

PYMOOSH FUNCTIONS

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This document serves as a reference to know what PyMoosh can do. **It lists all (user-facing) existing functions**, sorted by which file they are defined in.

It references what are their arguments and returned objects, as well as which Notebooks explain how they work (if there is one).

Throughout the manuscript, r and t refer to the amplitude reflection and transmission coefficients, while R and T refer to the intensity coefficients. TE (resp. TM) corresponds to the polarization where the electric field (resp. magnetic field) is perpendicular to the plain of incidence. As we define z to be the axis perpendicular to the interfaces, and xOz to be the plane of incidence, this means $E = E_y$ in TE polarization (resp. $H = H_y$ in TM).



1 `classes.py`

Main classes are defined here: Material, Structure, Beam, Window.

They are explained in `Basic_tutorials/PyMoosh_Basics`.

Name	Necessary Arguments	Returns	Comments
Material	mat	Material object	Used to define structures
<code>material.get_permittivity</code>	wavelength	Permittivity (complex) at wavelength	
<code>material.get_permeability</code>	wavelength	Permeability (complex) at wavelength	
Structure	materials, layer_type, thickness	Structure object	Used for most functions
<code>structure.plot_stack</code>	wavelength	None	Creates and shows a schematic of the structure
Beam	wavelength, incidence, polarization, horizontal_waist	Beam object	Used to compute fields
Window	width, beam_relative_position, horizontal_pixel_size, vertical_pixel_size	Window object	Used to compute fields

2 core.py

Most common functions are defined here, computing fields; r, t, R, T coefficients, and absorption in a structure.

Most are explained in Basic_tutorials/PyMoosh_Basics.

Name	Necessary Arguments	Returns	Comments
absorption	structure, wavelength, incidence, polarization	absorption, r, t, R, T	
field	structure, beam, window	Ey (TE) or Hy (TM)	
fields	structure, beam, window	Ey, Hx, Hz (TE) or Hy, Ex, Ez (TM)	
coefficient	structure, wavelength, incidence, polarization	r, t, R, T	just a wrapper, calls coefficient_S by default
coefficient_S	structure, wavelength, incidence, polarization	r, t, R, T	S matrix formalism
absorption_S	structure, wavelength, incidence, polarization, layers	absorbtion, r, t, R, T	absorption is computed in given layers, all if layers is None

3 vectorized.py

Vectorized versions of the core functions, to compute coefficients along spectral or angular ranges.

They are used in Basic_tutorials/PyMoosh_Basics.

Name	Necessary Arguments	Returns	Comments
spectrum	structure, incidence, polarization, wl_min, wl_max, len_wl	wavelengths, r, t, R, T	Just a wrapper, calls S matrix formalism by default
spectrum_S	structure, incidence, polarization, wl_min, wl_max, len_wl	wavelengths, r, t, R, T	S matrix formalism
spectrum_A	structure, incidence, polarization, wl_min, wl_max, len_wl	wavelengths, r, t, R, T	Abélès matrix formalism
angular	structure, wavelength, polarization, theta_min, theta_max, len_an	angles, r, t, R, T	Just a wrapper, calls S matrix formalism by default
angular_S	structure, wavelength, polarization, theta_min, theta_max, len_an	angles, r, t, R, T	S matrix formalism
angular_A	structure, wavelength, polarization, theta_min, theta_max, len_an	angles, r, t, R, T	Abélès matrix formalism

4 `models.py`

Most common functions to define Material models. Beware of units! **Wavelengths are always expected in nm**, but other parameter units may vary depending on common use (typically, eV).

Their general idea is presented in `In-depth_examples/How_materials_works`

Name	Necessary Arguments	Returns	Comments
BrendelBormann	wavelength, f0, omega_p, Gamma0, f, omega, gamma, sigma	permittivity	An improvement on the Drude-Lorentz model
Drude	wavelength, omega_p, Gamma0	permittivity	Free electrons
Lorentz	wavelength, f, omega, gamma, eps	permittivity	Bound electrons
DrudeLorentz	wavelength, omega_p, Gamma0, f, omega, gamma	permittivity	Both!
ExpData	wavelength, wavelength_list, permittivity_list	permittivity	Interpolation on external data

5 `modes.py`

Functions to play around with guided modes.

They are detailed in `In-depth_examples/Finding_representing_guided_modes`.

Name	Necessary Arguments	Returns	Comments
<code>dispersion</code>	<code>alpha</code> , <code>structure</code> , <code>wavelength</code> , <code>polarization</code>	$1/ r $	Computes the reflection coefficient at given k_x (α)
<code>complex_map</code>	<code>structure</code> , <code>wavelength</code> , <code>polarization</code> , <code>real_bounds</code> , <code>imag_bounds</code> , <code>n_real</code> , <code>n_imag</code>	<code>re(n)</code> , <code>im(n)</code> , $1/ r(n) $	Computes dispersion on a 2D map of complex n values
<code>guided_modes</code>	<code>structure</code> , <code>wavelength</code> , <code>polarization</code> , <code>neff_min</code> , <code>neff_max</code>	$n(\lambda)$	Computes the index of modes at given wavelength
<code>follow_guided_modes</code>	<code>structure</code> , <code>wavelength_list</code> , <code>polarization</code> , <code>neff_min</code> , <code>neff_max</code>	$n(\lambda)$	Computes the evolution of the index of modes with varying wavelength
<code>profile</code>	<code>structure</code> , <code>n_eff</code> , <code>wavelength</code> , <code>polarization</code>	x , $E(x)$	Computes the field along the structure thickness, for the given index

6 photo.py

Functions dedicated to photovoltaic computations.

photo is detailed in [In-depth_examples/Photovoltaics](#)

Name	Necessary Arguments	Returns	Comments
solar	wavelength	$I(\lambda)$	Solar irradiance, converted in A/cm^2
photo	structure, incidence, polarization, wl_min, wl_max, active_layers, number_points	conversion_efficiency, total_current, total_current_max, wavelength_list, photon_density, total_absorbed	short circuit currents across several active layers are summed
opti_photo	structure, incidence, polarization, wl_min, wl_max, active_layers, number_points	id.	id. but vectorized
gx	structure, incidence, polarization, wl_min, wl_max, number_points	x, g(x)	Density of absorbed photons in $photon/s/m^2/nm$

7 `alt_methods.py`

Other matrix formalisms to compute the main coefficients.

They are explained in the PyMoosh tutorial paper, and tested in `In-depth_examples/Adapting_matrix_formalism`

Name	Necessary Arguments	Returns	Comments
<code>coefficient_A</code>	structure, wavelength, incidence, polarization	r, t, R, T	Abélès formalism
<code>coefficient_T</code>	structure, wavelength, incidence, polarization	r, t, R, T	Transfer Matrix formalism
<code>coefficient_DN</code>	structure, wavelength, incidence, polarization	r, t, R, T	Dirichlet-to-Neumann formalism
<code>coefficient_I</code>	structure, wavelength, incidence, polarization	r, t, R, T	Impedance formalism
<code>absorption_A</code>	structure, wavelength, incidence, polarization	absorption, r, t, R, T	Abélès formalism

8 `optim_algo.py`

Tried and tested optimization algorithms. We have a deep-dive tutorial on ArXiv (but cite the article paper) about how to optimize, and the use of these functions is shown (with differential evolution as an example) in `Basic_tutorials/Optimizing_with_PyMoosh`

Name	Necessary Arguments	Returns	Comments
<code>differential_evolution</code>	<code>f_cout</code> , <code>budget</code> , <code>X_min</code> , <code>X_max</code>	best, convergence	DE!
<code>bfgs</code>	<code>f_cout</code> , <code>npas</code> , <code>start</code>	best, <code>f_cout(best)</code>	Gradient descent
<code>QODE</code>	<code>f_cout</code> , <code>budget</code> , <code>X_min</code> , <code>X_max</code>	best, convergence	
<code>QNDE</code>	<code>f_cout</code> , <code>budget</code> , <code>X_min</code> , <code>X_max</code>	best, convergence	

9 `anisotropic.py`

Equivalent to the main functions but for structures including anisotropic materials.

They are presented in `In-depth_examples/Anisotropic_structure`

Name	Necessary Arguments	Returns	Comments
<code>coefficients_ani</code>	<code>structure</code> , <code>wl</code> , <code>theta_inc</code>	<code>t_pp</code> , <code>t_ps</code> , <code>t_sp</code> , <code>t_ss</code> , <code>r_pp</code> , <code>r_ps</code> , <code>r_sp</code> , <code>r_ss</code>	
<code>AniStructure</code>	<code>materials</code> , <code>layer_type</code> , <code>thickness</code> , <code>ani_rot_angle</code> , <code>ani_rot_axis</code>	<code>AniStructure</code> object	Used to compute the reflection and transmission coefficients
<code>AniMaterial</code>	<code>mat</code>	<code>AniMaterial</code> object	Used to define the structures
<code>AniMaterial.get_permittivity_ani</code>	<code>wavelength</code>	[<code>epsilon1</code> , <code>epsilon2</code> , <code>epsilon2</code>]	

10 non_local.py

Equivalent to the main functions but for structures including non-local (spatially dispersive) materials.

They are presented in In-depth_examples/Spatial_dispersion_aka_nonlocality

Name	Necessary Arguments	Returns	Comments
NLcoefficient	structure, wavelength, incidence, polarization	r, t, R, T	
NLStructure	materials, layer_type, thickness	NLStructure object	Used to compute the reflection and transmission coefficients
NLMaterial	mat	NLMaterial object	Used to define the structures
NLMaterial.get_permittivity	wavelength	epsilon	
NLMaterial.get_values_nl	wavelength	beta2, chi_b, chi_f, omega_p	

11 green.py

The function computing the green Dyadic function for a point source. Yes, it's all alone and sad but it didn't make sense to put it anywhere else. It is presented in In-depth_examples/Computing_Green_functions

Name	Necessary Arguments	Returns	Comments
green	structure, window, lam, source_interface	E	Returns the field resulting from the source