

# RITSAR

Douglas Macdonald

Rochester Institute of Technology

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# 1 Introduction

Below is a description of each of the functions contained in the RITSAR package. For installation instructions, reference the README file.

Each of the sections below represent a module contained within the RITSAR package and each subsection details how to use each function. The format for the title of each subsection is "function(parameters)." This is followed by a short description of what the function does. Next, each parameter is defined. Finally, a description of what the function returns is given.

As an example, lets say you wanted to simulate your own phase history. The function required for this is `simulate_phs` located in the `phsTools` module. Before this function can be used, the required parameters need to be defined. The descriptions of each required parameter can be found in section 2.1. The first parameter that needs to be defined is "platform", which is a Python dictionary. A description of each key that goes into platform is also given in section 2.1. Once platform is defined, the user then creates an array specifying the locations of each point reflector in the scene as well as the corresponding amplitudes. After all this is done the user can then create their own phase history, stored in the `phs` variable, by typing the following into the Python console:

```
phs = ritsar.phsTools.simulate_phs(platform, points, amplitudes)
```

To get started, "sim\_demo.py" in the `./examples` folder can be run. This demo uses RITSAR's basic point simulator to generate a synthetic phase history. The default setup processes an X-band collect using the polar format algorithm. The final lines of `main.py` can be commented/un-commented to use the backprojection algorithm. For a demonstration of the omega-k algorithm, OpenCV will be required. You will most likely also want to use a UHF setup. To do this, simply change `SARplatform` to `SARplatformUHF` in the demo and comment/un-comment the bottom lines of the file as necessary. The platform dictionary keys defined in `SARplatform` and `SARplatformUHF` were taken from Carrera [1].

Also included in the `./examples` folder are examples of how to process actual data. Full advantage is taken of Python's object oriented nature by including all system parameters in a platform object. As explained in the `phsRead` section, only a few steps are required to use the algorithms in this package for any arbitrary data set. They are:

1. read in the phase history and export as a NumPy array
2. read in the auxillary data
3. place auxillary data in a platform dictionary and export

The function descriptions below detail what parameters need to be defined in the platform dictionary for a given function. For an arbitrary data source, the most rigorous task will just be renaming/deriving parameters already in the auxillary data and placing them in a platform dictionary. The source code for `phsRead` can be used as a reference. Included in the source code are examples of how to translate auxillary data from a `.xml` file (required for DIRSIG), a `.mat` file (required for the AFRL data set), and a binary file (required for the Sandia data set) into a python dictionary readable by RITSAR. Once the platform dictionary is defined, no additional work is needed use the algorithms in this package. Simply specify an image plane using `imgTools.img_plane.dict()` 3.5 or manually using the guidelines in 3.1.

## 2 phsTools

### 2.1 simulate\_phs(platform, points, amplitudes)

This function is a basic phase history simulator. It takes a list of target locations and amplitudes and saves the demodulated signal to './phase\_history.npy'. It also outputs the signal to the function caller. The amplitude of the received signal for a perfect point reflector is assumed to be one arbitrary unit. A uniform beam pattern is assumed and  $1/r^4$  attenuation is not modeled.

#### Parameters:

- **platform** : dictionary  
dictionary that has the following keys defined:
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **chirprate** : float  
chirprate in  $Hz/s$
  - **f\_0** : float  
carrier frequency in Hz
  - **t** : float array [nsamples]  
demodulated fast-time centered at t=0. The array should go from  $(-nsamples/2, nsamples/2) \times \Delta t$  where,  $\Delta t = (A/D \text{ sampling rate})^{-1}$
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
- **points** : float array [npoints x 3]  
locations of each point referenced to scene center
- **amplitudes** : float array [npoints]  
reflectance of each point. May be a complex number.

#### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
resultant phase history

## 2.2 RVP\_correct(phs, platform)

Corrects Residual Video Phase using the formulation in Carrera Appendix C.

### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history to be corrected for residual video phase
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **chirprate** : float  
chirprate in  $Hz/s$
  - **delta\_r** : float  
spacing between fast time samples. Defined here as  $\frac{c}{2B_{IF}}$  where,  $B_{IF}$  is the bandwidth of the intermediate frequency (de-chirped) signal.  $B_{IF}$  is simply  $(t_{max} - t_{min}) \times chirprate$  where  $t$  is the demodulated fast time, as defined in section 2.1.

### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
resultant phase history, corrected for residual video phase

### 2.3 phs\_to\_const\_ref(phs, platform, upchirp)

This program converts a phase history that was demodulated using a pulse dependant range to scene center to a phase history that is demodulated using a fixed reference. The fixed reference is defined as the minimum range to scene center.

#### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history that was compensated to a point
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **npulses** : int  
number of pulses
  - **chirprate** : float  
chirprate in  $Hz/s$
  - **f\_0** : float  
carrier frequency in Hz
  - **t** : float array [nsamples]  
demodulated fast-time centered at  $t=0$ . The array should go from  $(-nsamples/2, nsamples/2) \times \Delta t$  where,  $\Delta t = (A/D \text{ sampling rate})^{-1}$
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)

#### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
resultant phase history, compensated to a line

### 2.4 reMoComp(phs, platform, center)

This is the re-motion compensation algorithm. It re-motion compensates the phase history to a new scene center. The "center" argument is the 3D vector (in meters) that points to the new scene center using the old scene center as the origin.

#### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history that was compensated to a point
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **k\_r** : float array [nsamples]  
spatial frequencies corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
- **center** : array [3]  
3D vector (in meters) that points to the new scene center using the old scene center as the origin

#### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
resultant phase history, compensated to new scene center

## 3 imgTools

In this section the `img_plane` dictionary is introduced. A sample `img_plane` dictionary generator is provided in `sarit/examples/dictionaries`.

### 3.1 polar\_format(`phs`, `platform`, `img_plane`, `taylor`)

This is the Polar Format algorithm. The phase history data as well as platform and image plane dictionaries are taken as inputs.

The phase history data is collected on a two-dimensional surface in k-space. For each pulse, a strip of this surface is collected. The first step in this program is to project each strip onto the (ku,kv) plane defined by the normal vector contained in the image plane dictionary. This will result in data that is unevenly spaced in (ku,kv). This unevenly spaced data is interpolated onto an evenly spaced (ku,kv) grid defined in the image plane dictionary. The interpolation is done along the radial direction first, then along the along-track direction. Further details of this method are given in both the Jakowitz and Carrera texts.

#### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **npulses** : int  
number of pulses
  - **f\_0** : float  
carrier frequency in Hz
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
  - **k\_r** : float array [nsamples]  
spatial frequencies corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$
  - **R\_c** : float  
vector to scene center at aperture center
- **img\_plane** : dictionary  
dictionary that has the following keys defined:
  - **n\_hat** : float array [3]  
unit normal vector to image plane
  - **k\_u** : float array [size(u)]  
image plane spatial frequencies along u axis. u is defined to lie along the slant range vector, projected onto the image plane.
  - **k\_v** : float array [size(v)]  
image plane spatial frequencies along v axis. v is defined to lie perpendicular to the slant range vector in the image plane.
- **taylor** : int  
attenuation factor for taylor window in dB. defaults to 43 dB.

#### Returns:

- **img** : complex array [npulses x nsamples]  
processed image

### 3.2 omega-k(phs, platform, taylor, upsample)

This is an omega-k algorithm based off of the algorithm prescribed in the Carrera text. Only the phase history and platform files are taken as inputs, an img\_plane dictionary is not required.

The input phase history needs to have been demodulated to a fixed reference. If demodulated to scene center, the included phs\_const\_ref file can do the conversion for you. A straight line flight path is also assumed.

The first step in the algorithm is to perform a 1D FT along azimuth. A matched filter is applied to the resultant data to perfectly compensate the range curvature of all scatterers having minimum range  $R_s$ . The default setting for  $R_s$  is the minimum range of scene center. To correct the range curvature for other scatterers, the data is mapped onto a new grid.

#### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
  - **k\_r** : float array [nsamples]  
spatial frequencies of demodulated signal corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$
  - **k\_y** : float  
spatial frequencies along flight path. A straight line flight path (or phase history motion compensated to a straight line flight path) are required.
- **taylor** : int  
attenuation factor for taylor window in dB. defaults to 43 dB.
- **upsample** : int  
upsampling factor, default is 6.

#### Returns:

- **img** : complex array [npulses x nsamples]  
processed image

### 3.3 backprojection(phs, platform, img\_plane, taylor, upsample, prnt)

This is the Backprojection algorithm. The phase history data as well as platform and image plane dictionaries are taken as inputs. The (x,y,z) locations of each pixel are required, as well as the size of the final image (interpreted as [size(v) x size(u)]).

#### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
  - **k\_r** : float array [nsamples]  
spatial frequencies corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$
  - **delta\_r** : float  
spacing between fast time samples. Defined here as  $\frac{c}{2B_{IF}}$  where,  $B_{IF}$  is the bandwidth of the intermediate frequency (de-chirped) signal.  $B_{IF}$  is simply  $(t_{max} - t_{min}) \times chirprate$  where  $t$  is the demodulated fast time, as defined in section 2.1.
- **img\_plane** : dictionary  
dictionary that has the following keys defined:
  - **u** : float array [len(u)]  
for the backprojection algorithm, u is essentially a placeholder that defines one of the dimensions of the output image. Only the length of u matters, not its contents.
  - **v** : float array [len(v)]  
for the backprojection algorithm, v is essentially a placeholder that defines one of the dimensions of the output image. Only the length of v matters, not its contents.
  - **pixel\_locs** : float array [3 x n\_pixels]  
flattened array specifying the 3-dimensional location of each pixel in the rectangular output array whose shape is [len(u) x len(v)]. Consequently,  $n\_pixels = len(u) \times len(v)$ . The locations do not have to lie in a plane and can be specified to an arbitrary surface that matches the local terrain. A rectangular output array with the location of each pixel specified is the only constraint.
- **taylor** : int  
attenuation factor for taylor window in dB. defaults to 43 dB.
- **upsample** : int  
upsampling factor, default is 6.
- **prnt** : bool  
option to print the pulse number currently being processed.

#### Returns:

- **img** : complex array [npulses x nsamples]  
processed image



### 3.4 DSBP(phs, platform, img\_plane, center, size, derate, taylor, ftype, V)

This is the Digital Spotlight Backprojection algorithm based on K. Dungan et. al.'s 2013 SPIE paper [2]. In addition to the normal arguments used for the backprojection algorithm, DSBP also requires the "center" and "size" arguments. The "center" argument is the 3D vector (in meters) that points to the new scene center using the old scene center as the origin. The "size" argument is a 2 element array. Each element describes the extent of the digitally spotlighted scene, in units of img\_plane pixels, along azimuth and range, respectively. The elements of size are assumed to be even numbers.

#### Parameters:

- **phs** : complex array [npulses x nsamples]  
complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
  - **freq** : float array [nsamples]  
temporal frequencies corresponding to each fast time return. For a linear FM signal, this is given by  $(f_0 + chirprate \times t)$
- **img\_plane** : dictionary  
dictionary that has the following keys defined:
  - **u** : float array [len(u)]  
for the backprojection algorithm, u is essentially a placeholder that defines one of the dimensions of the output image. Only the length of u matters, not its contents.
  - **v** : float array [len(v)]  
for the backprojection algorithm, v is essentially a placeholder that defines one of the dimensions of the output image. Only the length of v matters, not its contents.
  - **u\_hat** : float array [3]  
unit vector for u. Defined to lie along the slant range vector, projected onto the image plane.
  - **v\_hat** : float array [3]  
unit vector for v. Defined to lie perpendicular to the slant range vector in the image plane.
  - **du** : float array [len(u)]  
pixel spacing in range direction
  - **dv** : float array [len(v)]  
pixel spacing in cross-range direction
  - **pixel\_locs** : float array [3 x n\_pixels]  
flattened array specifying the 3-dimensional location of each pixel in the rectangular output array whose shape is [len(u) x len(v)]. Consequently,  $n\_pixels = len(u) \times len(v)$ . The locations do not have to lie in a plane and can be specified to an arbitrary surface that matches the local terrain. A rectangular output array with the location of each pixel specified is the only constraint.
- **center** : float array[3]  
vector (in meters) that points to the new scene center using the old scene center as the origin
- **size** : int array[2]  
describes the extent of the digitally spotlighted scene, in units of img\_plane pixels, along azimuth and range, respectively
- **derate** : float  
amount by which to increase spot size. default is 1.05

- **taylor** : int  
attenuation factor for taylor window in dB. defaults to 43 dB.
- **ftype** : string  
type of filter to use for decimation. Default is 'iir' and is recommended for large scenes. 'fir' can also be chosen and is recommended for relatively smaller scenes.
- **V** : string  
Describes if the spotlighted scene is large or small. The default is 'large', which optimizes the output grid for spotlighted scenes that are a large fraction of the original scene. If V is set to small (or anything other than 'large'), an output grid optimized for small scenes is created.

**Returns:**

- **img** : complex array [npulses x nsamples]  
processed image

### 3.5 **img\_plane\_dict(platform, res\_factor, n\_hat, aspect, upsample)**

This function defines the image plane parameters. The user specifies the image resolution using the `res_factor`. A `res_factor` of 1 yields a (u,v) image plane whose pixels are sized at the theoretical resolution limit of the system (derived using `delta_r` which in turn was derived using the bandwidth). The user can also adjust the aspect of the image grid. This defaults to `nsamples/npulses`.

'n\_hat' is a user specified value that defines the image plane orientation w.r.t. to the nominal ground plane.

**Parameters:**

- **platform** : dictionary  
dictionary that has the following keys defined:
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **R\_c** : float array [3]  
vector to scene center at aperture center
  - **delta\_r** : float array [nsamples]  
spacing between fast time samples. Defined here as  $\frac{c}{2B_{IF}}$  where,  $B_{IF}$  is the bandwidth of the intermediate frequency (de-chirped) signal.  $B_{IF}$  is simply  $(t_{max} - t_{min}) \times chirprate$  where  $t$  is the demodulated fast time, as defined in section 2.1.
- **res\_factor** : float  
image resolution in units of theoretical resolution size. This defaults to 1.0.
- **n\_hat** : float array [3]  
unit normal vector to image plane. Defaults to (0,0,1).
- **aspect** : float  
aspect ratio of range to cross range. Determines cross-range resolution by multiplying `res_factor` by `aspect`. Defaults to `nsamples/npulses`
- **upsamples** : bool  
Option to upsample number of pixels. The number of pixels in the range direction is set to a power of 2 that is one power of 2 higher than `nsamples`. If `nsamples` is already a power of 2, the number of range pixels is just set to `nsamples`. The same goes for `npulses` and number of cross-range pixels.

**Returns:**

- **img\_plane** : dictionary  
dictionary that has the following keys defined:
  - **n\_hat** : float array [3]  
unit normal vector to image plane. Defaults to (0,0,1).
  - **u** : float array [len(u)]  
pixel locations in range direction. u is defined to lie along the slant range vector, projected onto the image plane.
  - **v** : float array [len(v)]  
pixel locations in cross-range direction. v is defined to lie perpendicular to the slant range vector in the image plane.
  - **u\_hat** : float array [3]  
unit vector for u. Defined to lie along the slant range vector, projected onto the image plane.
  - **v\_hat** : float array [3]  
unit vector for v. Defined to lie perpendicular to the slant range vector in the image plane.
  - **du** : float array [len(u)]  
pixel spacing in range direction
  - **dv** : float array [len(v)]  
pixel spacing in cross-range direction
  - **k\_u** : float array [size(u)]  
image plane spatial frequencies along u axis. u is defined to lie along the slant range vector, projected onto the image plane.
  - **k\_v** : float array [size(v)]  
image plane spatial frequencies along v axis. v is defined to lie perpendicular to the slant range vector in the image plane.
  - **pixel\_locs** : float array [3 x n\_pixels]  
flattened array specifying the 3-dimensional location of each pixel in a rectangular output array whose shape is [len(u) x len(v)]. Consequently,  $n\_pixels = len(u) \times len(v)$ . The locations do not have to lie in a plane and can be specified to an arbitrary surface that matches the local terrain. A rectangular output array with the location of each pixel specified is the only constraint. If the image plane is specified in this manner, only a time domain algorithm (backprojection) can be used to process the data.

### 3.6 autoFocus(img, win, win\_params)

This program autofocuses an image using the Phase Gradient Algorithm. If the parameter win is set to auto, an adaptive window is used. Otherwise, the user sets win to 0 and defines win\_params. The first element of win\_params is the starting windows size. The second element is the factor by which to reduce it by for each iteration. This algorithm is based off [3].

#### Parameters:

- **img** : complex array [npulses x nsamples]  
complex image
- **win** : string  
if 'auto', an adaptive window is used. Otherwise, win\_params is used to create the window. Default is 'auto.'
- **win\_params** : list  
The first element of win\_params is the starting windows size. The second element is the factor by which to reduce it by for each iteration.

**Returns:**

- **img\_af** : complex array [npulses x nsamples]  
autofocused image

### 3.7 imshow(img, dB\_scale, extent)

This program displays the processed data in dB. The brightest point in the image is used as the reference and the user can define the scale for the intensity range.

**Parameters:**

- **img** : complex array [npulses x nsamples]  
complex image
- **dB\_scale** : array [2]  
decibel scale of image. Format is  $[dB_{min}, dB_{max}]$ . The reference signal is the brightest intensity value of the input image. In other words, the brightest resolution cell has a dB value of 0. Intensity values for all other cells are w.r.t. this value and are  $\leq 0$
- **extent** : array [4]  
defines the image boundaries in units of meters. Format is  $[x_{min}, x_{max}, y_{min}, y_{max}]$

**Returns:**

- **img** : array [npulses x nsamples]  
magnitude of input data displayed using the user defined decibel scale

## 4 phsRead

This section describes the RITSAR interfaces to various datasets. In a nutshell, the user supplies the directory containing the phase history and auxillary data. The functions below essentially perform 3 tasks:

1. read in the phase history and export as a NumPy array
2. read in the auxillary data
3. place auxillary data in a platform dictionary and export

As shown in the functions below, code can be written to handle many different data formats including .envi, .mat, .xml, and binary. As of now, RITSAR can interface with DIRSIG and AFRL Gotcha data. A function is also provided that processes a Sandia data set. These functions serve mostly to show that data in almost any format can be processed with the RITSAR toolset provided code is written that performs the three tasks listed above for a given format.

### 4.1 DIRSIG(directory)

This function reads in the DIRSIG xml data as well as the envi header file from the user supplied directory. The phs and a Python dictionary compatible with RITSAR are returned to the function caller. More information on DIRSIG's SAR capability can be found here: <http://www.dirsig.org/docs/new/radar.html>

#### Parameters:

- **directory** : string  
string containing directory of DIRSIG data to include the xml and envi header files.

#### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
the complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **f\_0** : float  
carrier frequency in Hz
  - **chirprate** : float  
chirprate in  $Hz/s$
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **B** : float  
bandwidth of transmitted signal
  - **B\_IF** : float  
bandwidth of intermediate frequency, demodulated signal.
  - **delta\_r** : float  
spacing between fast time samples. Defined here as  $\frac{c}{2B_{IF}}$  where,  $B_{IF}$  is the bandwidth of the intermediate frequency (de-chirped) signal.  $B_{IF}$  is simply  $(t_{max} - t_{min}) \times chirprate$  where  $t$  is the demodulated fast time, as defined in section 2.1.
  - **delta\_t** : float  
time between fast time samples. Defined as  $(A/D \text{ sampling rate})^{-1}$
  - **vp** : float  
velocity of platform in m/s

- **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
- **R\_c** : float array [3]  
vector to scene center at aperture center
- **t** : float array [nsamples]  
demodulated fast-time centered at t=0. The array should go from  $(-nsamples/2, nsamples/2) \times \Delta t$  where,  $\Delta t = (A/D \text{ sampling rate})^{-1}$
- **k\_r** : float array [nsamples]  
spatial frequencies of demodulated signal corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$
- **k\_y** : float  
spatial frequencies along flight path. A straight line flight path (or phase history motion compensated to a straight line flight path) are required.
- **metadata** : string array  
xml metadata

## 4.2 AFRL(directory, pol, start\_az, n\_az=3)

This function reads in the AFRL Gotcha \*.mat files from the user supplied directory and exports both the phs and a Python dictionary compatible with RITSAR.

### Parameters:

- **directory** : string  
string containing directory of DIRSIG data to include the xml and envi header files.
- **pol** : string  
polarization to be processed. For example, 'HH'.
- **start\_az** : int  
starting azimuth (0 to 360)
- **n\_az** : int  
number of azimuth files to process. Default is 3.

### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
the complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **f\_0** : float  
carrier frequency in Hz
  - **chirprate** : float  
chirprate in  $Hz/s$
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **B\_IF** : float  
bandwidth of intermediate frequency, demodulated signal.
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)

- **delta\_r** : float  
spacing between fast time samples. Defined here as  $\frac{c}{2B_{IF}}$  where,  $B_{IF}$  is the bandwidth of the intermediate frequency (de-chirped) signal.  $B_{IF}$  is simply  $(t_{max} - t_{min}) \times chirprate$  where  $t$  is the demodulated fast time, as defined in section 2.1.
- **R\_c** : float array [3]  
vector to scene center at aperture center
- **t** : float array [nsamples]  
demodulated fast-time centered at  $t=0$ . The array should go from  $(-nsamples/2, nsamples/2) \times \Delta t$  where,  $\Delta t = (A/D \text{ sampling rate})^{-1}$
- **af\_r** : float array [npulses]  
column vector containing the correction for R\_c
- **af\_ph** : float array [npulses]  
column vector containing the phase correction
- **k\_r** : float array [nsamples]  
spatial frequencies of demodulated signal corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$
- **(k\_y)** : float  
not currently available (due to circular flight path). may be added later when moComp\_to\_line is complete.

### 4.3 Sandia(directory)

This function reads in the Sandia \*.phs and \*.au2 files from the user supplied directory and exports both the phs and a Python dictionary compatible with RITSAR.

#### Parameters:

- **directory** : string  
string containing directory of DIRSIG data to include the

#### Returns:

- **phase\_history** : complex array [npulses x nsamples]  
the complex phase history
- **platform** : dictionary  
dictionary that has the following keys defined:
  - **f\_0** : float  
carrier frequency in Hz
  - **chirprate** : float  
chirprate in  $Hz/s$
  - **nsamples** : int  
number of fast time samples
  - **npulses** : int  
number of pulses
  - **B\_IF** : float  
bandwidth of intermediate frequency, demodulated signal.
  - **pos** : float array [npulses x 3]  
(x,y,z) coordinates of platform, referenced to scene center, which is located at (0,0,0)
  - **delta\_r** : float  
spacing between fast time samples. Defined here as  $\frac{c}{2B_{IF}}$  where,  $B_{IF}$  is the bandwidth of the intermediate frequency (de-chirped) signal.  $B_{IF}$  is simply  $(t_{max} - t_{min}) \times chirprate$  where  $t$  is the demodulated fast time, as defined in section 2.1.

- **R\_c** : float array [3]  
vector to scene center at aperture center
- **t** : float array [nsamples]  
demodulated fast-time centered at t=0. The array should go from  $(-nsamples/2, nsamples/2) \times \Delta t$  where,  $\Delta t = (A/D \text{ sampling rate})^{-1}$
- **k\_r** : float array [nsamples]  
spatial frequencies of demodulated signal corresponding to each fast time return. For a linear FM signal, this is given by  $\frac{4\pi}{c}(f_0 + chirprate \times t)$  added later when moComp\_to\_line is complete.

## References

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- [2] K. E. Dungan, L. A. Gorham, and L. J. Moore, “SAR digital spotlight implementation in MATLAB,” *Proceedings of SPIE*, vol. 8746, 2013.
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