdot \mu \cdot x \cdot z}{r^5} & \frac{3 \cdot \mu \cdot y \cdot z}{r^5} & \frac{\mu}{r^3} \left(\frac{3 \cdot z^2}{r^2} - 1 \right) & 0 & 0 & 0 & -\frac{z}{r^3} \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 \end{bmatrix} \quad (2.5)$$

$$H(t_{i+1}) = \begin{bmatrix} \frac{x-x_G}{\rho} & \frac{y-y_G}{\rho} & \frac{z-z_G}{\rho} & 0 & 0 & 0 & 0 \end{bmatrix}$$

The algorithm is indicated with the aid of Figure 2.1. Note how the gravitational parameter μ is updated every iteration.

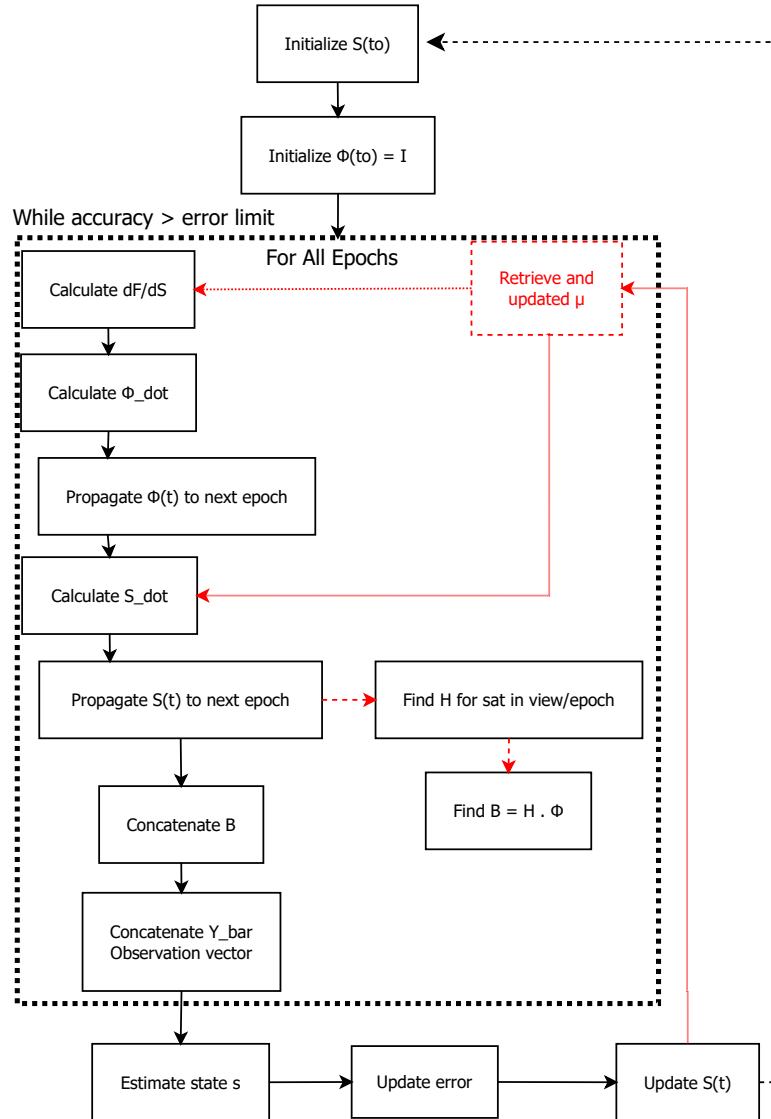


Figure 2.1: Flowchart illustrating steps undertaken to conduct LSQ for updated initial state vector including μ .

2.2. Results and comparison

The results of running the algorithm presented in Figure 2.1 is illustrated with the aid of Figure 2.2. Similar error limit of 10^{-3} is used, and the error is measured as the norm of the positional values of \hat{S} . The residual is again taken as the difference between observation vector \bar{Y} and the estimation \hat{Y} . Again SVD guided LSQ is used for parameter estimation. Interest reader is referred to follow the script presented in Appendix.

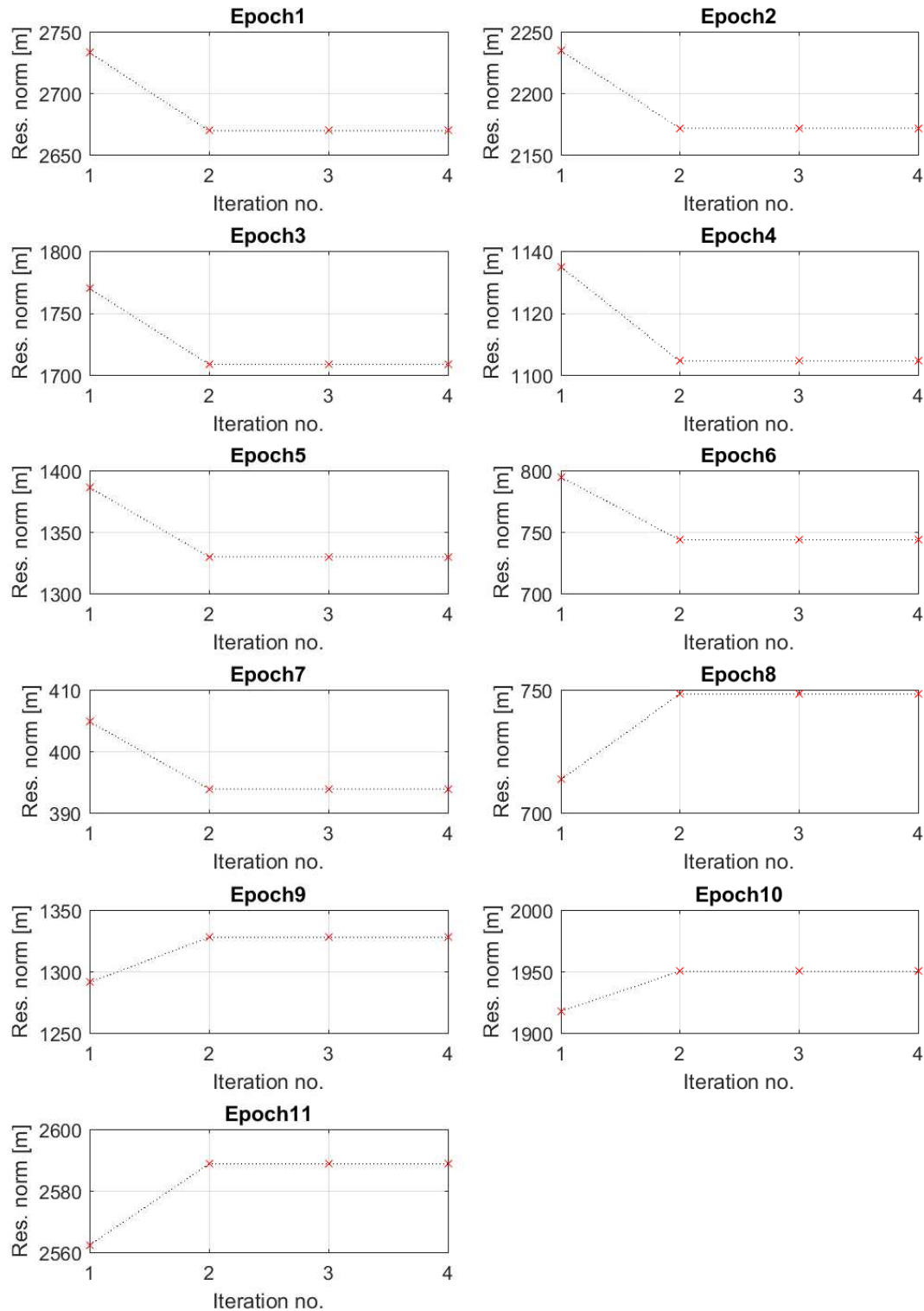


Figure 2.2: Evolution of epoch wise residual norm.

Comparing the results presented in Figures 1.1 and 2.1, both of the residual norms are seen to converge to the provided limit within 4 iterations. Table 2.1 illustrates the final residual norms for the case of 6 unknowns and 7 unknowns in the state vector. Overall, the final values for all epochs are quite close. Overall the residual

norms for state vector with 7 unknowns are lesser in magnitude compared to that of the 6 unknowns. But the difference is not significant.

Table 2.1: Residual behaviour at the end of iterations for unknown with 6 and 7 states.

Epoch	Residual for 6 unknowns [m]	Residual for 7 unknowns [m]
1	2671	2669
2	2172	2172
3	1709	1709
4	1104	1105
5	1333	1330
6	749.1	743.5
7	400	393.9
8	749.1	748.3
9	1328	1328
10	1951	1950
11	2591	2589

Table 2.2 is used to facilitate the comparison between the positional values from Assignment 2 and for different state vectors. Table 2.2 also illustrates the value of μ parameter obtained for state with 7 unknown parameters. The real value of μ for Earth is: 3.9860×10^{14} (up to 4 decimal places) ¹. Comparing this with the value of $\mu = 3.9451 \times 10^{14}$ presented in Table 2.2, the difference from the real value is 1.03%.

Comparing the final position of the satellite from assignment 2 with the final positions obtained in this assignment, it can be observed that, the x, y and z position are fairly close for the case of 6 and 7 unknown state parameters. However, they are different from the values obtained in Assignment 2. Norm of the [x, y, z] for Assignment 2 positional values is: 6.8150×10^6 m while that for states with 6 and 7 unknowns are: 6.8214×10^6 m and 6.8216×10^6 m consecutively. The differences in norm is quite significant (when measured in units of meter, 6400m and 6600m), this could be due to the fundamental differences in which the parameters are estimated. One conducts parameter estimation of $[\Delta x, \Delta y, \Delta z]$, while makes use of state transition matrix relations to conduct parameter estimation of $[\Delta x, \Delta y, \Delta z]$, but the primary initialiser is [x, y, z] along with $[\dot{x}, \dot{y}, \dot{z}]$ which are derived using based Euler integration technique with 10 time steps. Which might be considered to be quite significant.

Table 2.2: Initial and final states showing the change in position and velocities.

	Initial state	6 unknown parameters	7 unknown parameters
x [m]	1.939e06	2.0251e06	2.0251e06
y [m]	-5.838e06	-6.1198e06	-6.1199e06
z [m]	2.933e06	2.2314e06	2.2315e06
\dot{x} [m/s]	984.2	718.2100	719.6624
\dot{y} [m/s]	-3148.6	-2.4084e03	-2.4127e03
\dot{z} [m/s]	-6859.2	-7.1933e03	-7.1915e03
$\mu [m^3 s^{-2}]$	3e14	-	3.9451e14

¹Gravitational Parameter: https://en.wikipedia.org/wiki/Standard_gravitational_parameter

Influence of initial state vector on results

The aim of this chapter is to discuss the influence of initial state vector on the results and final estimated values. First the case of 6 unknown states is considered in Section 3.1 followed by the case of 7 unknown state in Section 3.2. The effect is observed in blocks of position, velocity and gravitational parameters.

3.1. Initial state vector with position and velocity.

Table 3.1 outlines the initial state vector along with initial state vectors of increased/decreased position and velocity components. For increased position, the x,y and z positions are increased by 1e06. Since until 1e05 no difference was observed neither in the no. iterations nor the final convergence values. Similarly, for decreased positions, the x,y, and z component are decreased by 1e06. The final values are outlined Table 3.2. Increasing/decreasing the position by 1e06 only impacts the no. of iteration, the rest stays unchanged. Quite surprising how this has no impact of the final values. For increased/decreased velocity, the velocity components are increased/decreased by 1e04 until any consequence is observed. The final values for this is also outline in Table 3.2. Once again, there is no consequence on the final value. Only consequence is in the number of iterations required for convergence to a desired accuracy.

Table 3.1: Different values of states for initial state vector.

	Initial state vector, s_1	Increased position, s_2	Decreased position, s_3	Increased velocity, s_4	Decreased velocity, s_5
x [m]	1.9393e06	2.9393e06	9.3925e05	1.9393e06	1.9393e06
y [m]	-5.8377e06	-4.8377e06	-6.8377e06	-5.8377e06	-5.8377e06
z [m]	2.9332e06	3.9332e06	1.9332e06	2.9332e06	2.9332e06
\dot{x} [m/s]	984.1881	984.1881	984.1881	1.0984e04	-9.0158e03
\dot{y} [m/s]	-3.1486e03	-3.1486e03	-3.1486e03	6.8514e03	-1.3149e04
\dot{z} [m/s]	-6.8592e03	-6.8592e03	-6.8592e03	3.1408e03	-1.6859e04

Table 3.2: Final values for different initial state vectors.

Final values	s_1	s_2	s_3	s_4	s_5
x [m]	2.0251e06	2.0251e06	2.0251e06	2.0251e06	2.0251e06
y [m]	-6.1198e06	-6.1198e06	-6.1198e06	-6.1198e06	6.1198e06
z [m]	2.2314e06	2.2314e06	2.2314e06	2.2314e06	2.2314e06
\dot{x} [m/s]	718.2100	718.2100	718.2100	718.2100	718.2100
\dot{y} [m/s]	-2.4084e03	-2.4084e03	-2.4084e03	-2.4084e03	-2.4084e03
\dot{z} [m/s]	-7.1933e03	-7.1933e03	-7.1933e03	-7.1933e03	-7.1933e03
No. of iterations	4	5	5	5	5

3.2. Initial state vector with position, velocity and gravitational parameter

Analysis similar to section 3.1 is carried out here, with the exception that now the state vector is of dimension 7 rather 6. Which is the result of embedding a gravitational parameter term μ . As mentioned earlier in Section 2.1, initial choice of μ alone has no consequence on the final results. It always converges to the same value

within the same number of iterations. Thus for this section, we still analyse the results of varying position and velocity components in the state vector. Interested reader can run the script provided in the Appendix or Git repository. Initial state vectors are similar to that presented in Table 3.1 with the exception of an additional state, $\mu = 3e14$. Initial conditions are tabulated in Table 3.3 while the final values are presented in Table 3.4.

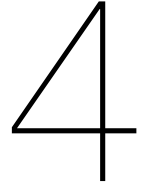
Table 3.3: Different values of states for initial state vector.

	Initial state vector, s_1	Increased position, s_2	Decreased position, s_3	Increased velocity, s_4	Decreased velocity, s_5
x [m]	1.9393e06	2.9393e06	9.3925e05	1.9393e06	1.9393e06
y [m]	-5.8377e06	-4.8377e06	-6.8377e06	-5.8377e06	-5.8377e06
z [m]	2.9332e06	3.9332e06	1.9332e06	2.9332e06	2.9332e06
\dot{x} [m/s]	984.1881	984.1881	984.1881	1.0984e04	-9.0158e03
\dot{y} [m/s]	-3.1486e03	-3.1486e03	-3.1486e03	6.8514e03	-1.3149e04
\dot{z} [m/s]	-6.8592e03	-6.8592e03	-6.8592e03	3.1408e03	-1.6859e04
μ [m^3s^{-2}]	3e14	3e14	3e14	3e14	3e14

Table 3.4: Final values for different initial state vectors.

Final values	s_1	s_2	s_3	s_4	s_5
x [m]	2.0251e06	2.0251e06	2.0251e06	2.0251e06	2.0251e06
y [m]	-6.1199e06	-6.1199e06	-6.1199e06	-6.1199e06	-6.1199e06
z [m]	2.2315e06	2.2315e06	2.2315e06	2.2315e06	2.2315e06
\dot{x} [m/s]	719.6224	719.6224	719.6224	719.6224	719.6224
\dot{y} [m/s]	-2.4127e03	-2.4127e03	-2.4127e03	-2.4127e03	-2.4127e03
\dot{z} [m/s]	-7.1915e03	-7.1915e03	-7.1915e03	-7.1915e03	-7.1915e03
μ [m^3s^{-2}]	3.9451e14	3.9451e14	3.9451e14	3.9451e14	3.9451e14
No. of iterations	4	5	5	5	5

To conclude, increasing the order of magnitude of increments added to position/velocity parts for both cases, increases the no. of iteration required to reach the same final value. However, increasing it higher than a limit causes the results to not converge anymore.



Standard deviation of gravitational parameter

This chapter focuses on calculating the standard deviation σ , of μ parameter. It is assumed that the observation is uncorrelated and that they have a standard deviation of 5m. Thus co-variance matrix of observation vector is identity matrix of size 87x87, with 5^2 on the diagonal. Once the information matrix is obtained the covariance of state vector can be related to the covariance matrix of the observation vector using the expression outlined in Equation 4.1.

$$P_{xx} = (A^T P_{yy}^{-1} A)^{-1}$$

Where the term on the left indicates the covariance matrix of the states.

Since P_{yy} is a diagonal matrix, the above expression be simplified to:

$$P_{xx} = 5^2 \cdot (A^T A)^{-1}$$

Where, A is the concatenated information matrix with full rank.

However, $A^T A$ results in loss of rank.

The loss of rank induces singularity warnings in Matlab and might result in numerical errors.

Unlike many other things in life, there is a quick fix for it.

Commutative law of matrix multiplication dictates the following iff both A and B are invertible:

$$(AB)^{-1} = B^{-1}A^{-1}$$

Since A is not a square matrix, it cannot be inverted. However, Moore–Penrose inverse can still be applied.

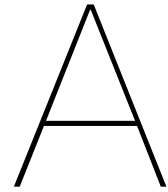
This results in the following relation:

$$P_{xx} = 5^2 \cdot A^+ (A^T)^+$$

Where the superscript " + " denotes it being a pseudo-inverse rather than an inverse.

(4.1)

Based on Equation 2.1 the standard deviation of the gravitation parameter σ_μ is obtained taking the square root of final diagonal element of the matrix P_{xx} . The final diagonal element of P_{xx} is 8.0273e+21. Taking the square root of this gives, $\sigma_\mu = 8.9595e+10$. Adding this standard deviation to the value of μ obtained via parameter estimation results in μ 3.9460e+14. The updated value of μ after estimation and inclusion of σ_μ is still 1.0045% off compared to the real value. However, it is definitely better than previous estimation which was 1.03% off. It is acceptable but accurate. Esp, since the assumption that observations are uncorrelated and have a standard deviation of 5m, is not entirely true and can easily be observed by checking the covariance matrix of P_{yy}



Matlab Script

Following script can be also be found at this Github Repository:

<https://github.com/Alixir/Satellite-Orbit-Determination>

```
1 % Single point positioning
2 % Wouter van der Wal
3 % Delft University of Technology
4
5 clear all
6 % close all
7 clc
8
9 % Fundamental GPS constants
10 % gpsconst;
11
12 dt = 10; % time step
13 number_of_epochs = 11;
14 order = 11; % order of interpolation of precise orbits GPS satellites
15 order_clk = 3; % order of interpolation of clock correction
16 t_orbit = 60*[-78:15:73]; % time vector for precise orbits for GPS satellite provided
    by IGS every 15 minutes
17 time = 0:dt:(number_of_epochs-1)*dt; % time vector for SWARM satellite
18 c = 299792458; % velocity of light m/sec
19 earth_radius = 6371; % in km
20 %% Read the IGS coordinates and clock corrections for all the GPSsatellites
21 [x_gps,y_gps,z_gps,clock_correction] = readsp3('igs19301.sp3');
22
23
24 %% Read the Swarm GPS data
25 fid_rinex = fopen('Data_SWARM_A.dat','r');
26 for ii = 1:15
27     line = fgets(fid_rinex);
28 end
29 % Read the observations
30 for epoch = 1:number_of_epochs
31     line = fgets(fid_rinex);
32     ft = sscanf(line,'%c %f %f %f %f %f %f %f %f');
33     no_sat(epoch) = ft(9);
34     si = 0;
35     for no = 1:no_sat(epoch)
36         si = si + 1;
37         line = fgets(fid_rinex);
38         ft = sscanf(line,'%c %f %f %f %f %f %f %f %f');
39         prns(epoch,no) = ft(2); % PRN number of satellite
40         n = prns(epoch,no);
41         C1(epoch,si) = ft(5);
42         % Interpolate the coordinates of the GPS satellites to the current epoch
43         [x_gps_interpolated(epoch,si),x_error(epoch,si)] = polint(t_orbit,x_gps(:,n),time(epoch),
    order);
44         [y_gps_interpolated(epoch,si),y_error(epoch,si)] = polint(t_orbit,y_gps(:,n),time(epoch),
    order);
45         [z_gps_interpolated(epoch,si),z_error(epoch,si)] = polint(t_orbit,z_gps(:,n),time(epoch),
    order);
46         [clock_correction_interpolated(epoch,si),t_error(epoch,si)] = polint(t_orbit,
    clock_correction(:,n),time(epoch),order_clk);
47     end
end
```



```

48 end
49 fclose(fid_rinex);
50 ccGPS = clock_correction_interpolated*c;
51 x_gps_interpolated = x_gps_interpolated*1e3; % Convert to meters
52 y_gps_interpolated = y_gps_interpolated*1e3;
53 z_gps_interpolated = z_gps_interpolated*1e3;
54 r_gps_interpolated = sqrt(x_gps_interpolated(:,1).^2+y_gps_interpolated(:,1).^2+z_gps_interpolated
(:,1).^2);
55 % Plot for checking
56 % figure; plot((r_gps_interpolated-6371e3)/1e3); title('altitude PRN 7 [km]')
57
58
59 %% Read the precise Swarm orbit
60 fid_obs = fopen('PreciseOrbit_SWARM_A.dat','r');
61 for ii = 1:22
62     line = fgets(fid_obs);
63 end
64 for epoch = 1:11
65     line = fgets(fid_obs);
66     line = fgets(fid_obs);
67     x_precise(epoch) = read_strval(line(5:18)); % These are in km from the center of the earth
- Ali
68     y_precise(epoch) = read_strval(line(19:32));
69     z_precise(epoch) = read_strval(line(33:46));
70     line = fgets(fid_obs);
71 end
72 fclose(fid_obs);
73 % Convert to spherical coordinates for checking
74 r_precise = sqrt(x_precise.^2+y_precise.^2+z_precise.^2);
75 % Plot for checking
76 % figure; plot(r_precise-6371); title('altitude Swarm [km]')
77 %% ===== START OF STUDENT SCRIPT =====
78 %% Author: Ali Nawaz
79 % Course: AE4872 Satellite Orbit Determination
80 % Homework Assignment 3: Least square estimation of an initial state vector
81 % and a dynamic parameter
82 %%=====
83 %% Data retrieval
84 xg = x_gps;
85 yg = y_gps;
86 zg = z_gps;
87 cc = clock_correction;
88
89 xgi = x_gps_interpolated;
90 ygi = y_gps_interpolated;
91 zgi = z_gps_interpolated;
92 rgi = r_gps_interpolated;
93 cci = clock_correction_interpolated;
94 ccf = ccGPS; % Final clock correction = cci*c
95
96 xp = x_precise;
97 yp = y_precise;
98 zp = z_precise;
99 rp = r_precise;
100
101 %% Part 1a
102 % Final values of the epochwise distance is stored in the variable
103 % s_final_rcc. Please refer to previous assignment or follow the github
104 % link. While the clock error is stored in del_t variable. Both of them are
105 % used to reconstruct the initial state vector.
106
107 % Loading the initial state vector from previous script
108 pos = load('s_final_rcc','-mat'); % Position in m
109 clock_errors = load('del_t','-mat'); % Clock errors in s
110
111 % s_final_rcc = pos.s_final_rcc(end,:); % Final x,y,z position for last epoch with clock correction
implemented
112 s_final_rcc = pos.s_final_rcc(1,:); % Final x,y,z position for last epoch with clock correction
implemented
113 final_clock_error = clock_errors.del_t(1,end); % Final receiver clock correction
114
115 % Basic backward Euler scheme to evaluate velocity terms in m/s
116 % vel = (pos.s_final_rcc(end,:) - pos.s_final_rcc(end-1,:))/dt;
117 vel = (pos.s_final_rcc(2,:) - pos.s_final_rcc(1,:))/dt;
118 % Initializing state vector: [ x, y, z, x_dot, y_dot, z_dot] with clock

```

```

119 % correction implemented
120 s0_buf = [ s_final_rcc , vel] + [ 0 0 0 0 0 0 ];
121
122 % Shifting initial vector slightly to visualize convergence.
123 % s0 = s0_buf + [ 1000 1000 1000 100 100 100];
124 s0 = s0_buf;
125 % s0 = [ 0.1 0.1 0.1 0.1 0.1 0.1];
126 % Standard Gravitational parameter Earth
127 mu = 3.986004418e14; % [ m^3 s^-2]
128 % Loop initializations
129 diff4 = norm(s0); % Current error resolution
130 iter4 = 0; % Current iteration step
131 diff_iter = [];
132 Svect = [];
133 res = [];
134 while diff4 >= 10^-3
135 % while iter4 <= 10
136     A4 = [];
137     y_bar4 = [];
138     iter4 = iter4 + 1;
139     n = 0;
140     phi = eye(6,6); % Initializing state transition matrix phi
141     S = s0;
142     Svect = [Svect; S];
143     % For all epochs
144     for e=1:epoch
145         clear H dx4 dy4 dz4 drho4;
146         r = norm(S(1,1:3)); % radial distance sqrt(x0, y0, z0) in meters
147         % Calculating dF/dS matrix for s0 = [ x, y, z, x_dot, y_dot, z_dot ]
148         dF_dS = [ 0 0 0 1 0 0; 0 0 0 0 1 0; 0 0 0 0 0 1; ...
149             (mu/r^3)*( (3*S(1)^2)/(r^2) - 1), 3*mu*S(1)*S(2)/r^5, 3*mu*S(1)*S(3)/r^5, 0,
150             0, 0; ...
151             3*mu*S(1)*S(2)/r^5, (mu/r^3)*( (3*S(2)^2)/(r^2) - 1), 3*mu*S(2)*S(3)/r^5, 0,
152             0, 0; ...
153             3*mu*S(1)*S(3)/r^5, 3*mu*S(2)*S(3)/r^5, (mu/r^3)*( (3*S(3)^2)/(r^2) - 1), 0, 0,
154             0];
155         % Calculating phi_dot
156         phi_dot = dF_dS*phi;
157         % Updating phi
158         phi = phi + phi_dot*dt;
159         % Calculating S_dot for s0 = [ x, y, z, x_dot, y_dot, z_dot ]
160         S_dot = [ S(4); S(5); S(6); -mu*S(1)/r^3; -mu*S(2)/r^3; -mu*S(3)/r^3 ]';
161         % Updating S
162         S = S + S_dot.*dt;
163
164         for s = 1:no_sat(e)
165             dx4(s) = S(1) - xgi(e,s); % x - x_gps for sth satellite in view
166             dy4(s) = S(2) - ygi(e,s); % y - y_gps for sth satellite in view
167             dz4(s) = S(3) - zgi(e,s); % z - z_gps for sth satellite in view
168
169             drho4(s) = sqrt(dx4(s)^2 + dy4(s)^2 + dz4(s)^2); % measured range for nth satellite in
170             view
171             H(s,:) = [ dx4(s)/drho4(s) dy4(s)/drho4(s) dz4(s)/drho4(s) 0 0 0]; % Assembling
172             epochwise information matrix.
173         end
174         B = H*phi; % New information matrix.
175         % Accumulating primary information matrix from all the epochs into
176         % one matrix
177         A4 = blkdiag(A4,B);
178         A4 = vertcat(A4,B);
179         % Accumulating observation for all epochs
180         % y_bar4 = [y_bar4;transpose(C1(e,1:no_sat(e)) + ccGPS(e,1:no_sat(e)) - drho_rcc) ] ;
181         % With inclusion of receiver clock correction from previous assignment
182         y_bar4 = [y_bar4;transpose(C1(e,1:no_sat(e)) + (ccGPS(e,1:no_sat(e))- final_clock_error*c)
183             - drho4) ] ;
184         n = n+3;
185     end
186
187 % rank check to check for information loss in the information matrix
188 rank4(iter4) = rank(A4);
189
190 % LSQ with the aid of SVD
191 [U4,S4,V4] = svd(A4,'econ'); % Singular value decomposition of A, to avoid singularity errors
192 or solution manifold due to rank deficit

```

```

187 S_inv_diag4 = [];
188 for nn=1:length(S4)
189     S_inv_diag4 = [S_inv_diag4, inv(S4(nn,nn))];
190 end
191 S_inv4 = diag(S_inv_diag4);
192
193 % Estimation of small letter case, s. s = x_hat4
194 x_hat4 = V4*S_inv4*U4'*y_bar4; % Parameter estimation via SVD aided LSQ
195 y_hat4 = A4*x_hat4; % Measured vector
196
197 res = [res, norm(y_bar4-y_hat4)];
198
199 nn = 0;
200 buf = [];
201 for ee = 1:epoch
202     buf = [buf, nn+1, sum(no_sat(1:ee))];
203     res4 = y_bar4(nn+1:sum(no_sat(1:ee))) - y_hat4(nn+1:sum(no_sat(1:ee))); % residual in
204     res_norm4(ee, iter4) = norm(res4); % Epoch wise norm of the residual vector of all satellites
205     nn = sum(no_sat(1:ee));
206 end
207 % Update error and state
208 diff4 = norm(x_hat4(1:3));
209 diff_iter = [diff_iter; diff4];
210 s0 = s0 + x_hat4';
211 if iter4 >= 100
212     disp('Iteration limit of 100 exceeds');
213     break
214 end
215 end
216
217 %% Plotting the behaviour of epochwise residual, inclusion of clock correction before batch LSQ for
218     all epochs
219 for ee = 1:epoch
220     figure(31)
221     subplot(6,2,ee)
222     plot(1:length(res_norm4(ee,:)), res_norm4(ee,:), 'rx');
223     hold on
224     plot(1:length(res_norm4(ee,:)), res_norm4(ee,:), 'k:');
225     hold off
226     title(['Epoch ' num2str(ee)]);
227     ylabel('Res. norm [m]');
228     xlabel('Iteration no. ');
229     grid on
230 end
231
232 nn = 0;
233 for ee = 1:epoch
234     yb = y_bar4(nn+1:sum(no_sat(1:ee)), 1)';
235     yh = y_hat4(nn+1:sum(no_sat(1:ee)), 1)';
236     nn = sum(no_sat(1:ee));
237     figure(32)
238     subplot(6,2,ee)
239     plot(1:length(yb), yb, 'r-x');
240     hold on
241     plot(1:length(yh), yh, 'b-*');
242     hold off
243     title(['Epoch ' num2str(ee)]);
244     ylabel('Range');
245     xlabel('Sat no. ');
246     grid on
247 end
248 %% Part b
249 % Final values of the epochwise distance is stored in the variable
250 % s_final_rcc. Please refer to previous assignment or follow the github
251 % link. While the clock error is stored in del_t variable. Both of them are
252 % used to reconstruct the initial state vector.
253
254 % Loading the initial state vector from previous script
255 pos = load('s_final_rcc', '-mat'); % Position in m
256 clock_errors = load('del_t', '-mat'); % Clock errors in s
257
258 % s_final_rcc = pos.s_final_rcc(end,:); % Final x,y,z position for last epoch with clock correction
259     implemented

```

```

258 s_final_rcc = pos.s_final_rcc(1,:); % Final x,y,z position for last epoch with clock correction
    implemented
259 final_clock_error = clock_errors.del_t(1,end); % Final receiver clock correction
260
261 % Basic backward Euler scheme to evaluate velocity terms in m/s
262 % vel = (pos.s_final_rcc(end,:) - pos.s_final_rcc(end-1,:))/dt;
263 vel = (pos.s_final_rcc(2,:) - pos.s_final_rcc(1,:))/dt;
264 % Initializing state vector: [ x, y, z, x_dot, y_dot, z_dot] with clock
265 % correction implemented
266 s0_buf = [ s_final_rcc, vel, 3e14] + [ 0 0 0 0 0 0 ];
267 % Shifting initial vector slightly to visualize convergence.
268 s0 = s0_buf ;
269 % s0 = [ 0.1 0.1 0.1 0.1 0.1 0.1];
270 % Standard Gravitational parameter Earth
271 % mu = 3.986004418e14; % [ m^3 s^-2]
272 % Loop initializations
273 diff4 = norm(s0); % Current error resolution
274 iter4 = 0; % Current iteration step
275 diff_iter = [];
276 Svect = [];
277 res = [];
278 while diff4 >= 10^-3
279 % while iter4 <= 10
280     A4 = [];
281     y_bar4 = [];
282     iter4 = iter4 + 1;
283     phi = eye(7,7); % Initializing state transition matrix phi
284     S = s0;
285     Svect = [Svect; S];
286     % For all epochs
287     mu = S(7);
288     for e=1:epoch
289         clear H dx4 dy4 dz4 drho4;
290         r = norm(S(1,1:3)); % radial distance sqrt(x0, y0, z0) in meters
291         % Calculating dF/dS matrix for s0 = [ x, y, z, x_dot, y_dot, z_dot ]
292         dF_dS = [ 0 0 0 1 0 0 0; 0 0 0 0 1 0 0; 0 0 0 0 0 1 0; ...
293             (mu/r^3)*( (3*S(1)^2)/(r^2) - 1), 3*mu*S(1)*S(2)/r^5, 3*mu*S(1)*S(3)/r^5, 0,
294             0, 0, -S(1)/r^3; ...
295             3*mu*S(1)*S(2)/r^5, (mu/r^3)*( (3*S(2)^2)/(r^2) - 1), 3*mu*S(2)*S(3)/r^5, 0,
296             0, 0, -S(2)/r^3; ...
297             3*mu*S(1)*S(3)/r^5, 3*mu*S(2)*S(3)/r^5, (mu/r^3)*( (3*S(3)^2)/(r^2) - 1), 0, 0,
298             0, -S(3)/r^3; ...
299             0, 0, 0, 0, 0, 0, 0];
300         % Calculating phi_dot
301         phi_dot = dF_dS*phi;
302         % Updating phi
303         phi = phi + phi_dot*dt;
304         % Calculating S_dot for s0 = [ x, y, z, x_dot, y_dot, z_dot ]
305         S_dot = [ S(4); S(5); S(6); -mu*S(1)/r^3; -mu*S(2)/r^3; -mu*S(3)/r^3; 0 ]';
306         % Updating S
307         S = S + S_dot.*dt;
308
309         for s = 1:no_sat(e)
310             dx4(s) = S(1) - xgi(e,s); % x - x_gps for sth satellite in view
311             dy4(s) = S(2) - ygi(e,s); % y - y_gps for sth satellite in view
312             dz4(s) = S(3) - zgi(e,s); % z - z_gps for sth satellite in view
313
314             drho4(s) = sqrt(dx4(s)^2 + dy4(s)^2 + dz4(s)^2); % measured range for nth satellite in
view
315             H(s,:) = [ dx4(s)/drho4(s) dy4(s)/drho4(s) dz4(s)/drho4(s) 0 0 0 0]; % Assembling
epochwise information matrix.
316         end
317         B = H*phi; % New information matrix.
318         % Accumalating primary information matrix from all the epochs into
319         % one matrix
320         A4 = vertcat(A4,B);
321         % Accumulating observation for all epochs
322         % y_bar4 = [y_bar4;transpose(C1(e,1:no_sat(e)) + ccGPS(e,1:no_sat(e)) - drho_rcc) ] ;
323         % With inclusion of receiver clock correction from previous assignment
324         y_bar4 = [y_bar4;transpose(C1(e,1:no_sat(e)) + (ccGPS(e,1:no_sat(e))- final_clock_error*c
- drho4) ) ] ;
325     end
326
327 % rank check to check for information loss in the information matrix

```

```

326 rank4(iter4) = rank(A4);
327
328 % LSQ with the aid of SVD
329 [U4,S4,V4] = svd(A4,'econ'); % Singular value decomposition of A, to avoid singularity errors
    or solution manifold due to rank deficit
330 S_inv_diag4 = [];
331 for nn=1:length(S4)
332     S_inv_diag4 = [S_inv_diag4, inv(S4(nn,nn))];
333 end
334 S_inv4 = diag(S_inv_diag4);
335
336 % Estimation of small letter case, s. s = x_hat4
337 x_hat4 = V4*S_inv4*U4'*y_bar4; % Parameter estimation via SVD aided LSQ
338 y_hat4 = A4*x_hat4; % Measured vector
339
340 res = [res, norm(y_bar4-y_hat4)];
341
342 nn = 0;
343 buf = [];
344 for ee = 1:epoch
345     buf = [buf, nn+1, sum(no_sat(1:ee))];
346     res4 = y_bar4(nn+1:sum(no_sat(1:ee))) - y_hat4(nn+1:sum(no_sat(1:ee))); % residual in
    meters
347     res_norm4(ee, iter4) = norm(res4); % Epoch wise norm of the residual vector of all satellites
348     nn = sum(no_sat(1:ee));
349 end
350 % Update error and state
351 diff4 = norm(x_hat4(1:3));
352 diff_iter = [diff_iter; diff4];
353 s0 = s0 + x_hat4';
354 if iter4 >= 100
355     disp('Iteration limit of 100 exceeds');
356     break
357 end
358 end
359
360 %% Plotting the behaviour of epochwise residual, inclusion of clock correction before batch LSQ for
    all epochs
361 for ee = 1:epoch
362     figure(33)
363     subplot(6,2,ee)
364     plot(1:length(res_norm4(ee,:)), res_norm4(ee,:), 'rx');
365     hold on
366     plot(1:length(res_norm4(ee,:)), res_norm4(ee,:), 'k:');
367     hold off
368     title(['Epoch ' num2str(ee)]);
369     ylabel('Res. norm [m]');
370     xlabel('Iteration no. ');
371     grid on
372 end
373
374 nn = 0;
375 for ee = 1:epoch
376     yb = y_bar4(nn+1:sum(no_sat(1:ee)), 1)';
377     yh = y_hat4(nn+1:sum(no_sat(1:ee)), 1)';
378     nn = sum(no_sat(1:ee));
379     figure(34)
380     subplot(6,2,ee)
381     plot(1:length(yb), yb, 'r-x');
382     hold on
383     plot(1:length(yh), yh, 'b-*');
384     hold off
385     title(['Epoch ' num2str(ee)]);
386     ylabel('Range');
387     xlabel('Sat no. ');
388     grid on
389 end
390
391 %% Part C Standard deviation estimation.
392 % Assumed standard deviation of observation vector
393 std_y = 5;
394 % Observation vector co-variance matrix
395 Pyy = (std_y^2)*eye(87,87);
396 % State co-variance matrix
397 Pxx = pinv(A4)*Pyy*pinv(A4)';

```

```
398 % Standard deviation of gravitational parameter
399 std_mu = sqrt(Pxx(end,end));
400 % mu combined with its standard deviation
401 mu_with_std = s0(7)+std_mu;
402 % Percent difference between estimated and actual mu.
403 mu_percent_diff = ((3.986004419*10^(14) - mu_with_std)/(3.986004419*10^(14)) ) * 100;
```