# Weißlichtspektren

Weißlichspektren herunterladen

```
#
# CHARACTERISING THE WITEC DETECTOR
#
# How does the detector response changes when changing the orientation of the lasers plane
of polarisation?

#
# Get some libraries and functions used for characterising the detector and plotting stuff
#
source("../bauteilCharakterisierung/charakterisierungDetektor_utilities.R")

# Fetch experimental data from elabFTW
# First try
detector.spectral <- GET.elabftw.bycaption(76, header=T, outputHTTP=T) %>% parseTimeSerie
s.elab(., header=F, sep="")
# Second try
detector.spectra2 <- GET.elabftw.bycaption(81, header=T, outputHTTP=T) %>% parseTimeSerie
s.elab(., header=F, sep="")
```

- Berechnen des gemittelten Spektrums
- Berechnen von absoluter und relativer Differenz zwischen Spektren und ihrem Mittel

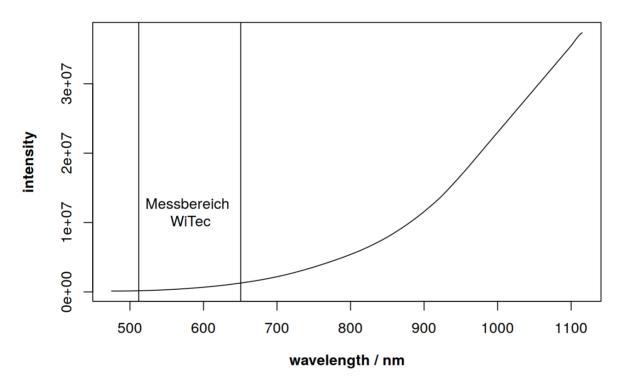
```
# Select one data set for evaluation
detector.spectra <- detector.spectra2
# PREPROCESS SPECTRA
# Vector normalisation and wavenumber conversion and mean calculation
# Wavelength of the WiTecs laser
laser.wavelength <- 514.624
detector.spectra <- lapply(detector.spectra, function(spec) {</pre>
  # Which columns contain the measured white lamp spectra?
  data.selector <- which(colnames(spec) %in% c("wavenumber", "wavelength", "mean") == F)</pre>
  # Convert raman shift in wavenumbers into absolute wavelength
  spec$wavelength <- 1/( 1/laser.wavelength - spec$wavenumber*1e-7 )</pre>
  # Vector normalisation of the spectra
  # spec[, data.selector] <- apply(spec[, data.selector], 2, function(spec) { spec / sum(s</pre>
pec^2) })
  # Compute mean spectrum and add it to the data.frame
  spec$mean <- rowMeans(spec[, data.selector])</pre>
  # Reorder data.frame
  spec <- spec[,c( which(colnames(spec) == "wavenumber"),</pre>
                   which(colnames(spec) == "wavelength"),
                   which(colnames(spec) == "mean"),
                    data.selector )]
  # Return
  return(spec)
})
# HOW DOES THE INFLUENCE OF THE POLARISATION CHANGE WITH THE WAVENUMBER?
detector.absDifference <- lapply(detector.spectra, function(spectra) {</pre>
  # Copy white lamp spectrum
  diffSpectra <- spectra</pre>
  # Compute the absolute difference between white lamp spectra and their mean spectrum
  diffSpectra[, -(1:3)] <- apply(diffSpectra[, -(1:3)], 2, function(spec) {spec - diffSpec
tra$mean})
  # Return result
  return(diffSpectra)
detector.relDifference <- lapply(detector.spectra, function(spectra) {</pre>
  # Copy white lamp spectrum
  diffSpectra <- spectra
  # Compute the relative difference between white lamp spectra and their mean spectrum
  diffSpectra[, -(1:3)] <- apply(diffSpectra[, -(1:3)], 2, function(spec) {(spec/diffSpect</pre>
ra$mean)-1})
  # Return result
  return(diffSpectra)
})
```

- Alle gemessenen Spektren übereinander legen
- Das Spektrum der Kalibrationslampe plotten

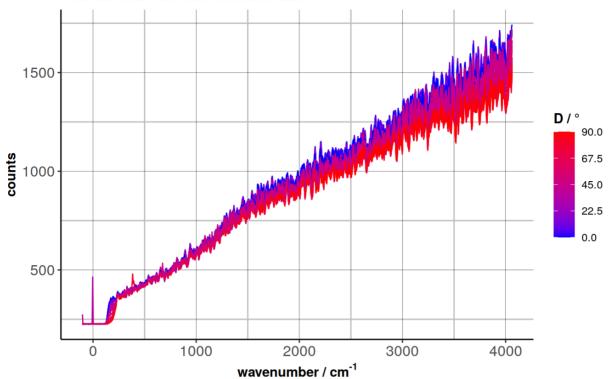
```
#
# PLOT
#

# Get the ideal white lamp spectrum
detector.whitelamp <- read.table(file = "../Weisslichtspektrum_Julian.txt", header = T)
# Plot ideal white lamp spectrum
detector.range <- detector.spectra[[1]]$wavelength[c(1, nrow(detector.spectra[[1]]))]
plot(detector.whitelamp, type = "l",
    main = "Spektrum Kalibrationslampe",
    xlab = expression(bold("wavelength / nm")),
    ylab = expression(bold("intensity")))
abline(v = detector.range)
text( mean(detector.range),
    y = mean(detector.whitelamp$Intensity), expression("Messbereich \n WiTec") )</pre>
```

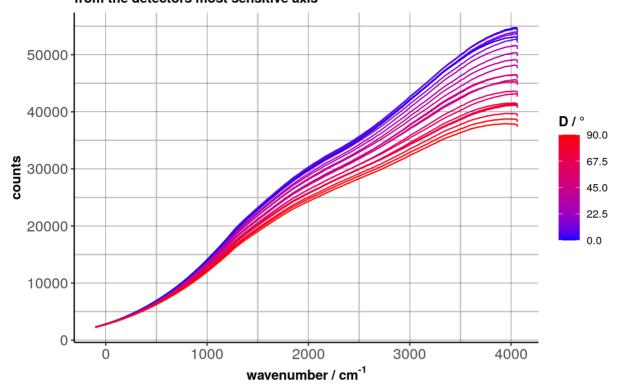
## Spektrum Kalibrationslampe



## **Detector Response for polarised white light (with microscope)**

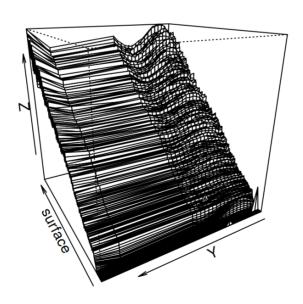


## Detector Response for polarised white light (without microscope the color gradient encodes the absolute deviation D of the linear polarisers position from the detectors most sensitive axis

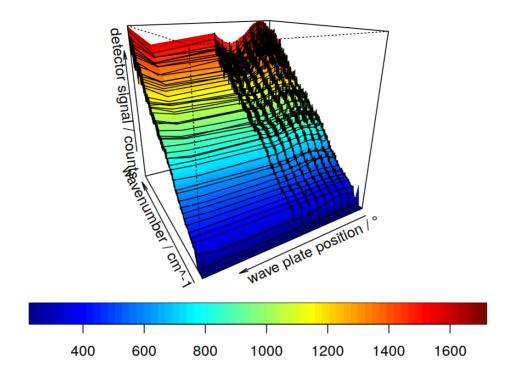


• Alle gemessenen Spektren als 3D-Plot

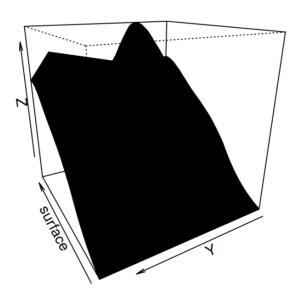
# Plot the WHITE LAMP SPECTRA in one 3d plot as 3D SURFACE
plot.detector.allSpectra(detector.spectra[[1]][,-c(2:3,21:24)], theta=240)



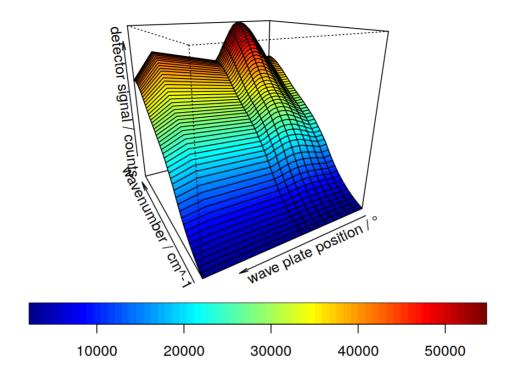
## The White Lamp Raman Spectra For Different Polarised Light



plot.detector.allSpectra(detector.spectra[[2]][,-c(2:3,21:24)], theta=240)

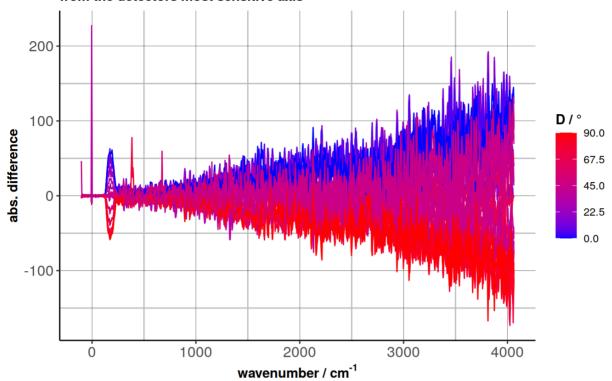


## The White Lamp Raman Spectra For Different Polarised Light



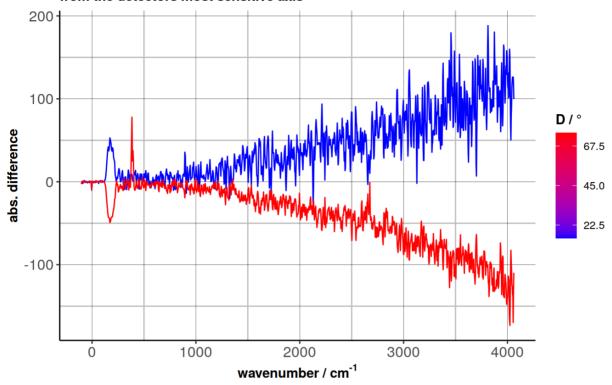
- Auftragen der Differenzspektren
- Von den Spektren, die mit dem Mikroskop gemessen wurden, werden die Spektren mit dem größten Unterschied gezeigt

## **Absolute Difference (with detector)**

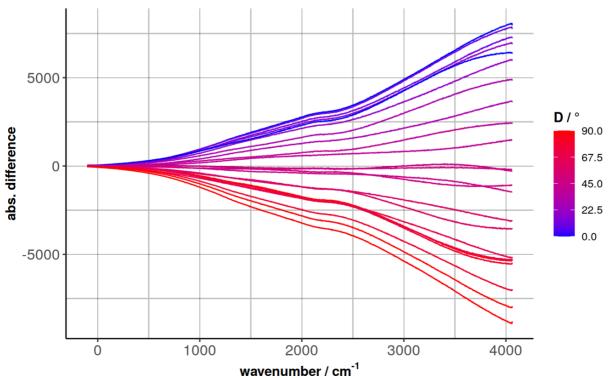


### **Absolute Difference (with detector)**

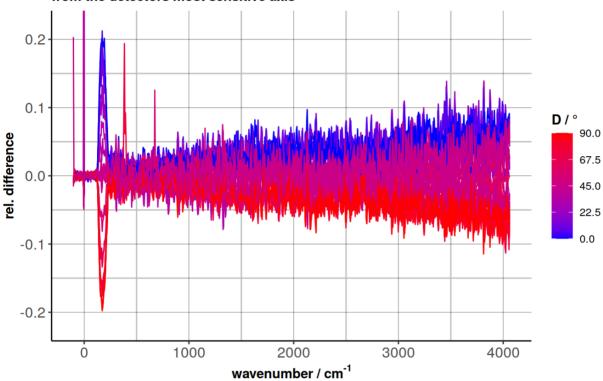
the color gradient encodes the absolute deviation D of the linear polarisers position from the detectors most sensitive axis



### **Absolute Difference (without detector)**

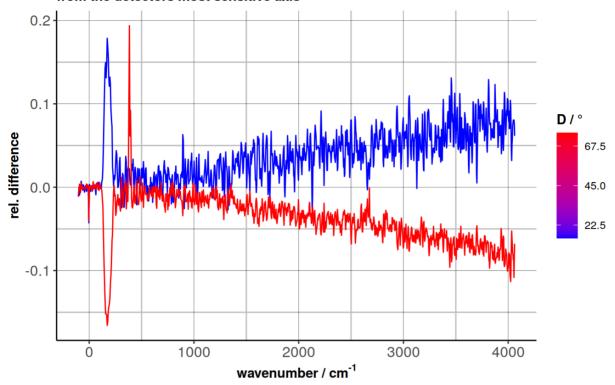


## Relative Difference (with detector)



## **Relative Difference (with detector)**

the color gradient encodes the absolute deviation D of the linear polarisers position from the detectors most sensitive axis



### **Relative Difference (without detector)**

