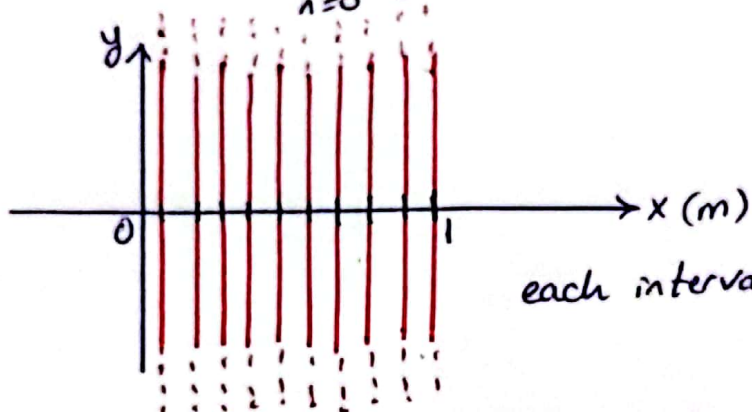


Describing charge distributions with delta functions

①

In 2D, a single delta function represents a line that "goes on forever". For example, $\delta(x - \frac{1}{10})$ is a line at $x = \frac{1}{10}$ that extends parallel to y-axis.

For $\lambda(x) = \sum_{n=0}^{10} q_0 n^2 \delta(x - \frac{1}{10})$, we have



each interval is separated by $\frac{1}{10} \text{ m}$.

②

$$Q = \int_{-\infty}^{\infty} \lambda(x) dx = \int_{-\infty}^{\infty} \sum_{n=0}^{10} q_0 n^2 \delta(x - \frac{1}{10}) dx = \sum_{n=0}^{10} q_0 n^2 \underbrace{\int_{-\infty}^{\infty} \delta(x - \frac{1}{10}) dx}_1$$
$$= q_0 \sum_{n=0}^{N=10} n^2 = q_0 \frac{N(N+1)(2N+1)}{6}$$

$$\text{For } N=10, \quad Q = 385 q_0$$

Spherical charge distributions are special

$$\textcircled{1} \quad \vec{\nabla} \cdot \vec{E} = \frac{\rho}{\epsilon_0} \Rightarrow \rho = \epsilon_0 \vec{\nabla} \cdot \vec{E}$$

In spherical coordinates :

$$\vec{\nabla} \cdot \vec{E} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 E_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta E_\theta) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} (E_\phi)$$

For the \vec{E} -field in this problem : $E_\theta = 0$ and $E_\phi = 0$.

$$\begin{aligned} \vec{\nabla} \cdot \vec{E} &= \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \cdot c r^n) = \frac{c}{r^2} \frac{\partial}{\partial r} (r^{n+2}) = \frac{c(n+2) r^{n+1}}{r^2} \\ &= c(n+2) r^{n-1} \end{aligned}$$

$$\text{Hence, } \rho = \epsilon_0 \vec{\nabla} \cdot \vec{E} = \epsilon_0 c(n+2) r^{n-1} \quad \text{for } n \neq -2$$

$\textcircled{2}$ For $n = -2$, $\rho = 0$ according to the equation above.

If $\rho = 0$, then we would not get $\vec{E} = c r^{-2} \hat{r}$

$$\text{For } n = -2 : \vec{\nabla} \cdot \vec{E} = c \vec{\nabla} \cdot \left(\frac{\hat{r}}{r^2} \right) = c 4\pi \delta^3(\vec{r})$$

$$\text{Hence, } \rho = 4\pi \epsilon_0 c \delta^3(\vec{r}) \quad \text{for } n = -2$$

Extra info: Actually not all values of n is physically possible. We need to have $\int_{\text{volume}} \rho(r) d\tau = \text{finite}$.

For $n \geq -2$, we have a finite total charge $\int \rho d\tau$.

However for $n < -2$, $\int \rho d\tau$ runs to infinity.

So only $n \geq -2$ cases are physically possible.

$$\textcircled{3} \quad E = \text{const.} \Rightarrow n = 0.$$

So $\rho = \epsilon_0 c(n+2) r^{n-1}$ becomes

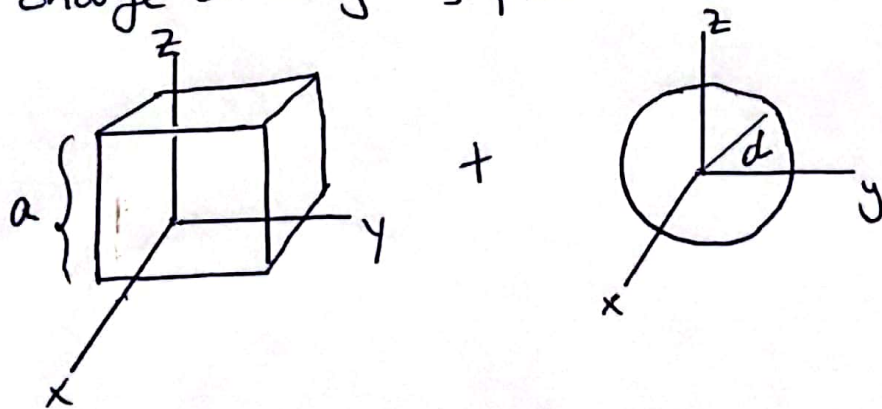
$$\rho(r) = \frac{2\epsilon_0 c}{r} \quad \text{for } n = 0.$$

Cube with a hole

- ①
- Gauss' law is always true.
 - In this problem, Gauss' law is partially useful. It can be used to find E-field due to the spherical hole for which we can use a spherical Gaussian surface.
 - A cube does not have spherical, cylindrical, or planar symmetry. So we cannot find a Gaussian surface with such symmetries.

② We can use superposition.

A cube with uniform charge density $+f$ with a spherical hole at its center is equivalent to a solid cube with charge density $+f$ plus a sphere with charge density $-f$

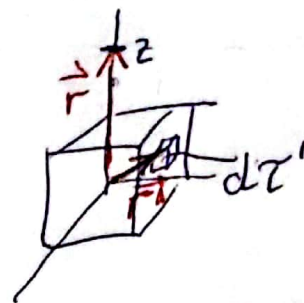


Let's use Gauss' law for the sphere of radius d by choosing a spherical Gaussian surface of radius z .

$$\oint_S \vec{E} \cdot d\vec{a} = E \oint_S da = E \cdot 4\pi z^2 = \frac{Q_{\text{enc}}}{\epsilon_0} = \frac{-f \frac{4}{3}\pi d^3}{\epsilon_0}$$

$$\text{So } \vec{E}_{\text{sphere}} = -\frac{f d^3}{3\epsilon_0 z^2} \hat{z}$$

For \vec{E}_{cube} use direct integration



$$d\tau = dx' dy' dz'$$

$$\vec{r} = z \hat{z}$$

$$\vec{r}' = x' \hat{x}' + y' \hat{y}' + z' \hat{z}'$$

$$\vec{r} = \vec{r} - \vec{r}' = -x' \hat{x}' - y' \hat{y}' + (z - z') \hat{z}'$$

$$\frac{\hat{r}}{r^2} = \frac{\vec{r}}{r^3} = \frac{-x' \hat{x}' - y' \hat{y}' + (z - z') \hat{z}'}{[x'^2 + y'^2 + (z - z')^2]^{3/2}}$$

$$\vec{E}_{\text{cube}} = \frac{1}{4\pi\epsilon_0} \int \rho \frac{\vec{r}}{r^3} d\tau' = \frac{\rho}{4\pi\epsilon_0} \int_{-\frac{a}{2}}^{\frac{a}{2}} \int_{-\frac{a}{2}}^{\frac{a}{2}} \int_{-\frac{a}{2}}^{\frac{a}{2}} \frac{-x' \hat{x}' - y' \hat{y}' + (z - z') \hat{z}'}{[x'^2 + y'^2 + (z - z')^2]^{3/2}} dx' dy' dz'$$

Actually by symmetry we can see that x and y components of \vec{E}_{cube} should vanish; i.e.

$$E_{x, \text{cube}} = \frac{\rho}{4\pi\epsilon_0} \underbrace{\int \int \int \left(\frac{-x'}{r^3} \right) dx' dy' dz'}_{\substack{\text{odd function} \\ \text{symmetric limits}}} = 0 \quad \text{and similarly } E_{y, \text{cube}} = 0.$$

$$\text{Hence } \vec{E}_{\text{cube}} = \frac{\rho}{4\pi\epsilon_0} \int_{-\frac{a}{2}}^{\frac{a}{2}} \int_{-\frac{a}{2}}^{\frac{a}{2}} \int_{-\frac{a}{2}}^{\frac{a}{2}} \frac{(z - z')}{[x'^2 + y'^2 + (z - z')^2]^{3/2}} dx' dy' dz' \hat{z}$$

$$\text{and } \vec{E}_{\text{total}} = \vec{E}_{\text{cube}} + \vec{E}_{\text{sphere}}$$