

# Satellite Orbit Control

## HW9

Almog Dobrescu

ID 214254252

February 7, 2025



## Contents

<b>1</b>	<b>Given</b>	<b>2</b>
1.1	Desired . . . . .	2
1.2	Limitations . . . . .	2
<b>2</b>	<b>Control</b>	<b>2</b>
2.1	East-West Control . . . . .	2
2.1.1	Drift Duo to axial asymmetry . . . . .	2
2.2	Drift due to The Tesseral Harmonics . . . . .	3
<b>3</b>	<b>The Results</b>	<b>4</b>
3.1	Part A . . . . .	4
3.2	Part C . . . . .	4

## List of Figures



# 1 Given

A communication satellite is placed in a geostationary orbit above longitude 30°E.

$$T = 8.6164 \cdot 10^4 [\text{sec}] \rightarrow a = 4.2164 \cdot 10^4 [\text{km}] \quad \lambda_n = 30^\circ \quad e = 0 \quad i = 0 \quad (1)$$

The longitude tolerance is  $\Delta\lambda = \pm 0.05^\circ$

## 1.1 Desired

## 1.2 Limitations

The thrust is only at the  $\pm y$  direction and:

$$a_{\max} = 1 \cdot 10^{-7} \left[ \frac{\text{km}}{\text{sec}^2} \right] \quad (2)$$

# 2 Control

## 2.1 East-West Control

### 2.1.1 Drift Duo to axial asymmetry

The gravity component normal to the radius changes the orbit period so the satellite moves toward the closest equilibrium point. The force component creates a tangential acceleration, which is equivalent to the angular acceleration of the longitude line underneath the satellite:

$$a\ddot{\lambda} = \vec{f} \cdot \vec{v} \quad (3)$$

The change in the semi-major axis and the longitude line over one cycle:

$$\begin{aligned} \frac{da}{dt} &= -12anJ_{22} \left( \frac{Re}{a} \right)^2 \sin(2(\lambda - \lambda_{22})) \\ \frac{d^2\lambda}{dt^2} &= -\frac{3n}{2a} \frac{da}{dt} \end{aligned} \quad (4)$$

Where:

$$\begin{aligned} J_{22} &= (C_{22}^2 + S_{22}^2)^{\frac{1}{2}} \\ \lambda_{22} &= \tan^{-1} \left( \frac{S_{22}}{C_{22}} \right) = -14.9^\circ \end{aligned} \quad (5)$$

After substitution:

$$\begin{aligned} \dot{a} &= 0.132 \sin(2(\lambda - \lambda_s)) \left[ \frac{\text{km}}{\text{day}} \right] \\ \ddot{\lambda} &= -0.00168 \sin(2(\lambda - \lambda_s)) \left[ \frac{\text{deg}}{\text{day}^2} \right] \end{aligned} \quad (6)$$

$\lambda_s$  is the closest stable equilibrium point 75° or 255°



Since the tolerance is small:

$$\ddot{\lambda} = -K \sin(2(\lambda_n - \lambda_s)) = \ddot{\lambda}_n$$

$$\lambda = \frac{1}{2\ddot{\lambda}_n} \left( \dot{\lambda} - \dot{\lambda}_0 \right)^2 + \frac{\dot{\lambda}_0}{\ddot{\lambda}_n} + \lambda_0 \quad (7)$$

The initial conditions of the satellite:

$$\lambda_0 = \lambda_n - \Delta\lambda \quad \dot{\lambda}_0 = -2 \operatorname{sgn}(\ddot{\lambda}_n) \sqrt{-\ddot{\lambda}_n \Delta\lambda} \quad (8)$$

The velocit pulse needes to change  $-\dot{\lambda}_0$  into  $\dot{\lambda}_0$ :

$$\Delta n = 2\dot{\lambda}_0 \quad (9)$$

The relation between the change in angular velocit and the semi-major axis:

$$n^2 = \frac{\mu}{a^3} \rightarrow \frac{\Delta a}{a} = -\frac{2}{3} \frac{\Delta n}{n}$$

$$\downarrow$$

$$\Delta v = \left| \frac{\mu}{2a^2 v} \Delta a \right| = \frac{a}{3} |\Delta n| = \frac{2a}{3} |\dot{\lambda}_0| \quad (10)$$

and the friquency of pulses:

$$t_m = 2 \left| \frac{\dot{\lambda}_0}{\ddot{\lambda}_n} \right| \quad (11)$$

The change in altitude between to corections:

$$\Delta a = 0.132 \sin(2(\lambda - \lambda_s)) t_m [\text{km}] \quad (12)$$

## 2.2 Drift due to The Tesseral Harmonics



### 3 The Results

#### 3.1 Part A

$$\begin{aligned}
 \ddot{\lambda}_n &= -0.00168 \sin(2(30^\circ - 75^\circ)) = 0.0017 \left[ \frac{\text{deg}}{\text{day}^2} \right] \\
 \dot{\lambda}_0 &= -2 \operatorname{sgn}(\ddot{\lambda}_n) \sqrt{-\ddot{\lambda}_n \Delta \lambda} = -0.0183 \left[ \frac{\text{deg}}{\text{day}} \right] \\
 &\quad \downarrow \\
 t_m &= 2 \left| \frac{\dot{\lambda}_0}{\ddot{\lambda}_n} \right| = 2 \left| \frac{-0.0183}{0.0017} \right| = 21.8218 [\text{day}] \\
 &\quad \downarrow \\
 \Delta a &= 0.132 \sin(2(30^\circ - 75^\circ)) t_m = -2.8805 [\text{km}]
 \end{aligned} \tag{13}$$

#### 3.2 Part C