Sentinel-2 Agriculture

Design Definition File

Algorithm Theoretical Basis Document for L3 monthly composite product









Milestone	Milestone 2
Version	1.0
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1 Introduction

This document provides the detailed processing model of a Monthly Synthesis processor, that delivers L3A products. Level 3A products provide the weighted average of the cloud free pixels observed by Sentinel-2 every month. This weighted average can be based on a synthesis duration which is longer that one month.

This document explains how the Monthly Synthesis processor should be defined. The way the Monthly Synthesis processor is launched is described in the Sen2Agri System specification (DRxx). The System Specification addresses how to do the selection of the L2A products, and how long they are awaited if turn out not to be available. An important parameter of the syntheses is the composite duration, however this parameter is handled by the Sen2Agri System and not by the L3A processor.

In order to limit the amount of L2A products to keep on line in the disk system, the L3A processor has been defined as a recurrent operation. When the fist L2A Product necessary to produce the L3A product becomes available, the L3A product is initialised. It is then updated each time a new L2A product belonging to the list of L2A products to be processed is obtained.

2 Monthly synthesis product

The monthly synthesis product contains the following information:

- ♦ The weighted average surface reflectance for all Sentinel-2 bands available in the L2A products, at their original resolution (10m or 20m). The Sentinel-2 bands at 60m resolution are not provided in the L2A product, as these bands are useful for atmospheric correction but not aimed at observing the surface.
- Flags:
 - * one byte per pixel, 10m resolution
 - * values:
 - * no-data: pixel was never observed during compositing period
 - * cloud: pixel is always cloudy or within cloud shadow
 - \star snow: pixel is always covered by snow for the available cloud free observations
 - \star water: pixel is always covered by water for the available cloud free observations
 - \star land: pixel was free from cloud, snow and water at least once during composite period
- ♦ Dates:
 - * a weighted average of dates used in the synthesis is provided for each pixel
 - \ast to be provided only at 20m resolution

A table of the available dates is provided in the L3A metadata file

3 Monthly synthesis processor

For a given tile, the monthly synthesis processor works in 3 steps:

- ♦ a pre-processing step, which for each date transforms the L2A in a L2A* product
 - * to normalize directional effects for Sentinel-2 (A & B)
 - * to resample LANDSAT 8 to the same resolution as the corresponding Sentinel-2 band.
- a processing step:
 - * the synthesis product is a recurrent process that updates the product date after date
 - * the processor is run each time a new L2A* is available
 - * it uses as input a L2A* and the current L3A product
 - * if L3A product is not available, it is created



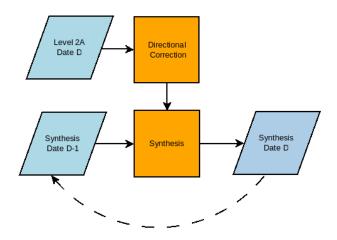


Figure 1: Diagram of synthesis processor

3.1 Pre-processing

This part is a pre-processor to be executed before the synthesis product processing. It uses a Level 2A as input, and delivers a pre-processed Level 2A referred hereafter as L2A*.

- ♦ For Sentinel-2, L2A* provides reflectances after directional correction (section 3.1.2)
- ♦ As LANDSAT-8 does not provide the viewing angles (yet?) for each pixel in its L1T product, it is not possible to apply the directional correction to LANDSAT 8. For LANDSAT 8, L2A* are L2A resampled to the resolution of the corresponding band of Sentinel-2 (see 3.1.3).

In both cases, the L2A* metadata and masks are identical to the L2A metadata and masks with only the resolution information changed in the case of LANDSAT-8.

3.1.1 Directional model function

Variables and parameters

Input variable	role
b	spectral band
NDVI	NDVI
$(\theta_s, \theta_v, \delta\phi)$	Solar zenith Angle, Viewing zenith Angle, Relative Az-
	imuth

Output variable	role
=mod-dir $=$	value of directional model for band b and for the input
	angles

parameter	role	default value
V0, V1	Coefficient for the volume scattering function, as a function	
	of spectral band	
R0, R1	Coefficient for the roughness scattering function, as a func-	
	tion of spectral band	

Pseudocode

- \diamond For each spectral band b
 - * \$mod-dir($\theta_s, \theta_v, \delta\phi$,b) = 1 + (V0(b)+V1(b)*NDVI).F_V($\theta_s, \theta_v, \delta\phi$) + (R0(b)+R1(b)*NDVI).F_R($\theta_s, \theta_v, \delta\phi$)
 - * /a python code to compute F_V and F_R is provided in Annex, chapter 5/

3.1.2 Directional Correction

The directional correction consists in applying a model to each pixel, except of this pixel is classified as cloud, cloud shadow, water or snow. The model depnds on the sun and observation angles and on the NDVI of the pixel.



Variables and parameters

Input variable	role
ρ	reflectance for each pixel and each spectral band
Angles from L2A Image	Solar zenith Angle, Viewing zenith Angle, Relative Az-
$(\theta_s, \theta_v, \delta\phi)$	imuth
Cloud-Shadow Mask	Mask of clouds and shadows
Water Mask	Mask of water pixels
Snow Mask	Mask of snow pixels

Pseudocode

- \diamond Compute NDVI at 10m resolution => NDVI_{10m}
- \diamond Resample it without aliasing to 20 m => NDVI_{20m}
- \diamond Compute $(\theta_s, \theta_v, \delta\phi)$ for each pixel and each band according to Sentinel-2 L1C product Users Manual. /The viewing angles are provided as a grid per Sentinel-2 spectral band and per detector (there are 12 detectors per band). Sentinel-2 metadata also provide a map of the detectors which were used to observe each pixel. The method consist in finding the detector corresponding to the pixel to process /
- ♦ Apply directional correction to each pixel except water and snow pixels, using directional model (section 3.1.1)
- \diamond For each pixel and each band :
 - * if pixel is not cloud free, shadow free, water and snow free

$$\star \ \rho(\theta_s, 0, 0) = \rho(\theta_s, \theta_v, \delta\phi) * \frac{mod - dir(\theta_s, 0, 0)}{mod - dir(\theta_s, \theta_v, \delta\phi)}$$

3.1.3 Resampling at S2 Resolution

Variables and parameters

Input variable	role
S2-Bands(L8-bands)	S2-Band corresponding to every L8 band (see annex 4)
$Res_S 2(L8 - bands)$	Resolution of each S2-band
ρ	reflectance for each pixel and each L8 spectral band

Output variable	role
$\rho*$	reflectance value for each pixel, for each L8 spectral band
	with a corresponding S2 spectral band
Cloud-Shadow Mask	Mask of clouds and shadows
Water Mask	Mask of water pixels

Pseudocode

- \diamond for each 18 spectral band b_{L8}
 - * resample the L8 band to the resolution $Res(S2-bands(b_{L8}))$
- \diamond resample cloud-shadow mask and snow mask at 10 and 20m resolution

3.2 Synthesis Update with new product

The Synthesis update is made of 3 successive steps :

- \diamond Resample cloud and snow masks to 10m resolution
- \diamond Compute individual weights for each pixel in the image (3.2.2)
- ♦ Update synthesis



3.2.1 Update synthesis

This module updates the synthesis using the recurrent expression of a weighted average. A weighted average is only computed for land and water pixels. Cloud and Snow pixels are discarded from the weighted average, unless no land or water pixel was observed so far in the composite window. If it is the first iteration of the synthesis update, an initial syntheses with no-data pixels everywhere must be created. Land and Water are considered the same class to avoid issues with rice, which is flooded at the early stages.

The weighted average is computed with a weight counter per $S2_{\rm band}$. This is not ideal as it adds a data volume equivalent to that of surface reflectance. If only Sentinel-2 (A and B) were used, it would be possible to use one counter for the 20m bands and a counter for the 10m bands (but that would not allow to account for possible failures in one of the bands of one of the satellites.

In order to reduce the data volume, the weighted average date and flag are provided for the red band only. It is assumed there will always be a red band and a blue band, which are required to choose the reflectances

Variables and parameters

input variable	role
S2-Bands(L8-bands)	S2-Band corresponding to every L8 band L2A* product
,	
$date_N$	date from input product
ρ_N*	reflectance value for each pixel, and each spectral band from
	input L2A* product
Cloud-Shadow Mask	Mask of clouds and shadows
Water Mask	Mask of water pixels
Snow Mask	Mask of snow pixels
L3A product	(if already created)
$ \mathbf{W}_{N-1} $	Weight for each pixel obtained so far
\overline{date}_{N-1}	weighted average date for L3A product so far
$ \bar{\rho}_{N-1} $	weighted average reflectance value so far, for each pixel, and
	each spectral band
$\operatorname{flag}_{N-1}$	status of each L3A pixel : cloud, water , snow

Output variable	role
L3A product	
$\mid \mathrm{W}_N$	Weight counter for each pixel and for each band
\overline{date}_N	weighted average date for L3A product so far
$ar ho_N$	weighted average reflectance value so far, for each pixel, and
	each spectral band
$flag_N$	status of each L3A pixel : cloud, water , snow

Pseudocode initialization if necessary

- ♦ if L3A product does not exist
 - * for each pixel
 - ★ flag_{N-1}=no-data
 - $\star \overline{date}_{N-1} = no data$
 - \star for each S2 band
 - $\bar{\rho}_{N-1} = no data$
 - $W_{N-1} = 0$

update with current product

- \diamond compute weights for L2A product of date N => w_N per band and pixel
- \diamond for resolutions res in (10m, 20m):
 - * for each pixel
 - * if pixel not cloud/shadow, snow, water (nominal case)
 - $flag_N = land$
 - for each S2 band with res resolution
 - if band is available in the case of LANDSAT 8, some bands are not available



$$- \bar{\rho}_N = \frac{W_{N-1}.\bar{\rho}_{N-1} + w_N.\rho_N}{W_{N-1} + w_N}$$

$$- \overline{date}_N = \frac{W_{N-1}.\overline{date}_{N-1} + w_N.date_N}{W_{N-1} + w_N}$$

$$- W_N = W_{N-1} + w_N$$

$$\circ \text{ else :}$$

$$- \bar{\rho}_N = \rho_{n-1}$$

$$- W_N = W_{N-1}$$

$$- \text{ if band is red}$$

$$\cdot \overline{date}_N = \overline{date}_{N-1}$$

- \star else : degraded case
 - if pixel is snow or water *replace the reflectance value
 - \circ if pixel is water
 - $flag_N$ =water
 - o if pixel is snow
 - $flag_N = snow$
 - o for each S2 band with res resolution
 - if band is available
 - · if $W_{N-1} == 0$ *pixel never observed without cloud, water or snow
 - $\cdot \ \bar{\rho}_N = \rho$
 - $W_N = 0$
 - \cdot if band is red
 - $\cdot \ \overline{date}_N = date_N$
 - · else: *pixel already observed cloud free, keep the previous weighted average
 - $\cdot \ \bar{\rho}_N = \bar{\rho}_{N-1}$
 - $W_N = W_{N-1}$
 - \cdot if band is red
 - $\cdot \ \overline{date}_N = \overline{date}_{N-1}$
 - · $flag_N = land$
 - else: *band not available, keep previous values
 - $\cdot \ \bar{\rho}_N = \bar{\rho}_{N-1}$
 - $\cdot W_N = W_{N-1}$
 - if pixel is cloud or shadow *pixel never observed cloud snow or water free
 - \circ if $flag_{N-1}$ is no-data *replace no_{data} with cloud
 - for each S2 band with res resolution
 - · if band is available
 - $\cdot \ \bar{\rho}_N = \rho$
 - $W_N = 0$
 - · if band is red
 - $\cdot \overline{date}_N = date_N$
 - · $flag_N = cloud$
 - · else
 - $\cdot \ \bar{\rho}_N = no data$
 - $W_N = 0$
 - \circ if flag_{N-1} is cloud or shadow *replace value only if new reflectance in the blue is smaller
 - if $\rho(\text{blue}) < \bar{\rho}_{\ell} N (blue)$
 - for each S2 band with res resolution
 - \cdot if band is available
 - $\cdot \ \bar{\rho}_N = \rho$
 - $W_N = 0$
 - · if band is red
 - $\cdot \overline{date}_N = date_N$
 - · flag_N=cloud
 - \cdot else

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$$\bar{\rho}_N = \bar{\rho}_{n-1}$$

$$W_N = 0$$

 \circ else:

- $flag_N = flag_{N-1}$
- for each S2 band with res resolution
 - $\bar{\rho}_N = \bar{\rho}_{n-1}$
 - $\cdot \ \overline{date}_N = \overline{date}_{N-1}$
 - $W_N = 0$

3.2.2 Compute weights

Weight on Sensors The idea here is to be able to merge different sensors in the synthesis product (for instance, LANDSAT 8 and Sentinel-2 (A or B) in the same synthesis. As the information brought by each sensor is not equivalent, different weights may be used depending on the sensor.

Input/Output variables, parameters

input	role
current date sensor	LANDSAT 8 or Sentinel-2

output	role
W_{Sensor}	Weight on sensor (one value per product)

parameter	role	default value
$W_{Sensor-S2A}$	Weight for Sensor S2A	1
$W_{Sensor-S2B}$	Weight for Sensor S2A	1
$W_{Sensor-L8}$	Weight for Sensor L8	0.33

Pseudocode

- $\diamond\,$ If current date sensor is Sentinel-2A
 - * $W_{Sensor}=W_{Sensor_S2A}$
- ♦ If current date sensor is Sentinel-2B
 - $*\ W_{Sensor}{=}W_{Sensor_S2B}$
- ♦ else
 - * W_{Sensor}=W_{Sensort.8}

Weight on clouds

Input/Output variables, parameters

input	role
Cloud-mask	Cloud and shadows mask from L2A product at 20m reso-
	lution

output	role
W_{Cloud}	Weight on distance to cloud (one value per pixel and per
	resolution, 10m, 20m)

parameter	role	default value
coarse resolution	coarse resolution for quicker convolution	240m
sigma-large-cloud-dist	standard deviation of gaussian filter for distance to large clouds	
sigma-small-cloud-dist	standard deviation of gaussian filter for distance to small clouds	

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Pseudocode

- ♦ binarize cloud mask (O if no cloud no shadow, 1 if either cloud or shadow)
- \diamond undersample to lower resolution (240m), with a bicubic resampling, without aliasing.
- ⋄ compute distances at low resolution
 - * $Dist_{LargeCloud,LowRes} = gaussian-filter(binarized low res cloud mask, standard deviation = sigma_{large-cloud-dist})$
 - * $Dist_{SmallCloud,LowRes} = gaussian-filter(binarized low res cloud mask, standard deviation = sigma-small-cloud-dist)$
 - * For each resolution (10m, 20m):
 - * resample at full resolution (10 and 20m)
 - \star Dist_{LargeCloud,FullRes} = bilinear-oversampling(Dist_{LargeCloud,LowRes})
 - \star Dist_{SmallCloud,FullRes} = bilinear-oversampling(Dist_{LargeCloud,LowRes})
 - \star for each pixel:
 - compute weight
 - W_{Cloud}= $(1 dist_{LargeCloud}) * (1 dist_{LargeCloud})$

Weight on Date The idea here is to give more weight to the dates close to the central date of the synthesis time window, and less weight to the edges of the time window. However, large difference in weights would result in not taking into account the dates at the edges of the synthesis window, unless they are the only one available in the synthesis window. As a result, we decided to weight the images by 1 in the middle of the synthesis window, and by 0.5 at the edges.

Input/Output variables, parameters

input variable	role
$date_{L2A}$	L2A date, expressed in days
$date_{L3A}$	L3A date, expressed in days
Δ_{Max}	Half Synthesis period (Days)

output variable	role
W_{date}	Weight on date (one value per product xxx)

parameter	role	default value
W_{Min}	Minimum weight at edge of synthesis time window	0.5

Pseudocode abs() is the absolute value

$$W_{date} = 1 - \frac{abs(date_{L2A} - date_{L3A})}{\Delta_{Max}}.(1 - W_{min})$$

Weight on AOT

Inputs, Outputs, Parameters

input	role
AOT	AOT values from L2A product at 20m resolution

output	role
W_{Cloud}	Weight on distance to cloud (one value per pixel and per
	resolution, 10m, 20m)

parameter	role	default value
W_{AOTMin}	min weight depending on AOT	
W_{AOTMax}	max weight depending on AOT	
AOT_{Max}	maximum value of the linear range for weights w.r.t AOT	

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Pseudocode

- $\diamond\,$ resample AOT at 10m resolution
- ♦ for each resolution (10m, 20m):
 - * for each pixel :
 - \star if (AOT(pix) \leq max_{AOT}):
 - $W_{AOT} = W_{AOTMin} + (W_{AOTMax} W_{AOTMin}) * (1. \frac{AOT(l,p)}{AOT_{Max}})$
 - \star else :
 - $W_{AOT} = W_{AOT}$

Compute total weight (product of individual weights)

Inputs, Outputs, Parameters

input	role
L2A	L2A product from date N

output	role
W_N	Weight for date N (one value per pixel and per resolution,
	10m, 20m)

Pseudocode

- \diamond Compute Weight on sensor (W_{sensor})
- ♦ Compute Weight on date (W_{date})
- ♦ Compute weight on AOT (W_{AOT})
- ♦ Compute Weight on clouds (W_{cloud})
- ♦ For each resolution (10m, 20m)
 - * For each pixel
 - \star W_N= W_{AOT} .W_{date}. W_{cloud}. W_{sensor}

4 Annex 1: Table of correspondance between LANDSAT 8 and Sentinel-2 bands

S2 Band	L8 band	wavelength (nm)	S2 Resol. (m)	provided in L3A
1	1	450	60	No
2	2	490	10	Yes
3	3	560	10	Yes
4	4	670	10	Yes
5	_	705	20	Yes
6	_	740	20	Yes
7	_	780	20	Yes
8	_	820	10	Yes
8a	5	865	20	Yes
9	_	940	60	No
10	6	1340	60	No
11	7	1650	20	Yes
12	8	2200	20	Yes



5 Annex 2: directional model

```
#This python_code provides the Ross_Thick- Li Sparse directional model
import math as m
import numpy as np
import pylab as p
class angles:
    #constructor
    def __init__(self,theta_s,phi_s,theta_v,phi_v):
        self.theta_s=theta_s*m.pi/180
        self.theta_v=theta_v*m.pi/180
        self.phi=(phi_s - phi_v)*m.pi/180
        #if self.phi < 0 :</pre>
            #self.phi=self.phi + 2*m.pi
    #function delta
    def delta(self):
        delta=m.sqrt(m.tan(self.theta_s)*m.tan(self.theta_s) + m.tan(self.theta_v)*m.tan(self.theta_v)
        return delta
    #Air Mass
    def masse(self):
        masse=1/m.cos(self.theta_s)+1/m.cos(self.theta_v)
        #print masse
        return masse
    #Function xsi
    def cos_xsi(self):
        cos_xsi=m.cos(self.theta_s)*m.cos(self.theta_v) + m.sin(self.theta_s)*m.sin(self.theta_v)*m.cos
        return cos_xsi
    def sin_xsi(self):
        x=self.cos_xsi()
        sin_xsi=m.sqrt(1 - x*x)
        return sin_xsi
    def xsi(self):
        xsi=m.acos(self.cos_xsi())
        return xsi
    xsi_0=1.5*m.pi/180.
    #Function t
    def cos_t(self):
        trig=m.tan(self.theta_s)*m.tan(self.theta_v)*m.sin(self.phi)
        d=self.delta()
        coef=1 #Coef=1 looks good, but Bréon et Vermote use Coef=2
        cos_t=min(max(coef/self.masse()*m.sqrt(d*d + trig*trig),-1),1)
        return cos_t
    def sin_t(self):
        x=self.cos_t()
        sin_t=m.sqrt(1 - x*x)
        return sin_t
    def t(self):
        #print 'theta_v %f cos_t %F'%(self.theta_v*180/m.pi,self.cos_t())
        t=m.acos(self.cos_t())
        return t
```



```
#function FV Ross_Thick, V stands for Volume
def FV(self):
    FV=self.masse()/m.pi*(self.t() - self.sin_t()*self.cos_t() - m.pi) + (1+self.cos_xsi())/2/m.cos
    return FV

#function FR Li-Sparse, R stands for Roughness
def FR(self):
    A=1/(m.cos(self.theta_s)+m.cos(self.theta_v))

    FR=4/3./m.pi*A*((m.pi/2-self.xsi())*self.cos_xsi()+self.sin_xsi())*(1+1/(1+self.xsi()/self.xsi_
    return FR

def dir_mod(self,kV,kR):
    rho=1 +kV*self.FV() + kR*self.FR()
    return rho
```