Computational Fluid Dynamics HW2

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Nomenclature

Δx	size of each cell in the domain
$\Delta ilde{t}$	normalized step size in time
$\Delta \tilde{x}$	normalized step size in space
γ	ratio of specific heats
κ	coefficient of thermal conductivity
μ	coefficient of viscosity
ρ	fluid density
c_p	constant specific heat capacity for a constant pressure
c_v	constant specific heat capacity for a constant volume
E	inviscid convective vector
e	total energy
E_{ν}	viscous convective vector
L	characteristic length
p	pressure
Q	conservation state space
R	gas constant
T	temperature
t	time
u	fluid velocity
x	spatial coordinate
x_F	x coordinate of the end of the domain
x_i	x coordinate of the i-th cell
Far-Aw	vay Properties
κ_{∞}	coefficient of thermal conductivity far away
μ_{∞}	coefficient of viscosity far away
$ ho_{\infty}$	density far away
a_{∞}	speed of sound far away
M_{∞}	mach number far away

pressure far away

 p_{∞}



 T_{∞} temperature far away

Matrices

 $\tilde{\Lambda}$ normalized eigenvalues matrix

 \tilde{A} normalized jacobian matrix of E w.r.t. Q

 \tilde{P} normalized jacobian matrix of E_{ν} w.r.t. Q

 \tilde{R} normalized jacobian matrix of E_{ν} w.r.t. Q_x

 \tilde{T} normalized eigenvectors matrix

Dimensionless Numbers

 Pr_{∞} Prandtl number far away

 $Re_{L\infty}$ Reynolds number with respect to L far away

1 Problem Definition

1.1 Governing Equations

Consider the one-dimensional Navier-Stokes Equations:

$$\frac{\partial Q}{\partial t} + \frac{\partial E}{\partial x} = \frac{\partial E_{\nu}}{\partial x} \tag{1}$$

Where:

$$Q = \begin{pmatrix} \rho \\ \rho u \\ e \end{pmatrix}, \quad E = \begin{pmatrix} \rho u \\ p + \rho u^2 \\ (e + p) u \end{pmatrix}, \quad E_{\nu} = \begin{pmatrix} 0 \\ \tau_{xx} \\ u\tau_{xx} - q_x \end{pmatrix} = \begin{pmatrix} \frac{4}{3}\mu \frac{\partial u}{\partial x} \\ \frac{4}{3}\mu u \frac{\partial u}{\partial x} + \kappa \frac{\partial T}{\partial x} \end{pmatrix}$$

$$p = (\gamma - 1) \left(e - \frac{1}{2}\rho u^2 \right), \qquad T = \frac{p}{\rho R},$$

$$\mu = 1.458 \cdot 10^{-6} \frac{T^{\frac{3}{2}}}{T + 110.4}, \quad \kappa = 2.495 \cdot 10^{-3} \frac{T^{\frac{3}{2}}}{T + 194}$$

$$R = c_p - c_v, \quad \gamma = \frac{c_p}{c_v}$$

$$(2)$$

The constants are:

- $\gamma = 1.4$ for air under standard atmospheric conditions
- R = 287.0 for air

1.2 Physical Domain

The physical domain is a tube extended between x = 0.2 and x = 1.0. At both ends there are impermeable walls.

1.3 Initial Conditions

The initial conditions are shown in Fig.1:

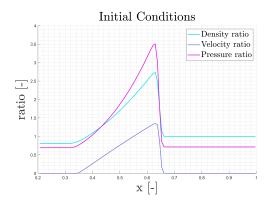


Figure 1: Initial conditions



1.4 Boundary Conditions

On each side of the tube there is an adiabatic, solid wall boundary conditions.

$$u_{(x=0.2)} = u_{(x=1.0)} = 0 \qquad \left\| \frac{\partial p}{\partial x} \right|_{x=0.2} = \left. \frac{\partial p}{\partial x} \right|_{x=1.0} = 0 \qquad \left\| \frac{\partial T}{\partial x} \right|_{x=0.2} = \left. \frac{\partial T}{\partial x} \right|_{x=1.0} = 0$$

2 Normalizing The Navier-Stokes Equations

Since the initial conditions are normalized, there is a need to normalize the N-S equations. We will use the following normalizations:

$$\rho = \rho_{\infty}\tilde{\rho}, \quad u = a_{\infty}\tilde{u}, \quad p = \gamma p_{\infty}\tilde{p}, \quad T = \gamma T_{\infty}\tilde{T}, \quad x = L\tilde{x}, \quad t = \frac{L}{a_{\infty}}\tilde{t}, \quad \mu = \mu_{\infty}\tilde{\mu}, \quad \kappa = \kappa_{\infty}\tilde{\kappa}$$
(3)

The normalization of the temperature was chosen to cancel out the γ in the normalization of the pressure:

$$p = \rho RT$$

$$\gamma p_{\infty} \tilde{p} = \rho_{\infty} \tilde{\rho} R \gamma T_{\infty} \tilde{T}$$

$$\tilde{p} = \tilde{\rho} \tilde{T}$$
(4)

The pressure normalization can be written also as

$$p = \gamma p_{\infty} \tilde{p} = \gamma \rho_{\infty} R T_{\infty} \tilde{p} = \rho_{\infty} a_{\infty}^{2} \tilde{p} \tag{5}$$

From equations 2 and 5 we can derive the normalization for the energy:

$$e = \frac{p}{\gamma - 1} + \frac{1}{2}\rho u^{2}$$

$$e = \frac{\rho_{\infty} a_{\infty}^{2} \tilde{p}}{\gamma - 1} + \frac{1}{2}\rho_{\infty} \tilde{\rho} a_{\infty}^{2} \tilde{a}^{2}$$

$$e = \rho_{\infty} a_{\infty}^{2} \left(\frac{\tilde{p}}{\gamma - 1} + \frac{1}{2} \tilde{\rho} \tilde{a}^{2}\right)$$

$$e = \rho_{\infty} a_{\infty}^{2} \tilde{e}$$

$$(6)$$

After substituting the normalizations in the N-S equations we get:

$$\frac{\partial}{\partial \frac{L}{a_{\infty}}\tilde{t}} \begin{pmatrix} \rho_{\infty}\tilde{\rho} \\ \rho_{\infty}a_{\infty}\tilde{\rho}\tilde{u} \\ \rho_{\infty}a_{\infty}^{2}\tilde{\rho}\tilde{u} \end{pmatrix} + \frac{\partial}{\partial L\tilde{x}} \begin{pmatrix} \rho_{\infty}a_{\infty}\tilde{\rho}\tilde{u} \\ \rho_{\infty}a_{\infty}^{2}\tilde{p} + \rho_{\infty}a_{\infty}^{2}\tilde{\rho}\tilde{u}^{2} \\ \rho_{\infty}a_{\infty}^{3}\left(\tilde{e} + \tilde{p}\right)\tilde{u} \end{pmatrix} = \frac{\partial}{\partial L\tilde{x}} \begin{pmatrix} 0 \\ \frac{4}{3}\mu_{\infty}a_{\infty}\tilde{\mu}\frac{\partial\tilde{u}}{\partial L\tilde{x}} \\ \frac{4}{3}\mu_{\infty}a_{\infty}\tilde{\mu}\frac{\partial\tilde{u}}{\partial L\tilde{x}} + \frac{\kappa_{\infty}a_{\infty}^{2}}{R}\tilde{\kappa}\frac{\partial\tilde{T}}{\partial L\tilde{x}} \end{pmatrix}$$

Rearranging:

$$\frac{\rho_{\infty}a_{\infty}}{L}\frac{\partial}{\partial \tilde{t}}\begin{pmatrix} \tilde{\rho} \\ a_{\infty}\tilde{\rho}\tilde{u} \\ a_{\infty}^{2}\tilde{e} \end{pmatrix} + \frac{\rho_{\infty}a_{\infty}}{L}\frac{\partial}{\partial \tilde{x}}\begin{pmatrix} \tilde{\rho}\tilde{u} \\ a_{\infty}\tilde{p} + a_{\infty}\tilde{\rho}\tilde{u}^{2} \\ a_{\infty}^{2}\left(\tilde{e} + \tilde{p}\right)\tilde{u} \end{pmatrix} = \frac{\mu_{\infty}}{L^{2}}\frac{\partial}{\partial \tilde{x}}\begin{pmatrix} \frac{4}{3}a_{\infty}\tilde{\mu}\frac{\partial\tilde{u}}{\partial\tilde{x}} \\ \frac{4}{3}a_{\infty}\tilde{\mu}\frac{\partial\tilde{u}}{\partial\tilde{x}} + \frac{\kappa_{\infty}a_{\infty}^{2}}{\mu_{\infty}R}\tilde{\kappa}\frac{\partial\tilde{T}}{\partial\tilde{x}} \end{pmatrix} (8)$$



Dividing the second equation by a_{∞} , the third equation by a_{∞}^2 , and the whole set of equations by $\frac{\rho_{\infty}a_{\infty}}{L}$ we get:

$$\frac{\partial}{\partial \tilde{t}} \begin{pmatrix} \tilde{\rho} \\ \tilde{\rho}\tilde{u} \\ \tilde{e} \end{pmatrix} + \frac{\partial}{\partial \tilde{x}} \begin{pmatrix} \tilde{\rho}\tilde{u} \\ \tilde{p} + \tilde{\rho}\tilde{u}^{2} \\ (\tilde{e} + \tilde{p})\tilde{u} \end{pmatrix} = \frac{\mu_{\infty}}{L\rho_{\infty}a_{\infty}} \frac{\partial}{\partial \tilde{x}} \begin{pmatrix} 0 \\ \frac{4}{3}\tilde{\mu}\frac{\partial \tilde{u}}{\partial \tilde{x}} \\ \frac{4}{3}\tilde{\mu}\tilde{u}\frac{\partial \tilde{u}}{\partial \tilde{x}} + \frac{\kappa_{\infty}}{\mu_{\infty}R}\tilde{\kappa}\frac{\partial \tilde{T}}{\partial \tilde{x}} \end{pmatrix} \tag{9}$$

The Reynolds number and the mach number far away are defined as:

$$M_{\infty} = \frac{u_{\infty}}{a_{\infty}} \quad Re_{L_{\infty}} = \frac{\rho_{\infty} u_{\infty} L}{\mu_{\infty}}$$

$$\downarrow \qquad \qquad \downarrow$$

$$\frac{\mu_{\infty}}{L \rho_{\infty} a_{\infty}} = \frac{M_{\infty}}{Re_{L_{\infty}}}$$
(10)

The Prandtl number far away is defined as:

$$Pr_{\infty} = \frac{c_{p}\mu_{\infty}}{\kappa_{\infty}}$$

$$\frac{\kappa_{\infty}}{\mu_{\infty}R} = \frac{c_{p}}{Pr_{\infty}(c_{p} - c_{v})} = \frac{\gamma}{Pr_{\infty}(\gamma - 1)}$$
(11)

Substituting into the normalized N-S equations:

$$\frac{\partial \tilde{Q}}{\partial \tilde{t}} + \frac{\partial \tilde{E}}{\partial \tilde{x}} = \frac{M_{\infty}}{Re_{L_{\infty}}} \frac{\partial \tilde{E}_{\nu}}{\partial \tilde{x}}$$
(12)

Where:

$$\tilde{Q} = \begin{pmatrix} \tilde{\rho} \\ \tilde{\rho}\tilde{u} \\ \tilde{e} \end{pmatrix}, \quad \tilde{E} = \begin{pmatrix} \tilde{\rho}\tilde{u} \\ \tilde{p} + \tilde{\rho}\tilde{u}^{2} \\ (\tilde{e} + \tilde{p})\tilde{u} \end{pmatrix}, \quad \tilde{E}_{\nu} = \begin{pmatrix} 0 \\ \frac{4}{3}\tilde{\mu}\frac{\partial \tilde{u}}{\partial \tilde{x}} \\ \frac{4}{3}\tilde{\mu}\tilde{u}\frac{\partial \tilde{u}}{\partial \tilde{x}} + \frac{\gamma}{Pr_{\infty}(\gamma - 1)}\tilde{\kappa}\frac{\partial \tilde{T}}{\partial \tilde{x}} \end{pmatrix}$$
(13)

The normalized Navier-Stokes equations are:

$$\frac{\partial \tilde{Q}}{\partial \tilde{t}} + \frac{\partial \tilde{E}}{\partial \tilde{x}} = \frac{M_{\infty}}{Re_{L\infty}} \frac{\partial \tilde{V}_{1}}{\partial \tilde{x}}$$
(14)

Where:

$$\tilde{V}_1 = \tilde{V}_{1\left(\tilde{Q}, \tilde{Q}_x\right)} = \tilde{E}_{\nu}$$

3 The Computational Domain

3.1 Discretization

The physical domain $[x_I, x_F]$ is discretized into N equispaced cells. The size of each cell is there for:

$$\Delta x = \frac{x_F - x_I}{N} = \frac{L}{N} \tag{15}$$

so the x coordinate of the i-th cell x_i is:

$$x_i = x_I + \frac{1}{2}\Delta x + \Delta x \cdot (i-1)$$
 when starting from $i = 1$ (16)



3.2 Boundary Conditions

In order to set the boundary conditions on the edge faces we will define ghost cells that will be calculated like so:

$$\begin{array}{rcl}
 u_{(i=0)} & = & -u_{(i=1)} \\
 u_{(i=N+1)} & = & -u_{(i=N)}
 \end{array}
 \tag{17}$$

in order to maintain velocity zero on the boundary and like so:

$$T_{(i=0)} = T_{(i=1)}$$

 $T_{(i=N+1)} = T_{(i=N)}$
(18)

In order to maintain adiabatic boundary conditions. Since the gradient of the pressure on the wall is zero, we get:

$$p_{(i=0)} = p_{(i=1)}$$

 $p_{(i=N+1)} = p_{(i=N)}$
(19)

From equations 2, 18, and 19 we can conclude:

$$\begin{array}{rcl}
\rho_{(i=0)} & = & \rho_{(i=1)} \\
\rho_{(i=N+1)} & = & \rho_{(i=N)}
\end{array}$$
(20)

and from equations 2, 17, 19, and 20 we can conclude:

$$e_{(i=0)} = e_{(i=1)}$$

 $e_{(i=N+1)} = e_{(i=N)}$ (21)

4 The Numerical Schemes

4.1 Jacobian Matrices of The Navier-Stokes Equations

$$\frac{\Delta Q}{\Delta t} = -\left(\frac{\partial \tilde{E}}{\partial \tilde{x}} - \frac{\partial \tilde{V}_1}{\partial \tilde{x}}\right)^{n+1}$$

$$\frac{\Delta Q}{\Delta t} = -\left(\underbrace{\frac{\partial \tilde{E}}{\partial \tilde{Q}}}_{\tilde{A}} \frac{\partial \tilde{Q}}{\partial \tilde{x}} - \frac{M_{\infty}}{Re_{L\infty}} \left(\underbrace{\frac{\partial \tilde{V}_1}{\partial \tilde{Q}}}_{\tilde{P}} \frac{\partial \tilde{Q}}{\partial \tilde{x}} + \underbrace{\frac{\partial \tilde{V}_1}{\partial \tilde{Q}_x}}_{\tilde{R}} \frac{\partial \tilde{Q}_x}{\partial \tilde{x}}\right)\right)^{n+1} \tag{22}$$

Where:

$$\tilde{A} = \begin{pmatrix}
0 & 1 & 0 \\
\frac{\gamma - 3}{2}\tilde{u}^{2} & (3 - \gamma)\tilde{u} & \gamma - 1 \\
-\frac{\gamma\tilde{e}\tilde{u}}{\tilde{\rho}} - (\gamma - 1)\tilde{u}^{3} & \frac{\gamma\tilde{e}}{\tilde{\rho}} - \frac{3(\gamma - 1)\tilde{u}^{2}}{2} & \gamma\tilde{u}
\end{pmatrix}$$

$$\tilde{P} - \tilde{R}_{x} = -\frac{1}{\rho} \begin{pmatrix}
0 & 0 & 0 \\
-\tilde{u}\left(\frac{4}{3}\tilde{\mu}\right)_{x} & \left(\frac{4}{3}\tilde{\mu}\right)_{x} & 0 \\
-\tilde{u}^{2}\left(\frac{4}{3}\tilde{\mu}\right)_{x} & \tilde{u}\left(\frac{4}{3}\tilde{\mu}\right)_{x} & 0
\end{pmatrix}$$

$$\tilde{R} = -\frac{1}{\rho} \begin{pmatrix}
0 & 0 & 0 \\
-\tilde{u}^{2}\left(\frac{4}{3}\tilde{\mu}\right)_{x} & \tilde{u}\left(\frac{4}{3}\tilde{\mu}\right)_{x} & 0
\end{pmatrix}$$

$$\frac{4}{3}\tilde{u}\tilde{\mu}} & -\frac{4}{3}\tilde{\mu} & 0 \\
\left(\frac{4}{3}\tilde{\mu} - \alpha\frac{\tilde{\kappa}}{c_{v}}\right)\tilde{u}^{2} + \alpha\frac{\tilde{\kappa}}{c_{v}}\frac{\tilde{e}}{\tilde{\rho}} & -\left(\frac{4}{3}\tilde{\mu} - \alpha\frac{\tilde{\kappa}}{c_{v}}\right)\tilde{u} & -\alpha\frac{\tilde{\kappa}}{c_{v}}\right)$$

and α is:

$$\alpha = \frac{\gamma}{Pr_{\infty} \left(\gamma - 1\right)}$$

4.2 Linearization In Time

4.2.1 \tilde{E}_i^{n+1} Estimation

$$\tilde{E}_{i}^{n+1} = \tilde{E}_{i}^{n} + \underbrace{\frac{\partial \tilde{E}}{\partial \tilde{Q}} \Big|_{i}^{n}}_{\tilde{A}_{i}^{n}} \Delta \tilde{Q}_{i}^{n} + \text{H.O.T}$$

$$\tilde{E}_{i}^{n+1} = \tilde{E}_{i}^{n} + \tilde{A}_{i}^{n} \Delta \tilde{Q}_{i}^{n}$$
(24)

4.2.2 $\tilde{V_1}_i^{n+1}$ Estimation

$$\tilde{V}_{1,i}^{n+1} = \tilde{V}_{1,i}^{n} + \underbrace{\frac{\partial \tilde{V}_{1}}{\partial \tilde{Q}} \bigg|_{i}^{n}}_{\tilde{Q}_{i}} \Delta \tilde{Q}_{i}^{n} + \underbrace{\frac{\partial \tilde{V}_{1}}{\partial \tilde{Q}_{x}} \bigg|_{i}^{n}}_{\tilde{R}_{i}^{n}} \Delta \tilde{Q}_{x_{i}}^{n} + \text{H.O.T}$$

$$\tilde{V}_{1,i}^{n+1} = \tilde{V}_{1,i}^{n} + \tilde{P}_{i}^{n} \Delta \tilde{Q}_{i}^{n} + \tilde{R}_{i}^{n} \Delta \tilde{Q}_{x_{i}}^{n}$$
(25)

The difficulty stems from the fact that the solution vector is $\Delta \tilde{Q}$ and not $\Delta \tilde{Q}_x$. This can be solved by a linearization of the term $\Delta \tilde{Q}_x$ which can be conducted using the following relation:

$$\frac{\partial \left(\tilde{R}\Delta\tilde{Q}\right)_{i}^{n}}{\partial \tilde{x}} = \frac{\partial \tilde{R}_{i}^{n}}{\partial \tilde{x}}\Delta\tilde{Q}_{i}^{n} + \tilde{R}_{i}^{n}\frac{\partial \Delta\tilde{Q}_{i}^{n}}{\partial \tilde{x}} = \frac{\partial \tilde{R}}{\partial \tilde{x}}\Delta\tilde{Q} + \tilde{R}_{i}^{n}\Delta\tilde{Q}_{x_{i}}^{n}$$

$$\psi$$

$$\tilde{V}_{1,i}^{n+1} = \tilde{V}_{1,i}^{n} + \left(\tilde{P} - \tilde{R}_{x}\right)_{i}^{n} \Delta \tilde{Q}_{i}^{n} + \frac{\partial}{\partial \tilde{x}} \left(\tilde{R}\Delta \tilde{Q}\right)_{i}^{n} \tag{26}$$

4.3 First Order Approximate Riemann Roe Method

4.4 First Order Steger-Warming – Explicit

A is a Diagonalizable matrix and can be written as:

$$\tilde{A} = \tilde{T}\tilde{\Lambda}\tilde{T}^{-1}$$

$$1 \qquad \frac{\tilde{\rho}}{2\tilde{a}} \qquad -\frac{\tilde{\rho}}{2\tilde{a}}$$

$$\tilde{u} \qquad \frac{\tilde{\rho}}{2\tilde{a}}(\tilde{u} + \tilde{a}) \qquad -\frac{\tilde{\rho}}{2\tilde{a}}(\tilde{u} - \tilde{a})$$

$$\frac{\tilde{u}^2}{2} \quad \frac{\tilde{\rho}}{2\tilde{a}}\left(\frac{\tilde{u}^2}{2} + \tilde{u}\tilde{a} + \frac{\tilde{a}^2}{\gamma - 1}\right) \quad -\frac{\tilde{\rho}}{2\tilde{a}}\left(\frac{\tilde{u}^2}{2} - \tilde{u}\tilde{a} + \frac{\tilde{a}^2}{\gamma - 1}\right)$$

$$\tilde{\Lambda} = \begin{pmatrix} \tilde{u} & 0 & 0 \\ 0 & \tilde{u} + \tilde{a} & 0 \\ 0 & 0 & \tilde{u} - \tilde{a} \end{pmatrix}$$

$$\tilde{T}^{-1} = \begin{pmatrix} 1 - \frac{\gamma - 1}{2}\frac{\tilde{u}^2}{\tilde{a}^2} & (\gamma - 1)\frac{\tilde{u}^2}{\tilde{a}^2} & -\frac{\gamma - 1}{\tilde{a}^2} \\ \frac{1}{\tilde{\rho}\tilde{a}}\left((\gamma - 1)\tilde{u}^2 - \tilde{u}\tilde{a}\right) & \frac{1}{\tilde{\rho}\tilde{a}}\left(\tilde{a} - (\gamma - 1)\tilde{u}\right) & \frac{\gamma - 1}{\tilde{\rho}\tilde{a}} \\ -\frac{1}{\tilde{\rho}\tilde{a}}\left((\gamma - 1)\tilde{u}^2 + \tilde{u}\tilde{a}\right) & \frac{1}{\tilde{\rho}\tilde{a}}\left(\tilde{a} + (\gamma - 1)\tilde{u}\right) & -\frac{\gamma - 1}{\tilde{\rho}\tilde{a}} \end{pmatrix}$$

Where:

$$\tilde{a} = \sqrt{\frac{\gamma \tilde{p}}{\tilde{\rho}}}$$

Let the Λ^{\pm} matrix be defined as:

$$\tilde{\Lambda}^{\pm} = \begin{pmatrix} \frac{\tilde{u} \pm |\tilde{u}|}{2} & 0 & 0\\ 0 & \frac{\tilde{u} + \tilde{a} \pm |\tilde{u} + \tilde{a}|}{2} & 0\\ 0 & 0 & \frac{\tilde{u} - \tilde{a} \pm |\tilde{u} - \tilde{a}|}{2} \end{pmatrix}$$
(28)

Where the matrix $\tilde{\Lambda}^+$ contains only positive eigenvalues and the matrix $\tilde{\Lambda}^-$ contains only negative eigenvalues.

Define:

$$\tilde{A}^{+} \triangleq \tilde{T}\tilde{\Lambda}^{+}\tilde{T}^{-1}
\tilde{A}^{-} \triangleq \tilde{T}\tilde{\Lambda}^{-}\tilde{T}^{-1} \Rightarrow \begin{vmatrix} \tilde{A} & = \tilde{A}^{+} + \tilde{A}^{-} \\ |\tilde{A}| & \triangleq \tilde{A}^{+} - \tilde{A}^{-} \end{vmatrix}$$
(29)

Assuming a perfect gas, the flux vector $\tilde{E}_{(Q)}$ is a homogeneous function of degree one in \tilde{Q} , meaning:

$$\forall \alpha \quad \tilde{E}_{(\alpha \tilde{Q})} = \alpha \tilde{E}_{(\tilde{Q})}$$

The homogeneity allows to rewrite the flux vector \tilde{E} as:

$$\tilde{E} = \tilde{A}\tilde{Q} = \left(\tilde{A}^{+} + \tilde{A}^{-}\right)\tilde{Q} = \underbrace{\tilde{A}^{+}\tilde{Q}}_{\tilde{E}^{+}} + \underbrace{\tilde{A}^{-}\tilde{Q}}_{\tilde{E}^{-}} = \tilde{E}^{+} + \tilde{E}^{-}$$
(30)

There is a discontinuities and deference between \tilde{E}^+, \tilde{E}^- . To eliminate the discontinuities and guarantee a smooth transition through critical points (sonic points or stagnation points), a blending function is introduced together with a blending parameter ε . An appropriate choice of the blending parameter has to be chosen.

$$\tilde{\lambda}^{+} = \frac{\tilde{\lambda} + \left| \tilde{\lambda} \right|}{2} \qquad \qquad \tilde{\lambda}^{+'} = \frac{\tilde{\lambda} + \sqrt{\tilde{\lambda}^{2} + \varepsilon^{2}}}{2}$$

$$\Rightarrow \qquad \qquad \tilde{\lambda}^{-} = \frac{\tilde{\lambda} - \left| \tilde{\lambda} \right|}{2} \qquad \qquad \tilde{\lambda}^{-'} = \frac{\tilde{\lambda} - \sqrt{\tilde{\lambda}^{2} + \varepsilon^{2}}}{2}$$
(31)

Rewriting the conservation law from of the N-S equations:

$$\frac{\partial \tilde{Q}}{\partial \tilde{t}} = -\frac{\partial \tilde{E}^{+}}{\partial \tilde{x}} - \frac{\partial \tilde{E}^{-}}{\partial \tilde{x}} + \frac{M_{\infty}}{Re_{L\infty}} \frac{\partial \tilde{V}_{1}}{\partial \tilde{x}}$$
(32)

A simple, explicit, first order (in space and time) scheme in delta from is obtained using:

$$\Delta \tilde{Q}_{i}^{n} = -\frac{\Delta \tilde{t}}{\Delta \tilde{x}} \left(\nabla \tilde{E}_{i}^{+n} + \Delta \tilde{E}_{i}^{-n} - \frac{M_{\infty}}{Re_{L_{\infty}}} \delta \tilde{V}_{1,i}^{n} \right)$$
(33)

And advancing the solution by:

$$\tilde{Q}_i^{n+1} = \Delta \tilde{Q}_i^n + \tilde{Q}_i^n \tag{34}$$

4.5 First Order Steger-Warming – Implicit

The Implicit Steger-Warming scheme starts from:

$$\frac{\Delta \tilde{Q}_{i}^{n}}{\Delta \tilde{t}} = -\frac{\partial \tilde{E}_{i}^{n+1}}{\partial \tilde{x}} + \frac{M_{\infty}}{Re_{L\infty}} \frac{\partial \tilde{V}_{1,i}^{n+1}}{\partial \tilde{x}}$$

$$\Delta \tilde{Q}_{i}^{n} = -\Delta \tilde{t} \frac{\partial}{\partial \tilde{x}} \left(\tilde{E}_{i}^{n} + \tilde{A}_{i}^{n} \Delta \tilde{Q}_{i}^{n} \right) + \Delta \tilde{t} \frac{M_{\infty}}{Re_{L\infty}} \frac{\partial}{\partial \tilde{x}} \left(\tilde{V}_{1,i}^{n} + \left(\tilde{P} - \tilde{R}_{x} \right)_{i}^{n} \Delta \tilde{Q}_{i}^{n} + \frac{\partial}{\partial \tilde{x}} \left(\tilde{R} \Delta \tilde{Q} \right)_{i}^{n} \right) \tag{35}$$

Rearranging in delta form:

$$\left(I + \Delta \tilde{t} \left(\frac{\partial}{\partial \tilde{x}} \left[\tilde{A} - \left(\tilde{P} - \tilde{R}_{x}\right)\right]_{i}^{n} - \frac{\partial^{2} \tilde{R}_{i}^{n}}{\partial \tilde{x}^{2}}\right)\right) \Delta \tilde{Q}_{i}^{n} = RHS_{i}^{n}$$

$$RHS_{i}^{n} = -\Delta \tilde{t} \frac{\partial}{\partial \tilde{x}} \left(\tilde{E}^{+} + \tilde{E}^{-} - \frac{M_{\infty}}{Re_{L_{\infty}}} \tilde{V}_{1}\right)_{i}^{n}$$

$$\left(I + \frac{\Delta \tilde{t}}{\Delta \tilde{x}} \left(\left[\nabla \tilde{A}_{i}^{+n} + \Delta \tilde{A}_{i}^{-n} - \delta \left(\tilde{P} - \tilde{R}_{x}\right)\right]_{i}^{n} - \frac{\delta^{2}}{\Delta \tilde{x}} \tilde{R}_{i}^{n}\right)\right) \Delta \tilde{Q}_{i}^{n} = RHS_{i}^{n}$$

$$RHS_{i}^{n} = -\frac{\Delta \tilde{t}}{\Delta \tilde{x}} \left(\nabla \tilde{E}^{+} + \Delta \tilde{E}^{-} - \frac{M_{\infty}}{Re_{L_{\infty}}} \delta \tilde{V}_{1}\right)_{i}^{n}$$
(37)